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# Phase Relations of Peralkaline Silicic Magmas and Petrogenetic Implications

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*The phase relationships of three peralkaline rhyolites from the Kenya Rift have been established at 150 and 50 MPa, at oxygen fugacities of  $NNO - 1.6$  and  $NNO + 3.6$  ( $\log fO_2$  relative to the Ni–NiO solid buffer), between 800 and 660°C and for melt  $H_2O$  contents ranging between saturation and nominally anhydrous. The stability fields of fayalite, sodic amphiboles, chevkinite and fluorite in natural hydrous silicic magmas are established. Additional phases include quartz, alkali feldspar, ferrohedenbergite, biotite, aegirine, titanite, montdorite and oxides. Ferrohedenbergite crystallization is restricted to the least peralkaline rock, together with fayalite; it is replaced at low melt water contents by ferrichterite. Riebeckite–arfvedsonite appears only in the more peralkaline rocks, at temperatures below 750°C (dry) and below 670°C at  $H_2O$  saturation. Under oxidizing conditions, it breaks down to aegirine. In the more peralkaline rocks, biotite is restricted to temperatures below 700°C and conditions close to  $H_2O$  saturation. At 50 MPa, the tectosilicate liquidus temperatures are raised by 50–60°C, and that of amphibole by 30°C. Riebeckite–arfvedsonite stability extends down nearly to atmospheric pressure, as a result of its F-rich character. The solidi of all three rocks are depressed by 40–100°C compared with the solidus of the metaluminous granite system, as a result of the abundance of F and Cl. Low  $fO_2$  lowers solidus temperatures by at least 30°C. Comparison with studies of metaluminous and peraluminous felsic magmas shows that plagioclase crystallization is suppressed as soon as the melt becomes peralkaline, whatever its CaO or volatile contents. In contrast, at 100 MPa and  $H_2O$  saturation, the liquidus temperatures of quartz and alkali feldspar are not significantly affected by changes in rock peralkalinity, showing that the incorporation of water in peralkaline melts diminishes the depression of liquidus temperatures in dry peralkaline silicic melts compared with dry metaluminous or peraluminous varieties. At 150 MPa, pre-eruptive melt  $H_2O$  contents range from 4 wt % in the least peralkaline rock to nearly 6 wt % in the two more*

*peralkaline compositions, in broad agreement with previous melt inclusion data. The experimental results imply magmatic  $fO_2$  at or below the fayalite–quartz–magnetite solid buffer, temperatures between 740 and 660°C, and melt evolution under near  $H_2O$  saturation conditions.*

KEY WORDS: peralkaline; rhyolite; phase equilibria

## INTRODUCTION

Since the field recognition of peralkaline igneous rocks by Foerstner (1881), much effort has been expended in the petrological, geochemical and experimental fields to understand their petrogenesis (e.g. Washington, 1913, 1914a, 1914b; Bowen, 1937; Zies, 1960; Carmichael, 1962; Carmichael & MacKenzie, 1963; Bailey & Schairer, 1964, 1966; Thompson & MacKenzie, 1967; Ewart *et al.*, 1968; Noble, 1968; Nicholls & Carmichael, 1969; Bailey & Macdonald, 1970; Gibson, 1972; Macdonald & Bailey, 1973; Bailey, 1974; Barberi *et al.*, 1974, 1975; Villari, 1974; Roux & Varet, 1975; Bailey & Cooper, 1978; Harris, 1983; Mahood, 1984; Mungall & Martin, 1995; Bohron & Reid, 1997). For the vast majority of natural occurrences, peralkaline felsic rocks can be explained by fractional crystallization of mildly alkalic or transitional basalts at both low  $fH_2O$  and low  $fO_2$  (e.g. Harris, 1983; Caroff *et al.*, 1993). Perhaps the best example of such a mechanism is the Boina volcanic rock suite in the Afar Rift, Ethiopia (Barberi *et al.*, 1975). Other studies have shown, however, that this simple

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petrogenetic scheme may not apply everywhere. Davies & Macdonald (1987), Macdonald *et al.* (1987) and Black *et al.* (1997), for instance, have conclusively shown, mainly on geochemical grounds, that the peralkaline rhyolites cropping out near Naivasha in the Kenyan Rift Valley (the Greater Olkaria Volcanic Complex, GOVC) cannot be the daughter magmas of the associated basalts, and they proposed an origin by partial melting of the crust triggered by alkali-bearing volatiles. Mahood *et al.* (1990) proposed that pantellerite rhyolites at the type locality of Pantelleria are the product of a two-step process, whereby partial melting of alkali gabbros yields trachyte melts whose subsequent fractionation ends in the strongly peralkaline rhyolites [but see Civetta *et al.* (1998)].

All studies up to 1988 agreed as to the dry, or at least the very H<sub>2</sub>O-poor, character of peralkaline rhyolites, in particular when compared with their metaluminous equivalents. The reasons for believing so (Bailey & Macdonald, 1987) include the nearly anhydrous state of all peralkaline obsidians, their high Cl content, thought to be incompatible with the exsolution of a hydrous vapour phase, and the fact that phase equilibrium experiments performed under dry conditions better reproduce the crystallization sequence in the rocks than those under wet conditions (Bailey, 1974). Another rarely considered factor is that peralkalinity is accompanied by a significant fall in liquidus temperatures relative to metaluminous rhyolites. This was apparent in the pioneering studies of Schairer & Bowen (1955, 1956), who worked out the phase relationships in the dry synthetic Na<sub>2</sub>O–Al<sub>2</sub>O<sub>3</sub>–SiO<sub>2</sub> and K<sub>2</sub>O–Al<sub>2</sub>O<sub>3</sub>–SiO<sub>2</sub> ternaries and found eutectic temperatures as low as 740°C at 1 bar on the peralkaline sides. Addition of iron to the Na<sub>2</sub>O–Al<sub>2</sub>O<sub>3</sub>–SiO<sub>2</sub> system does not change the trend of falling temperatures (Bailey & Schairer, 1966), with even lower eutectic temperatures, around 728°C, in the SiO<sub>2</sub>-oversaturated field of the Na<sub>2</sub>O–Al<sub>2</sub>O<sub>3</sub>–SiO<sub>2</sub>–Fe<sub>2</sub>O<sub>3</sub> system. The implication is that there are no thermal limitations to generating dry silicic peralkaline melts by direct melting of crust under dry conditions (given a suitable protolith), a conclusion in sharp contrast to metaluminous or peraluminous systems where dry melting requires temperatures of ~1000°C (Bowen, 1937), i.e. temperatures unlikely to be reached in the crust, even in extensional tectonic settings.

Melt inclusion studies of peralkaline rhyolites, however, have found significant amounts of dissolved H<sub>2</sub>O (1.4–5 wt %, Ascension: Harris, 1986; Pantelleria: Kovalenko *et al.*, 1988; Lowenstern & Mahood, 1991; Fantale, Ethiopia: Webster *et al.*, 1993; GOVC, Kenya: Wilding *et al.*, 1993; Mayor Island, New Zealand: Barclay *et al.*, 1996). Although this conflicts with previously available phase equilibrium constraints (Bailey, 1974), it is worth stressing that our knowledge of the physico-chemical conditions of the evolution of peralkaline silicic magmas is still

surprisingly poor. For instance, because of the scarcity of coexisting Fe–Ti oxide minerals in peralkaline rocks, pre-eruptive temperatures are poorly constrained. The few quantitative estimates available (Carmichael, 1967, 1991; Bizouard *et al.*, 1980; Mahood, 1981; Wolff & Wright, 1981; Conrad, 1984; Novak & Mahood, 1986; Crisp & Spera, 1987; Jørgensen, 1987; Ghiorso & Sack 1991) indicate pre-eruptive temperatures in the range 700–950°C at *f*O<sub>2</sub> generally close to the fayalite–magnetite–quartz buffer (FMQ).

Available experimental work is of little use in constraining the pre-eruptive conditions of peralkaline magmas (Webster *et al.*, 1993). Most studies have been performed in synthetic systems under either dry (Bailey & Schairer, 1966) or water-saturated conditions (Carmichael & MacKenzie, 1963; Thompson & MacKenzie, 1967). The same problem applies to the few experimental studies on natural peralkaline rhyolites (Bailey & Cooper, 1978). In addition, in none of these studies was the *f*O<sub>2</sub> carefully controlled. Experiments at H<sub>2</sub>O saturation have all been performed under very oxidizing conditions which seriously contrast with the rock record (Carmichael, 1962; Nicholls & Carmichael, 1969; Sutherland, 1974). Thus, the phase equilibria of peralkaline magmas are poorly known compared with their metaluminous or peraluminous silicic counterparts (e.g. Clemens & Wall, 1981; Clemens *et al.*, 1986; Webster *et al.*, 1987; Conrad *et al.*, 1988; Scaillet *et al.*, 1995; Martel *et al.*, 1998, 1999; Dall'Agnol *et al.*, 1999; Scaillet & Evans, 1999), which seriously hinders any attempt at constraining their evolution. This work intends to fill this gap by determining the phase relationships of three natural peralkaline rhyolites under controlled *f*O<sub>2</sub> and *f*H<sub>2</sub>O conditions. Although our main goal is to provide quantitative estimates of pre-eruptive conditions, the comprehensive phase equilibrium data provided should prove useful in determining the pre-eruptive conditions of other mildly to moderately peralkaline silicic rocks.

## ROCKS STUDIED

The three natural rhyolites used come from the Greater Olkaria Volcanic Complex in the Naivasha area of the Kenyan Rift Valley, where the geology is well known (Clarke *et al.*, 1990) and where there has been extensive petrological and geochemical work (Bailey & Macdonald, 1987; Davies & Macdonald, 1987; Macdonald *et al.*, 1987; Black *et al.*, 1997; Marshall *et al.*, 1998). The three rocks are phenocryst-poor obsidians; bulk compositions are listed in Table 1. They are comendites in the classification scheme of Macdonald (1974), with agpaite indices [(Na + K)/Al, in moles] ranging from 1.05 to 1.36. They represent the two extremes of the geochemical

range displayed by the comendite suite at Olkaria (Macdonald *et al.*, 1987). Sample ND 002 is only marginally peralkaline and has <5% (volume) phenocrysts of alkali feldspar, clinopyroxene, fayalite, magnetite, zircon, apatite and chevkinite. The glassy matrix has the same albitic index as the bulk rock (Table 1). Sample BL 575 is the most peralkaline specimen (NK/A = 1.36). It has 5% phenocrysts of alkali feldspar, quartz, riebeckite–arfvedsonite, aenigmatite, biotite, ilmenite and chevkinite. BL 575 contains some spherulites with a radial arrangement of riebeckite laths and aggregates of quartz and alkali feldspar. Microphenocrysts of riebeckite–arfvedsonite also occur in the matrix glass. Sample SMN 49 is chemically similar to BL 575, being only slightly more SiO<sub>2</sub> rich and less peralkaline (Table 1). Its phenocryst assemblage differs from that of BL 575 in the occurrence of fluorite (Marshall *et al.*, 1998) and the absence of riebeckite and aenigmatite. Also, it has lower Fe<sub>2</sub>O<sub>3</sub>/FeO than BL 575 (Table 1). In both BL 575 and SMN 49, the matrix glass is slightly more peralkaline than the bulk rock (Table 1). For all three samples, microprobe analyses of the matrix glass yield totals close to 100%, implying water contents <1 wt %. This is confirmed by Fourier transform IR spectroscopy (FTIR) analyses carried out on sample ND 002, whose matrix glass has 1800 ppm dissolved H<sub>2</sub>O (Wilding *et al.*, 1993). FTIR analyses performed by Wilding *et al.* (1993) on samples belonging to the same suite as BL 575 and SMN 49 yielded similar results.

