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Tyrone O. Rooney, Barry Hanan, David W. Graham, Tanya Furman, Janne Blichert-Toft, et al.. Upper Mantle Pollution during Afar Plume-Continental Rift Interaction. Journal of Petrology, 2012, 53 (2), pp.365-389. 10.1093/petrology/egr065 . insu-00673476

# HAL Id: insu-00673476 https://insu.hal.science/insu-00673476

Submitted on 20 Feb 2013

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# Upper Mantle Pollution during Afar Plume-Continental Rift Interaction

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### RECEIVED JULY 2, 2010; ACCEPTED NOVEMBER 9, 2011 **ADVANCE ACCESS PUBLICATION DECEMBER 18, 2011**

New Pb, Sr, Nd, Hf, and He isotope data for Quaternary basalts, erupted from Debre Zeyit, Butajira, and the Wonji Fault Belt of the Main Ethiopian Rift, show systematic mixing relationships involving three distinct mantle sources. The Pb, Sr, Nd, and Hf isotopic arrays converge in a specific region of isotopic multi-space where they define the composition of the Afar mantle plume (centered about  ${}^{206}Pb/{}^{204}Pb = 19.5$ ,  ${}^{87}Sr/{}^{86}Sr = 0.7035$ ,  $\varepsilon_{Nd} = +4.6$ ,  $\varepsilon_{Hf} =$ +9.3,  ${}^{3}He/{}^{4}He > 15 R_{A}$ ). This plume end-member has an identical composition to that observed previously in oceanic basalts. The distinct isotopic arrays for the various volcanic areas in the Main Ethiopian Rift vary spatially in a systematic manner, and may be viewed as pseudo-binary mixing arrays. This further suggests that the Afar mantle plume interacts with the local continental lithosphere and upper mantle asthenosphere (mid-ocean ridge basalt-like source) through an ordered sequence of mixing events. Simple mixing models require that the mass proportions of continental lithosphere and upper mantle involved in magma generation must be nearly constant within each volcanic area, but that the proportion of plume material decreases regularly with distance southwestward along the Main Ethiopian Rift, away from the central axis of the plume. This systematic behavior means that continental lithosphere can become detached and mixed into the shallow mantle prior to the flow of upwelling plume material beneath the developing rift system. Detachment and mixing into the asthenosphere during continental rift evolution is an important process for producing the range of ambient upper mantle compositions sampled by mid-ocean ridge volcanism away from island hotspots.

KEY WORDS: Ethiopian Rift; Afar mantle plume; continental lithosphere; Pb, Sr, Nd, Hf, He isotopes

#### INTRODUCTION

Mantle plumes, anomalously hot and buoyant upwellings from the deep mantle, can strongly influence the chemical composition of basalts erupted at spreading ocean ridges (Morgan, 1972; Schilling, 1973a, 1985; Hanan et al., 1986; Detrick et al., 2002) as well as those in continental settings (Morgan, 1971; Burke & Wilson, 1972; Watts, 1976; Crough, 1978; Courtney & White, 1986; Campbell & Griffiths, 1990; McNutt & Judge, 1990; Griffiths & Campbell, 1991; Farnetani & Richards, 1994; Ingle et al., 2002; Ernst & Buchan, 2003; Jones & White, 2003; Camp & Ross, 2004; Pik et al., 2006; Saunders et al., 2007; Hanan et al., 2008; Camp & Hanan, 2008; Shervais & Hanan, 2008; Graham et al., 2009). Much of the spatial variation of Pb, Nd, Sr, and He isotopes in mid-ocean ridge basalts (MORB) from the South Atlantic appears to have been produced during broad radial dispersion of isotopically distinct

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plume material into the asthenosphere during rifting and formation of the ocean basin, prior to when the plumes were over-ridden (in the fixed hotspot reference frame; Duncan, 1981) by the migrating Mid-Atlantic Ridge (Hanan et al., 1986; Graham et al., 1992). Thus, one explanation for the observed range of upper mantle compositions is that it reflects isotopic heterogeneity introduced by hotspots derived from distinct source reservoirs in the deep mantle (e.g. Storey et al., 1989). However, geochemical studies of MORB and continental basalts also indicate that the upper mantle has been polluted by continental lithosphere and/or lower crust (Arndt & Goldstein, 1989; Kamenetsky et al., 2001; Hanan et al., 2004; Carlson et al., 2005; Escrig et al., 2005; Class & le Roex, 2006; Goldstein et al., 2008). The spectrum of isotopic signatures observed in ocean island (hotspot) basalts, ranging between EM I, EM II and HIMU (Zindler & Hart, 1986), is also observed in basalts and xenoliths that sample the continental lithosphere and lower continental crust (e.g. McKenzie & O'Nions, 1995). Therefore, a second explanation for the origin of isotopic heterogeneity in the upper mantle is that it reflects direct pollution and mixing with detached continental lithosphere. Here we address these two different possibilities (deep vs shallow sources of mantle isotopic heterogeneity) through isotopic characterization and modelling of the sources involved in plume-lithosphere interaction along the Main Ethiopian Rift. The Main Ethiopian Rift is the type example of a modern continental rift in a plume-influenced environment (Falcon et al., 1970; Girdler, 1970; Schilling, 1973b), and provides a natural laboratory for studying the interaction of a mantle plume with the continental lithosphere.

The Red Sea, Gulf of Aden, and Main Ethiopian Rift meet in the Afar triangle, where they delineate three arms of a rift system that began at ~30 Ma over the upwelling Afar mantle plume (Fig. 1; Hofmann *et al.*, 1997). The Main Ethiopian Rift is the least evolved of the three triple junction arms, is about 18 Myr younger than the Red Sea and Gulf of Aden rifts, and is thus not a primary feature of the Afar plume, but is instead related to plate kinematics (Wolfenden *et al.*, 2004). The Main Ethiopian Rift began to propagate north from the Turkana region, a zone previously rifted in the Mesozoic and again in the Cenozoic, at ~11 Ma, coincident in time with the propagation of the Kenya Rift (Ebinger *et al.*, 2000).

Previous studies have described the isotopic diversity of the Main Ethiopian Rift magmas in terms of mixing between plume, depleted upper mantle, and continental lithosphere (Deniel et al., 1994; Barrat et al., 1998; Trua et al., 1999; Rogers et al., 2000; Furman, 2007), but only in very broad terms. Although the timing, extent and mechanisms of mixing between these mantle reservoirs are central to the development of an integrated geodynamic model of the rifting process, they have remained poorly

understood. Schilling (1973b) and Schilling et al. (1992) reported rare earth element and Nd–Sr–Pb isotopic compositions of basalts dredged from along the Sheba Ridge in the Gulf of Aden and its extension into the Gulf of Tadjoura, and of subaerial basalts from the Asal Rift in eastern Afar (Fig. 1). This dataset provided an initial framework for interpreting melting processes associated with interaction of the Afar plume with the continental lithosphere and an emerging ocean basin. The mixing relationships (Schilling et al., 1992) define three mantle source end-members interpreted as Pan-African continental lithosphere, the Afar mantle plume, and the depleted upper mantle (MORB asthenosphere source).

In this study we refine and extend the original three-component mixing model of Schilling *et al.* (1992) to the Butajira, Debre Zeyit, and Wonji Fault Belt regions of the Main Ethiopian Rift, three extensional zones that together form a 250 km transect along the Main Ethiopian Rift (Fig. 1, Table 1). We focus on Quaternary basalts having ages <1 Ma, and present new Pb, Nd, Sr, Hf, and He isotope results, as well as new Hf isotope data for the same Gulf of Aden sample suite previously analyzed for Pb, Nd, and Sr isotopes by Schilling *et al.* (1992; Table 2). Collectively, we use these new data to quantitatively characterize the distribution of mantle source reservoirs along the Main Ethiopian Rift portion of the East African Rift, their relative contribution to volcanism, and their sequence of mixing.

