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# The carbon dioxide solubility in alkali basalts: an experimental study

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## Abstract:

Experiments were conducted to determine CO<sub>2</sub> solubilities in alkali basalts from Vesuvius, Etna and Stromboli. The basaltic melts were equilibrated with nearly pure CO<sub>2</sub> at 1200°C under oxidizing conditions, and at pressures ranging from 269 to 2060 bars. CO<sub>2</sub> solubility was determined by FTIR measurements. The results show that alkalis have a strong effect on the CO<sub>2</sub> solubility and confirm and refine the relationship between the compositional parameter  $\Pi$  devised by Dixon (1997) and the CO<sub>2</sub> solubility. A general thermodynamic model for CO<sub>2</sub> solubility in basaltic melts is defined for pressures up to 2 kbars. The new CO<sub>2</sub> solubility model yields saturation pressures lower by as much as 50% relative to existing models when applied to volatile-rich alkali basalts.

**Key words:** CO<sub>2</sub> solubility, alkali basalts, experiments, thermodynamics

## ***Introduction***

Dissolved volatiles in silicate melts dramatically influence the physical properties of magmas, such as density, viscosity, vesiculation, and largely control eruptive styles. Volatiles exsolve from melts into bubbles as magmas rise toward the earth's surface and depressurize (e.g., Blank et al., 1993). Such a process may give rise to explosive eruptions, leading to important human and material damages, and may have a major impact on the global Earth climate (Robock 2000). Water is generally the most abundant volatile in terrestrial magmas, CO<sub>2</sub> being the second most common species, notably in mafic compositions. Thus, understanding the thermodynamics of CO<sub>2</sub> in silicate melts can provide important constraints on the behaviour of basaltic reservoirs, the ensuing eruptive processes, and ultimately on the evolution of planetary degassing. To model basalt degassing during volcanic eruptions, a detailed knowledge of the mixing properties of volatile components in both fluid and melt is required since, because of its low solubility, carbon dioxide is usually a major component of the gases escaping from magmas as they rise through the crust.

Experiments on CO<sub>2</sub> solubility in basaltic melts have been performed mostly at high pressures where dissolved CO<sub>2</sub> contents may reach several wt% (>10 kbar, Blank and Brooker, 1994, Morizet et al., 2002). At lower pressures, efforts have mainly focused on silicic liquids and only a few studies have been performed on CO<sub>2</sub> solubility in basaltic melts (e.g., Pawley et al., 1992, Dixon et al., 1995, Jendrzewski et al., 1997; Behrens et al., 2009). As pointed out by Stolper and Holloway (1988), this gap of knowledge reflects in part the small influence of CO<sub>2</sub> on solid-liquid phase equilibria at low pressures, due to its low solubility in most silicate melts under such conditions. Another reason is the technical difficulties associated with the need to work at the high temperature conditions required for mafic compositions (e.g., Berndt et al., 2002; Di Carlo et al., 2006; Behrens et al., 2009), in particular in the mid pressure range (2-6 kbar) where solid pressure media, such as piston cylinders, are not routinely operated. More specifically, the low viscosity of mafic melts requires them to be fast quenched to prevent their crystallisation during cooling. This demands the implementation of appropriate experimental designs in gas pressure vessels (Holloway et al., 1992; Roux and Lefevre, 1992) and, for H<sub>2</sub>O-bearing basalts, such devices may still fail to properly quench liquids to glasses (e.g., Behrens et al., 2009). As a consequence of these technical difficulties, the number of high quality solubility data in the low to mid pressure range has remained limited. In particular, alkali basalt compositions have received limited attention. Petrological studies indicate, however, that such alkali-rich compositions are volatile-rich, in particular in arc

settings (Dixon et al., 1997; Métrich et al., 2001; Bertagnini et al., 2004). Experimental data show indeed that alkali-rich basalt compositions tend to dissolve more CO<sub>2</sub> than those alkali-poor (Dixon, 1997; Botcharnikov et al., 2005; Behrens et al., 2009), a key role being played by potassium (Behrens et al., 2009) but the exact compositional control of such an increase is still unclear.

In an effort to decipher the role of alkalis on CO<sub>2</sub> solubility, in this study we report measurements of CO<sub>2</sub> solubility obtained in several basalts of varying alkalinity coming from Italian active volcanoes at 1200°C, at pressures ranging from 250 up to 2000 bar, and under oxidizing conditions. CO<sub>2</sub> contents in quenched glasses from vapour-saturated experiments were measured by infrared spectroscopy. Solubility measurements give upper bounds on the amounts of CO<sub>2</sub> that can dissolve in magmas at various temperatures and pressures. A companion paper presents the thermodynamics of water in these silicate melts. Our aim is to produce a set of internally consistent data on the solubility of C-O-H-S volatiles under moderately oxidizing conditions so as to be ultimately able to quantitatively model the degassing of alkali basalts, such as recently done for silicic compositions by Burgisser and Scaillet (2007).

## ***Experimental techniques***

### **Starting materials, capsules and charges**

Three alkali basalts from Vesuvius, Etna and Stromboli (Lesne et al. this issue) were selected for our investigation. The natural samples were powdered and melted in a Pt crucible at atmospheric pressure, at 1400°C during 3 hours in order to get homogeneous and nominally anhydrous glasses. The resulting glasses were analyzed by electron microprobe. Their compositions are listed in Table 1. Experiments were conducted in parallel on those three alkali basalts.

As for experiments on water solubility, small glass chips were loaded in capsules to prevent bubble formation. About 60 to 100 mg of starting material was used for each experiment. Capsules (2 cm height, 2.5 mm inner and 2.9 mm outer diameters) made of platinum were used. CO<sub>2</sub> was added as silver oxalate (Ag<sub>2</sub>C<sub>2</sub>O<sub>4</sub>, density 5g/cm<sup>3</sup>). Although silver does dissolve in basaltic melt, its effect on CO<sub>2</sub> solubility has been shown to be negligible (Pan et al., 1991). Capsules were sealed by electric arc. They were then put in an oven for several hours and reweighed to check for leaks.

## **Equipment and experiments**

Equilibrium experiments between a fluid (CO<sub>2</sub>-rich fluids, see below) and a silicate liquid phase were conducted in an internally heated pressure vessel (IHPV) operated in a vertical position and equipped with a drop quench set up (see Di Carlo et al., 2006). Experiments were performed at 1200±5°C, at the following pressures: 269, 524, 1013, 1062, 1128, 2059 ±25 bars, under oxidizing conditions. Experiments lasted between 4 to 8 hours. Based on previous experimental works performed using similar procedures, the equilibrium between the basaltic liquid and the fluid phase is considered to be reached under these conditions (e.g., Dixon et al, 1995; Botcharnikov et al, 2005; Behrens et al., 2009). Phase equilibrium experiments performed with the same duration and P-T-H<sub>2</sub>O-CO<sub>2</sub> conditions (Di Carlo et al., 2006; Pichavant et al. 2009) on a Strombolian basalt had no problem in attaining crystal-liquid equilibrium. Since the kinetics of melt-fluid equilibrium is faster than that of crystal-liquid one, we conclude that our run durations were long enough to attain equilibrium. In support of this, major element analyses performed on two or three glass chips on microprobe yielded the same results within analytical error. The same was observed with FTIR analyses carried out on several double polished glass pieces (2 to 3).

Any use of sensor was not possible because of the short duration of the experiments. All the runs were done using the intrinsic  $fH_2$  of the vessel, as measured by Pichavant et al. (2009) for experiments done with the same apparatus, without hydrogen added, at 1150°C, at 2010 and 880 bars (runs 22 and 23). We assume that the  $fH_2$  measured in those experiments is equivalent to the  $fH_2$  in our runs, then the  $fO_2$  of each charge was calculated from the dissociation reaction of water, using  $fH_2O$  as calculated from the dissolved water content in the glass (see below). It appears that the  $fO_2$  calculated for all charges, averages at NNO+1.9, with an error of 0.6 log unit. All performed experiments for this study are reported in Table 2.

## **Analytical techniques**

Analytical techniques used in this study are the same as those used in the accompanying paper (Lesne et al.). Major elements were analysed with a SX50 CAMECA electron microprobe (BRGM-ISTO, Orléans, France). Water contents and its speciation were determined with FTIR analyses. All of these methods are described in detail in Lesne et al. (this issue). The total carbon dissolved in quenched glasses was determined by using transmission IR spectroscopy on doubly polished glass chips, with thicknesses varying between 60 and 270 µm. One sample (4#4) had very low CO<sub>2</sub> contents, which required the use of a thick double polished glass (370 µm).. Sample thicknesses were measured using an optical microscope to

within 3  $\mu\text{m}$ . To determine the total carbon dissolved in the glass, a microchamber attachment to the FTIR Nicolet<sup>TM</sup> Magna Spectrometer 760 together with a white source, a KBr beamsplitter and a MCT-A detector (mercury, cadmium tellurium) were used. Analytical conditions for each analysis were 128 scans and a 2  $\text{cm}^{-1}$  resolution. The diameter of the analyzed spot was 100  $\mu\text{m}$ . The background was acquired in air at each measurement. For each sample, a minimum of 2 spots were made, in order to check the homogeneity of the sample.

Carbon can be dissolved in silicate melts either as molecular  $\text{CO}_2$  or as carbonate groups  $\text{CO}_3^{2-}$  (e.g., Fine and Stolper, 1985, 1986, Morizet et al., 2002). The intensity of the 2350  $\text{cm}^{-1}$  is attributed to the  $\nu_3$  antisymmetric stretch of molecular  $\text{CO}_2$  (Fine and Stolper, 1985; Fine and Stolper, 1986). The peaks at 1515  $\text{cm}^{-1}$  and 1435  $\text{cm}^{-1}$  correspond to an antisymmetric stretch of distorted  $\text{CO}_3^{2-}$  groups (Fine and Stolper, 1986). Figure 1 shows a typical IR spectrum obtained on our glasses which shows that there is a double peak at 1515  $\text{cm}^{-1}$  and 1435  $\text{cm}^{-1}$ , whilst none is present at 2350  $\text{cm}^{-1}$ . Thus, the carbon dissolved in the melt is present as carbonate groups ( $\text{CO}_3^{2-}$ ) only, in agreement with previous works on basaltic compositions (e.g., Pan et al., 1991; Behrens et al., 2009).