## EXPERIMENTAL AND ANALYTICAL TECHNIQUES

### Preparation of charges

Hydrothermal experiments were performed using rock powders as starting material. Therefore, given the near-aphyric character of the natural rhyolites, most experiments in this study were of crystallization type (e.g. Clemens & Wall, 1981; Pichavant, 1987; Conrad *et al.*, 1988; Scaillet *et al.*, 1995; Martel *et al.*, 1998; Dall'Agnol *et al.*, 1999; Scaillet & Evans, 1999). Charges consisted of Au capsules (length 1.5–1.2 cm, thickness 0.02 cm, inner diameter 0.29–0.27 cm), first loaded with deionized water, then with silver oxalate as a source of CO<sub>2</sub> for H<sub>2</sub>O-undersaturated experiments, and finally with the rock powder. Fluids of various H<sub>2</sub>O/CO<sub>2</sub> ratio were used as a practical means of varying the water activity (*a*H<sub>2</sub>O) at any given temperature (*T*) and pressure (*P*) but also because geological evidence points to a role for CO<sub>2</sub> in the petrogenesis of felsic peralkaline magmas in the Kenya rift (Bailey & Macdonald, 1987; Macdonald *et al.*, 1993). The weight proportion of H<sub>2</sub>O + CO<sub>2</sub> [i.e. (H<sub>2</sub>O + CO<sub>2</sub>)/(H<sub>2</sub>O + CO<sub>2</sub> + silicate)] in most charges was maintained in the range 7–10 except for charge numbers

1–7, in which this parameter was around four (in the following,  $X_{H_2O_m}$  refers to the initial mole fraction of H<sub>2</sub>O in the fluid phase loaded to the capsule). A single series of experiments was also performed with only H<sub>2</sub>O added in amounts below those required for saturation. Nominally dry charges were also prepared simply by loading the rock powder into Au capsules. These charges contain in fact 0.2–0.4 wt % water, owing to the presence of residual H<sub>2</sub>O in the matrix glass and adsorbed water. Upon loading, capsules were sealed by arc-welding, immersed in a 120°C oil bath, and then weighed to check for leaks. After completion of an experiment, capsules were reweighed to ensure closed-system behaviour at *P* and *T*.

### Equipment

All experiments were performed in standard cold-seal pressure vessels with or without semi-permeable H<sub>2</sub> membranes (respectively designed CSPV-H<sub>2</sub> and CSPV). Total pressure is known to within 2 MPa. Temperatures were continuously recorded by external unsheathed type-K thermocouples and are known to ±8°C. Experiments performed with an H<sub>2</sub> membrane were terminated by removing the vessel from the furnace and allowing it to cool in air. To minimize exsolution, most experiments (charges with numbers higher than 12) were terminated using a small ventilator to accelerate cooling and by maintaining isobaric conditions down to 150°C. Experiments performed without H<sub>2</sub> membranes at 150 MPa were drop quenched. For each composition, phase equilibria were established at two pressures, 50 and 150 MPa, and under two contrasting redox conditions, in the temperature range 660–800°C. Run durations varied according to temperature, between 7 days at around 800°C and 51 days at or below 700°C. A total of 158 phase equilibrium experiments was performed.

### *f*H<sub>2</sub> control

Two types of vessels were used to perform either reduced (CSPV-H<sub>2</sub>) or oxidized (CSPV) experiments. In the former, the hydrogen fugacity (*f*H<sub>2</sub>) was read continuously via H<sub>2</sub> membranes connected to a transducer with an uncertainty of 0.01 MPa (Scaillet *et al.*, 1992; Schmidt *et al.*, 1995; Scaillet & Evans, 1999). Under oxidizing conditions the *f*H<sub>2</sub> of the vessel was determined by using the dependence of the ferric/ferrous ratio of the three bulk compositions on *f*H<sub>2</sub>. This relationship was calibrated by annealing the rock powders at 150 MPa, 800°C, H<sub>2</sub>O saturation and various *f*H<sub>2</sub>, as monitored with H<sub>2</sub> membranes. Volumetric titration of the ferrous iron content of the run products yields the relationship between FeO content and *f*H<sub>2</sub> specific to each composition. Subsequent

Table 1: Bulk-rock and matrix glass compositions (wt %)

Sample no.:	ND 002		BL 575		SMN 49	
	bulk <sup>1</sup>	matrix glass 6	bulk <sup>2</sup>	matrix glass 5	bulk <sup>2</sup>	matrix glass 6
<i>n</i> : <sup>3</sup>						
SiO <sub>2</sub>	75.2	75.92 (21)4	72.64	73.24 (15)	74.17	73.47 (43)
TiO <sub>2</sub>	0.17	0.11 (6)	0.18	0.14 (2)	0.17	0.12 (5)
Al <sub>2</sub> O <sub>3</sub>	12.11	11.99 (12)	10.30	10.25 (16)	10.89	10.49 (14)
Fe <sub>2</sub> O <sub>3</sub>	0.83	—	1.39	—	1.98	—
FeO	1.06	—	2.75	—	2.21 (2)	—
FeO <sub>tot</sub>	1.81	1.70 (16)	4.00	4.09 (22)	3.98	3.82 (23)
MnO	0.04	0.02 (3)	0.06	0.05 (5)	0.06	0.10 (8)
MgO	0.07	0.03 (3)	0.04	0.01 (1)	0.00	0.02 (1)
CaO	0.44	0.34 (3)	0.10	0.05 (4)	0.29	0.06 (4)
Na <sub>2</sub> O	4.59	4.43 (16)	5.68	5.64 (10)	5.81	5.68 (11)
K <sub>2</sub> O	4.73	4.85 (8)	4.35	4.89 (7)	4.34	4.68 (20)
F	0.37	0.38 (8)	1.00	1.11 (11)	1.00	1.03 (11)
Cl	0.22	0.17 (2)	0.50	0.50 (5)	0.50	0.50 (3)
LOI	0.06	—	0.79	—	—	—
Total	99.80	98.78	98.59	98.43	100.15	99.73
Sr	6	—	1	—	—	—
Rb	290	—	668	—	741	—
Zr	439	—	2198	—	2666	—
NK/A (mole)	1.05	1.05	1.36	1.42	1.31	1.37
FeO/FeO + Fe <sub>2</sub> O <sub>3</sub> (wt %)	0.56	—	0.66	—	0.53	—

<sup>1</sup>From Macdonald *et al.* (1987).

<sup>2</sup>BL 575 and SMN 49 bulk analyses from Black *et al.* (1997) and Marshall *et al.* (1998), respectively, except for FeO determined by wet chemistry in this study, and for F and Cl, determined by EMPA of corresponding supraliquidus glasses.

<sup>3</sup>Number of analyses.

<sup>4</sup>Numbers in parentheses indicate 1 SD of replicate analyses in terms of smallest units cited.

determination of the ferrous content of H<sub>2</sub>O-saturated charges around 800°C thus gives the *f*H<sub>2</sub> of that run. The *f*H<sub>2</sub> varies among the vessels, between 0.011 and 0.004 MPa; these values correspond to *f*O<sub>2</sub> 3.6–4.4 log units above that of the Ni–NiO solid buffer (NNO + 3.6 to NNO + 4.4) at H<sub>2</sub>O saturation. The *f*H<sub>2</sub> obtained at that temperature for a given vessel was considered constant for that vessel and used for subsequent *f*O<sub>2</sub> calculations at lower temperatures. The *f*O<sub>2</sub> of two vessels used without an H<sub>2</sub> membrane was also measured using solid NiPd sensors (Taylor *et al.*, 1992; Pownceby & O'Neill, 1994). Values of NNO + 3 to NNO + 5.1 were obtained, straddling the haematite–magnetite equilibrium.

### *f*O<sub>2</sub> determination

The *f*O<sub>2</sub> of fluid-saturated charges was calculated using the equation (see Scaillet & Evans, 1999)

$$\log fO_2(\text{charge}) = \log fO_2(\text{H}_2\text{O saturation}) - 2\log(XH_2O_{in}). \quad (1)$$

For experiments performed at high *f*H<sub>2</sub>, the vast majority of *f*O<sub>2</sub> values fall in the range NNO – 1 to NNO – 2, with an average at NNO – 1.67 ± 0.4. For experiments performed at low *f*H<sub>2</sub>, most of the calculated *f*O<sub>2</sub> are in the range NNO + 3 to NNO + 4, with an average at NNO + 3.63 ± 0.44. Thus, for simplicity, average values of NNO – 1.6 and NNO + 3.6 are used in the following to characterize runs performed under reduced and oxidized conditions, respectively. At *X*H<sub>2</sub>O values < 0.3, however, the *f*O<sub>2</sub> starts to decrease significantly relative to that at H<sub>2</sub>O saturation, by several orders of magnitude for extremely H<sub>2</sub>O-poor charges (see Webster *et al.*, 1987). For nominally dry charges the amount of H<sub>2</sub>O present in the fluid phase, if any, depends on the amount of both adsorbed and dissolved H<sub>2</sub>O of the rock powder, which can vary between charges and cannot be evaluated rigorously. For instance, the *f*O<sub>2</sub>

of nominally dry charges calculated using equation (1) assuming a fluid phase composition of  $X_{H_2O} = 0.01$  is around NNO for dry charges at low  $f_{H_2}$ , or 5 log units below it (NNO - 5) for dry charges at  $f_{H_2} = 20\text{--}30$  bar. The same calculation with  $X_{H_2O} = 0.1$  yields  $f_{O_2}$  of NNO + 2 to NNO - 3, respectively. This shows that, if extremely dry, runs performed at low  $f_{H_2}$  may have  $f_{O_2}$  values approaching those of  $H_2O$ -saturated charges at high  $f_{H_2}$ .

For experiments carried out with CSPV- $H_2$  at  $H_2O$  saturation, the calculated  $f_{O_2}$  values are accurate to within 0.03 log units whereas those calculated for  $H_2O$ -undersaturated charges must be seen as maxima, as equation (1) overestimates the abundance of  $H_2O$  in the fluid and neglects the departure from ideality of  $H_2O\text{--}CO_2$  fluid. For experiments performed in CSPV, the uncertainty of  $f_{O_2}$  is larger. In runs for which  $f_{H_2}$  has been determined (runs at  $\sim 800^\circ C$ ), we evaluate the uncertainty to be around  $\pm 0.6$  unit log. For runs performed in CSPV at lower temperatures, the error on our calculated  $f_{O_2}$  can be much larger, possibly approaching 1 log unit. We consider, however, that all the runs performed in CSPV have their  $f_{O_2}$  within the range NNO + 3 to NNO + 5 (excluding dry runs).

### Analytical methods

Run products were characterized by polarized light microscopy, and by electron microprobe analysis (EMPA) and scanning electron microscopy (SEM). Analytical conditions for EMPA were accelerating voltage 15 kV, sample current 6 nA, counting time 10 s on peak for all elements and a focused beam for minerals. For glasses the beam was defocused to 5  $\mu m$ . Na and K were analysed first and a ZAF correction procedure was applied.  $Na_2O$  migration under the probe beam was assessed using two sets of hydrous metaluminous and peralkaline (SMN 49) rhyolitic glass standards of known  $H_2O$  contents (Scaillet & Evans, 1999). At the beginning of each analytical session the two set of standards were analysed, the metaluminous being used for retrieving the  $Na_2O$  correction factor to be applied to composition ND 002, whereas the peralkaline set was used for calculating that specific to compositions BL 575 and SMN 49.

Melt  $H_2O$  contents ( $H_{2O_{melt}}$ ) of drop-quenched glasses at 800 and 750°C were estimated using the by-difference method (Devine *et al.*, 1995) after correction for alkali migration. The precision of melt water contents is estimated to be 0.5 wt %. Experiments performed at lower temperatures display textural evidence of fluid exsolution during quench (sub-micrometric bubbles). The melt water content of these charges was estimated using the linear correlations between  $X_{H_2O_{in}}$  and melt  $H_2O$  content obtained in drop-quenched charges at 800 and 750°C.

These correlations were also used to calculate the melt water content of near-solidus charges in which the melt composition could not be microprobe analysed. For nominally anhydrous charges the melt  $H_2O$  content was calculated assuming perfectly incompatible behaviour for both  $H_2O$  and F during crystallization and using a bulk  $H_2O$  content for the three starting rock powders of 0.4 wt %.