# BACKGROUND

# Sample selection

A subset of the most primitive Main Ethiopian Rift samples (6-10 wt % MgO) previously analyzed for major and trace elements (Rooney et al., 2005, 2007, 2011; Furman et al., 2006; Rooney, 2010) were selected for Hf, Sr, Nd, Pb (n=15), and He (n=12) isotopic analysis. The major element and trace element data for two samples, WFB-1026 and WFB-N-14, that were analyzed but not reported by Furman et al. (2006), are reported in Appendix Table Al. Hf isotope data are also reported for a suite of basalts dredged from the Sheba Ridge axis in the Gulf of Aden and its extension into the Gulf of Tadjoura, and for subaerial basalts from the Asal Rift of eastern Afar. The Pb, Sr, and Nd isotope data for these samples were reported by Schilling et al. (1992), and are the basis for the three-component mixing model proposed therein. The analytical details for the Main Ethiopian Rift and Gulf of Aden isotope data are given in the legends of Tables 1–3. The complete dataset is provided as Supplementary Data (available for downloading at http://www.petrology. oxfordjournals.org/).

Crustal contamination indicators (e.g. La/Nb and Ce/Pb) for the samples selected for isotopic analysis and modeling show no evidence for significant crustal assimilation

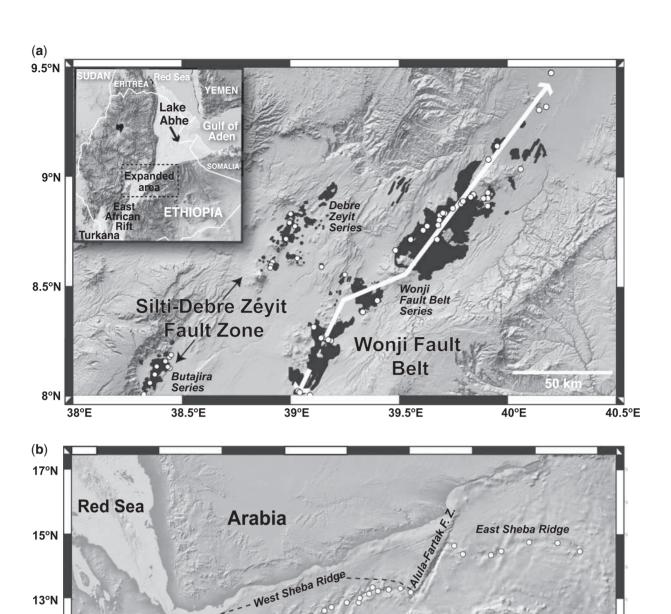


Fig. 1. (a) Map of the Main Ethiopian Rift study area with Quaternary sample locations shown along the Wonji Fault Belt, Debre Zeyit, and Butajira extensional zones. The Debre Zeyit, Butajira, and Wonji Fault Belt zones contain widely (20–40 km) spaced felsic volcanoes and calderas. Samples are from smaller basaltic cinder cones, associated flows, and occasional fissure eruptions that are interspersed between these volcanoes. The inset shows the study area location relative to the Afar triple junction. (b) Map of the Gulf of Aden showing the sample locations of Schilling et al. (1992). Backgrounds are reflectance images of East Africa and Yemen from the NASA/JPL Shuttle Radar Tomography Mission (USGS, 2004). The topography and bathymetry are derived from Geomapapp (http://www.geomapapp.org; Ryan et al., 2009).

49°E

51°E

100 km

55°E

53°E

**Gulf of** 

Aden

**Africa** 

47°E

Gulf of Tadjoura

45°E

11°N

9°N

Afar

43°E

41°E

Table 1: Location and isotope data for basalts from the Main Ethiopian Rift

Location	Lat. (°N)	Long. (°E)	Distance*	<sup>206</sup> Pb/ <sup>204</sup> Pb	$2\sigma\sqrt{n}$	<sup>207</sup> Pb/ <sup>204</sup> Pb	<b>2</b> σ√ <i>n</i>	<sup>208</sup> Pb/ <sup>204</sup> Pb	2σ√ <i>n</i>	<sup>87</sup> Sr/ <sup>86</sup> Sr	<b>2</b> σ√ <i>n</i>	<sup>143</sup> Nd/ <sup>144</sup> Nd	<b>2</b> σ√ <i>n</i>	$\epsilon_{Nd}$	<sup>176</sup> Hf/ <sup>177</sup> Hf	2σ√ <i>n</i>	$\epsilon_{Hf}$
																	—
Debre Zeyit DZ-1003	8-84	39.00	377	18-1518	6	15-5978	7	38-3801	10	0.704539	2†	0.512631	6	<b>−</b> 0·1	0.282839	6	2.4
	8.80								19				7	3.0	0.282839	О	2.4
DZ-1004	8.80	39.00	380	18·5096	10	15·5891	10 9	38.6038	27	0.703833	14	0.512794	7	3·0 2·4			
DZ-1005 DZ-1007	8.72	39·00 38·99	380 388	18·3151 18·4088	10 7	15·5782 15·5723	5	38·4726 38·5149	19	0·703905 0·703909	8† 5*	0·512763 0·512784	7	2.4	0.282997	4	8.0
							5 7		13		5†		-			-	
DZ-1008	8.72	38.98	387	18·4949	9	15·5716	-	38.5552	21	0.703810	2†	0.512808	8	3.3	0.283019	3	8.7
DZ-1009	8.59	39.15	386	18.5007	8	15·5735	7	38.5330	17	0.704173	7† 7±	0.512805	6	3·3 2·7	0.283039	4	9.4
DZ-1013	8-62	38.92	401	18-3893	6	15.5742	6 7	38-4825	18	0.703978	7† 1÷	0.512775	8		0.000007		9.0
DZ-1014	8-60	38-91	402	18·5816	8	15·5714	/	38.6030	17	0.703825	1†	0.512850	6	4.1	0.283027	4	9.0
Butajira	0.40	00.45	470	10.5100	0	45 5504	-	00.4404	10	0.700050	47	0.540000	0	0.0			
BJ-1053	8-19	38.45	470	18-5166	6	15.5594	5	38-4494	12	0.703652	17	0.512803	9	3.2	0.000007		
BJ-1051	8-16	38-43	475	18-4916	9	15.5519	8	38-4200	19	0.703631	10	0.512839	6† -	3.9	0.283007	4	8.3
BJ-1043	8-14	38-44	476	18-4300	12	15.5519	11	38-3816	28	0.703719	8	0.512804	7	3.2	0.282993	5	7.8
BJ-1049	8-14	38.39	480	18-2712	9	15.5433	8	38-1914	20	0.703582	8	0.512790	6	3.0			
BJ-1044	8.10	38.38	483	18.5186	7	15.5534	6	38-4340	13	0.703647	11	0.512837	6	3.9	0.282997	4	8.0
BJ-1045	8.06	38.36	488	18-5223	7	15.5602	6	38-4697	16	0.703606	14	0.512820	7† -	3.6	0.282973	4	7.1
BJ-1048	8.00	38-33	495	18-1741	10	15.5389	13	38·1730	22	0.703786	15	0.512808	7	3.3			
Wonji Fault B	•																
WFB-1026	9.48	40.20	238	18-9194	6	15.5803	5	38-8592	13	0.703681	9†	0.512896	6	5.0			
WFB-1023	9.14	39-95	284	18·7524	5	15·5837	4	<i>38</i> · <i>7819</i>	11	0.704102		0.512845	7	4.0	0.283049	3	9.8
WFB-1027	9.48	40-20	238	18·8272	3	<i>15</i> · <i>5807</i>	4	38-8958	7	0.703722		0.512890	28	4.9	0.283075	4	10.7
WFB-E99-2	8-91	39.82	313	18-6022	12	15.5854	13	38-7011	25	0.704117		<i>0</i> · <i>512826</i>		3.7			
WFB-N-18	8-91	39.83	313	18·7317	6	<i>15</i> ·5969	6	<i>38</i> · <i>8345</i>	17	0.703961		<i>0</i> · <i>512827</i>		3.7	0.283019	4	8.7
WFB-1037	8.86	39.75	323	18·5571	4	15·5840	4	<i>38</i> · <i>6318</i>	9	0·704251		0.512796	13	3.1	0.283009	5	8.4
WFB-KO-12	8.82	39.68	330	18·6577	10	15·5934	8	<i>38</i> · <i>8260</i>	23	<i>0</i> · <i>704303</i>		<i>0</i> ·512793		3.0			
WFB-KO-4B	8.82	39.68	330	18·8011	6	<i>15</i> · <i>5960</i>	5	<i>38</i> · <i>8717</i>	17	<i>0</i> ·704015		0.512829		3.7			
WFB-N-20	8.83	39.69	329	18·8109	6	15·5923	5	<i>38</i> · <i>8730</i>	17	0.703996							
WFB-1035	8.80	39.68	332	18·6267	4	<i>15</i> ·5847	4	<i>38</i> · <i>74</i> 97	8	0.704198		<i>0</i> ·512795	16	3.1	0.283007	5	8.3
WFB-N-22	8.72	39.68	340	18·9449	14	15·5948	16	38-8959	56	<i>0</i> ·703768		<i>0</i> ·512866		4.4	0.283056	3	10.0