In this work, following Behrens et al. (2009), we measured carbonate absorbance by subtracting from the sample spectrum that of a compositionally similar but volatile-free glass (ie the starting dry glass) scaled to the same thickness of the unknown. This enables to model the background and measure the peak height of the 1515  $\text{cm}^{-1}$  wavenumber.

The total carbon content of glasses was calculated with the Beer-Lambert law (Fine and Stolper 1985;1986), by measuring the absorbance of the 1515  $\text{cm}^{-1}$  band:

$$C_{\text{CO}_2} = \frac{A_{1515} * M_{\text{CO}_2}}{d * \rho * \epsilon_{1515}} \quad (1)$$

Where  $C_{\text{CO}_2}$  is the total carbon content in the glass (in wt%),  $A_{1515}$  is the absorbance of the 1515  $\text{cm}^{-1}$  band,  $d$  is the thickness of the sample (in cm),  $\rho$  is the density in  $\text{g.l}^{-1}$ ,  $\epsilon_{1515}$  is the molar extinction coefficient ( $\text{l.mol}^{-1}.\text{cm}^{-1}$ ), and  $M_{\text{CO}_2}$  is the molecular weight of  $\text{CO}_2$ .

Glass densities, were calculated using the method of Richet et al. (2000), considering only the amount of water dissolved in the glass, assuming that the small quantities of  $\text{CO}_2$  dissolved do not affect density to any measurable extent (Lange, 1994, Bourgue and Richet, 2001).

The choice of appropriate molar extinction coefficients is crucial to measure correctly dissolved carbon contents (e.g. Behrens et al., 2009). However, we do not have bulk carbon data, and thus it is not possible to obtain molar extinction coefficients for each basaltic

compositions studied here. Consequently, two choices are offered: either we arbitrarily take the same value for the three basaltic compositions, referring to Fine and Stolper (1986) ( $\epsilon_{1515} = 375 \pm 20 \text{ l.mol}^{-1}.\text{cm}^{-1}$ , calculated for basaltic glasses), or we calculate the extinction coefficient for each basaltic melt using the method of Dixon and Pan (1995). These authors found a linear dependency of the molar extinction coefficient with the Na/(Na+Ca) molar ratio, in the range 0.25 to 0.5. Values of molar extinction coefficients determined by Dixon and Pan's method yield  $381 \text{ l.mol}^{-1}.\text{cm}^{-1}$  for Vesuvius,  $327 \text{ l.mol}^{-1}.\text{cm}^{-1}$  for Etna and  $367 \text{ l.mol}^{-1}.\text{cm}^{-1}$  for Stromboli samples. The calculated extinction coefficients for both Vesuvius and Stromboli are within 2% of the Fine and Stolper (1986) value, whilst that for Etna is lower by 13%, which reflects its higher Na content. Behrens et al. (2009) determined the extinction coefficient of their phonotephrite melt to be  $308 \pm 110 \text{ l.mol}^{-1}.\text{cm}^{-1}$ , whilst the value retrieved with the Dixon and Pan (1995) method is  $370 \text{ l.mol}^{-1}.\text{cm}^{-1}$ . Although both values agree within error, the difference observed suggests that the Dixon and Pan method may overestimate extinction coefficients of strongly potassic compositions by up to 20%, which translates into a similar overestimates of carbon contents calculated with the Beert-Lambert law. In the following, owing to the lack of absolute method to measure dissolved C content, and considering the significant errors associated to existing measurements of extinction coefficients for basaltic compositions, we first present the results obtained by using the same extinction coefficient for all compositions, and latter discuss the effect arising from the use of composition specific molar extinction coefficients .

In addition to the uncertainty on extinction coefficients, one of the largest sources of errors in measuring the total carbon dissolved in the melt by FTIR arises from the method chosen to model the background. The method used here is somewhat subjective and leads to an error in absorbance estimated to be  $\pm 0.015$ , which propagates into an error of  $\pm 50 \text{ ppm}$  of the  $\text{CO}_2$  dissolved in the glass, regardless of the  $\text{CO}_2$  contents. Thicknesses were measured to within  $3 \mu\text{m}$  which induces an error of 2% on the total  $\text{CO}_2$  contents dissolved in the melt. The maximum error on the density calculations is equal to 20%. Altogether we estimate that the  $\text{CO}_2$  content reported in this work are known to within 20%.

## **Results**

The fluid-saturated charges have been checked after each experiment: capsules were first weighed to check any volatile loss during experiment, then they were punctured. Any escape of bubbles/gas and difference of weigh between before and after piercing is a proof of fluid-saturated experiment. All the data presented in this study are fluid saturated.

## Macro and microscopic observations

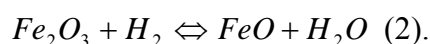
Run products were translucent to brownish glasses, becoming darker and opaque when quench crystals ( $<5\mu\text{m}$ ) were present (i.e., in experiments performed with more than 4 capsules). Most quenched glasses contain a few microbubbles whose abundance does not exceed 0.1 % in volume. Larger bubbles ( $<0.5\text{mm}$  of diameter), indicative of fluid saturation at P and T, are located at the melt/capsule interface, as shown by the presence of large open smooth cavities at the surface of the glass blocks.

## Major elements and H<sub>2</sub>O contents

The compositions of quenched glasses are reported in the Table 3. Analytical errors were less than 0.2% for all oxides, except for SiO<sub>2</sub> (0.4 to 0.6%) and for Fe<sub>tot</sub> (less than 0.4%). Apart from Fe<sub>tot</sub>, the concentrations of major oxides do not differ appreciably from those of the starting materials (Table 1).

For experiments carried out in platinum capsules, the iron loss is around 30%, whereas the iron loss of experiments carried out in Au<sub>75</sub>Pd<sub>25</sub> capsules is less than 10% (Table 3).

Despite that no water was added to the capsules, the water contents of quenched glasses vary between 0.70 and 1.30 wt% without any obvious relationships with pressure or melt composition. Several sources of water can be proposed: it could come from the silver oxalate which contains some water (up to 5 wt%) and could not be dried enough, or from adsorption of atmospheric water on the starting glasses. However, the most probable source of H<sub>2</sub>O is from reduction of ferric into ferrous iron (e.g., Holloway et al., 1992), inasmuch as our starting glasses are oxidized, which enables the operation of the following reaction:



Note that, in addition to the above reaction, reduction of ferrous iron via alloying with the capsule wall will also produce some water.

## CO<sub>2</sub> contents

Dissolved CO<sub>2</sub> contents in glasses are reported in Table 3. We stress that, as the quenched glasses also contain dissolved H<sub>2</sub>O, the measured CO<sub>2</sub> contents do not represent pure CO<sub>2</sub> solubilities. The data show linear relationships between dissolved carbon and pressure for each composition (figure 2). Data from Dixon et al. (1995) and Behrens et al. (2009) are also shown for comparison. Figure 2 shows clearly that there is a strong dependence of the carbon dissolved in the melt on melt composition. Alkali-rich compositions dissolve more CO<sub>2</sub> than



those alkali-poor. However, to rigorously compare CO<sub>2</sub> solubilities of the different melt compositions, one has to determine the fugacities of CO<sub>2</sub>.

### Determination of the fugacities of CO<sub>2</sub>.

As the H<sub>2</sub>O of quenched glasses was determined in each charge, and as solubilities laws of water were defined for each composition in the companion paper (Lesne et al.), we can calculate the  $f_{H_2O}$  of each charge, using the following empirical relationships derived from the experimental data base of Lesne et al. (this issue):

$$f_{H_2O}^{1200^\circ C} = 126.54 \times (wt\%H_2O)^{1.69} \quad \text{for Vesuvius} \quad (3);$$

$$f_{H_2O}^{1200^\circ C} = 104.98 \times (wt\%H_2O)^{1.83} \quad \text{for Etna;} \quad (4);$$

$$f_{H_2O}^{1200^\circ C} = 128.49 \times (wt\%H_2O)^{1.73} \quad \text{for Stromboli.} \quad (5).$$

In the above equations  $wt\%H_2O$  is the wt% amount of total water dissolved in the quenched glasses. Once the  $f_{H_2O}$  is known, we can derive the mole fraction of H<sub>2</sub>O ( $X_{H_2O}^f$ ) in the coexisting fluid by using the standard relationships:

$$f_i = \gamma_i * P_i \quad (6),$$

where  $\gamma_i$  is the fugacity coefficient of species  $i$ , and  $P_i$  its partial pressure.