### Attainment of equilibrium

The successful crystallization of all the naturally occurring minerals apart from aenigmatite shows that no major nucleation problems were encountered in this study. At 800°C and 150 MPa, melting of the three  $H_2O$ -saturated samples was complete after 10 days. A few relict minerals were observed in some charges, mainly at temperatures below 750°C and at low  $X_{H_2O}$  (Table 2), but their modal abundance is <0.5%, implying that the non-reactive part of the system was minimal. Experimental work performed in dry and  $H_2O$ -saturated peralkaline systems has shown that equilibrium conditions are attained during run durations of the order of a week (Carmichael & Mackenzie, 1963; Thompson & Mackenzie, 1967; Bailey *et al.*, 1974; Bailey & Cooper, 1978). In the present work, run duration was at least 10 days at 800°C and up to 8 weeks at temperatures below 700°C (Table 2). Previous researchers carrying out phase equilibrium experiments in metaluminous or peraluminous silicic systems have concluded that bulk equilibrium was reached for run durations similar to those applied in the present study (e.g. Clemens & Wall, 1981; Pichavant, 1987; Webster *et al.*, 1987; Holtz *et al.*, 1992; Scaillet *et al.*, 1995; Dall'Agnoli *et al.*, 1999). These lines of evidence argue for the attainment of near-equilibrium conditions in the experimental charges. Probe analyses show that in some runs, contamination by Ni from the vessel occurred, with the Ni content of ferromagnesian phases of a single charge varying in an erratic way from below detection up to several weight percent, suggesting that Ni was not instrumental in crystal growth. The  $FeO_{tot}$  content of glasses of supra-liquidus charges at low  $f_{O_2}$  is  $\sim 6\%$  lower than the starting bulk composition, showing that iron loss toward the Au capsules was minimal.

## RESULTS

### Phases and textures

Phases grown in our experiments are alkali feldspar, quartz, plagioclase, hedenbergite, magnetite, ilmenite, chevkinite, biotite, fayalite, zircon, fluorite, titanite, aegirine, ferrichterite and riebeckite-arfvedsonite.

Table 2: Experimental results on Naivasha rhyolites

	<i>P</i> (bar)	<i>T</i> (°C)	<i>P</i> <sub>H<sub>2</sub></sub> (bar)	Duration (h)	XH <sub>2</sub> O <sub>m</sub> <sup>1</sup>	log <i>f</i> O <sub>2</sub> <sup>2</sup> (ΔNNO)	H <sub>2</sub> O <sup>3</sup> (wt %)	Run products <sup>4</sup>
<b>Sample ND 002</b>								
<i>150 MPa, reduced</i>								
1	1540	797	26	255	1.00	-1.19	4.87	gl, vp
2	1556	788	25	378	0.90	-1.26	4.47	gl, vp
3	1556	788	25	378	0.79	-1.37	2.82	gl, vp, af, qtz, hd, ox
4	1505	788	23	354	0.72	-1.40	2.70	gl, vp, af, qtz, hd, fay, ox
5	1505	788	23	354	0.60	-1.56	—	gl, vp, af, qtz, hd, fay ox
6	1608	738	20	617	1.00	-1.05	4.56	gl, vp, hd, ox
7	1608	738	20	617	0.90	-1.14	4.18	gl, vp, hd, af, qtz, ox
8	1480	699	18	690	1.00	-1.12	4.87	gl, vp, hd, ox, (af)
9	1480	699	18	690	0.89	-1.22	4.23	gl, vp, hd, af, qtz, ox
22	1490	788	25	306	0.50	-1.80	2.49	gl, vp, hd, af, qtz, ox
31 <sup>5</sup>	1558	693	24	1000	0.79	-1.55	—	vp, hd, af, qtz, znr, ox, richt
33	1558	693	24	1000	0.58	-1.82	2.42	gl, vp, af, qtz, bt, fay, ox
37	1505	661	30	951	1.00	-1.66	4.87	gl, vp, af, qtz, fay, bt, fl, ox
42 <sup>6</sup>	1515	758	22	959	0.10	-3.13	0.88	gl, vp, qtz, af, fay, bt, fl, ox, richt
43 <sup>6</sup>	1515	729	22	959	0.10	-3.20	1.08	gl, vp, qtz, af, fay, bt, fl, ox
44 <sup>6</sup>	1515	773	22	959	0.10	-3.10	0.96	gl, vp, qtz, af, fay, fl, ox, richt
<i>150 MPa, oxidized</i>								
12	1550	794	0.04	407	1.00	4.39	4.84	gl, vp, ox
13	1550	794	0.04	407	0.90	4.30	4.14	gl, vp, ox
14	1550	794	0.04	407	0.80	4.20	3.75	gl, vp, ox
15	1550	794	0.04	407	0.68	4.06	3.52	gl, vp, hd, af, pl, ox
16	1550	794	0.04	407	0.59	3.93	2.08	gl, vp, hd, af, qtz, ox, fl
17	1520	752	0.04	502	1.00	4.29	4.66	gl, vp, ox, chevsk
18	1520	752	0.04	502	0.89	4.19	4.22	gl, vp, hd, ox, (af), tit
19	1520	752	0.04	502	0.80	4.10	3.93	gl, vp, hd, af, ox, tit
20	1520	752	0.04	502	0.68	3.96	3.00	gl, vp, hd, af, qtz, ox, tit
21	1520	752	0.04	502	0.61	3.86	2.59	gl, vp, af, qtz, ox, fl
23	1550	696	0.04	1222	1.00	4.17	4.87	gl, vp, hd, af, ox, tit
24	1550	696	0.04	1222	0.90	4.08	4.29	gl, vp, hd, af, qtz, ox
25	1550	696	0.04	1222	0.79	3.97	3.64	gl, vp, hd, af, qtz, ox, fl
26	1550	696	0.04	1222	0.67	3.83	2.94	vp, hd, af, qtz, ox
27	1550	696	0.04	1222	0.59	3.72	2.47	vp, hd, af, qtz, ox
38 <sup>6</sup>	1500	784	0.05	1138	0.10	2.35	1.04	gl, vp, af, qtz, ox, fl
<i>50 MPa, reduced</i>								
10	503	794	14	423	1.00	-1.48	3.72	gl, vp, hd, ox
11	503	794	14	423	0.80	-1.67	2.81	gl, vp, hd, af, ox
28	520	790	15	617	0.70	-1.82	2.19	gl, vp, hd, af, ox
29	520	790	15	617	0.61	-1.94	3.40	gl, vp, hd, af, qtz, ox, ap
30	520	790	15	617	0.50	-2.11	2.50	gl, vp, hd, af, qtz, ox
34	566	756	13	867	1.00	-1.36	2.58	gl, vp, hd, af, ox
35	566	756	13	867	0.81	-1.54	3.11	gl, vp, hd, af, ox, qtz
36	566	756	13	867	0.63	-1.76	—	vp, af, qtz, ox
40	524	688	20	1138	0.82	-2.08	—	vp, af, qtz, bt, rbk-arf, fl
41	524	688	20	1138	0.63	-2.31	—	vp, af, qtz, ox
<b>Sample BL 575</b>								
<i>150 MPa, reduced</i>								
1	1540	798	26	255	1.00	-1.19	5.79	gl, vp
2	1556	798	25	378	0.91	-1.23	5.38	gl, vp
3	1556	798	25	378	0.81	-1.33	4.16	gl, vp
4	1505	798	23	354	0.71	-1.41	2.52	gl, vp, ox
5	1505	798	23	354	0.59	-1.54	2.79	gl, vp, ox
6	1608	746	20	617	1.00	-1.03	4.75	gl, vp, ox
7	1608	746	20	617	0.90	-1.12	4.67	gl, vp, ox, (qtz)
8	1480	700	18	690	1.00	-1.12	5.79	gl, vp, ox
9	1480	700	18	690	0.91	-1.20	5.07	gl, vp, ox
22	1490	788	25	306	0.47	-1.86	2.45	gl, vp, ox, (qtz)
23	1490	791	25	306	0.37	-2.06	0.77	gl, vp, qtz, af, ox
24	1490	791	25	306	0.28	-2.30	0.01	gl, vp, qtz, af

	<i>P</i> (bar)	<i>T</i> (°C)	<i>P</i> <sub>H<sub>2</sub></sub> (bar)	Duration (h)	XH <sub>2</sub> O <sub>in</sub> <sup>1</sup>	log <i>f</i> O <sub>2</sub> <sup>2</sup> (ΔNNO)	H <sub>2</sub> O <sup>3</sup> (wt %)	Run products <sup>4</sup>
33	1558	676	24	1000	0.78	-1.61	4.03	gl, vp, qtz, af, rbk-arf, chev
34	1558	676	24	1000	0.65	-1.77	3.00	gl, vp, qtz, af, rbk-arf, fl
35	1558	676	24	1000	0.51	-1.98	1.88	gl, vp, qtz, af, rbk-arf, fl
47	1505	661	30	951	1.00	-1.66	5.79	gl, vp, rbk-arf, bt, ox, fl, (qtz)
48	1505	702	30	951	0.63	-1.94	2.83	gl, vp, qtz, af, rbk-arf, fl
49	1505	702	30	951	0.45	-2.23	1.41	gl, vp, qtz, af, rbk-arf
50	1505	702	30	951	0.31	-2.56	0.29	gl, vp, qtz, af, rbk-arf
56 <sup>6</sup>	1515	758	22	959	0.10	-3.13	0.80	gl, vp, qtz, af, fl
57 <sup>6</sup>	1515	729	22	959	0.10	-3.20	1.28	gl, vp, qtz, af, rbk-arf, fl
58 <sup>6</sup>	1515	773	22	959	0.10	-3.10	0.67	gl, vp, qtz, af
<i>150 MPa, oxidized</i>								
12	1550	805	0.11	407	1.00	3.61	5.59	gl, vp, ox
13	1550	805	0.11	407	0.87	3.49	4.94	gl, vp, ox
14	1550	805	0.11	407	0.79	3.40	4.36	gl, vp, ox
15	1550	805	0.11	407	0.68	3.27	3.00	gl, vp, ox, (qtz)
16	1550	805	0.11	407	0.57	3.12	2.36	gl, vp, ox, qtz
17	1520	757	0.11	502	1.00	3.50	5.98	gl, vp, ox
18	1520	757	0.11	502	0.90	3.41	5.60	gl, vp, ox
19	1520	757	0.11	502	0.78	3.28	4.14	gl, vp, ox
20	1520	757	0.11	502	0.67	3.15	3.18	gl, vp, ox, (af)
21	1520	757	0.11	502	0.59	3.04	2.69	gl, vp, qtz, af, ox, fl
25	1550	703	0.11	1222	1.00	3.39	5.79	gl, vp, ox
26	1550	703	0.11	1222	0.87	3.27	4.75	gl, vp, ox
27	1550	703	0.11	1222	0.78	3.17	4.03	gl, vp, qtz, af, ox
28	1550	703	0.11	1222	0.67	3.04	3.16	gl, vp, qtz, af, ox
29	1550	703	0.11	1222	0.59	2.93	2.52	gl, vp, qtz, af, ox, fl
36	1550	678	0.05	955	1.00	3.97	5.79	gl, vp, ox, chev, fl
37	1550	678	0.05	955	0.91	3.89	5.07	gl, vp, ox, chev, fl, qtz, af
38	1550	678	0.05	955	0.80	3.78	4.19	gl, vp, ox, qtz, af, aeg, fl
39	1550	678	0.05	955	0.62	3.56	2.76	gl, vp, ox, qtz, af, aeg, fl
40	1550	678	0.05	955	0.46	3.30	1.48	vp, ox, fl, qtz, af, aeg, znr
41	1520	700	0.11	955	0.51	2.78	1.88	gl, vp, qtz, af, aeg, fl
42	1400	786	0.04	955	0.51	3.71	1.85	gl, vp, qtz, af, ox, fl
43	1400	786	0.04	955	0.39	3.48	1.98	gl, vp, qtz, af, ox
51 <sup>6</sup>	1500	784	0.05	1138	0.10	2.35	0.88	gl, vp, qtz, af, ox, fl
<i>50 MPa, reduced</i>								
10	563	798	14	423	1.00	-1.38	3.25	gl, vp, ox
11	563	798	14	423	0.81	-1.56	3.29	gl, vp, ox
30	520	798	15	617	0.69	-1.83	2.76	gl, vp, ox
31	520	798	15	617	0.62	-1.92	2.68	gl, vp, ox, (qtz)
32	520	797	15	617	0.49	-2.12	2.25	gl, vp, ox
44	566	737	13	867	1.00	-1.38	3.32	gl, vp, ox
45	566	737	13	867	0.79	-1.59	3.83	gl, vp, ox, qtz, af, fl
46	566	737	13	867	0.57	-1.87	2.56	gl, vp, ox, qtz, af
53	524	686	20	1138	1.00	-2.10	—	gl, vp, qtz, af, rbk-arf, ox, fl
54	524	686	20	1138	0.82	-2.27	—	gl, vp, qtz, af, rbk-arf, ox, fl
55	524	686	20	1138	0.82	-2.62	—	gl, vp, qtz, af, rbk-arf, ox
<b>Sample SMN 49</b>								
<i>150 MPa, reduced</i>								
1	1540	797	26	162	1.00	-1.19	5.73	gl, vp
2	1556	796	25	378	0.89	-1.25	3.94	gl, vp
3	1556	796	25	378	0.76	-1.39	3.39	gl, vp
4	1505	796	23	354	0.71	-1.41	2.85	gl, vp, qtz
5	1505	796	23	354	0.55	-1.55	2.24	gl, vp, qtz, af
6	1608	744	20	617	1.00	-1.03	5.29	gl, vp
7	1608	744	20	617	0.90	-1.13	3.52	gl, vp, qtz
8	1480	703	18	690	1.00	-1.11	5.73	gl, vp
9	1480	703	18	690	0.86	-1.20	5.19	gl, vp, qtz
22	1490	796	25	306	0.50	-1.79	2.99	gl, vp, qtz