(continued)

Table 1: Continued

Location	Lat. (°N)	Long. (°E)	Distance* (km)	<sup>206</sup> Pb/ <sup>204</sup> Pb	2σ√ <i>n</i>	<sup>207</sup> Pb/ <sup>204</sup> Pb	2σ√ <i>n</i>	<sup>208</sup> Pb/ <sup>204</sup> Pb	2σ√ <i>n</i>	<sup>87</sup> Sr/ <sup>86</sup> Sr	2σ√ <i>n</i>	<sup>143</sup> Nd/ <sup>144</sup> Nd	2σ√ <i>n</i>	$\epsilon_{Nd}$	<sup>176</sup> Hf/ <sup>177</sup> Hf	2σ√ <i>n</i>	ε <sub>Hf</sub>
WFB-N-15	8.78	39-63	337	18-6799	6	15·5883	5	38-8788	15	0.704203		0.512828		3.7			
WFB-1036	8.84	39.71	327	18-8083	4	15·5896	4	<i>38-8616</i>	10	0.703949		0.512827	8	3.7	0.283008	5	8.4
WFB-N-21	8.78	39-68	334	18-6241	12	15·5856	12	38·7938	36	<i>0</i> ·704228		0.512821		3.6	0.283009	4	8.4
WFB-E99-20	8.38	39-33	393	18-6940	9	15·5908	7	<i>38</i> · <i>7907</i>	20	0.704117							
WFB-N-01	8.39	39-33	392	18·5185	4	15·5892	4	38-6860	15	<i>0.704485</i>		0·512758		2.3	0.282976	4	7.2
Mojo																	
1010	8.56	39-23	384	17.8821	8	15.5875	8	38-2677	17	0.705289	5†	0.512551	6	<b>−1·7</b>			

<sup>\*</sup>Radial distance from Lake Abhe, Afar (11.25°N, 41.42°E).

Prior to dissolution rock chips were cleaned with distilled 2 N HCI, 1 N HBr, and ultrapure H<sub>2</sub>O. Sr, Nd, and Pb were separated for analysis following the procedures of Hanan & Schilling (1989) and Blichert-Toft et al. (1997) for Hf. Sr isotopes were measured on a VG Sector 54, seven-collector thermal ionization mass spectrometer in multi-dynamic mode. Hf, Nd, and Pb were analyzed on a Nu Plasma HR multi-collector inductively coupled plasma mass spectrometry (MC-ICP-MS) system using a Nu Instruments desolvating nebulizer equipped with a 100 µl min<sup>-1</sup> nebulizer (DSN100). The <sup>87</sup>Sr /<sup>86</sup>Sr isotope ratios were normalized to <sup>88</sup>Sr/<sup>86</sup>Sr = 0.1194 to correct for mass fractionation, and the Sr isotopic data are reported relative to NIST SRM 987 = 0.710250. The Nd and Hf isotope ratios were corrected for instrumental mass fractionation and machine bias by applying a discrimination factor determined by bracketing sample runs with analyses of the SDSU AMES Nd standard using a value of 0.512130 and the JMC 475 standard (every two or three samples) with <sup>176</sup>Hf/<sup>177</sup>Hf of 0.282160. Hf isotope sample data are normalized to  $^{179}$ Hf/ $^{177}$ Hf = 0.7325 and reported relative to JMC 475 = 0.282160. The CHUR value for  $^{176}$ Hf/ $^{177}$ Hf of 0.282772 of Blichert-Toft & Albarède (1997) was used to calculate the  $\epsilon_{Hf}$  values. Sample data were normalized to  $^{146}Nd/^{144}Nd = 0.7219$  and are reported relative to the laboratory value of  $^{143}Nd/^{144}Nd = 0$ ·512130 for the SDSU AMES Nd standard. The measured value of La Jolla Nd at SDSU = 0·511844. Pb was separated on two anion exchange columns, a 500 μl followed by a 200  $\mu$ l column to maximize the Pb separation for optimum ICP-MS analyses. The lead yields were determined to be  $\sim$ 100%, therefore the columns do not have a fractionation effect. The in-run mass discrimination for Pb was determined with NIST SRM 997 TI with 205TI/203TI = 2 3889 and NIST SRM 981 (Todt. 1996) using the method of White et al. (2000). The Pb isotope ratios were corrected for machine bias and drift by bracketing the sample analyses with analyses of the NIST SRM 981 (every two samples) corrected for fractionation. Total procedural blanks were: <25 pg Hf, <90 pg Pb, <200 pg Nd, <250 pg Sr, No blank corrections were applied to the data because they are insignificant. The reported uncertainties are 2 standard errors of the mean  $(2\sigma\sqrt{n})$  for the in-run precision, except for some Nd and Sr samples (†), where the uncertainties represent  $2\sigma\sqrt{n}$  for 2-3 replicate analyses, and thus provide a measure of sample reproducibility. The NIST SRM 987 standard averaged  $^{87}$ Sr/ $^{86}$ Sr = 0.710257  $\pm$  8 2 $\sigma$  $\sqrt{n}$ , n = 16, AMES Nd standard averaged  $^{143}$ Nd/ $^{144}$ Nd = 0.512118  $\pm$  2 2 $\sigma$  $\sqrt{n}$ , n = 22, JMC475 standard averaged 0.282151  $\pm$  0.000001 2 $\sigma$  $\sqrt{n}$ , n = 21, and NIST SRM 981 averaged  $^{206}$ Pb/ $^{204}$ Pb = 16.9361  $\pm$  0.0004.  $^{207}$ Pb/ $^{204}$ Pb = 15.4914  $\pm$  0.0004.  $^{208}$ Pb/ $^{204}$ Pb =  $38.6981 \pm 0.0012 \ 2\sigma \sqrt{n}$ , n = 17.