Since

$$P_i = X_i * P_{tot} \quad (7), \text{ where } P_{tot} \text{ is total pressure,}$$

we obtain

$$f_i = \gamma_i * X_i * P_{tot} \quad (8).$$

Assuming that the fluid is made of only H<sub>2</sub>O and CO<sub>2</sub> species (under oxidized conditions CO<sub>2</sub> has been shown to be the only C-bearing species in the C-O-H system, Pawley et al., 1992), it follows:

$$X_{H_2O} + X_{CO_2} = 1 \quad (9),$$

and so

$$X_{CO_2} = 1 - X_{H_2O} \quad (10).$$

In this work, we use the  $\gamma_{CO_2}$  and  $\gamma_{H_2O}$  values of Dixon et al. (1995). These fugacity coefficients correspond to those of pure species at P and T. Strictly speaking, the fugacity coefficient of any species  $i$  in a complex fluid is dependent on the composition of that fluid, in addition to P and T, and this can be accounted for via the use of a modified Redlich Kwong equation of state with appropriate mixing rules between different fluid species (Holloway,

1987; Ferry and Baumgartner, 1987). However, in the low pressure and high temperature range of interest here, application of the rules of real mixing of real fluid has only a trivial effect on the calculated fugacity coefficients (less than 1 %, see Holloway, 1977), relative to the assumption of ideal mixing of real fluids (Lewis and Randall rule) which has thus been used here for simplicity. Application of the empirical relationships 2-4 assumes also that dissolved CO<sub>2</sub> does not affect the thermodynamics of H<sub>2</sub>O solution in the selected melts. Considering the low concentration in CO<sub>2</sub> of our synthetic melts, this assumption holds true to a first approximation. The reverse (ie H<sub>2</sub>O affects the CO<sub>2</sub> solubility) might not be true, however, as discussed below. Using such an approach, the fugacity of CO<sub>2</sub> of any run product can be calculated in a way that ensures internal consistency between fluid species fugacities:

$$f_{CO_2} = \gamma_{CO_2} \times X_{CO_2} \times P_{tot} \quad (11).$$

The relationships between dissolved CO<sub>2</sub> ( $X_{CO_2}^m$  being the mole fraction of CO<sub>2</sub> dissolved calculated following Silver and Stolper (1985) on the basis of one oxygen) and  $f_{CO_2}$  are shown in figure 3.

Each trend can be fitted to a second order polynomial equation, in mole fractions:

$$f_{CO_2}^{1200^\circ C} = 8.57 \times 10^6 \times X_{CO_2}^m{}^{1.26}; \quad R^2 = 0.99 \text{ for Vesuvius} \quad (12)$$

$$f_{CO_2}^{1200^\circ C} = 1.14 \times 10^7 \times X_{CO_2}^m{}^{1.23}; \quad R^2 = 0.93 \text{ for Etna} \quad (13)$$

$$f_{CO_2}^{1200^\circ C} = 3.49 \times 10^6 \times X_{CO_2}^m{}^{1.04}; \quad R^2 = 0.99 \text{ for Stromboli} \quad (14)$$

or in ppm:

$$f_{CO_2}^{1200^\circ C} = 1.93 \times 10^{-1} \times CO_2^{melt1.26}; \quad R^2 = 0.99 \text{ for Vesuvius} \quad (15)$$

$$f_{CO_2}^{1200^\circ C} = 4.06 \times 10^{-1} \times CO_2^{melt1.22}; \quad R^2 = 0.93 \text{ for Etna} \quad (16)$$

$$f_{CO_2}^{1200^\circ C} = 1.62 \times CO_2^{melt1.04}; \quad R^2 = 0.99 \text{ for Stromboli} \quad (17)$$

$f_{CO_2}$  versus  $CO_2^{melt}$  relationships for MORB (Dixon et al., 1995) is also shown on figure 3: the CO<sub>2</sub> solubility varies non-linearly with fluid composition. Again, the dependence of the total dissolved carbon with melt composition is apparent. It should be stressed that equations 12-17 are essentially valid for the selected melt compositions having dissolved water contents similar to those of the present work and over the pressure range investigated (1-2 wt% and up to 2 kb). On figure 3 we have also plotted the data of Behrens et al (2009), calculating the  $f_{CO_2}$  using their measured  $X_{CO_2}$  and equation (11). Despite some scatter, in part related to

quench effects arising from the excessively low viscosity of phonotephrite melt, the data display a similar exponential pattern, being located to the right of the Vesuvius trend, which is in keeping with the much stronger potassic character of the phonotephrite composition (7.52 wt% K<sub>2</sub>O) used by Behrens et al. (2009) relative to any of our melt compositions.

## **Discussion**

### *Effect of alkalis*

Our results show clearly that alkalis have a positive effect on CO<sub>2</sub> solubility in basaltic melts as shown by previous studies (e.g., Dixon 1997; Behrens et al. 2009).

At any investigated pressure, the tephrite from Vesuvius dissolves 1.4 more CO<sub>2</sub> than the Etnean basalt, and 1.7 more CO<sub>2</sub> than the Strombolian basalt, and up to 2 times more CO<sub>2</sub> than a MORB (Dixon et al., 1995). This general trend is confirmed and extended by the data of Behrens et al. (2009). This effect is illustrated in figure 4: for each isobar, there exists a positive relationship between CO<sub>2</sub> solubility and the alkalinity of the melt.

These results were obtained using the same extinction coefficients for the three basaltic compositions. If we consider now different extinction coefficients for our three basaltic melt compositions, using the method of Dixon and Pan (1995), we obtain similar results to those obtained with the molar extinction coefficient of Fine and Stolper (1986) for VES-9 (2% lower) and PST-9 (2% higher). Only for ETN-1 are the results obtained with the Dixon and Pan method different, compared to those obtained with the constant extinction coefficient. This is simply explained by the fact that Dixon and Pan (1995) method takes into account of the Ca and Na contents, the concentration in the latter element being higher in the Etnean composition relative to the two other basaltic compositions. However, regardless of which extinction coefficient is used (constant or specific to each composition), the same effect of alkalis is observed with pressure: VES-9 dissolves more CO<sub>2</sub> than ETN-1 and then than PST-9. We thus conclude that alkalis (Na<sub>2</sub>O+K<sub>2</sub>O) have broadly a positive effect on the CO<sub>2</sub> solubility in basaltic melts (see also King and Holloway, 2002). In the following, we consider the CO<sub>2</sub> solubility results obtained by using the molar extinction coefficient of by Fine and Stolper (1986).

### *Refinement of the model of Dixon (1997)*

Dixon (1997) used experimentally measured solubilities of CO<sub>2</sub> in alkali basalts (composition from a MORB to a leucitite) to propose a method to rationalise the influence of melt

composition on CO<sub>2</sub> solubility. Dixon (1997) defined a compositional parameter  $\Pi$  which takes into account the roles of Si<sup>4+</sup>, Al<sup>3+</sup>, Ca<sup>2+</sup>, K<sup>+</sup>, Na<sup>+</sup>, Mg<sup>2+</sup> and Fe<sup>2+</sup> cation fractions:

$$\Pi = -6.50(\text{Si}^{+4} + \text{Al}^{+3}) + 20.17(\text{Ca}^{+2} + 0.8\text{K}^{+1} + 0.7\text{Na}^{+1} + 0.4\text{Mg}^{+2} + 0.4\text{Fe}^{+2}) \quad (19)$$

With this factor, the solubility of CO<sub>2</sub> in alkali-rich basalts can be predicted at 1200°C and 1000 bars. CO<sub>2</sub> solubilities for our three compositions at 1200°C and 1000 bars are plotted against the factor  $\Pi$  and show a good agreement with the linear relationship predicted by Dixon (1997) (figure 6). Also represented is the composition of Behrens et al (2009), obtained from linear extrapolation to 1 kbar of the results obtained by these authors at 2 and 5 kbar (which yields at 1 kbar a CO<sub>2</sub> content of 767 ppm).

Overall, the available experimental data base supports the parameterisation between CO<sub>2</sub> solubility and the  $\Pi$  factor proposed by Dixon (1997) at 1200°C and 1000 bar as a useful tool for the description of CO<sub>2</sub> solubility in basalt melts of variable alkalinity. At 1 kbar and 1200°C, the extended data base yields the following equation:

$$\text{CO}_2 \text{ (ppm)} = 920 (\pm 1)\Pi - 17 (\pm 53), \quad R^2 = 0.89 \quad (20)$$

CO<sub>2</sub> solubilities at 1000 bars calculated with equation (20) are slightly lower (<10% ) than the CO<sub>2</sub> solubilities calculated by using the former regression coefficients of Dixon (1997). Calculating the CO<sub>2</sub> solubility with the new regression, and comparing to actual CO<sub>2</sub> contents, differences are less than 15% for both VES-9 and PST-9, whilst for the Etna composition the back calculated values may differ by up to 30%. Despite such an uncertainty equation (20) can be used to predict to a first approximation the CO<sub>2</sub> solubility of basaltic magmas of different alkali contents at 1 kbar and 1200°C. Although there are not enough data to test its validity at pressures other than 1 kb, we can only assume that such a correlation holds true up to 2 kbar.