Table 2: continued

	<i>P</i> (bar)	<i>T</i> (°C)	<i>P</i> <sub>H<sub>2</sub></sub> (bar)	Duration (h)	XH <sub>2</sub> O <sub>in</sub> <sup>1</sup>	log <i>f</i> O <sub>2</sub> <sup>2</sup> (ΔNNO)	H <sub>2</sub> O <sup>3</sup> (wt %)	Run products <sup>4</sup>
23	1490	796	25	306	0.40	-1.98	2.45	gl, vp, qtz, af
32	1558	678	24	1000	0.79	-1.59	4.58	gl, vp, qtz, af, rbk-arf, chevk
33	1558	678	24	1000	0.69	-1.71	4.04	gl, vp, qtz, af, rbk-arf, fl, mondt
34	1558	678	24	1000	0.50	-1.99	2.99	gl, vp, qtz, af, rbk-arf, fl
44	1505	661	30	951	1.00	-1.66	—	gl, vp, bt, fl, chevk
45	1505	702	30	951	0.62	-1.96	3.65	gl, vp, qtz, af, ox
46	1505	702	30	951	0.50	-2.14	2.99	gl, vp, qtz, af, ox
47	1505	702	30	951	0.27	-2.68	1.73	gl, vp, qtz, af, rbk-arf, ox
58 <sup>6</sup>	1515	758	22	959	0.10	-3.13	1.12	gl, vp, qtz, af, fl
59 <sup>6</sup>	1515	729	22	959	0.10	-3.20	1.42	gl, vp, qtz, af, rbk-arf, fl
60 <sup>6</sup>	1515	773	22	959	0.10	-3.10	0.96	gl, vp, qtz, af
<i>150 MPa, oxidized</i>								
12	1550	775	0.05	407	1.00	4.17	5.93	gl, vp
13	1550	775	0.05	407	0.89	4.07	3.56	gl, vp
14	1550	775	0.05	407	0.75	3.92	3.75	gl, vp
15	1550	775	0.05	407	0.67	3.82	3.52	gl, vp
16	1550	775	0.05	407	0.54	3.64	2.77	gl, vp, qtz, af, chevk
17	1520	725	0.05	502	1.00	4.05	5.68	gl, vp, ox
18	1520	725	0.05	502	0.88	3.94	4.91	gl, vp, ox, chevk
19	1520	725	0.05	502	0.79	3.85	4.11	gl, vp, qtz, ox
20	1520	725	0.05	502	0.68	3.72	4.22	gl, vp, qtz, fl, ox
21	1520	725	0.05	502	0.63	3.65	3.66	gl, vp, qtz, af, ox
24	1550	679	0.05	1222	1.00	3.94	5.73	gl, vp, ox, chevk
25	1550	679	0.05	1222	0.87	3.82	5.02	gl, vp, qtz, af, aeg, ox
26	1550	679	0.05	1222	0.78	3.72	4.53	gl, vp, qtz, af, aeg, ox
27	1550	679	0.05	1222	0.67	3.59	3.93	gl, vp, qtz, af, aeg, ox, fl
28	1550	679	0.05	1222	0.59	3.48	3.49	gl, vp, qtz, af, aeg, ox
35	1520	700	0.11	955	0.78	3.12	4.91	gl, vp, qtz, af, aeg
36	1520	700	0.11	955	0.66	2.97	3.99	gl, vp, qtz, af, aeg, fl
37	1520	700	0.11	955	0.55	2.87	3.27	gl, vp, qtz, af, aeg
38	1520	700	0.11	955	0.50	2.73	2.99	gl, vp, qtz, af, aeg
39	1400	786	0.05	955	0.39	3.29	2.78	gl, vp, qtz, af
40	1400	786	0.05	955	0.23	2.83	2.41	gl, vp, qtz, af, fl
48 <sup>6</sup>	1500	784	0.05	1138	0.10	2.16	1.45	gl, vp, qtz, af, aeg, ox, fl
50	1500	731	0.05	598	1.00	4.09	5.68	gl, vp, ox
51 <sup>7</sup>	1500	731	0.05	598	fa	—	4.98	gl, vp, ox
52 <sup>7</sup>	1500	731	0.05	598	fa	—	4.36	gl, vp, ox, qtz, (af)
53 <sup>7</sup>	1500	731	0.05	598	fa	—	3.28	gl, vp, ox, qtz, af, fl
54 <sup>6</sup>	1500	731	0.05	598	0.01	0.09	1.64	gl, vp, qtz, af, rbk-arf, fl
<i>50 MPa, reduced</i>								
10	563	802	14	423	1.00	-1.37	2.49	gl, vp, (fl), (bt)
11	563	802	14	423	0.80	-1.57	2.52	gl, vp, (fl)
29	520	798	15	617	0.64	-1.89	2.77	gl, vp
30	520	797	15	617	0.56	-2.01	2.70	gl, vp
31	520	797	15	617	0.50	-2.11	3.06	gl, vp
41	566	732	13	867	1.00	-1.39	3.32	gl, vp
42	566	732	13	867	0.82	-1.56	3.42	gl, vp, qtz, af
43	566	732	13	867	0.57	-1.88	3.02	gl, vp, qtz, af, fl
55	524	679	20	1138	1.00	-1.92	—	gl, vp, qtz, af, fl
56	524	679	20	1138	0.83	-2.08	—	gl, vp, qtz, af, rbk-arf
57	524	679	20	1138	0.55	-2.44	—	gl, vp, qtz, af, rbk-arf

<sup>1</sup>XH<sub>2</sub>O<sub>in</sub> = H<sub>2</sub>O/(H<sub>2</sub>O + CO<sub>2</sub>) loaded in the capsule (in moles). For nominally anhydrous experiments, the final fluid phase composition is arbitrarily set at XH<sub>2</sub>O = 0.1 except for charge 54 of composition SMN 49.

<sup>2</sup>ΔNNO = log *f*O<sub>2</sub>(experiment) - log *f*O<sub>2</sub> (NNO; Pownceby & O'Neill, 1994).

<sup>3</sup>Water contents of glasses determined using the by-difference method. Numbers in italics are calculated as explained in the text.

<sup>4</sup>Phase abbreviations as given by Kretz (1983): gl, glass; vp, fluid; af, alkali feldspar; qtz, quartz; hd: ferrohedenbergite; fay, fayalite; ox, oxides; fl, fluorite; chevk, chevkinitite; tit, titanite; richt, ferrorichterite; rbk-arf, riebeckite-arfvedsonite solid solution; aeg, aegirine; mondt, montdorite; ap, apatite. Parentheses are for relict phases.

<sup>5</sup>Leaked capsule.

<sup>6</sup>Nominally anhydrous charge.

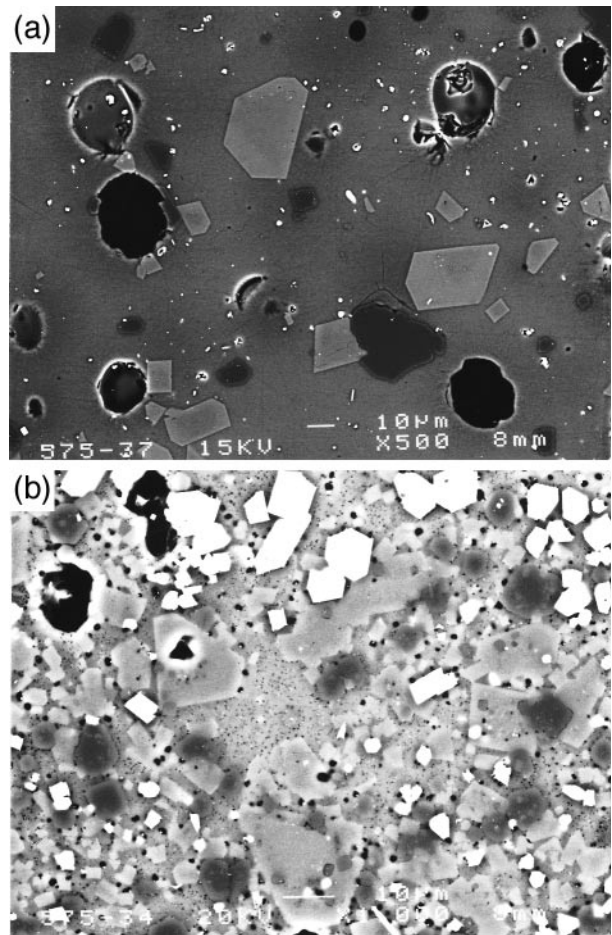
<sup>7</sup>Fluid-absent charge.

Montdorite, a tetrasilicic mica (Robert & Maury, 1979), was identified in one charge from SMN 49 (Table 2). The only phase in the natural samples not identified in this study is aenigmatite.

Minerals display euhedral shapes with sizes of 5–30  $\mu\text{m}$ , except for the Na-amphibole and aegirine, which crystallize as stubby prisms over 100  $\mu\text{m}$  long, and for oxides whose size rarely exceeds 5  $\mu\text{m}$ . Close to the liquidus, crystallization is marked by the appearance of a few large individuals (Fig. 1a), whereas at lower temperatures, nucleation is more efficient than growth, giving rise to a large number of smaller crystals (Fig. 1b). Although relict phases are very rare (<0.5% by volume of the crystal), they are clearly distinguished on the basis of their larger size (>100  $\mu\text{m}$ ), their rounded shape (alkali feldspar, quartz), dissolution textures at their borders (alkali feldspar), a reaction zone surrounding the crystal (hedenbergite) or the presence of a rim compositionally different from the core. In the last texture, the rim in all cases has the same composition as small euhedral crystals that grew during the experiment. Rounded relicts of tectosilicates are observed only in cases where the charge is outside, yet close to, the stability field of the corresponding phase. All charges, including those nominally anhydrous, display evidence of fluid saturation in the form of evenly dispersed bubbles with sizes varying between 10 and 100  $\mu\text{m}$ , some containing cubic minerals. This population of equilibrium bubbles is clearly distinguishable from the micrometric to sub-micrometric bubbles present in charges that experienced slow quench rates (Fig. 1b). The reason for nominally anhydrous melts to be fluid saturated may be the presence of Cl and F, elements that promote early fluid saturation compared with the purely hydrous condition (e.g. Webster, 1992, 1997).