‡The analyses in italics for the Wonji Fault Belt were published previously (Furman *et al.*, 2006). The Pb isotopes for the Furman *et al.* (2006) Wonji Fault Belt data were analyzed at SDSU under the same conditions as reported here. We have added the fourth decimal place and errors (from the SDSU data archives) to be consistent with the new Main Ethiopian Rift data. The analytical methods and procedures for Nd and Sr have been reported by Furman *et al.* (2006). The Nd isotope ratios for these samples have been renormalized to the SDSU La Jolla = 0·511844 and the Sr isotope ratios are normalized to NIST SRM 987 = 0·710250.

Table 2: Location and Hf isotope data for the Gulf of Aden basalts

Sample	Latitude (°N)	Longitude (°E)	Distance* (km)	<sup>176</sup> Hf/ <sup>177</sup> Hf	2σ√ <i>n</i>	$\epsilon_{Hf}$
Asal Rift						
DJ-74	11.60	42.48	122	0.283122	6	12-4
DJ-42	11.60	42.49	122	0.283161	7	13.8
DJ-26	11.60	42.50	123	0.283130	5	12.7
DJ-18	11.60	42.51	124	0.283115	5	12-1
DJ-9	11.60	42.51	125	0.283141	4	13-0
DJ-5	11.60	42.52	126	0.283144	5	13-2
DJ-2	11.60	42.52	126	0.283141	4	13.0
Gulf of Tadjoura	a					
VD-11	11.51	42.85	158	0.283070	5	10.5
V-60 g	11.75	43.02	183	0.283077	6	10.8
VD-39	11.59	43·10	187	0.283122	10	12-4
V-55	11.87	43-27	213	0.283083	5	11.0
V3307 41D-1	11.94	43-41	229	0.283098	5	11.5
Tadjoura Trougi	h					
V3307 42D-1 g	11.94	43.69	259	0.283134	5	12.8
V3307 44D-2g	12.08	43.93	289	0.283263	6	17-4
V3307 46D-1 g	12.06	44.71	370	0.283230	6	16-2
West Sheba Rid	dge					
V3307 47D-1 g	12.07	44.83	382	0.283186	8	14-6
V3307 48D-2g	12.04	45.21	422	0.283138	8	12-9
V3307 49D-1	12.04	45.51	454	0.283008	9	8.3
V3307 50D-1 g	12·10	45.85	492	0.283105	7	11.8
V3307 51D-1 g	12·16	46-29	539	0.283126	7	12·5
V3307 52D-1 g	12.27	46.71	587	0.283274	7	17-8
V3307 53D-1	12-42	47·10	632	0.283304	6	18-8
V3307 54D-1 g	12.42	47-25	648	0.283225	5	16-0
V3307 55D-2g	12-47	47.84	712	0.283270	7	17-6
V3307 56D-1 g	12-65	48-27	762	0.283262	6	17-3
V3306 39D-1 g	12.75	48-49	788	0.283246	5	16-8
V3307 57D-1 g	12.90	48.99	844	0.283284	7	18-1
V3307 58D-1	12.90	49-36	883	0.283309	12	19-0
V3307 59D-1 g	13.04	49-39	890	0.283253	7	17-0
V3307 60D-1	13·16	49-69	925	0.283302	6	18-7
V3307 62D-1 g	13.36	49.81	942	0.283315	7	19-2
V3307 61D-2g	13.26	49-97	958	0.283277	6	17-9
V3307 63D-3 g	13.29	50.32	995	0.283295	7	18.5
V3307 64D-3 g	13.34	50-69	1036	0.283339	7	20.1
V3307 65D-1 g	13-21	51.02	1068	0.283349	8	20.4
East Sheba Ridg	ge					
V3306 35D-1	14-65	52-41	1254	0.283293	6	18-4
V3307 66D-1 g	14.39	52.68	1275	0.283163	7	13.8

(continued)

Table 2: Continued

Sample	Latitude (°N)	Longitude (°E)	Distance* (km)	<sup>176</sup> Hf/ <sup>177</sup> Hf	2σ√ <i>n</i>	ε <sub>Hf</sub>
V3306 25D-1	14-48	53-91	1406	0.283134	5	12.8
V3306 24D-1 g	14.75	54.80	1508	0.283187	6	14.7
V3307 68D-1 g	14.74	55.71	1603	0.283189	8	14.7
V3307 69D-1 g	14-47	56-42	1672	0.283201	7	15-2

\*Radial distance from Lake Abhe, Afar.

Hf isotope data reported for the Gulf of Aden sample suite were analyzed by MC-ICP-MS (Plasma 54) at the Ecole Normale Supérieure in Lyon following the procedures of Blichert-Toft et~al.~(1997). The JMC-475 Hf standard, which was run every two samples, averaged  $0\cdot282160\pm0\cdot000010$  during the analytical run sessions of the Gulf of Aden samples. The total procedural Hf blank was less than 20 pg. The CHUR value for  $^{176}{\rm Hf/}^{177}{\rm Hf}$  of  $0\cdot282772$  of Blichert-Toft and Albarède (1997) was used to calculate the  $\epsilon_{\rm Hf}$  values. The Pb, Sr, and Nd isotope compositions have been reported by Schilling et~al.~(1992).

(Rooney et al., 2005, 2007) and the samples fall largely within the range observed globally for mantle-derived basalts (Hofmann et al., 1986, 2005). We also examined the relationship between K/P and Ti/Yb and the Nd and Sr isotope ratios for the Main Ethiopian Rift lavas. Concurrent increases in K/P with increasing <sup>87</sup>Sr/<sup>86</sup>Sr (or decreasing 143Nd/144Nd or 176Hf/177Hf) are indicative of silicic crustal assimilation (Hart et al., 1989). We observed no consistent relationships between Sr, Nd, Hf, or Pb isotope ratios and K/P. The Main Ethiopian Rift lavas have Ti/Yb ratios of about 6000-10000 (except for sample 1003, which has Ti/Yb = 4500 and contains mafic crustal xenoliths; Rooney et al., 2005). Low K/P < 6.5 accompanied by high Ti/Yb (>6000) for the Main Ethiopian Rift lavas rule out significant crustal contamination (Leeman & Hawkesworth, 1986; Van Calsteren et al., 1986; Hart et al., 1989). Furthermore, the He isotope data presented below show that the signature of the Afar plume, as well as the upper mantle, dominates the He isotope variations in the Main Ethiopian Rift basalts selected for this study, which indicates that any potential crustal contamination is minor. The lack of crustal assimilation in these magma suites is consistent with focused recent volcanism in the Wonji Fault Belt and Silti-Debre Zevit Fault Zone (e.g. WoldeGabriel et al., 1990; Ebinger & Casey, 2001; Casey et al., 2006). Magmas have re-utilized earlier conduits, incorporating xenocrysts and cumulate xenoliths from the walls, with apparently little interaction with the continental lithosphere on their way to the surface (Rooney et al., 2005).