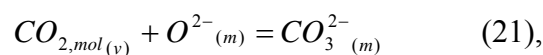
### *Thermodynamic model*

In this section we address the thermodynamics of CO<sub>2</sub> solution in our basaltic melts, using essentially the same approach as Dixon et al. (1995). Recent work on CO<sub>2</sub> solubility on a variety of silicate melts have emphasized that the anhydrous major element composition has relatively little effect on CO<sub>2</sub> solubility up to 2 kbar (Botcharnikov et al., 2005; 2006). In contrast, it has been shown that dissolution of water may profoundly affect the CO<sub>2</sub> solubility trends in a variety of silicate liquids (Mysen, 1976), the case being made strongly for andesitic to dacitic compositions (Jakobsson, 1997; King and Holloway, 2002; Behrens et al., 2004; Botcharnikov et al., 2005). In particular, experiments show that water presence

enhances the solubility of CO<sub>2</sub>, leading to strong departure from Henrian behaviour, notably at high pressures (>2 kb), in agreement with earlier suggestion by Mysen (1976). Owing to water production upon reduction of ferric/ferrous iron, none of our melts is anhydrous, which means that we lack the benchmark needed to evaluate whether introduction of water in our basaltic melts indeed affects their CO<sub>2</sub> solubility in the investigated pressure range. However, Behrens et al. (2009) observed a near linear relationships between dissolved CO<sub>2</sub> and XCO<sub>2</sub> at 2 kbar for their phonotephrite composition, suggesting small departure from Henrian behaviour at this, and presumably below it, pressure. Similarly, the work of Dixon et al. (1995) investigated pressures below 1 kb, and noted an essentially Henrian behaviour of CO<sub>2</sub> solubility, despite that the water content of their quenched melts was in the range 1-2 wt%, as in our charges. Comparison with their data, obtained by equilibrating a MORB melt with an CO<sub>2</sub>-H<sub>2</sub>O vapour, with those of Stolper and Holloway (1988) and Jendrzewsky et al. (1997), which were obtained under drier conditions (H<sub>2</sub>O<0.6 wt%) than ours, shows that all data sets fall on the same  $f\text{CO}_2$  vs CO<sub>2</sub><sup>melt</sup> trend. This indicates that the presence of 1-2 wt% H<sub>2</sub>O does not affect significantly the systematics of CO<sub>2</sub> solubility in basaltic melts up to an  $f\text{CO}_2=3000$  bar at least, which is the maximum  $f\text{CO}_2$  reached in our experiments, or that such an effect is not detectable with the analytical tools used to determine dissolved CO<sub>2</sub>. In contrast, the few data available at 5 kb for basalts equilibrated with mixed H<sub>2</sub>O-CO<sub>2</sub> fluids show a significant departure from linear behaviour between dissolved CO<sub>2</sub> and  $f\text{CO}_2$  (Botcharnikov et al., 2005; Behrens et al., 2009), which has been interpreted as resulting from the depolymerising effects of OH- groups which enhances CO<sub>2</sub> solubility (Mysen, 1976, see also Behrens et al. 2009).

In view of the above, we have modelled our experimental data using the same approach as Dixon et al. (1995), keeping in mind that our model is essentially calibrated on hydrous basalts with dissolved water contents up to 2 wt% at pressures below 3 kb. Application of the equations derived below to wetter melts may result in overestimated pressures of fluid saturation.

The solubility of carbon dioxide in basaltic liquids can be described by the following simple reaction (e.g. Stolper and Holloway, 1988):



where  $\text{O}^{2-}_{(m)}$  represents any reactive oxygen species,  $\text{CO}_{2,mol(v)}$  is the molecular carbon dioxide in the vapor phase, and  $\text{CO}_3^{2-}_{(m)}$  is the carbonate groups dissolved in the melt. The equilibrium constant at T and P is equal to:

$$K(P,T) = \frac{a_{CO_3^{2-}}^m(P,T)}{a_{O^{2-}}^m * f_{CO_2}(P,T)} \quad (22).$$

Assuming that  $O^{2-}_{(m)}$  and  $CO_3^{2-}_{(m)}$  mix ideally (Stolper et al., 1997), the activity of the melt species can be substituted by their mole fractions, so that:

$$K(P,T) = \frac{X_{CO_3^{2-}}^m(P,T)}{X_{O^{2-}}^m * f_{CO_2}(P,T)} \quad (23).$$

Then, standard thermodynamic manipulations (e.g., Holloway and Blank, 1994) show that the pressure and temperature effects on K are given by:

$$K(P,T) = K_0(P_0, T_0) \times \exp \left\{ - \int_{P_0}^P \frac{\Delta V_r^{0,m}}{RT} dP + \int_{T_0}^T \frac{\Delta H_{H_2O}^0(T, P_0)}{RT^2} dT \right\} \quad (24),$$

where  $K_0(P_0, T_0)$  is the equilibrium constant at  $P_0 = 1000$  bar and  $T_0 = 1200^\circ\text{C}$  the reference pressure and temperature,  $\Delta V_r^{0,m}$  is the volume change of the condensed components of reaction 15:

$$\Delta V_r^{o,m} = \left( V_{CO_3^{2-}}^{o,m} \right) - \left( V_{O^{2-}}^{o,m} \right) \quad (25),$$

where  $\left( V_{CO_3^{2-}}^{o,m} \right)$  and  $\left( V_{O^{2-}}^{o,m} \right)$  are the molar volume of species of the melt in their standard states.

$\Delta H_{H_2O}^0$  is the heat of solution of a mole of  $CO_2$  into the melt. Since we did not investigate specifically the temperature effect, an appropriate value of  $\Delta H_{H_2O}^0$  (-13.1 kJ/mol) is that of basanite which has a composition intermediate between tholeiitic basalt and nephelinite (Holloway and Blank, 1994).

According to reaction 24, a plot of  $\ln \left\{ \frac{f_{CO_2}^0}{X_{CO_2, tot}} \right\}$  versus  $\frac{(P - P_0)}{RT}$  obtained from  $CO_2$  solubility

data of experiments at constant temperature, results in a straight line whose slope gives  $\Delta V_r^{o,m}$  ( $X_{CO_2, tot}$  is the mole fraction of total  $CO_2$  contents dissolved in the melt; figure 7).

Values calculated for VES-9, PST-9 and ETN-1 are similar to values found in previous works (Table 4): from 21.5  $\text{cm}^3/\text{mol}$  for leucitite (Thibault and Holloway, 1994), to 28  $\text{cm}^3/\text{mol}$  for tholeiite (Stolper and Holloway, 1988). There is no clear dependence between  $\Delta V_r^{o,m}$  and the  $\Pi$  parameter, however. Considering the assumptions and uncertainties associated to the determination of  $\Delta V_r^{o,m}$  with the above method we can only say that in most basaltic liquids  $\Delta V_r^{o,m}$  is close to  $25 \pm 3 \text{ cm}^3/\text{mol}$ . It is also important to note that the values of  $\Delta V_r^{o,m}$

calculated here are for moderately hydrous basalts only and should not be interpreted as an accurate measurement of  $\Delta V_r^{o,m}$  (e.g., Zhang, 1999).

Then, knowing values of  $\Delta V_r^{o,m}$ , values of  $\ln K_0$  can be calculated. Values found are similar to those calculated in previous works (Table 4). The influence of melt composition on the solubility of CO<sub>2</sub> is borne out by the variability of  $K^\circ$  constants determined in this and other studies (Holloway and Blank, 1994). A plot of  $\ln K^\circ$  versus  $\Pi$  (Fig. 8) reveals that there exists a good linear relationship between these two parameters.

Linear regression of the data set yields the following equation:

$$\ln K^\circ = 0.893 \Pi - 15.247 \quad (26)$$

with a correlation coefficient  $R^2 = 0.891$ .

The above parameterisation of the  $K^\circ$  coefficient offers a simple model of calculating the solubility of CO<sub>2</sub> in basaltic melts with varying alkali contents. The results of back calculations of the CO<sub>2</sub> contents of our experimental quenched glasses performed with this model are shown against measured CO<sub>2</sub> values (Fig. 9). The model appears to retrieve observed CO<sub>2</sub> contents to within 10% relative.

### **Application to the calculation of saturation pressures for alkali-rich basaltic compositions**

Having determined solubility the laws of H<sub>2</sub>O (Lesne et al., this issue) and CO<sub>2</sub> in the three basaltic melts from Vesuvius, Etna and Stromboli, we are now able to calculate saturation pressures of melt inclusions analysed in related volcanic products. By determining the total water and CO<sub>2</sub> dissolved in the melt inclusions, we are able to calculate fugacities and partial pressures of those species, and then the saturation pressures at which volatiles were entrapped in the melt inclusion (ignoring the role of other volatiles, in particular sulphur and halogens).

For Vesuvius, we use the data of Marianelli et al. (1999, 2005) and Fulignati et al. (2004) to calculate saturation pressures. In those previous works, the quantities of dissolved water and carbon dioxide were determined via FTIR measurements. They considered the OH stretching vibration band at 3535 cm<sup>-1</sup> for measuring the water content, which was not used for measuring water content in our experimental glasses. CO<sub>2</sub> contents were calculated using the Dixon and Pan (1995) relationship (Marianelli et al., 2005) and using  $\epsilon_{1515} = 365$  l.mol<sup>-1</sup>.cm<sup>-1</sup> (Fulignati et al., 2004) and  $\epsilon_{1515} = 380$  l.mol<sup>-1</sup>.cm<sup>-1</sup> (Cioni, 2000). Those values of  $\epsilon_{1515}$  are not so different from that of Fine and Stolper (1986) used in this study, and translate into a maximum difference of 3% on the total carbon dissolved in the melt. Such a difference in CO<sub>2</sub> content will not induce a significant effect on the saturation pressures calculated. Both

Marianelli et al. (1999, 2005) and Fulignati et al. (2004) used VolatileCalc to calculate saturation pressures. Comparison of their findings with our own model shows that previous calculated pressures were overestimated by 50% on average (figure 10a).

For Etna we use the recent data of Métrich et al. (2004) and Spilliaert et al. (2006) who calculated saturation pressures with the VolatileCalc model (Newman and Lowenstern, 2002). To determine the CO<sub>2</sub> content of melt inclusions, Spilliaert et al. (2006) used a LPS nuclear microprobe, whilst for water they used a CAMECA IMS3f ion microprobe. Métrich et al. (2004) determined, instead, water and carbon contents by using the FTIR technique. As for Vesuvius, Métrich et al. (2004) determined water contents by using the total water band assignment and determined CO<sub>2</sub> content by using the molar extinction coefficient calculated from the Dixon and Pan (1995) relationship. Thus, CO<sub>2</sub> contents in melt inclusions and in our glasses are comparable, whereas the method used to calculate water content is not the same which could induce some differences in calculated water fugacities. Our calculated range of saturation pressures varies from 200 to 4200 bars, which is broadly similar to the range of pressures calculated by Métrich et al. (2004) and Spilliaert et al. (2006) (figure 10b).