### Phase relationships

Phase relationships for the three compositions are shown in Figs 2–4. At 150 MPa, phase relations are shown in  $T$ –melt  $\text{H}_2\text{O}$  content projections. This type of projection could not be established at 50 MPa and only the  $T$ – $X\text{H}_2\text{O}_{\text{in}}$  projection is shown. No attempt was made to establish the stability fields of zircon or apatite. The small size prevented in some cases clear identification of the type of Fe–Ti oxide. As a rule, magnetite grew in ND 002 whereas ilmenite or a Ti-bearing haematite crystallized in both BL 575 and SMN 49. Ilmenite is present, however, in some run products of ND 002, especially those whose  $f\text{O}_2$  is significantly below  $\text{NNO} - 1$ . The nominally anhydrous runs performed under high and low  $f\text{H}_2$  are also plotted. Despite the fact that their  $f\text{O}_2$  may have been significantly below the  $f\text{O}_2$  of  $\text{H}_2\text{O}$ – $\text{CO}_2$ -bearing charges held under similar  $f\text{H}_2$  conditions, both sets of

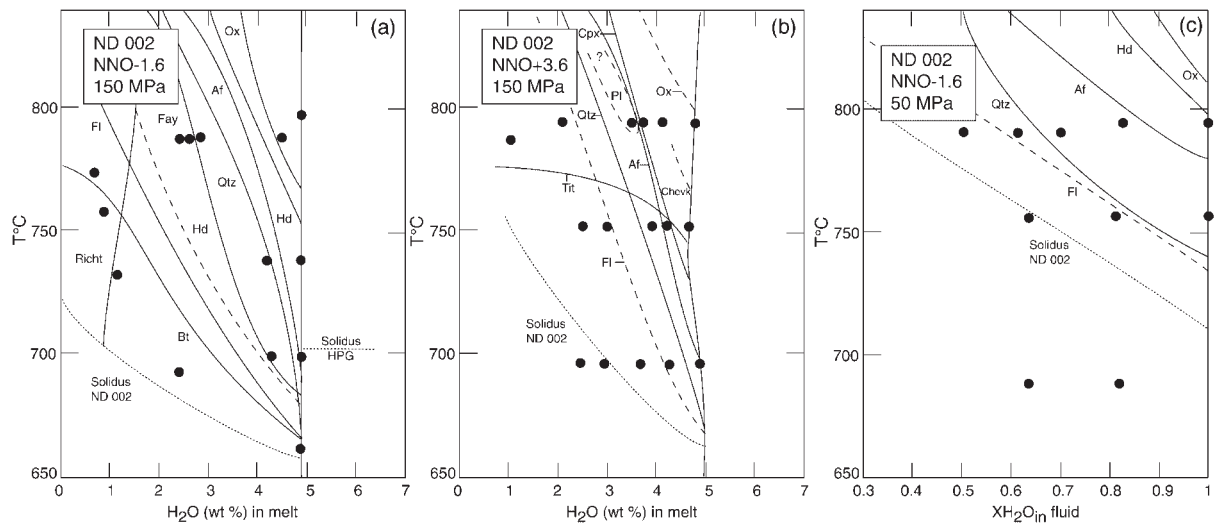


**Fig. 1.** Scanning electron micrographs of run products. In both pictures quartz has a dark grey to black shade compared with alkali feldspar. (a) Coexisting quartz and alkali feldspar crystals and glass in a near-liquidus charge of BL 575. Bright minerals are oxides. (b) Quartz, alkali feldspar and riebeckite–arfvedsonite (bright equant grains) and glass assemblage. (Note the presence of micron to sub-micron sized bubbles scattered throughout the charge, interpreted to occur during the quench.) The 10  $\mu\text{m}$  large void in the upper left corner represents a former equilibrium fluid bubble.

charges display consistent phase assemblages, except for charge 54 in composition SMN 49 whose results conflict with the remainder. This point is discussed further below.

### ND 002

At 150 MPa under reduced conditions ( $\text{NNO} - 1.6$ ), magnetite and ferrohedenbergite are the liquidus phases, closely followed by alkali feldspar, quartz and then fayalite (Fig. 2a). At  $\text{H}_2\text{O}$  saturation, the crystallization temperatures of oxide and ferrohedenbergite are estimated to be close to 760°C, whereas alkali feldspar precipitates near 700°C, followed by quartz near 660°C. At melt  $\text{H}_2\text{O}$  content lower than 2–3 wt % and at temperatures higher than 700°C, hedenbergite is no longer stable and the main stable ferromagnesian silicate is fayalite,



**Fig. 2.** Phase relationships of ND 002. Mineral abbreviations are as recommended by Kretz (1983). The curves are labelled with minerals lying inside their respective stability fields, and are shown dashed when inferred. (a) 150 MPa and NNO - 1.6; (b) 150 MPa and NNO + 3.6; (c) 50 MPa and NNO - 1.6. Solidus HPG is the solidus of the subaluminous haplogranite system at 150 MPa (Johannes & Holtz, 1996).

coexisting with ilmenite. At low  $H_2O_{melt}$  (<2 wt %) and temperatures above 730°C, a ferro-richteritic amphibole is present. Below 700°C amphibole is present only in a subsolidus charge (charge 31, Table 2), and the amphibole-out curve has a rather steep slope. In the dry charge at 729°C (charge 44, Table 2), a former hedenbergite phenocryst was found rimmed by fayalite, ilmenite and fluorite, suggesting that these phases replace pyroxene at low  $H_2O_{melt}$ . Finally, the scarce occurrence or absence of hedenbergite in charges at temperatures below 700°C suggests that its stability field may not extend down to the solidus in ND 002 in the water-rich part of the diagram. Thus, we have drawn a hedenbergite-out curve that is crossed either as temperature falls or as melt  $H_2O$  decreases. Biotite was identified only below 700°C at  $H_2O$  saturation, or in dry charges at higher temperatures, which suggests a restricted stability for  $fO_2$  at or below NNO - 1.6, close to the solidus. Similarly, fluorite is present only in strongly crystallized charges or below 700°C at  $H_2O$  saturation and its stability field lies close to the solidus. Although not looked for systematically, chevkinite was not found. No attempt was made to locate the solidus precisely, but it is below 660°C at  $H_2O$  saturation (Table 2), which is 40°C lower than the solidus of the haplogranite system at 150 MPa (Johannes & Holtz, 1996). Melt  $H_2O$  contents at saturation are slightly lower than 5 wt %, being similar to those found in the metaluminous haplogranite system (Johannes & Holtz, 1996) at similar  $P$  and  $T$ .

Increased  $fO_2$  does not alter the order of crystallization of the major phases (Fig. 2b) but it yields, at  $H_2O$  saturation, magnetite crystallization above 800°C, whereas clinopyroxene saturation occurs slightly below

750°C. Although the liquidus of the alkali feldspar at  $H_2O$  saturation may be higher by  $\sim 10^\circ\text{C}$ , over most of the investigated domain and within the error in the melt  $H_2O$  determination, the crystallization temperatures of the tectosilicates remain unchanged relative to that at low  $fO_2$ . Biotite, fayalite and amphibole were not identified. In contrast, titanite is present in most of the 750°C charges and possibly in all the 700°C charges, and the titanite-in curve is drawn flat at temperatures close to 750°C. As at low  $fO_2$ , fluorite is present in some crystal-rich charges close to the solidus. Although data are lacking, its stability domain extends to higher temperatures, as seen under low  $fO_2$ . Chevkinite was identified in only one charge at  $H_2O$  saturation and 752°C, and its stability field is shown dashed with a shape similar to that seen in BL 575 and SMN 49 at high  $fO_2$ . The solidus seems to be at significantly higher temperature, by  $\sim 30^\circ\text{C}$ , than under reduced conditions. This is apparent when comparison is made between charge 33, which contains at least 10% residual glass by volume as estimated from scanning electron micrographs, and charge 27, which has no clear textural evidence of residual glass, despite the fact that both charges have similar  $XH_2O_{in}$  values and were run at almost identical temperatures. Finally, plagioclase was identified (by EMPA) in charge 15 run at  $\sim 790^\circ\text{C}$  under relatively dry conditions. This is the only instance of plagioclase crystallization in this study. Charges bracketing charge 15 do not contain plagioclase and it does not appear at lower temperature, implying a closing of the plagioclase stability field, if any, at low temperature.

Decreasing the pressure to 50 MPa under low  $fO_2$  does not affect the relative order of crystallization, although

fayalite has not been identified (Fig. 2c). The liquidus temperatures of all phases at  $H_2O$  saturation are significantly raised, by  $\sim 60^\circ C$ , for the tectosilicates. Alkali feldspar-in and quartz-in curves are more clearly separated from each other, suggesting that these two phases are not equally sensitive to changes in pressure. Biotite and Na-amphibole have been identified only in some subsolidus charges at  $\sim 690^\circ C$  (Table 2), but their stability in the presence of liquid cannot be ruled out. The solidus at  $H_2O$  saturation is probably close to  $700^\circ C$ , that is,  $60^\circ C$  lower than that of the haplogranite system at 50 MPa. Melt  $H_2O$  contents at saturation range from 2.6 to 3.7 wt %, although the lower figure is probably correct (Table 2).

#### BL 575

This composition is a typical comendite with higher  $FeO_{total}$  and  $Na_2O$  contents than ND 002. At 150 MPa under reduced conditions, an oxide is the liquidus phase, crystallizing at  $750^\circ C$  and  $H_2O$  saturation (Fig. 3a). It is followed by alkali feldspar and quartz, crystallizing simultaneously over the  $T-H_2O_{melt}$  range explored. Their temperature of appearance at  $H_2O$  saturation is slightly higher than  $650^\circ C$ . Fluorite was found mainly in low-temperature charges, where it usually crystallizes at lower temperature than amphibole, except at  $H_2O$  saturation. Under dry conditions, the maximum thermal stability of fluorite does not exceed  $775^\circ C$ . At  $H_2O_{melt} > 3$  wt %, Na-amphibole crystallization is restricted to temperatures below  $700^\circ C$ . Its crystallization at  $H_2O$  saturation occurs slightly above  $660^\circ C$ . Under nominally dry conditions its thermal stability extends to near  $740^\circ C$ . Biotite is present only at the lowest temperature investigated and at  $H_2O$  saturation, where it is a near-liquidus phase together with amphibole and fluorite. Its liquidus curve displays a positive slope in the  $T-H_2O_{melt}$  projection, but the exact shape of this boundary is not yet well constrained. The chevkinite stability field seems to be restricted to low melt  $H_2O$  contents (Table 2), but is also poorly constrained. Melt  $H_2O$  contents at saturation are slightly lower than 6 wt %, being 1 wt % higher than those of ND 002. The solidus has not been located but is definitely below  $650^\circ C$  at  $H_2O$  saturation, possibly close to, or even below,  $600^\circ C$ .

Increasing  $fO_2$  by 5 log units increases the thermal stability of Fe-Ti oxide (Fig. 3b). Below  $750^\circ C$ , quartz and alkali feldspar appear to be insensitive to variations in  $fO_2$ . Fluorite crystallizes at  $\sim 680^\circ C$  at water saturation, and its liquidus roughly parallels that of the tectosilicates over the temperature range investigated. In contrast to low  $fO_2$ , its thermal stability seems largely to exceed  $800^\circ C$ . Chevkinite may have a slightly wider stability field relative to low- $fO_2$  conditions. Under  $H_2O$  saturation it is the second phase to precipitate after oxide

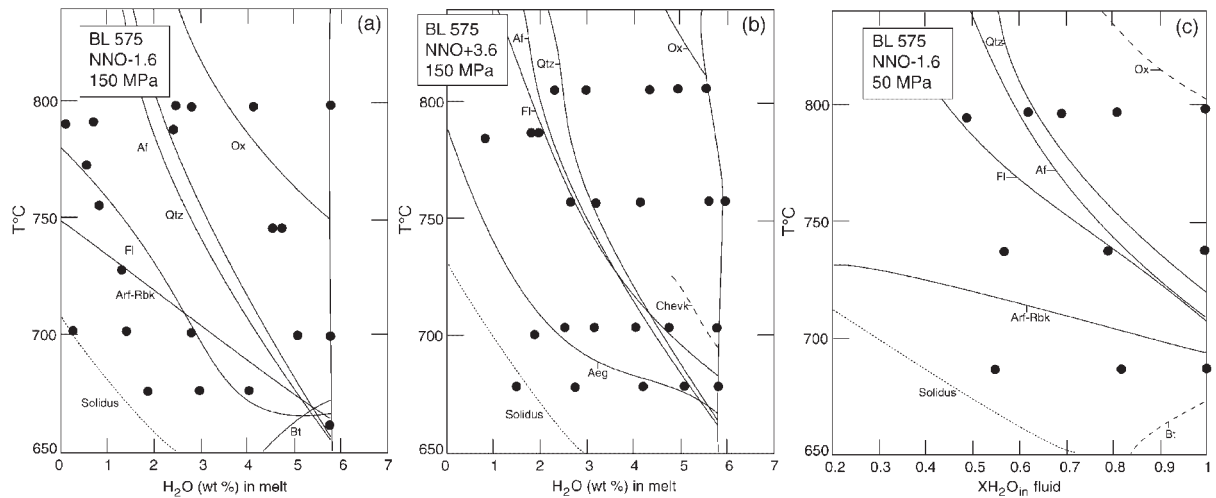
at  $\sim 690^\circ C$ . Its thermal stability when  $H_2O_{melt}$  decreases is again not well constrained. Both Na-amphibole and biotite are absent at high  $fO_2$ . The former is replaced by aegirine whose thermal stability also increases as  $H_2O_{melt}$  decreases. Solidus conditions were approached at  $678^\circ C$  in charge 40 ( $X_{H_2O_{in}} = 0.46$ ); no glass is seen on SEM images. Comparison with results at low  $fO_2$  (charge 35 at  $676^\circ C$  and  $X_{H_2O_{in}} 0.51$  has  $>20\%$  glass in volume) again suggests that a rise in  $fO_2$  produces a rise in solidus temperatures, as in ND 002. Melt  $H_2O$  contents required for quartz and alkali feldspar saturation at any given temperature are identical to those at low  $fO_2$ .