Table 3: Location and He isotope data for basalts from the Main Ethiopian Rift

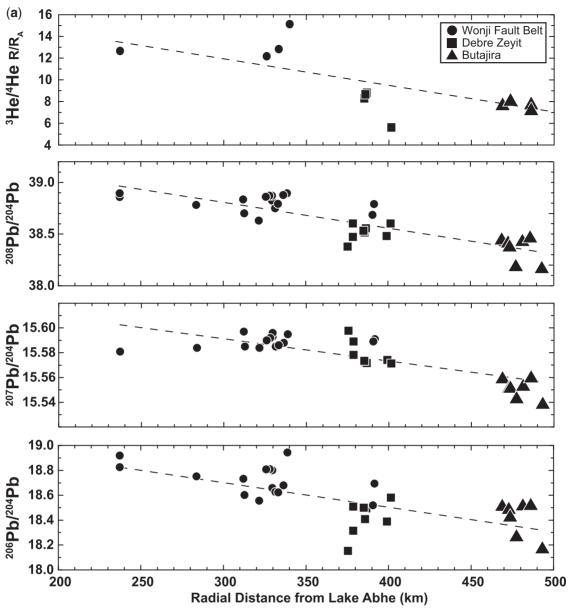
Sample	Lat. (°N)	Long. (°N)	Distance (km)	Phase	Wt (mg)	<sup>3</sup> He/⁴He R/R <sub>A</sub>	<b>2</b> σ√ <i>n</i>	[He]
Debre Zeyit								
DZ-1014	8.60	38-91	402	срх	598-1	5-61	0.34	1·66E - 09
DZ-1009	8.59	39-15	386	olivine	442-1	8-30	0.40	2·59E - 09
DZ-1007	8.72	38-99	388	olivine	744-6	8-66	0.14	1·12E - 08
DZ-1008	8.72	38-98	387	olivine	581·5	8.80	0.16	1·06E - 08
Butajira								
BJ-1045	8.06	38-36	488	olivine	422-4	7.52	0.30	4·22E - 09
BJ-1045	8.06	38-36	488	срх	420.0	7.53	0.16	1·14E - 08
BJ-1053	8-19	38.45	470	срх	541.4	7.70	0.24	4·91E - 09
BJ-1043	8.14	38.44	476	olivine	430-4	8-11	0.36	2·55E - 09
Wonji Fault Belt								
WFB-1027	9.48	40-20	238	olivine	428-6	12-66	0.98	7·49E - 10
WFB-1036	8.84	39.71	327	olivine	631-4	12-16	0.24	8·02E - 09
WFB-N-21	8.78	39-68	334	olivine	537.8	12.82	0.26	6·40E - 09
WFB-N-14	8.76	39-61	340	olivine	871-4	15.12	0.30	3·13E - 09

Helium isotope analyses were carried out by *in vacuo* crushing of mineral separates from the Ethiopian basalts. Analytical procedures have been described elsewhere (Graham *et al.*, 1998). The reported uncertainties are 2 standard errors, computed from the quadrature sum of uncertainties associated with the sample analysis, air standard analyses and blank determinations. Distance is measured radially in kilometers.

# Isotope signatures, mantle components, and mixing

The isotopes of Pb, Sr, Nd, Hf, and He are powerful tracers of mantle dynamics, because mantle-derived basalts fall into distinct subdivisions based on their radiogenic isotope compositions (White & Hofmann, 1982; Zindler & Hart, 1986; Hofmann, 1997). These subdivisions are related to distinct processes (pelagic vs terrigenous sediment subduction, hydrothermal alteration, etc.) as well as to distinct tectonic or petrogenetic environments (e.g. mid-ocean ridge, oceanic plateau, island arc, etc.). As one example, DUPAL signatures (Hart, 1984) may arise from the recycling of oceanic and/or continental lithosphere, ranging from ancient (EM I) to recent (EM II) in age, may be associated with, respectively, ancient and recent sediments, and may include metasomatized lithosphere (e.g. HIMU) (Zindler & Hart, 1986; Mahoney et al., 1989, 1992, 1996; McKenzie & O'Nions, 1995; Rehkämper & Hofmann, 1997; Lustrino, 2005; Willbold & Stracke, 2010). The source of MORB is best constrained in terms of its location in the mantle, and corresponds to the trace-element depleted upper mantle MORB source (White & Hofmann, 1982; Zindler & Hart, 1986). The common isotopic component C, defined by the convergence of trends in Pb isotope compositional space (at relatively radiogenic Pb;  $^{206}$ Pb/ $^{204}$ Pb =  $19.6 \pm 0.4$ ) for MORB from the Atlantic,

Pacific and Indian Oceans, has been attributed to the presence of recycled oceanic lithosphere in the mantle source region of many MORB (Hanan & Graham, 1996). In contrast, the mantle component FOZO, defined by the convergence of global ocean island basalt isotopic trends, has been attributed to the presence of deep, relatively primitive mantle (Hart et al., 1992; Hauri et al., 1994), although this interpretation has since been modified (Stracke et al., 2005). Mixing models involving these mantle components and end-members, however, often cannot define a unique source for the basalts. In particular, for continental rift basalts it is difficult to distinguish a deep mantle component derived by melting recycled lithosphere from shallower melting involving old subcontinental lithosphere. Although this is a challenging problem, the systematics of mixing during petrogenesis can be deciphered using a multi-isotopic approach. For example, the order of mixing between three end-members may be revealed through the combined use of <sup>208</sup>Pb-<sup>207</sup>Pb-<sup>206</sup>Pb isotopes (Hanan et al., 1986; Hanan & Schilling, 1997; Douglass & Schilling, 2000). Systematic spatial and temporal variations in this isotope space provide further constraints on the mixing sequence. Mixing models typically use exactly chosen values for end-member compositions to avoid the mathematical complexity that arises from the mixing of heterogeneous end-members. However, mantle sources



**Fig. 2.** (a) Variation of  $^{206}$ Pb/ $^{204}$ Pb ( $r^2 = 0.48$ ),  $^{207}$ Pb/ $^{204}$ Pb ( $r^2 = 0.60$ ),  $^{208}$ Pb/ $^{204}$ Pb ( $r^2 = 0.83$ ), and  $^{3}$ He/ $^{4}$ He R/R<sub>A</sub> ( $r^2 = 0.53$ ), and (b)  $^{87}$ Sr/ $^{86}$ Sr, ε<sub>Nd</sub>, and ε<sub>Hf</sub> in Quaternary lavas of the Main Ethiopian Rift (Wonji Fault Belt, Debre Zeyit, and Butajira) as a function of radial distance from the inferred present-day location of the Afar plume axis in the Lake Abhe region (Table 1). (c) Variation of <sup>87</sup>Sr/<sup>86</sup>Sr (r<sup>2</sup>=0·18), ε<sub>Hf</sub>  $(r^2=0.14)$ ,  $\epsilon_{\rm Nd}$   $(r^2=0.10)$ , and  $\epsilon_{\rm Nd}^{206}$  Pb/ $\epsilon_{\rm Nd}^{204}$  Pb  $(r^2=0.036)$  as a function of radial distance from the Afar plume center for the Asal Rift, Gulf of Tadjoura, Tadjoura Trough and the Sheba Ridge. The prominent shift in  $\epsilon_{\rm Hf}$  across the Alula-Fartak fracture zone as the East Sheba Ridge steps northward towards Yemen should be noted. The trend lines [in (a) and (c); dashed black lines] represent the least-squares linear regression for the Main Ethiopian Rift [from (a) and (b)]. The 'by eye' trends for the Gulf of Aden data are shown by a thin dashed line in (c). The Gulf of Aden Hf isotope data are from Table 2 and the Pb, Sr, and Nd isotope data are from Schilling et al. (1992). Also shown are the Main Ethiopian Rift data from (a) and (b) and Table l. Negative distances signify samples from along the Gulf of Aden, positive distances indicate samples along the East African Rift. E- and W-Torus refer to the position of the torus plume head beneath the Main Ethiopian Rift and Gulf of Aden as proposed by Schilling et al. (1992). Debre Zeyit sample DZ-1003 is labeled.

are, by their very nature, heterogeneous to varying degrees, and, hence, simple mixing models assuming homogeneous end-members should be viewed only as approximations; they cannot be expected to account for all the natural variability observed in most sample suites. The approach that we adopt below also involves a multi-component mixing model with fixed end-member compositions. Nevertheless, we impose rigorous mathematical and statistical tests to evaluate the ability of the model to account for the observed geochemical variability