For Stromboli, we refer to the data of Métrich et al. (2001) and Bertagnini et al. (2003) who calculated saturation pressures of melt inclusions by using Papale's (1999) model. Both determined water and total carbon contents in the melt inclusions by using FTIR and the relationship of Dixon and Pan (1995) to determine  $\epsilon_{1515}$ . Métrich et al. (2001) used extinction coefficient values for the 5200 and 4500 cm<sup>-1</sup> band assignments of Dixon et al. (1995), whereas Bertagnini et al. (2003) used the 3535 cm<sup>-1</sup> band assignment. Thus, the data of Métrich et al. (2001) are directly comparable to ours, contrary to the water contents of Bertagnini et al. (2003). However, as CO<sub>2</sub> largely predominates over H<sub>2</sub>O when calculating saturation pressures, comparison between our calculations and those of previous works can still be done faithfully. Métrich et al. (2001) and Bertagnini et al. (2003) found a range of saturation pressures from 2800 up to more than 4000 bars, with a cluster at 3500 bars. Both studies do not tabulate the calculated pressures, which prevents us from drawing a comparative plot such as done for Vesuvius and Etna. Our results yield a range of saturation pressures from 1500 to 3000 bars, with a cluster at 2500 bars, which is 1000 bars less than the cluster found by Bertagnini et al. (2003).

It thus appears that VolatileCalc overestimates the saturation pressures by 70% for Vesuvius basalt, by 26% for Stromboli and by 14% for Etnean basalt. Botcharnikov et al (2005) have already shown that VolatileCalc does not work properly at pressures significantly higher than 2 kb, tending also to overestimate fluid saturation pressures for common basalt-



andesite-rhyolite melts. The present study shows that a further limitation of this model is its failure in correctly predicting saturation pressures for alkali-rich mafic melts. A similar conclusion was reached by Behrens et al. (2009). For those melts, it appears that more complex models, such as that of Papale (1999), are needed to better account for the possible interactions between H<sub>2</sub>O and CO<sub>2</sub> in silicate melts (see for instance Tamic et al., 2001), in addition to other compositional effects.

## **Conclusions**

Experiments at 1200°C and pressures from 250 to 2086 bar have been performed to determine the solubilities of CO<sub>2</sub> in alkali basaltic melts.

1. Results of CO<sub>2</sub> contents presented in this study were measured by infrared measurements. Without any bulk carbon data, we used the molar extinction coefficient given by Fine and Stolper (1986) for basaltic melts, as well as molar extinction coefficients calculated from the melt composition given by Dixon and Pan (1995). Clearly, there is still a large uncertainty surrounding the determination of extinction coefficients in basaltic glasses and their variations with alkali and alkali-earth elements, which considerably hamper an accurate determination of dissolved CO<sub>2</sub> in basalts other than MORBs. Future work should be directed at a better definition of such a relationships
2. Aside from this problem, our experiments, along with previous similar works, show that the solubility of any particular melt can be modelled in a simple way. For the three worked compositions, such simple relationships are:

$$f_{CO_2}^{1200^\circ C} = 8.57 * 10^6 * X_{CO_2}^m{}^{1.26}; \quad R^2 = 0.99 \text{ for Vesuvius} \quad (27)$$

$$f_{CO_2}^{1200^\circ C} = 1.14 * 10^7 * X_{CO_2}^m{}^{1.23}; \quad R^2 = 0.93 \text{ for Etna} \quad (28)$$

$$f_{CO_2}^{1200^\circ C} = 3.49 * 10^6 * X_{CO_2}^m{}^{1.04}; \quad R^2 = 0.99 \text{ for Stromboli} \quad (29).$$

3. Experimental results obtained on these basaltic compositions show a strong effect of alkalis on CO<sub>2</sub> behaviour in these melts, confirming previous studies (Dixon, 1997; Behrens et al., 2009): at any pressure of the investigated range, alkali-rich basalts dissolve up to 1.7 times more CO<sub>2</sub> than alkali-poor basalts, and up to 2 times more than a MORB.

4. Data obtained in this study provide a refinement of the model of Dixon (1997): the parameterisation between the CO<sub>2</sub> solubility at 1 kbar and 1200°C and the  $\Pi$  factor is confirmed, and the extended data base yields: CO<sub>2</sub> (ppm) = 920 ( $\pm 1$ ) $\Pi$  - 17 ( $\pm 53$ ),
5. Comparing our data with those obtained in previous works, it appears that there exists a compositional dependence for  $\ln K^\circ$ , which varies linearly with the  $\Pi$  parameter, therefore providing a basis for a simple thermodynamical modelling of CO<sub>2</sub> solubility in basaltic melts.
6. Lastly, application of our results to the calculation of volatile saturation pressures using dissolved H<sub>2</sub>O and CO<sub>2</sub> contents of primitive melt inclusions from the three studied volcanoes shows that previous saturation pressures of primitive magmas feeding these volcanoes were significantly overestimated, notably for the Vesuvius.

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## FIGURE CAPTIONS

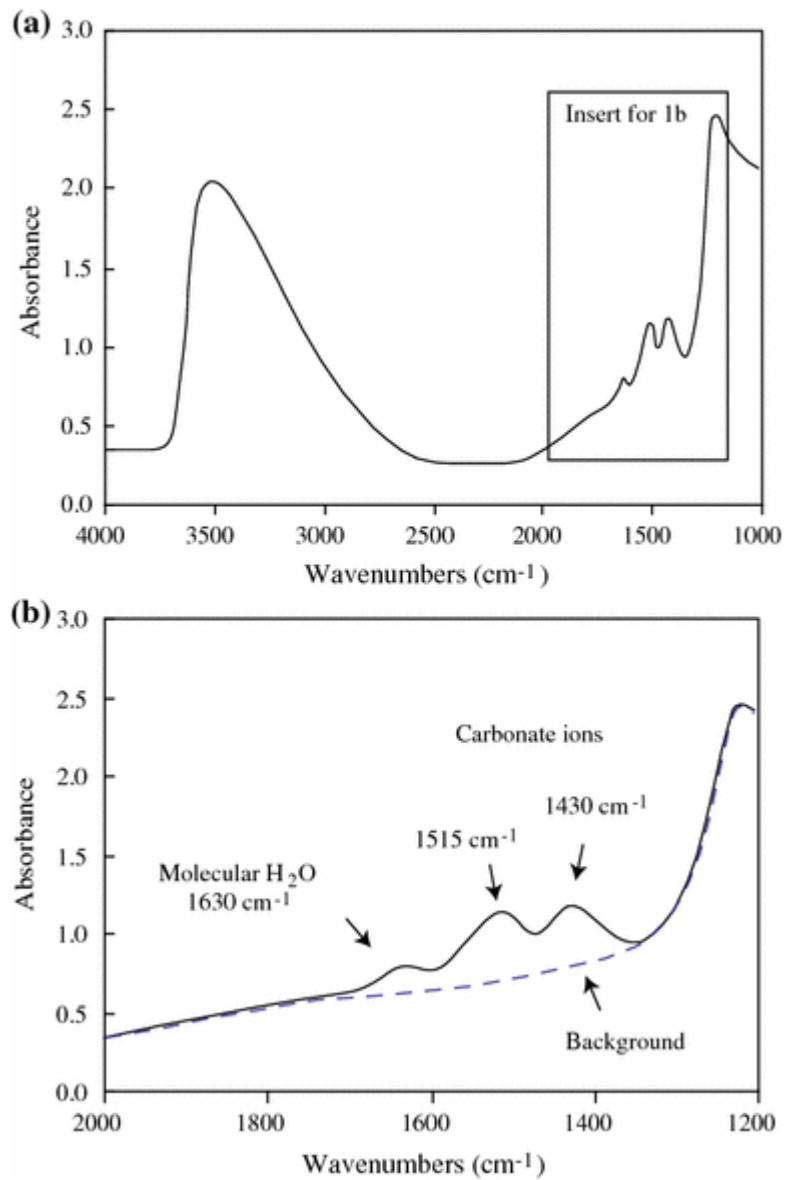


Figure 1:

Typical mid-IR spectrum of basaltic glass from experiment run1#4 (ETN-1) (thickness 133 $\mu$ m) showing the absorptions of carbonate bands (1515 cm<sup>-1</sup> and 1430 cm<sup>-1</sup>). Small peaks at 2350 cm<sup>-1</sup> are absorptions by atmospheric CO<sub>2</sub>. Background modelled as described in text.