Decreasing the pressure to 50 MPa raises the crystallization temperatures by  $\sim 50-60^\circ C$  (Fig. 3c), as for ND 002. Quartz appears before alkali feldspar but the less extensive coverage of  $T-H_2O_{melt}$  data prevents accurate location of the phase boundaries. Melt  $H_2O$  contents at saturation are in the range 3.2–3.5 wt % (Table 2). At  $H_2O$  saturation, the Na-amphibole stability field is higher by  $\sim 30^\circ C$  compared with that at 150 MPa. Biotite was not identified; if stable, it must crystallize below  $680^\circ C$ , as tentatively suggested in Fig. 3c.

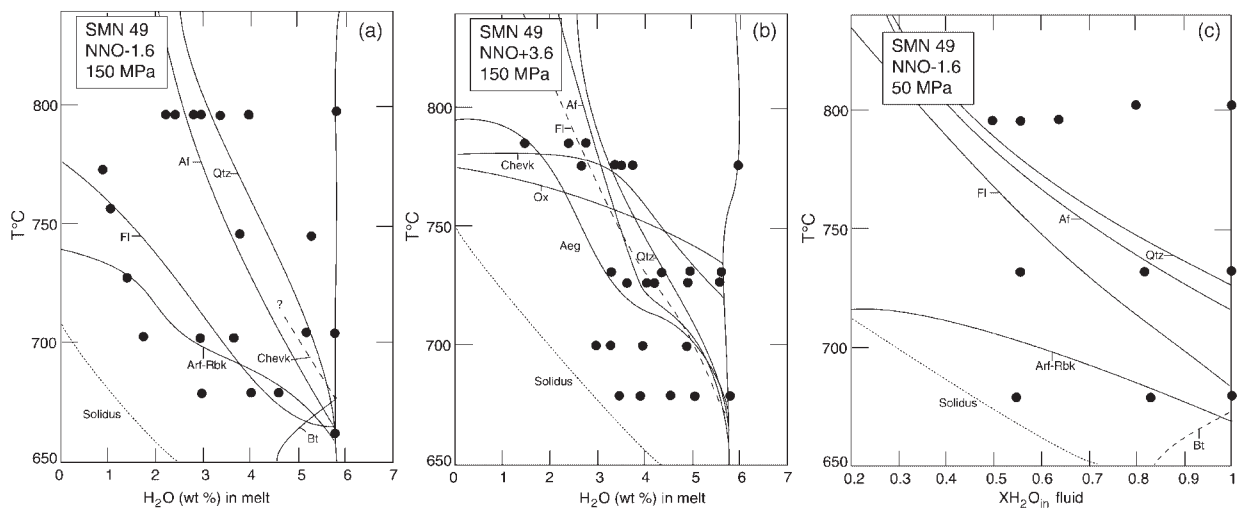
#### SMN 49

This composition is close to BL 575, differing only in a slightly higher  $SiO_2$  content (Table 1). The phase relations of SMN 49 are thus essentially the same as those of BL 575 (Fig. 4), except that at low  $fO_2$  no stability field for Fe-Ti oxides could be established. The liquidus curves of quartz and alkali feldspar are also more clearly separated from each other. Montdorite was recognized (by EMPA) in only one low- $fO_2$  and near-solidus charge (33, Table 2).

At high  $fO_2$  under  $H_2O$ -saturated conditions, Fe-Ti oxide crystallizes at slightly over  $730^\circ C$ , followed by chevkinite at  $\sim 720^\circ C$ . The thermal stability of chevkinite is significantly larger than that in BL 575. The dry charge at  $731^\circ C$  (54) yielded Na-amphibole and no pyroxene (Table 2). This difference is interpreted as due to a significantly lower  $fO_2$  relative to the  $H_2O-CO_2$ -bearing charges, possibly resulting from less adsorbed water in the powder loaded to charge 54. This charge has been ignored in establishing the phase relations at high  $fO_2$ . As for BL 575, a rise in  $fO_2$  produces a marked rise in the solidus temperature, best seen when comparing charges 38 and 46, which have the same  $T-H_2O_{melt}$  location but were run under different  $fO_2$  conditions (Table 2). Whereas charge 38 (high  $fO_2$ ) shows a minimal amount of residual glass in scanning electron micrographs, charge 46 (low  $fO_2$ ) displays a much larger fraction of residual glass. A further feature of interest is that the crystallization sequence produced at  $\sim 730^\circ C$  with the  $H_2O$ -undersaturated but  $CO_2$ -free charges



**Fig. 3.** Phase relationships of BL 575. Mineral abbreviations are as recommended by Kretz (1983). The curves are labelled with minerals lying inside their respective stability fields, and are shown dashed when inferred. (a) 150 MPa and NNO - 1.6; (b) 150 MPa and NNO + 3.6; (c) 50 MPa and NNO - 1.6.



**Fig. 4.** Phase relationships of SMN 49. Mineral abbreviations are as recommended by Kretz (1983). The curves are labelled with minerals lying inside their respective stability fields, and are shown dashed when inferred. (a) 150 MPa and NNO - 1.6; (b) 150 MPa and NNO + 3.6; (c) 50 MPa and NNO - 1.6.

(charges 50–53) is virtually identical, both in terms of phase assemblages and  $T$ -melt  $H_2O$  coordinates, to that obtained in  $CO_2$ -bearing charges (charges 17–21), or within the analytical error of our melt  $H_2O$  content determination technique.

At 50 MPa, at  $H_2O$  saturation, the crystallization of both quartz and alkali feldspar occurs around 725°C (Fig. 4c), or 60°C higher than at 150 MPa. The fluorite-in curve possibly has the same relative position as at 150 MPa. Melt  $H_2O$  contents at saturation are identical to those determined for BL 575 (Table 2). Amphibole stability is slightly expanded towards higher temperature.

Biotite is not present under the experimental conditions investigated at 50 MPa. If this mineral crystallizes in SMN 49 it must do so below 680°C.

## DISCUSSION

### Aenigmatite stability

Our phase equilibrium experiments successfully reproduced the phase assemblages of the rocks, apart from aenigmatite. Several factors may have contributed to this. The low modal abundance of aenigmatite in the rock

suggests that this phase may also have crystallized in such low amounts in the experiments that it was overlooked in the run products. However, phases present in equally low amounts were identified, e.g. biotite and fayalite. Previous experimental studies do not mention difficulties in growing aenigmatite under laboratory conditions similar to ours (Ernst, 1962; Thompson & Chisholm, 1969; Lindsley 1971). Rather, an inappropriate choice of one or more of the intensive parameters may have prevented aenigmatite crystallization. The experimental work of Ernst (1962) on a synthetic Na–Fe end-member amphibole has shown that Ti-free aenigmatite is unstable at pressures >90 MPa. Thompson & Chisholm (1969) synthesized Ti-bearing stoichiometric aenigmatite at 100 MPa, which suggests that Ti enhances the pressure stability of this phase. Although there are currently no data on aenigmatite stability at pressures >100 MPa, its stability may be pressure dependent, possibly explaining its absence in our experiments. On the other hand, Ernst (1962) and Lindsley (1971) showed that the occurrence of aenigmatite also depends on redox conditions, being enhanced at low  $fO_2$ . Ti-free aenigmatite is unstable at  $fO_2$  above that of the wüstite–magnetite solid buffer whereas the Ti-bearing variety is stable to an  $fO_2$  between NNO and FMQ. Although our lowest  $fO_2$  is in the correct range for stoichiometric Ti-bearing aenigmatite to crystallize, a small departure of the natural aenigmatite composition from the theoretical one of Lindsley (1971) may affect the  $fO_2$  stability of that phase in the rock. It is therefore possible that our imposed  $fO_2$  may not have been sufficiently low to stabilize aenigmatite, even at the lowest pressure investigated here.

### Comparison between ND 002, BL 575 and SMN 49

At 150 MPa, liquidus temperatures are higher in ND 002, consistent with the fact that it represents the least differentiated member of the GOVC rhyolite suite. However, comparing only the tectosilicates, which are >90% of the crystallizing assemblage, the liquidus temperature of ND 002 is only 30°C higher than those of BL 575 and SMN 49. The phase relations established at 50 MPa suggest that the positions of tectosilicate liquids are more sensitive to composition at decreasing pressures, this being most apparent for the liquidus temperature of alkali feldspar in ND 002, which is significantly higher than in the two more peralkaline compositions. As regards crystallization sequence, the liquidus tectosilicate in ND 002 is alkali feldspar, cotectic quartz–alkali feldspar crystallization is observed in BL 575, and quartz crystallizes before alkali feldspar in SMN 49. Although this may partly be due to anhydrous compositional variations, the influence of other volatiles is possible. The fact that  $CO_2$ -free and  $CO_2$ -bearing charges yield essentially the same

results in SMN 49 shows that this volatile has a limited solubility in peralkaline silicic melts and that it mainly acts as a diluting agent in the fluid phase, as documented for other silicic magmas at low pressures (e.g. Blank *et al.*, 1993). Differences in  $CO_2$  solubility cannot therefore explain the observed differences in liquidus minerals. In contrast, fluorine expands the primary volume of quartz in the haplogranite system (Manning, 1981). The combined effects of the slightly more silicic composition of SMN 49 compared with BL 575 and the fluorine complexing effect in aluminosilicate melts may explain why the quartz-in curve is more clearly separated from that of alkali feldspar in SMN 49 than in BL 575.

Montdorite has been identified in SMN 49 only but may be present in BL 575 as well. The stability domain of this phase needs refinement but our data agree with published occurrences of montdorite in magmatic rocks (e.g. Robert & Maury, 1979). Montdorite was first described in a comendite sill (Robert & Maury, 1979), where it occurs in the groundmass along with manganian arfvedsonite, quartz, anorthoclase, oxides and fluorite among other accessory phases.  $T$ – $fO_2$  conditions of crystallization obtained from two-oxide geothermometry were 760°C and NNO, pointing to low temperature and moderately reduced  $fO_2$  during montdorite crystallization. This estimate, together with the textural evidence and our experimental data, all suggest that montdorite occurs only late in the crystallization sequence of comendite magmas. On the basis of the relationships in SMN 49, we suggest that montdorite may replace biotite as the stable ferromagnesian potassic phase, when the residual liquid becomes strongly peralkaline.

It is well known that substitution of fluorine for hydroxyl groups in hydrous minerals such as amphibole can substantially raise their thermal stability, as shown, for instance, by Holloway & Ford (1975) for pargasite. The Na-amphibole produced in our work has a F content of ~2 wt % at  $H_2O$  saturation. Despite this, the thermal stability of this phase in comendite magma at  $H_2O$  saturation does not exceed 670°C, which is lower than the upper stability limit of pure OH-riebeckite in  $H_2O$  fluid (Ernst, 1962).

The fluxing effect of the volatiles is well illustrated using the phase relationships obtained on ND 002, which most closely approached solidus conditions in this study. In Fig. 2, the solidus of the haplogranite system at 150 MPa at  $H_2O$  saturation (Johannes & Holtz, 1996) is shown in the ND 002 phase diagram obtained at 150 MPa. The difference between the two solidi is at least 40°C. The reason for such a low solidus temperature in ND 002 is the presence of additional components relative to the metaluminous base composition of the haplogranite system, most notably Fe, F and Cl. Previous experimental studies have shown that these components depress solidus temperatures to some extent in granitic

systems (Wyllie & Tuttle, 1961; Abbott, 1981; Manning, 1981; Pichavant *et al.*, 1987; Webster *et al.*, 1987; Weidner & Martin, 1987).