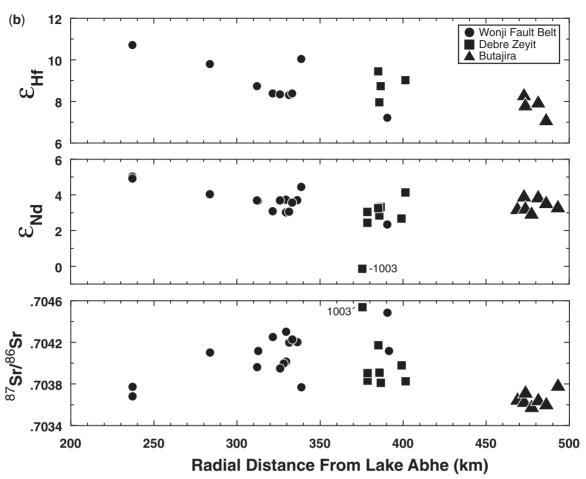


Fig. 2. Continued.

in the Main Ethiopian Rift sample suite. The details and mathematical development of this multi-component mixing model are well documented (e.g. Hanan *et al.*, 1986, 2000; Schilling *et al.*, 1992; Hanan & Schilling, 1997; Douglass & Schilling, 2000), and its algorithms and rationale have been given by Schilling *et al.* (1992), Albarède (1995) and Douglass & Schilling (2000) and hence are not repeated here.

### RESULTS

### Pb-Sr-Nd-Hf-He spatial isotope variations

The Pb–Sr–Nd–Hf–He isotope variations along the axis of the Main Ethiopian Rift, relative to the presumed location of the center of the Afar plume beneath the Lake Abhe region, are shown in Fig. 2a and b. The location of the plume center near Lake Abhe is based on a synthesis of independent structural, tectonic, magnetic, and geochemical data that all focus toward this region (Schilling, 1973b). The new Pb, Sr, Nd, and Hf isotope ratios for the Main

Ethiopian Rift lavas all become more radiogenic upon approaching the inferred plume center, except for <sup>87</sup>Sr/<sup>86</sup>Sr and <sup>207</sup>Pb/<sup>204</sup>Pb in two Wonji Fault Belt samples closest to the plume, which appear to be slightly low (Table 1; Fig. 2). In addition, <sup>143</sup>Nd/<sup>144</sup>Nd and <sup>176</sup>Hf/<sup>177</sup>Hf data for Butajira basalts are not quite as low as predicted by a monotonic linear trend (Fig. 2b). There is a general trend of increasing <sup>3</sup>He/<sup>4</sup>He towards Lake Abhe (Table 3). However, there is also significant variation on a shorter length scale; for example, the highest <sup>3</sup>He/<sup>4</sup>He in our sample suite (15·1 RA) occurs in a Wonji Fault Belt lava at ~340 km distance (sample WFB N-14), similar to what is observed for <sup>206</sup>Pb/<sup>204</sup>Pb (Fig. 2a). The bulk composition of sample N-14, which contains abundant plagioclase and olivine phenocrysts relative to the other samples, is unlikely to represent a magmatic liquid because it falls off the major element trends of the combined dataset (e.g. CaO/ Al<sub>2</sub>O<sub>3</sub> and MgO).

The similar profiles for Pb, Sr, Nd, Hf, and He isotopes are not expected, because mantle-derived lavas

NUMBER 2

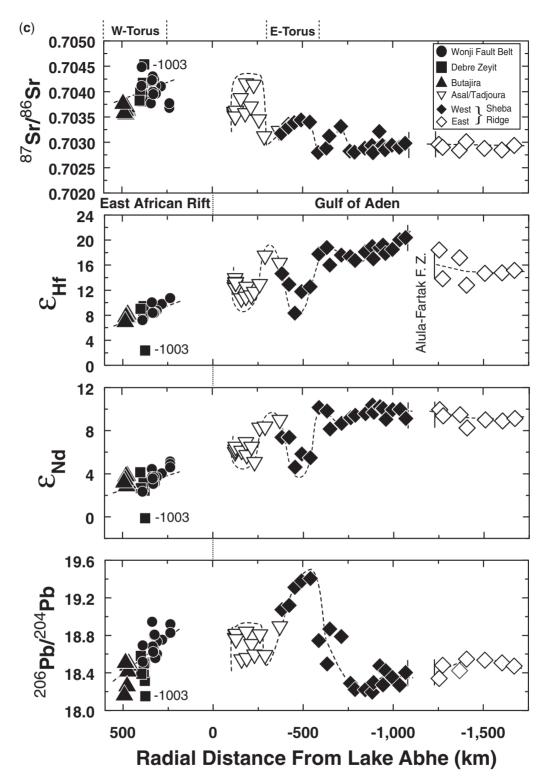


Fig. 2. Continued.

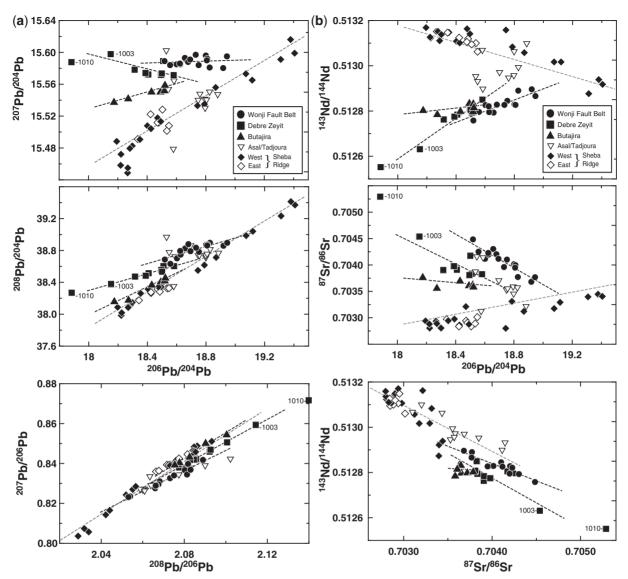


Fig. 3. (a) Covariation of <sup>208</sup>Pb/<sup>204</sup>Pb and <sup>207</sup>Pb/<sup>204</sup>Pb with <sup>206</sup>Pb/<sup>204</sup>Pb, and <sup>207</sup>Pb/<sup>206</sup>Pb with <sup>208</sup>Pb/<sup>206</sup>Pb in the Main Ethiopian Rift Quaternary and Gulf of Aden sample suites. The dashed lines are least-squares regression lines for the Gulf of Aden, Wonji Fault Belt, Debre Zeyit, and Butajira volcanic zones. (b) Covariation of <sup>143</sup>Nd/<sup>144</sup>Nd and <sup>87</sup>Sr/<sup>86</sup>Sr with <sup>206</sup>Pb/<sup>204</sup>Pb, and <sup>143</sup>Nd/<sup>144</sup>Nd with <sup>87</sup>Sr/<sup>86</sup>Sr in the Wonji Fault Belt, Debre Zeyit, and Butajira volcanic zones and Gulf of Aden sample suites. The dashed lines represent the least-squares linear regression lines of the datasets for the Main Ethiopian Rift and Gulf of Aden data. Debre Zeyit samples Mojo 1010 and DZ-1003 are labeled.

usually show broad anti-correlation between <sup>143</sup>Nd/<sup>144</sup>Nd (or <sup>176</sup>Hf/<sup>177</sup>Hf) and <sup>87</sup>Sr/<sup>86</sup>Sr (or <sup>206</sup>Pb/<sup>204</sup>Pb) reflecting long-term *in situ* radiogenic production. The signatures observed here, therefore, rather reflect recent mixing.