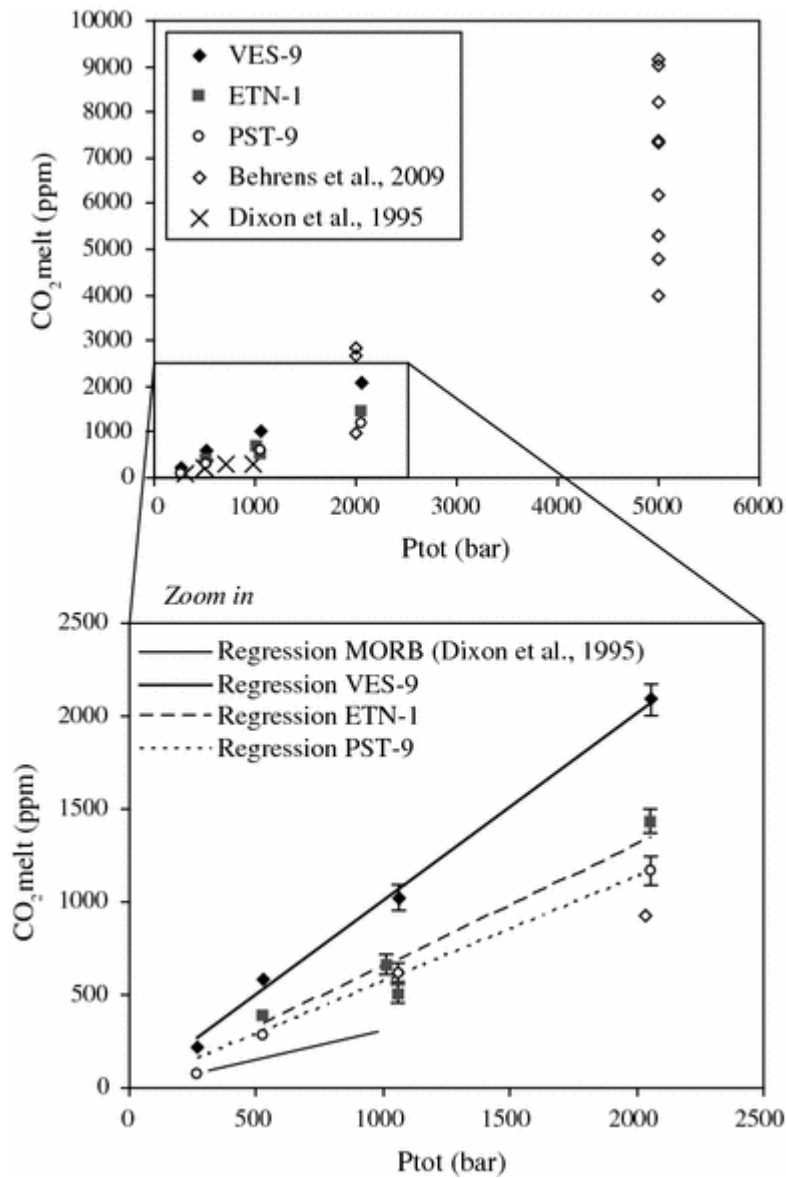


Figure 2:

Experimental results on  $CO_2$  contents in basaltic melts (Vesuvius, Etna, Stromboli and MORB from Dixon et al., 1995), versus experimental pressure.



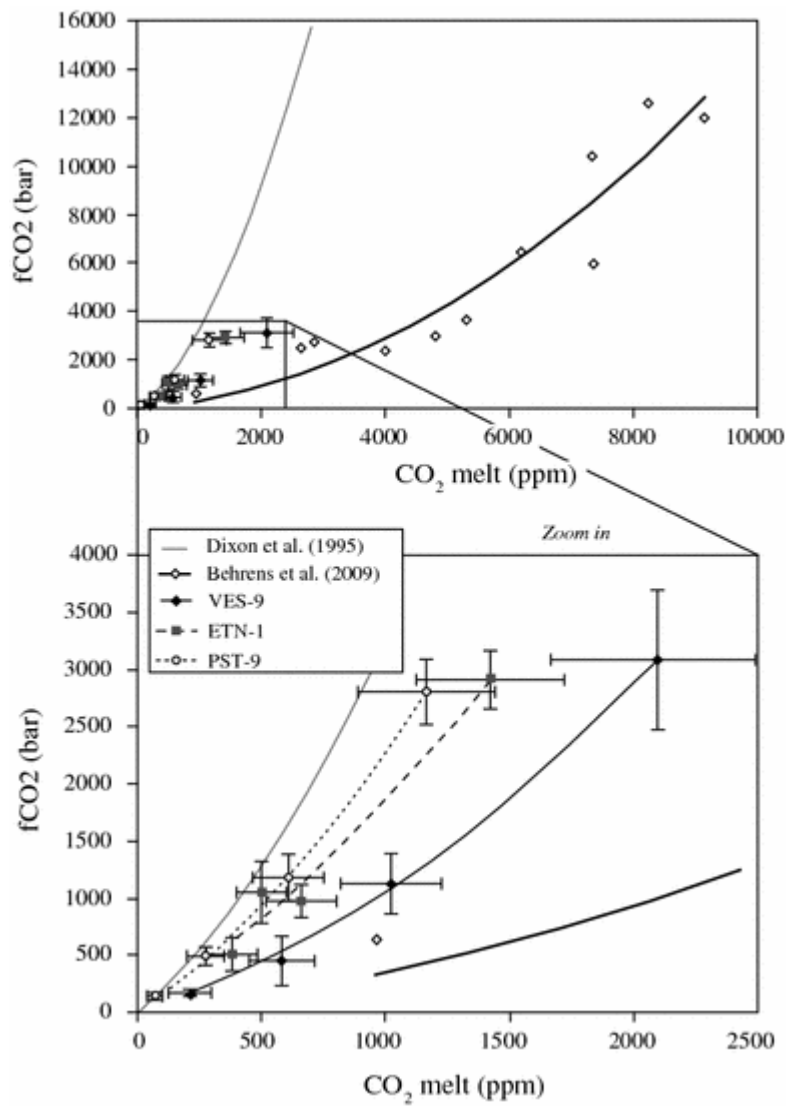


Figure 3:

Fugacity of CO<sub>2</sub> for VES-9, ETN-1 and PST-9, calculated with an MRK equation of state (Holloway, 1977), versus CO<sub>2</sub> dissolved in the melts in ppm. Model from Dixon et al. (1995) obtained on MORB are reported. Errors on  $f_{CO_2}$  are smaller than the symbols.

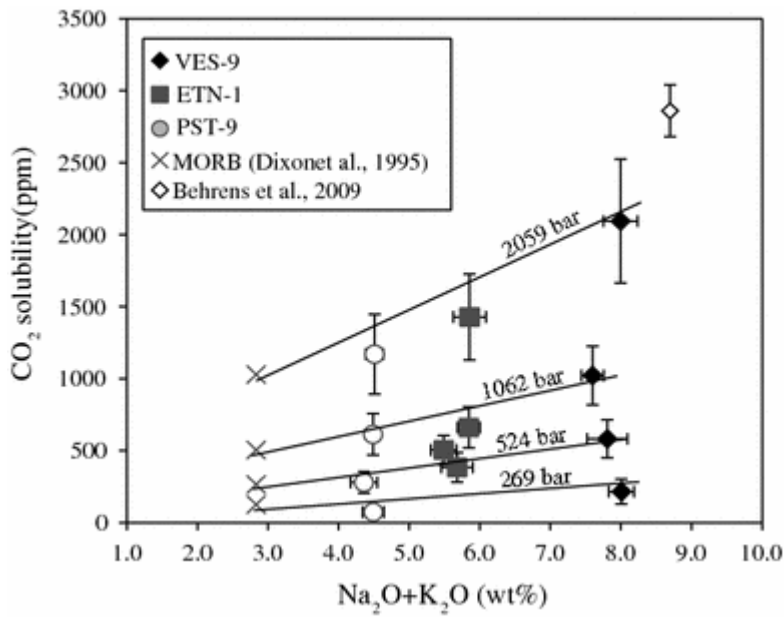


Figure 4:  
CO<sub>2</sub> solubility calculated at experimental pressures in basaltic melts against the alkalinity of the melt.

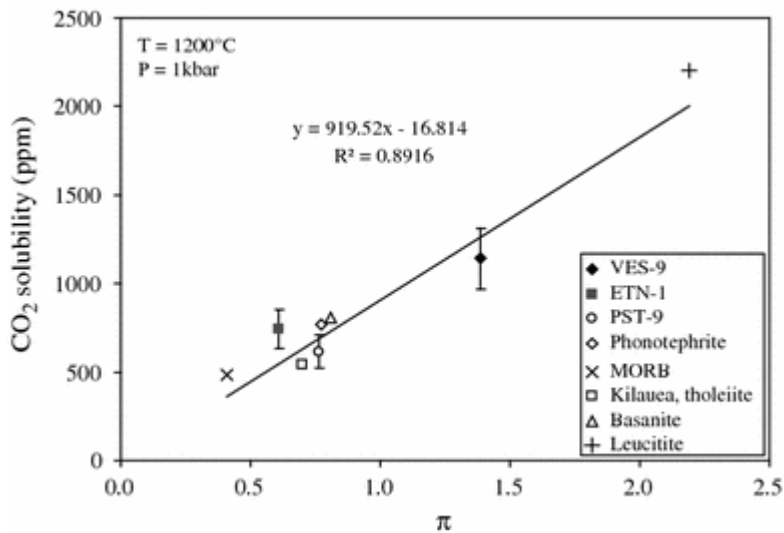


Figure 5:  
Saturation pressures calculated with VolatileCalc (Newmann and Lowenstern, 2002) for VES-9, ETN-1, PST-9, taking total SiO<sub>2</sub> measured in hydrous glass, at T = 1200°C, versus experimental pressures. Calculated saturation pressures of MORB (Dixon et al., 1995) were calculated and reported (crosses).

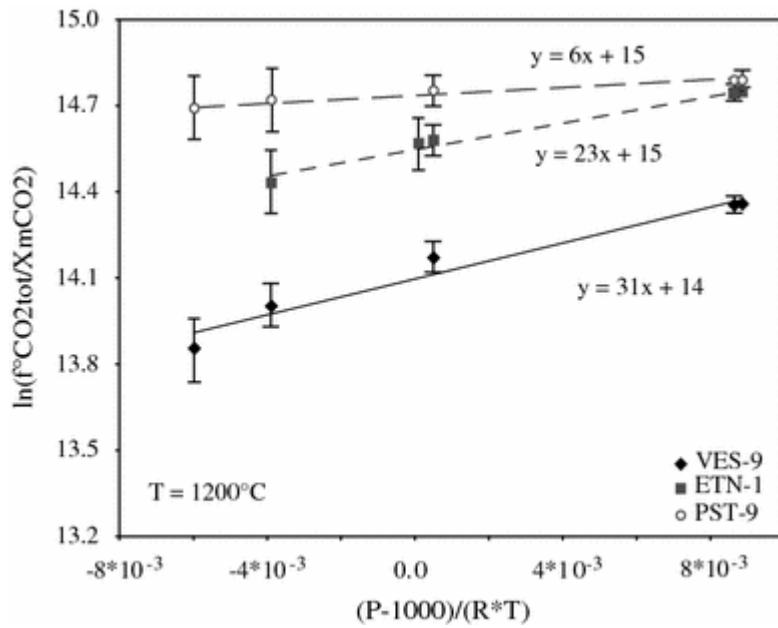


Figure 6:

CO<sub>2</sub> (ppm) solubility calculated for Vesuvius, Etna and Stromboli at 1200°C and 1 kbar, using solubility laws defined in this study, against the  $\Pi$  factor defined by Dixon (1997).