### Comparison with previous work

Only a few experimental studies have been performed on natural peralkaline silicic rocks, notably by Bailey and co-workers, who determined the liquidus relations, under both hydrous and anhydrous conditions, of several pantelleritic obsidians (Bailey & Cooper, 1978). The redox conditions of their H<sub>2</sub>O-saturated experiments were imposed by the vessel, and were presumably relatively oxidizing, above NNO + 1. The more peralkaline nature of their starting product, in combination with our three starting materials, allows us to establish the effect of peralkalinity on the near-liquidus phase relations of peralkaline magmas at H<sub>2</sub>O saturation and high *f*O<sub>2</sub>. In pantellerites at 150 MPa, quartz crystallizes at 650°C, whereas alkali feldspar is the liquidus tectosilicate appearing at ~690°C. Comparison with BL 575 and SMN 49 at high *f*O<sub>2</sub> (Figs 3b and 4b) shows that the alkali feldspar stability field is slightly expanded in pantellerites relative to comendites, whereas quartz preserves its position. In ND 002, the least peralkaline composition, quartz and alkali feldspar have liquidus temperatures equal to, or only marginally higher than, those in the other comendites and pantellerites. In pantellerites, the liquidus phase at high *f*O<sub>2</sub> is an acmitic pyroxene. It crystallizes at 760°C or ~100°C higher than in comendites at H<sub>2</sub>O saturation (Figs 3 and 4).

Bailey & Cooper (1978) did not perform H<sub>2</sub>O-saturated and low-*f*O<sub>2</sub> experiments, which prevents direct comparison with our results obtained at low *f*O<sub>2</sub>. Their low-*f*O<sub>2</sub> experiments (thought to be around FMQ) were nominally anhydrous. However, a qualitative agreement in the phase assemblage is observed between the two datasets. In pantellerites, aegirine is absent (but a sodic hedenbergite is reported) whereas a sodic amphibole crystallizes at ~750°C, i.e. 10–20°C higher than in the comendites under the driest conditions investigated in our study. Thus, it appears that the increase of agpaicity in peralkaline silicic magmas does not produce any appreciable effect on tectosilicate liquidus temperatures. In contrast, mineral phases typifying the peralkaline condition, such as aegirine and Na-amphibole, have their stability fields expanded when peralkalinity increases, by ~100°C from comendites to pantellerites in the case of aegirine.

Temperatures of the liquidus minima in synthetic peralkaline acid systems at 100 MPa (Carmichael & Mackenzie, 1963; Thompson & Mackenzie, 1967), both iron free and iron bearing, are in the range 715–690°C (Carmichael & Mackenzie, 1963), slightly higher than

those obtained in the present study. At 100 MPa and H<sub>2</sub>O saturation, composition ND 002 falls on the quartz–alkali feldspar cotectic at 700°C, whereas BL 575 and SMN 49 do so at 680°C (the values at 100 MPa for our three compositions have been obtained by linear interpolation of the data obtained at 50 and 150 MPa). Minimum cotectic temperatures in the haplogranite system are higher by 20–30°C, i.e. 720°C at 100 MPa (Tuttle & Bowen, 1958). The tight grouping of quartz–feldspar cotectic temperatures (in the range 680–720°C) in synthetic and natural systems encompassing a large range of agpaitic indices shows that peralkalinity does not induce a marked depression in liquidus temperatures of tectosilicates in H<sub>2</sub>O-saturated silicic magmas, which contrasts with the situation under dry conditions (Bailey & Schairer, 1966). The fact that the cotectic temperatures are slightly lower in this work than in the related synthetic systems may rather be attributed to the role of other volatiles (F, Cl). The addition of 1 wt % F to the H<sub>2</sub>O-saturated haplogranite system at 100 MPa produces a fall in the temperature of the minimum of 35°C (Pichavant & Manning, 1984). Such an effect is similar in magnitude to that observed in the present study for the two comendites with ~1 wt % F, whose cotectic temperature is lower by ~20°C than their synthetic F-free counterparts.

### Plagioclase stability in silicic magmas

Phase relationships have been established for both metaluminous (Webster *et al.*, 1987) and peraluminous (Bénard *et al.*, 1985; Pichavant *et al.*, 1987; London *et al.*, 1989) silicic rock compositions. Considering first the metaluminous rocks, Webster *et al.* (1987) worked on an F-enriched vitrophyre (1.2 wt % F). Apart from its higher Al<sub>2</sub>O<sub>3</sub> and lower FeO<sub>tot</sub> contents, the vitrophyre composition is similar to that of ND 002. At 100 MPa under H<sub>2</sub>O saturation, both compositions crystallize the tectosilicates at 700°C. However, in the vitrophyre, two feldspars crystallize, unlike the case in ND 002 (Fig. 2). This is despite the fact that both rock compositions have almost identical bulk CaO content (0.40 wt % for the vitrophyre), and the vitrophyre is slightly more potassic than ND 002 (5.1 and 4.7 wt %, respectively). This suggests that even a slight deficit in Al<sub>2</sub>O<sub>3</sub>, such as in ND 002, inhibits plagioclase crystallization at the expense of alkali feldspar and quartz in silicic magmas. This is in accord with general petrographic evidence; in a compilation of 104 analyses of comendites and pantellerites, only one rock (from Oraefajokull, Iceland) clearly contained plagioclase phenocrysts (Macdonald & Bailey, 1973). At all water contents quartz is the liquidus phase in the vitrophyre, closely followed by sanidine and then plagioclase, the reverse of that in ND 002 (Fig. 2). This relative crystallization order is opposite to that predicted

from the normative quartz content of the bulk rocks (higher in ND 002 than in the vitrophyre). As proposed by Webster *et al.* (1987), such a crystallization sequence in the vitrophyre probably results from fluorine expanding significantly the primary volume of quartz relative to coexisting feldspars in the haplogranite system (Manning, 1981). With respect to the minor phases, the vitrophyre lacks clinopyroxene and oxides but biotite displays a larger stability field than in ND 002. Both topaz and fluorite crystallize in the vitrophyre, yet neither phase approaches liquidus conditions (<50–75°C below), unlike fluorite in the two comendites studied here. Overall, it appears that plagioclase crystallization in Ca-poor silicic melts is not suppressed by the addition of these levels of fluorine, as it has crystallized in the F-rich vitrophyre and not in the F-poorer ND 002 sample.

For F-poor peraluminous silicic magmas, comparison of our results with those obtained by Bénard *et al.* (1985) shows that, in peraluminous melts, quartz and alkali feldspar crystallize at similar temperatures (710°C at H<sub>2</sub>O saturation and 100 MPa) to ND 002, our F-poor, least peralkaline composition. Unlike in ND 002, however, plagioclase precipitates at ~770°C at 100 MPa in the peraluminous system despite bulk calcium contents on average similar to, or even lower than, those of the rocks used in our study [between 0.38 and 0.54 wt % for the three natural compositions used by Bénard *et al.* (1985)]. Experimental work on various F, Li, B-enriched peraluminous granites (Pichavant *et al.*, 1987; London *et al.*, 1989) showed also that all three major tectosilicates are stable in these magmas. Again, despite bulk CaO contents of ~0.2 wt %, plagioclase is the liquidus tectosilicate, closely followed by quartz and alkali feldspar. All peraluminous silicic rocks lack fluorite, despite bulk F contents in the range 1.2–2.4 wt %. The above observations suggest therefore that lack of plagioclase crystallization in peralkaline magmas is directly linked to the intrinsic peralkaline character of the melt and not to an unusual deficit in CaO.

### Role of $fO_2$ on the phase relations of peralkaline magmas

The phase relations displayed in Figs 2–4 help in understanding the main effects of  $fO_2$  variations in peralkaline silicic magmas. For the three rock compositions, variations in  $fO_2$  do not significantly affect the liquidus curves of quartz and alkali feldspar. The main exception is plagioclase crystallization in ND 002, which apparently occurs under oxidizing conditions only. The alkali feldspar liquidus in ND 002 may also be slightly enhanced under high  $fO_2$ . Overall, this indicates that feldspar- and quartz-producing species in the melt are not affected by  $fO_2$ . As variations in  $fO_2$  primarily affect the  $Fe^{2+}/Fe^{3+}$

ratio of silicate melts, this in turn suggests that quartz and alkali feldspar melt components are not directly linked to iron in the melt structure. This contrasts with the experimental observations of Dall'Agnol *et al.* (1999) on a metaluminous silicic composition. By establishing the phase relationships under two widely different sets of redox conditions (NNO – 1 and NNO + 2.5), they showed that a rise in  $fO_2$  produces a significant increase in the thermal stability of the tectosilicates, the effect being most dramatic for the feldspars, especially plagioclase. This strong dependence of plagioclase stability on  $fO_2$  suggests an intimate association between the feldspar melt components and the iron species in iron-bearing metaluminous silicic melts. Such a relationship appears not to exist in relatively F-rich peralkaline silicic melts.

In contrast to the tectosilicates, all other phases crystallizing in the peralkaline rocks are sensitive to  $fO_2$  to various degrees. In all three compositions, an increase in  $fO_2$  is marked by an increase in the thermal stability of oxides, an effect similar to that seen in almost all Fe-bearing magmas. The accessory phase chevkinite also displays an expansion of its stability field as  $fO_2$  increases (Fig. 4), especially in SMN 49, where its stability rises by ~50°C when  $fO_2$  increases. Differences in the stability fields of chevkinite in BL 575 and SMN 49 are possibly related to different trace element contents of the bulk rocks but could be due also to difficulties in distinguishing chevkinite from oxide in BL 575, unlike in SMN 49. Fayalite in ND 002 and biotite in all three rocks apparently are stable only at low  $fO_2$ . The Ca-rich pyroxene is present in ND 002 whatever  $fO_2$  prevails but its thermal stability is lowered when  $fO_2$  increases. Titanite is present under oxidizing conditions only, as in metaluminous silicic magmas (e.g. Wones, 1989; Dall'Agnol *et al.*, 1999) and its crystallization may in part explain the diminished stability field for the Ca-pyroxene at high  $fO_2$ , as both phases compete for Ca in liquid. One of the most significant changes produced by the increase in  $fO_2$  is the replacement in the low-temperature range of Na-amphibole by aegirine. The fact that both phases display similar stability fields suggests that they are related to each other through a redox reaction (Deer *et al.*, 1978). Given that the amphibole is strongly enriched in F, and that aegirine accommodates no F in its structure, amphibole breakdown must release F. This F release may in part explain the slightly enhanced stability field of fluorite under oxidizing conditions.

The last important feature with respect to  $fO_2$  concerns its influence upon the solidus. Although non-systematic, the data reported here clearly show that peralkaline silicic magmas at low  $fO_2$  can have solidus temperatures significantly lower than at high  $fO_2$ . We estimate that the lowering of solidus temperatures as a result of decreasing  $fO_2$  from NNO + 3.6 to NNO – 1.6 is at



least 30°C and could be as high as 50°C, the magnitude of the effect increasing with the bulk iron content. Clearly, this relies on the solution properties of iron in silicate melts and their dependence on  $fO_2$ . One implication is that crystal fractionation in peralkaline silicic magmas will be enhanced by low  $fO_2$ . Another is that a peralkaline protolith will start melting at a lower temperature under low  $fO_2$  rather than high  $fO_2$ . This redox melting mechanism has been already proposed by Bailey (1974) for mafic systems.