The variations in Nd–Sr–Pb–Hf isotope compositions and La/Sm ratios along the Gulf of Aden spreading center define an 800 km long gradient eastward from the Afar region, based on samples from the Asal Rift, Gulf of Tadjoura, Tadjoura Trough, and along the Sheba Ridge to about 50°E (Schilling *et al.*, 1992; Table 2). The Asal Rift and Gulf of Tadjoura represent a nascent stage of rift

evolution, characterized by diffuse extension. In contrast, the more developed Sheba Ridge is approaching steady-state seafloor spreading. The differences in the rift development stage are reflected in the Pb, Sr, Nd, and Hf isotope characteristics of the basalts. The West Sheba Ridge shows an enrichment peak for Pb, Sr, Nd, and Hf isotopes and large ion lithophile elements (LILE) centered around 46°E, about 500 km east of the presumed plume center (Table 2). East of 48°E, Pb, Sr, Nd, and Hf isotope compositions show much less variation. The  $^{206}$ Pb/ $^{204}$ Pb,  $^{87}$ Sr/ $^{86}$ Sr,  $\varepsilon_{\rm Nd}$ , and  $\varepsilon_{\rm Hf}$  display similar spatial variation

(Fig. 2c) consisting of well-defined alternating peaks and troughs. Although the peaks and troughs appear to be slightly out of phase, there is no apparent fracture zone control on the Pb, Sr, and Nd isotope compositions (Fig. 2; Schilling et al., 1992). However, the new Hf isotope data show a step at the Alula-Fartak fracture zone toward less radiogenic, more continental-like signatures, as the ridge jumps northeastward toward Yemen. The Asal Rift and Gulf of Tadjoura show a much smaller enrichment spike in Pb, Nd, and Hf isotope ratios relative to the Sheba Ridge. However, <sup>206</sup>Pb/<sup>204</sup>Pb is more radiogenic,  $\epsilon_{\rm Nd}$  and  $\epsilon_{\rm Hf}$  are less radiogenic than observed in the LILE-depleted basalts of the East Sheba Ridge and all overlap the range of the West Sheba Ridge samples. The range in <sup>87</sup>Sr/<sup>86</sup>Sr for the Asal Rift and Gulf of Tadjoura data is distinctly higher than observed along the entire Sheba Ridge. The anomaly at 46°E was previously attributed to the interaction of the Sheba Ridge with a torus-like mantle plume (Schilling et al., 1992). In the torus plume model, the Gulf of Tadjoura and Asal Rift overlie a mantle region located to the inside of a plume-head torus, where the mantle source is proposed to be dominated by Pan-African continental lithosphere and the upper mantle asthenosphere (Fig. 2c; Schilling et al., 1992). The plume torus overlies the enrichment peaks in radiogenic Pb and Sr and corresponding troughs in  $\varepsilon_{Nd}$  and  $\varepsilon_{Hf}$  located on the Sheba Ridge at about 500 km from the plume center. In the Schilling et al. (1992) model, the opposite side of the torus would be located at a similar distance from the plume center, along the southwestern Main Ethiopian Rift.

#### Pb-Pb isotope correlations

The new data for the Main Ethiopian Rift were collected from three spatially resolvable sections of the Main Ethiopian Rift: Debre Zeyit, Butajira, and the Wonji Fault Belt. The Debre Zevit and Butaiira volcanic fields are two major zones of volcanism and tectonism separated by a 50 km gap that comprises the northernmost portion of the Silti-Debre Zeyit Fault Zone (WoldeGabriel et al., 1990; Rooney et al., 2011). The most distinctive Quaternary magmatic feature in the Ethiopian Rift is the Wonji Fault Belt, which extends from around Lake Abhe into southern Ethiopia (Mohr, 1967). The three sample populations (Debre Zevit, Butajira, and Wonji Fault Belt) have different, but overlapping ranges in Pb isotope ratios, and appear to form distinct arrays in Pb-Sr-Nd-Hf multiisotope space (Figs 3a, b and 4; Table 4). This raises the question of whether the three spatially distinct datasets can be treated as a single population, or whether they constitute three distinct sub-populations.

Least-squares linear regressions of the three datasets show that their trends are significantly different at the 95% confidence level for both the <sup>206</sup>Pb/<sup>204</sup>Pb vs <sup>208</sup>Pb/<sup>204</sup>Pb and <sup>207</sup>Pb/<sup>204</sup>Pb binary projections (Table 4).

The three sub-populations (Wonji Fault Belt, Debre Zevit, and Butajira) define linear arrays in Pb isotopic space, with  $r^2$  values of 0.68, 0.93, and 0.97, respectively, in the  $^{208}\text{Pb}/^{204}\text{Pb}$  vs  $^{206}\text{Pb}/^{204}\text{Pb}$  diagram (Fig. 3a). In the <sup>207</sup>Pb/<sup>204</sup>Pb vs <sup>206</sup>Pb/<sup>204</sup>Pb diagram the trends for Debre Zeyit and Butajira have  $r^2$  values of 0.43 and 0.89, respectively. The Wonji Fault Belt data form a near horizontal array with  $r^2 = 0.02$ . The  $r^2$  statistic measures how significantly the slope of a fitted line differs from zero, and is not necessarily a measure of good fit. To further test the validity of subdividing the data on the basis of spatial location, we performed analysis of the covariance of the Pb-Pb least-squares regression lines for each of the sub-populations against the combined datasets of the other two areas. The rationale for this comparison is that if all the data fall along the same regression, then removing a subset of the data should not affect the pooled regression. We tested the null hypothesis that the least-squares regression for a sub-population was the same as the regression for the other two groups combined. The F-test results for  $^{208}\text{Pb}/^{204}\text{Pb}$  and  $^{207}\text{Pb}/^{204}\text{Pb}$  vs  $^{206}\text{Pb}/^{204}\text{Pb}$ , using the slopes and intercepts of the least-squares linear regressions for the three sub-populations (Wonji Fault Belt, Debre Zeyit, and Butajira) relative to the data subsets containing the other two groups, allow rejection of the null hypothesis at the 95% confidence level (Table 5). Therefore, these three sample suites, from spatially different volcanic belts, have distinct Pb isotope correlations, or trends, and define three statistically distinct sub-populations that make up the complete Main Ethiopian Rift sample suite of this study.

### Nd-Hf isotope correlations

In  $\epsilon_{Nd}$  vs  $\epsilon_{Hf}$  isotope space (Fig. 5a) the Main Ethiopian Rift data lie along and above the mantle array, trending between the Afar plume, as represented by the C field, and compositions typical of ancient lithosphere. In contrast, the new data from the Sheba Ridge axis in the Gulf of Aden and its extension into the Gulf of Tadjoura and the Asal Rift plot at more radiogenic values of  $\varepsilon_{Hf}$  and ε<sub>Nd</sub>, similar to Indian Ocean MORB with predominantly positive values of  $\Delta \epsilon_{Hf}$  rather than negative  $\Delta \epsilon_{Hf}$  values such as those of most Pacific MORB (Fig. 5;  $\Delta \varepsilon_{Hf}$  is the vertical displacement from the mantle array in  $\epsilon_{Hf}$  units for a data point in the  $\epsilon_{\rm Nd}$  vs  $\epsilon_{\rm Hf}$  diagram). Both the Main Ethiopian Rift and the Gulf of Tadjoura-Asal Rift data have distinct high  $\Delta^{208} \text{Pb}^{/204} \text{Pb}$  (where  $\Delta^{208} \text{Pb}^{/204} \text{Pb}$  is relative to the <sup>208</sup>Pb/<sup>204</sup>Pb of the Sheba Ridge array at any given <sup>206</sup>Pb/<sup>204</sup>Pb; Fig. 3a). These observations are consistent with the different tectonic settings. The Sheba Ridge data reflect an origin dominated by plume-oceanic lithosphere interaction (e.g. Schilling et al., 1992), whereas the Main Ethiopian Rift and the Gulf of Tadjoura-Asal Rift data indicate strong influence from a continental lithosphere component, in line with a continental rift setting.