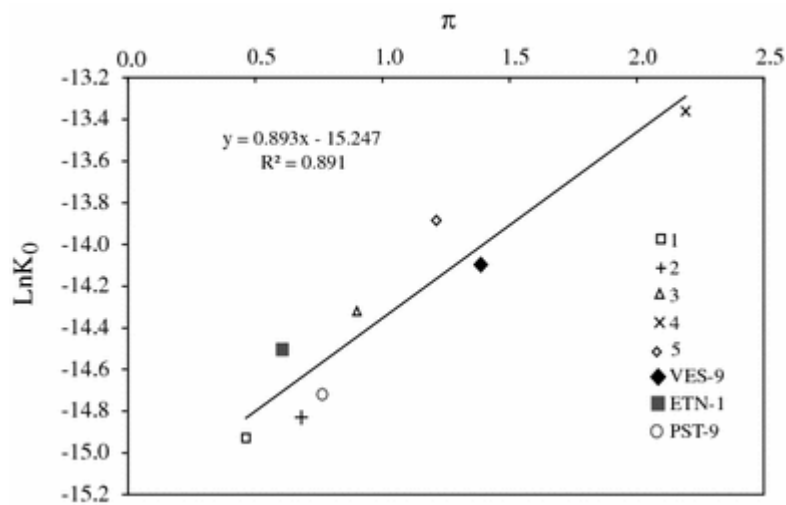


Figure 7:

Determination of  $\Delta V_r^{o,m}$  by fitting  $\ln \left\{ \frac{f_{CO_2}^0}{X_{CO_2,tot}} \right\}$  versus  $\frac{(P - P_0)}{RT}$ .

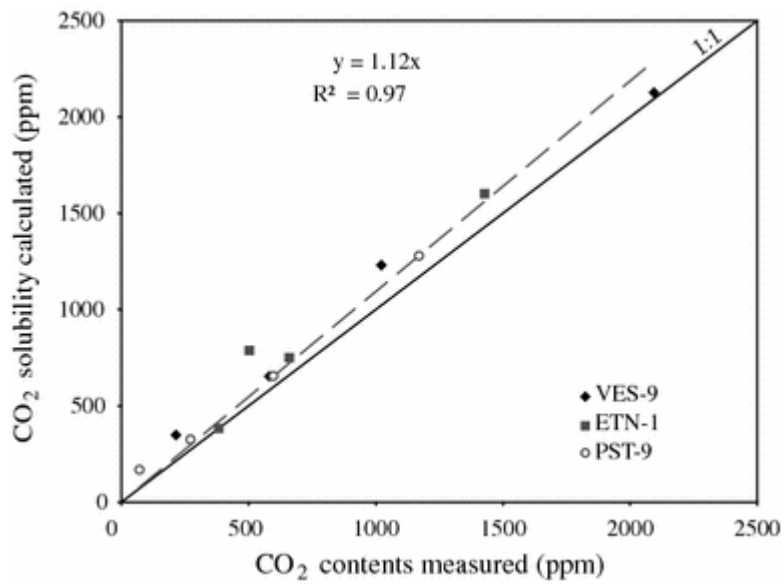


Figure 8:

LnK<sup>o</sup> versus Π parameter (Dixon 1997).

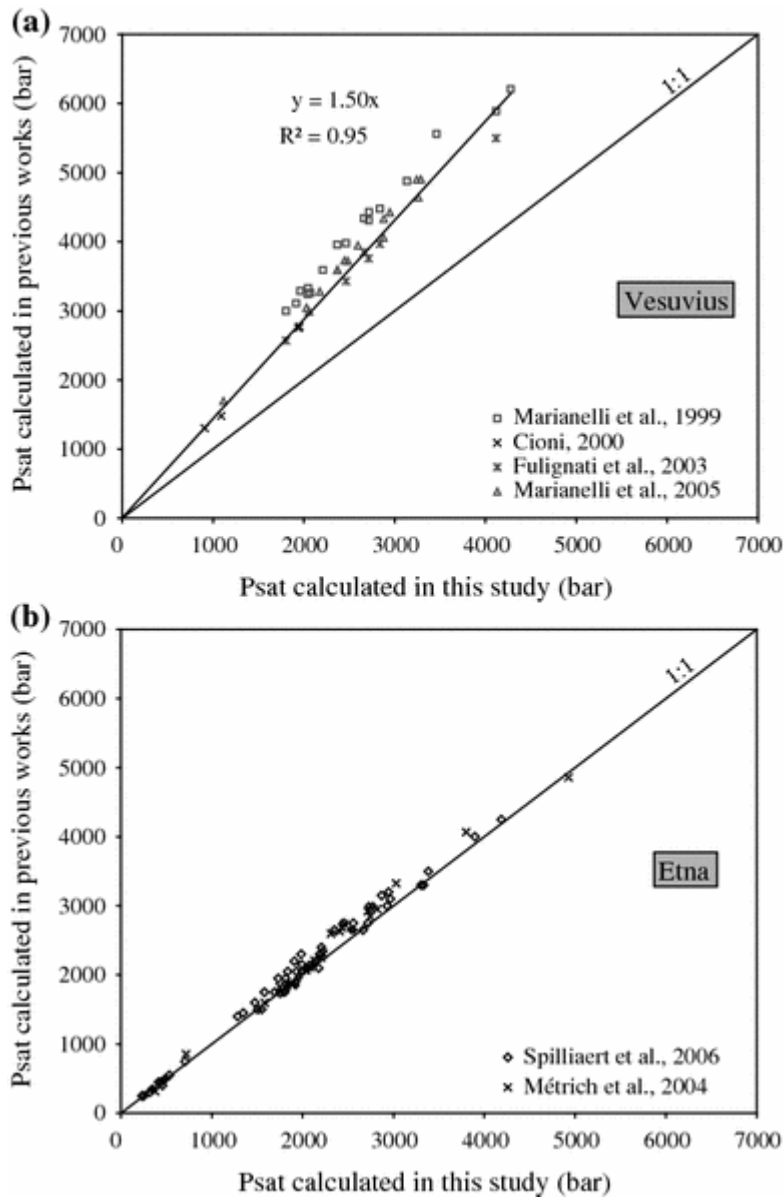


Figure 9:

Comparison between measured and calculated  $\text{CO}_2$  content with the thermodynamical parameters calculated in this study.

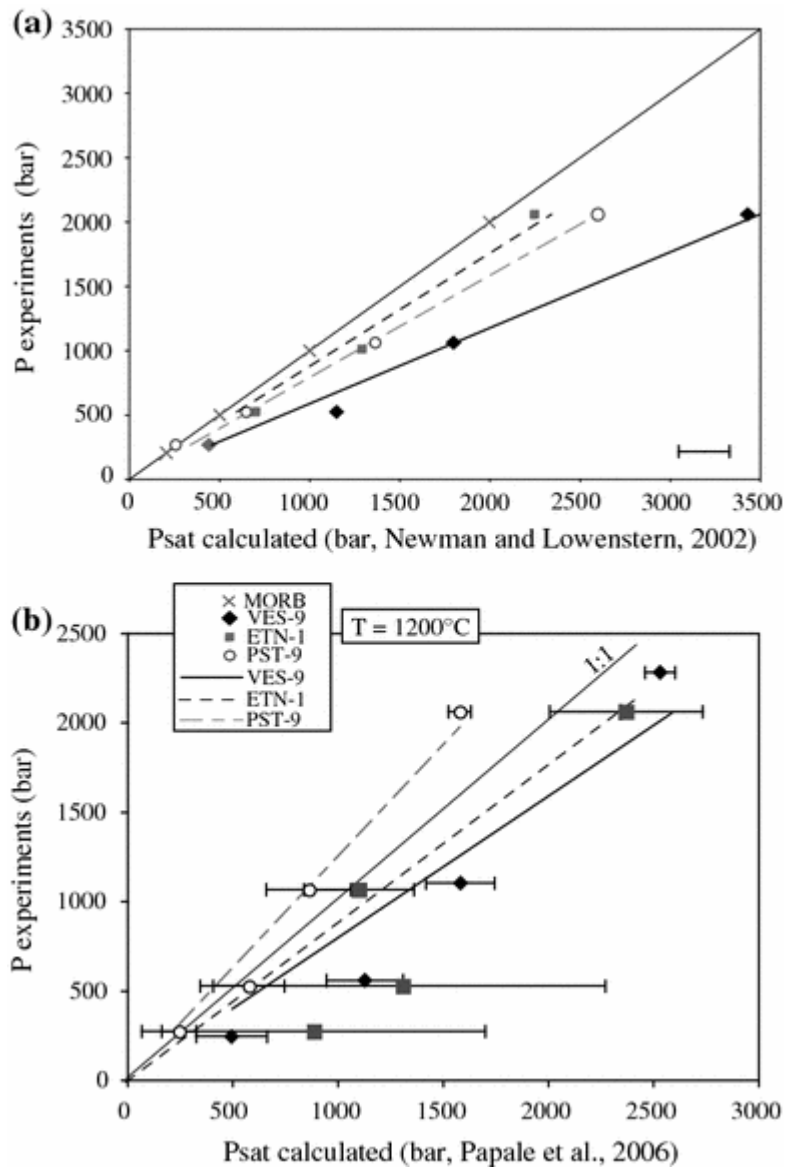


Figure 10:

Saturation pressures calculated by using data from melt inclusions of (a) Vesuvian products from Marianelli et al (1999), Cioni (2000), Fulignati et al. (2004) and Marianelli (2005) and of (b) Etnean products Métrich et al. (2004) and Spilliaert et al. (2006), compared to the saturation pressures calculated by using our solubility laws.