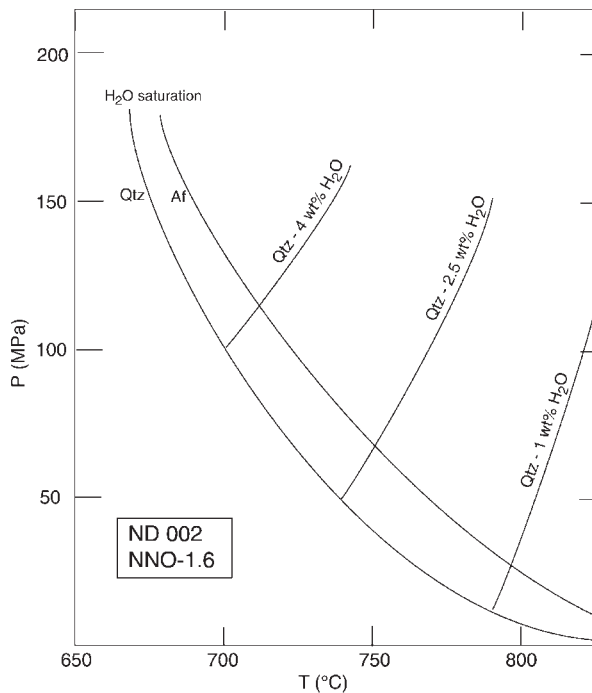
### Role of pressure

Variation in pressure between 50 and 150 MPa in our study does not produce any fundamental change in the phase relationships, apart perhaps for aenigmatite stability. In particular, the relative order of crystallization of alkali feldspar and quartz remains unaltered. Anhydrous pantellerites, on the other hand, exhibit a change in their liquidus phase with pressure, quartz replacing alkali feldspar at pressures >100 MPa (Bailey *et al.*, 1974). In pantellerites, the stability of Na-amphibole extends down to near-atmospheric pressure (Bailey & Cooper, 1978) and is weakly sensitive to pressure. Such behaviour is here confirmed for the riebeckite–arfvedsonite variety, showing furthermore that, in comendite magma, not only is amphibole stable at low pressure but its thermal stability increases with decreasing pressure, albeit slightly in the pressure range investigated. The thermal stability of amphibole at  $H_2O$  saturation at 50 MPa, where the melt  $H_2O$  content is 3.3 wt %  $H_2O$ , is similar to that at 150 MPa for a melt  $H_2O$  content of 3 wt %. Thus, it can be expected that at  $P < 50$  MPa, the thermal stability of amphibole may still increase to reach the maximum observed under anhydrous conditions at 150 MPa, 730–740°C. In the least peralkaline composition, amphibole replaces clinopyroxene at low water activity, the opposite behaviour to the amphibole–pyroxene relationship observed in all metaluminous melts so far worked on (e.g. Martel *et al.*, 1999; Scaillet & Evans, 1999). Amphibole stability in peralkaline magmas is thus fundamentally different from that in calc-alkaline, silicic to intermediate magmas, where amphibole crystallization demands melt water contents >4 wt % (e.g. Scaillet & Evans, 1999), and where it is hardly stable at pressures below 50–100 MPa (e.g. Sato *et al.*, 1999). The contrast is probably partly related to the fact that amphiboles in peralkaline magmas are more F rich. Thus, the crystallization of amphibole in peralkaline silicic magmas does not appear to be a reliable indicator of water pressure. Rather, its occurrence in mildly peralkaline rhyolites mainly indicates that both low- $fO_2$  and low-temperature conditions prevailed in the magma chamber. Biotite does not follow the same pattern, as this phase is

not stable at low pressure, at least under near-liquidus conditions. Biotite crystallization in comendite magmas at near-liquidus conditions implies pressures of at least 150 MPa.

Our experimental data help constrain the mineralogical evolution, mostly the behaviour of quartz, alkali feldspar and amphibole, of an ascending peralkaline magma under various  $T$ – $H_2O_{in}$  melt conditions. For this purpose, the stability curves of some critical phases in a  $P$ – $T$  projection are shown in Figs 5 and 6. In ND 002 at  $H_2O$  saturation (Fig. 5), the appearance curves of all major crystallizing phases (quartz, feldspar, clinopyroxene and magnetite) display negative slopes, which implies that, if ascent takes place at constant temperature under equilibrium conditions, extensive crystallization of the magma should occur. In contrast, if the magma starting conditions are below  $H_2O$  saturation, the appearance curves display a positive slope, down to the pressure at which the magma reaches  $H_2O$  saturation (Fig. 5). As illustrated for quartz in ND 002, the slope of the liquidus increases progressively as the bulk  $H_2O$  content of the magma decreases. A magma starting at 150 MPa, 740°C, with 4 wt %  $H_2O$ , will soon dissolve quartz upon rising, but it will do so only until 100 MPa, where it reaches  $H_2O$  saturation. Beyond, quartz crystallization will take place again. In contrast, a magma ascending from 150 MPa, 790°C and 2.5 wt %  $H_2O$  will intersect the quartz saturation boundary only at very low pressures, around 10 MPa. The alkali feldspar liquidus is reached at slightly higher pressures, at around 25 MPa. A similar line of reasoning shows that compositions BL 575 and SMN 49, both starting at 150 MPa, 760°C and 3 wt %  $H_2O$ , will reach the cotectic condition at ~25 MPa (Fig. 6). The relatively weak dependence of amphibole thermal stability on pressure implies that this phase will persist upon rising, unlike in calc-alkaline magmas. Instead, amphibole crystallization may take place at very shallow levels where dehydration of the magma occurs, thus raising amphibole thermal stability. This probably explains the common occurrence of amphibole microlites in the groundmass of volcanic rocks. For instance, an  $H_2O$ -saturated magma residing at 150 MPa and 700°C, i.e. outside the amphibole stability field, will upon ascent and concomitant degassing cross the amphibole liquidus at 40 MPa and end its ascent with abundant microlites, but no phenocrysts, of amphibole (Fig. 6).

This general picture is similar overall to that for other types of silicic magmas (Johannes & Holtz, 1996). It indicates that, to preserve their near-liquidus character up to near-surface conditions, peralkaline magmas must be  $H_2O$  undersaturated, unless ascent rates are fast enough to prevent chemical reaction. The fact that many peralkaline rocks display dissolution textures, such as rounded or embayed quartz phenocrysts (Sutherland, 1974; Barclay *et al.*, 1996), may potentially be related



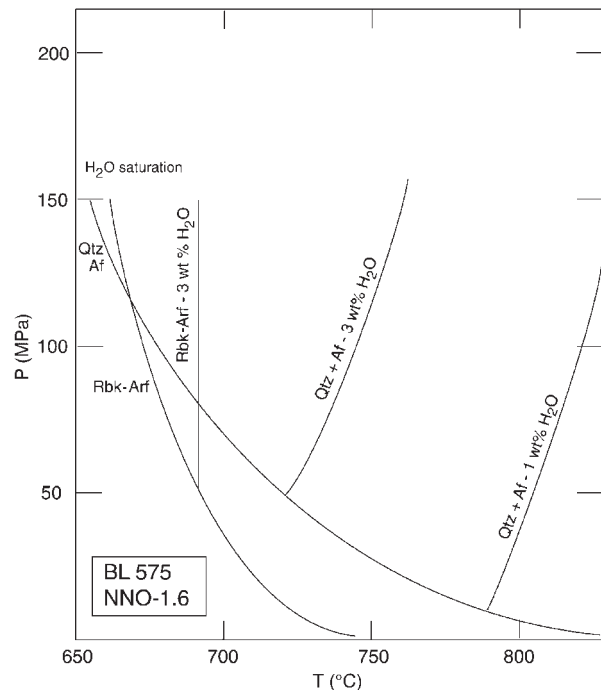
**Fig. 5.**  $P$ - $T$  stability curves for quartz and alkali feldspar in ND 002 at  $H_2O$  saturation. The stability curves of quartz for a magma containing 1, 2.5 and 4 wt %  $H_2O$  are also shown. The curves are labelled with minerals lying inside their respective stability fields.

to the  $H_2O$ -undersaturated character of the ascending magma. According to our data, ascent durations of  $<2$  days are implied for a magma stored at 150 MPa, 830°C, with  $\sim 3$  wt %  $H_2O$ , if quartz phenocrysts are to be preserved.

### Pre-eruptive conditions at Naivasha

First-order constraints on pre-eruptive conditions can be made by direct comparison of the phase equilibria established here with the phase assemblages observed in the rocks. This assumes that, in the latter, bulk crystal-melt equilibrium prevailed shortly before eruption and was not subsequently altered during extrusion. The lack of conspicuous zoning of the major phases and of evidence for xenocrysts suggests that this is the case (Macdonald *et al.*, 1987).

Let us consider first ND 002 (Fig. 2). The absence of titanite indicates a pre-eruptive  $fO_2$  close to or below FMQ. The stable coexistence of ferrohedenbergite and fayalite implies that pre-eruptive conditions were at the overlap in their stability fields. If the magma was at water saturation, then temperature is constrained by the liquidus of alkali feldspar, at  $\sim 700^\circ\text{C}$ . Significantly lower temperatures are precluded by the hedenbergite-out reaction and by the appearance of quartz in the experiments. At 50 MPa coprecipitation of alkali feldspar,



**Fig. 6.**  $P$ - $T$  stability curves for quartz-alkali feldspar and riebeckite-arfvedsonite in BL 575 at  $H_2O$  saturation. The stability curves of quartz-alkali feldspar for a magma containing 1 and 3 wt %  $H_2O$ , and for riebeckite-arfvedsonite for 3 wt %  $H_2O$ , are also shown. The curves are labelled with minerals lying inside their respective stability fields.

hedenbergite and magnetite (+ fayalite) is achieved at around  $780^\circ\text{C}$  at  $H_2O$  saturation ( $\sim 2.5$  wt %  $H_2O$ ). Temperatures lower than  $740^\circ\text{C}$  are ruled out by quartz crystallization. On the other hand, lower melt water contents would require still higher temperatures. Melt water contents of melt inclusions in some Olkaria rocks reach 3.4 wt % (Wilding *et al.*, 1993), similar to those obtained here, whatever the pressure.

The presence of Na-amphibole at near-liquidus conditions in the natural BL 575 constrains the pre-eruptive temperature to have been lower than  $700^\circ\text{C}$  and, along with the absence of aegirine, suggests that the  $fO_2$  was close to, or below, FMQ. Ignoring aenigmatite, the phase assemblage is restricted to a narrow  $T$ -melt water domain, with melt water contents of  $\sim 5.5$  wt % and temperature below the intersection point of the biotite-in and Na-amphibole-in curves, which occurs slightly above  $670^\circ\text{C}$ . The estimate of melt water content is significantly higher than the maximum so far recorded in melt inclusions in quartz phenocrysts at Olkaria (Wilding *et al.*, 1993), which might indicate that 150 MPa for the pre-eruptive pressure of the storage zone is too high. A decrease in pressure would, nevertheless, not raise significantly the pre-eruption temperature estimate, as

this parameter is constrained by the occurrence of Na-amphibole whose stability is rather independent of pressure. However, a decrease in pressure removes amphibole from the liquidus (Fig. 6) and biotite has not been found at 50 MPa, both facts arguing against a pressure significantly lower than 150 MPa.

Sample SMN 49 lacks both aenigmatite and Na-amphibole as phenocrysts but has biotite, in addition to quartz and alkali feldspar. This implies a very restricted set of pre-eruptive temperatures and melt water contents, basically in the coldest and wettest part of the investigated domain, corresponding to the overlap between the stability fields of biotite, quartz and alkali feldspar but above that of Na-amphibole. At 150 MPa, a pre-eruptive temperature of 660°C and H<sub>2</sub>O-saturated conditions are implied. Again, as for BL 575, considering the highest melt H<sub>2</sub>O content found in melt inclusions (Wilding *et al.*, 1993), a pressure of 150 MPa would seem to be too high. At 50 MPa, coprecipitation of quartz and alkali feldspar is reached at 720°C at H<sub>2</sub>O saturation (~3.3 wt % H<sub>2</sub>O), but at this pressure and temperature biotite is not a liquidus phase.

Our experiments confirm the idea that peralkaline magmas may contain significant amounts of dissolved water, yet our estimates are nearly twice those from melt inclusion analyses (Wilding *et al.*, 1993). The phase equilibrium constraints detailed above are robust and we believe that this mismatch arises from a problem in interpreting the melt inclusion data. Such a difference could arise from a variety of factors, such as post-eruptive H<sub>2</sub>O diffusive loss in the trapped melt inclusions or inclusion formation during ascent, i.e. at pressures lower than the pre-eruptive one, either of these processes being driven by, or occurring during, open-system degassing during eruption. Evidence for open-system degassing is in fact suggested by H/D isotope data (Wilding *et al.*, 1993). Restored pre-eruptive melt H<sub>2</sub>O contents using measured  $\delta D$  and the matrix H<sub>2</sub>O values of glassy volcanic rocks at Olkaria have yielded values up to 5.7 wt % (Wilding *et al.*, 1993), i.e. identical to the H<sub>2</sub>O contents at saturation at 150 MPa and in agreement with our phase equilibrium constraints. We conclude that melt inclusion data at Olkaria provide only a minimum estimate of the amount of water present in the magma chamber before eruption.

In summary, the pre-eruptive intensive parameters derived in this study are characterized by last equilibration at an  $fO_2$  at, or possibly below, FMQ, temperatures well below 800°C, and water-rich conditions. For the two more peralkaline samples, the pressure of the storage zone is probably >50 MPa, as indicated by the occurrence of near-liquidus biotite in the rocks. It cannot be much higher than 150 MPa, however, as increasing pressures will further depress the liquidus of

tectosilicates, whereas that of amphibole will be correspondingly less affected, making it difficult to preserve the liquidus phase assemblage displayed by the rocks. Such a pressure for magma storage is also indicated by independent geochemical evidence (Davies & Macdonald, 1987).

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