## Tables

Table 1 Bulk rock compositions

<b>Oxides</b>	<b>VES-9<sup>a</sup></b>	<b>ETN-1<sup>b</sup></b>	<b>PST-9<sup>c</sup></b>
SiO <sub>2</sub>	48.67	47.95	49.82
TiO <sub>2</sub>	0.97	1.67	0.81
Al <sub>2</sub> O <sub>3</sub>	14.72	17.32	15.94
FeO	7.62	10.24	7.71
MnO	0.14	0.17	0.20
MgO	6.82	5.76	8.07
CaO	12.94	10.93	12.81
Na <sub>2</sub> O	1.82	3.45	2.32
K <sub>2</sub> O	5.63	1.99	1.92
P <sub>2</sub> O <sub>5</sub>	0.66	0.51	0.40
Total	100.00	100.00	100.00

<sup>a</sup>Natural alkali basalt from VES-9

<sup>b</sup>Natural alkali basalt from ETN-1

<sup>c</sup>Natural alkali basalt from PST-9

Table 2 Experimental conditions and FTIR results for H<sub>2</sub>O and CO<sub>2</sub> dissolved in experimental glasses

Run	Composition	Capsule	wt% H <sub>2</sub> O loaded	wt% CO <sub>2</sub> loaded	% FeO kept <sup>a</sup>	wt% H <sub>2</sub> O <sub>tot</sub> <sup>b</sup>	CO <sub>2</sub> <sup>c</sup> (ppm)	CO <sub>2</sub> <sup>d</sup> (ppm)
Experiment 1: 2,059 bars, duration: 4 h								
Run1#2	VES-9	Pt	0	3.93	68	0.80 ± 0.40	2,094 ± 431	2,061 ± 431
Run1#4	ETN-1	Pt	0	4.15	79	1.49 ± 0.13	1,429 ± 299	1,638 ± 299
Run1#6	PST-9	Pt	0	4.35	70	1.58 ± 0.16	1,170 ± 277	1,196 ± 277
Experiment 2: 1,013 bars, duration: 8 h								
Run2#8	ETN-1	Pt	0	1.91	70	1.55 ± 0.23	661 ± 142	757 ± 142
Experiment 3: 524 bars, duration: 4 h								
Run3#4	VES-9	Pt	0	3.93	83	0.82 ± 0.37	582 ± 132	572 ± 132
Run3#5	ETN-1	Pt	0	3.94	83	0.73 ± 0.20	383 ± 103	440 ± 103
Run3#6	PST-9	Pt	0	3.50	93	0.71 ± 0.10	279 ± 75	285 ± 75
Experiment 4: 269 bars, duration: 4 h								
Run4#4	VES-9	Pt	0	1.24	69	0.85 ± 0.08	215 ± 87	212 ± 87
Run4#5	ETN-1	Pt	0	1.57	81	0.95 ± 0.10	n.d.	n.d.
Run4#6	PST-9	Pt	0	1.00	72	0.98 ± 0.07	73 ± 31	75 ± 31
Experiment 19: 1,062 bars, duration: 5 h								
Run19#1	VES-9	Pt	0	3.16	74	1.02 ± 0.23	1,022 ± 204	1,006 ± 204
Run19#2	ETN-1	Pt	0	2.74	74	1.51 ± 0.39	504 ± 101	578 ± 101
Run19#3	PST-9	Pt	0	2.59	71	0.91 ± 0.15	613 ± 144	626 ± 144

<sup>a</sup>Proportion of FeO kept in the glass relative to the initial glass FeO concentration, calculated as  $100 \cdot \text{FeO}_{\text{glass}} / \text{FeO}_{\text{starting glass}}$ , with  $\text{FeO}_{\text{starting glass}}$  from Table 1

<sup>b</sup>H<sub>2</sub>O determined by NIR-FTIR measurements (see text for explanation)



$^{13}\text{C}$  determined by MIR-FTIR measurements and calculated with Beer–Lambert law, using  $\epsilon_{1515}$  from Fine and Stolper (1986)

$^{13}\text{C}$  determined by MIR-FTIR measurements and calculated with Beer–Lambert law, using  $\epsilon_{1515}$  from Dixon and Pan (1995)



Table 3 Experimental results obtained on basalt from (a) Vesuvius, (b) Etna and (c) Stromboli

$P_{tot}$ (bar)	Run	Na <sub>2</sub> O + K <sub>2</sub> O	Density <sup>a</sup> (g/l)	% FeO kept <sup>b</sup>	wt% H <sub>2</sub> O <sub>mol</sub> <sup>c</sup>	wt% OH <sup>c</sup>	wt% H <sub>2</sub> O <sub>tot</sub> <sup>c</sup>	CO <sub>2</sub> (ppm) <sup>d</sup>	e	f	g	g	h	h	h	h	i
(a)																	
2,059	Run1#2	7.99	2,696	68.36	0.36	0.44	0.80	2,094	87	3,128	0.041	0.959	1.61	0.73	1.76	0.17	-5.991
		0.24	3		0.07	0.01	0.16	431	17	612	0.008						
524	Run3#4	7.81	2,692	83.10	0.27	0.55	0.82	582	91	482	0.174	0.826	1.65	0.55	2.20	0.05	-5.952
		0.28	2		0.12	0.02	0.37	132	40	217	0.078						
269	Run4#4	8.01	2,692	69.32	0.27	0.58	0.85	215	96	183	0.359	0.641	1.71	0.53	2.35	0.02	-5.903
		0.18	1		0.02	0.04	0.08	87	9	30	0.048						
1,062	Run19#1	7.60	2,689	73.71	0.36	0.66	1.02	1022	131	1,165	0.124	0.876	2.05	0.72	2.67	0.08	-5.634
		0.16	2		0.07	0.06	0.23	204	29	265	0.028						
(b)																	
2,059	Run1#4	5.86	2,680	79.40	0.69	0.81	1.49	1429	216	2,926	0.103	0.897	2.99	1.37	3.22	0.12	-5.198
		0.23	1		0.06	0.02	0.13	299	19	256	0.009						
1,013	Run2# 8	5.84	2,679	69.84	0.80	0.74	1.55	661	231	967	0.229	0.771	3.09	1.61	2.97	0.05	-5.141
		0.16	2		0.10	0.06	0.23	142	34	145	0.034						
524	Run3#5	5.68	2,694	82.50	0.34	0.38	0.73	383	58	519	0.111	0.889	1.47	0.69	1.55	0.03	-6.341
		0.22	1		0.09	0.02	0.20	103	16	145	0.031						
269	Run4#5	5.57	2,690	81.42	0.55	0.41	0.95	n.d.	95	183	0.356	0.644	1.92	1.10	1.64	n.d.	-5.909
		0.18	1		0.04	0.03	0.10		10	31	0.051						
1,062	Run19#2	5.49	2,679	73.92	0.74	0.77	1.51	504	222	1,050	0.210	0.790	3.03	1.48	3.09	0.04	-5.175
		0.18	4		0.19	0.04	0.39	101	57	272	0.054						
(c)																	
2,059	Run1#6	4.51	2,678	69.61	0.66	0.92	1.58	1,170	285	2,820	0.136	0.864	3.16	1.32	3.69	0.10	-4.959

$P_{\text{tot}}$ (bar)	Run	Na <sub>2</sub> O + K <sub>2</sub> O	Density <sup>a</sup> (g/l)	% FeO kept <sup>b</sup>	wt% H <sub>2</sub> O <sub>mol</sub> <sup>c</sup>	wt% OH <sup>c</sup>	wt% H <sub>2</sub> O <sub>tot</sub> <sup>c</sup>	CO <sub>2</sub> (ppm) <sup>d</sup>	e	f	g	g	h	h	h	h	i
		0.12	1		0.06	0.03	0.16	277	28	286	0.014						
524	Run3#6	4.36	2,694	93.22	0.30	0.41	0.71	279	71	504	0.136	0.864	1.43	0.61	1.66	0.02	-6.165
		0.19	1		0.04	0.01	0.10	75	10	81	0.021						
269	Run4#6	4.49	2,689	71.69	0.32	0.66	0.98	73	124	152	0.465	0.535	1.98	0.65	2.65	0.01	-5.678
		0.15	1		0.06	0.02	0.17	31	22	33	0.091						
1,062	Run19#3	4.49	2,691	71.07	0.34	0.57	0.91	613	109	1,192	0.103	0.897	1.83	0.69	2.28	0.05	-5.791
		0.10	0		0.05	0.04	0.15	144	18	205	0.018						

Table 4 Best-fit thermodynamic parameters for Eq. 17

Composition	$\Delta V^0$	$\ln K_0$	$P_0$ (bar)	$T_0$ (°C)
VES-9	31	$-14.10 \pm 0.03$	1,000	1,200
ETN-1	23	$-14.55 \pm 0.00$	1,000	1,200
PST-9	6	$-14.74 \pm 0.01$	1,000	1,200

	V carbonate				
Thibault et al. (1994)	Leucitite	22			
Pan et al. (1991)	Tholeiite	$23 \pm 1.03$	$-14.83 \pm 0.1$	1,000	1,200
Stolper and Holloway (1988)	Tholeiite	$33 \pm 0.5$	$-15.15 \pm 0.06$	1,000	1,200
Stolper et al. (1987)	NaAlSi <sub>3</sub> O <sub>8</sub>	25			
Spera and Bergman (1980)	Andesite CA	34			
Spera and Bergman (1980)	Tholeiite K1921	33			
Spera and Bergman (1980)	Olivine Melilite OM1	23			