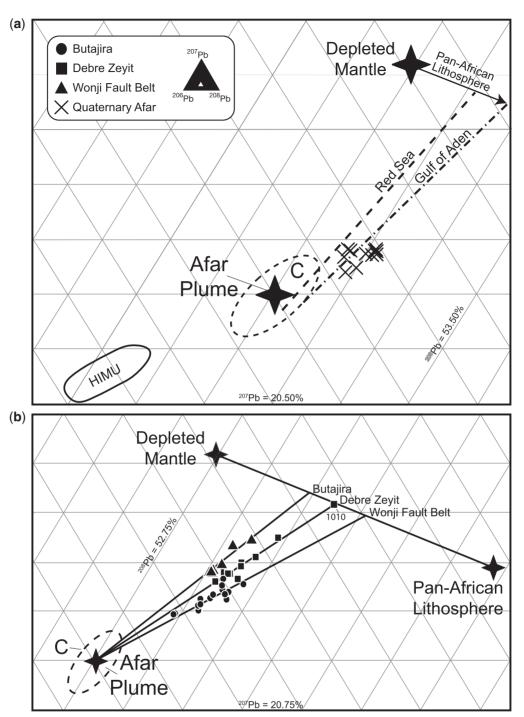


Fig. 4. (a) Triangular Pb isotope plot (after Hanan et al., 1986) of the relative abundances of <sup>206</sup>Pb, <sup>207</sup>Pb, and <sup>208</sup>Pb in Quaternary lavas from Afar and the northern Main Ethiopian Rift (Deniel et al., 1994). Only lavas having MgO > 7 wt % are included, to limit the potential variation related to shallow crustal processes. The least-squares trend for the Red Sea and Gulf of Aden are shown for comparison. Data for the Red Sea are from Altherr et al. (1988), Rogers (1993), and Volker et al. (1993, 1997), and for the Gulf of Aden from Schilling et al. (1992). (b) Triangular Pb isotope plot (after Hanan et al., 1986) of the relative abundances of <sup>206</sup>Pb, <sup>207</sup>Pb, and <sup>208</sup>Pb in lavas from the Main Ethiopian Rift and the model end-members depleted mantle, Pan-African lithosphere, and the Afar plume. Basalts from Butajira, Debre Zeyit, and the Wonji Fault Belt define three distinct pseudo-binary mixing arrays that converge at the composition of the Afar plume, which lies within the range of compositions for the common component 'C' deduced from oceanic basalts (Hanan & Graham, 1996). The depleted ends of the mixing arrays are characterized by three ratios of depleted mantle to Pan-African lithosphere, each of which is constant over the range of mixing seen in the three arrays. The proportion of the Afar plume is largest for the Wonji Fault Belt basalts (~240–440 km from the Afar plume center), smallest for the Butajira basalts (~470–500 km distant), and intermediate for the Debre Zeyit basalts (~370–400 km distant). Sample Mojo-1010 is labeled. The ternary grid is drawn at 0·25% intervals. The ternary grid interval is 0.25%.

Table 4: Characteristics of slopes fitted to the linear arrays in Pb isotope space for Main Ethiopian Rift magmas

Pb isotopes Region Test Value for Value for other single belt two belts pooled 206-208 WFB Slope 0.54 0.69 Intercept 28.75 25.67 DΖ Slope 0.52 1.12 28.87 17.90 Intercept B. Slope 0.78 0.88 Intercept 22.23 24.25 207-206 WFB 0.01 0.00 Slope 15.55 Intercept 15.49 DΖ Slope -0.05 0.08 Intercept 16.44 14.12 BJ Slope 0.05 0.01 Intercept 14.58 15.34

Differences between the slopes and intercepts can be utilized to examine differences and similarities between the groups (BJ, Butajira; DZ, Debre Zeyit; WFB, Wonji Fault Belt) using statistical methods (*F*-test; see Table 5).

## Hf-Pb isotope correlations

In  $\epsilon_{\rm Hf}$  vs  $^{206}{\rm Pb}/^{204}{\rm Pb}$  space the West Sheba Ridge array extends between the fields for the most depleted MORB from the Southwest Indian Ridge and the C-like Afar plume. In contrast, the Main Ethiopian Rift data extend between the Pan-African lithosphere and the Afar plume end-member (Fig. 5b). The Asal Rift and Tadjoura data have affinity to, and overlap with, the East Sheba Ridge data in a region of the diagram between the Main Ethiopian Rift and the West Sheba Ridge arrays, displaced toward the Pan-African lithosphere component. Overall, the combined dataset, consisting of the Main Ethiopian Rift and the Gulf of Aden samples, has Indian Ocean-like Hf and Pb isotope signatures, distinct from the Pacific asthenosphere MORB source (Fig. 5c) and consistent with a continental-like mantle source component.

#### DISCUSSION

# Afar plume-continental lithospheredepleted mantle mixing model

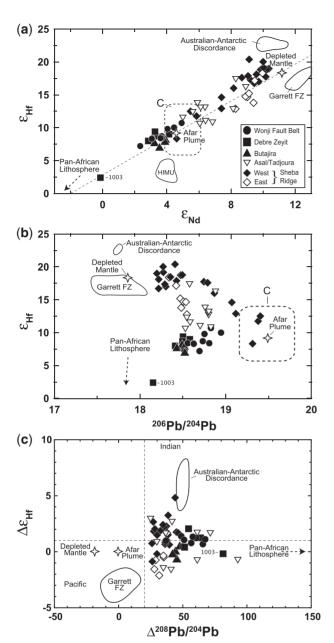
The Arabian–Nubian shield represents Precambrian microplate accretion of inter-continental arcs, slivers of back-arc oceanic lithosphere, and older Precambrian terranes reactivated during Pan-African thermotectonic events (Schilling *et al.*, 1992; Be'eri-Shlevin *et al.*, 2010). The Pan-African sub-continental lithosphere is of variable mean age, origin, and isotopic composition; there are very

Table 5: F-test analysis of the similarities of the WFB, Butajira and Debre Zevit in Pb isotope space

Pb isotopes	Region	Test	<i>F</i> -value	Degrees of freedom	F-crit. (95% confidence)	Is the group different from pooled?
206-208	WFB	Slope	0.58	27	4.2	No
		Elevation	19-33	28	4.2	Yes
	DZ	Slope	5.94	27	4.2	Yes
		Elevation	2.43	28	4.2	No
	BJ	Slope	0.37	27	4.2	No
		Elevation	60.2	28	4.2	Yes
207-206	WFB	Slope	0.01	27	4.2	No
		Elevation	9.79	28	4.2	Yes
	DZ	Slope	15.68	27	4.2	Yes
		Elevation	5.44	28	4.2	Yes
	BJ	Slope	2.9	27	4.2	No
		Elevation	67.48	28	4.2	Yes

The primary question posed was whether each of the three recognized geographical groups (BJ, Butajira; DZ, Debre Zeyit; WFB, Wonji Fault Belt) were distinctive on the basis of their isotopic results. As discussed in the text, Pb isotopes form distinctive linear arrays and hence can be used to examine differences and similarities between the groups. To establish whether each of the groups is distinctive from the dataset as a whole, we examined the slope and elevation of the arrays generated by  $^{206}$ Pb/ $^{204}$ Pb $^{208}$ Pb/ $^{204}$ Pb and  $^{206}$ Pb/ $^{204}$ Pb $^{207}$ Pb/ $^{204}$ Pb. The F-test specifically asked: Are the Wonji Fault Belt, Debre Zevit and Butajira samples distinctive from the remaining dataset in terms of slope and elevation? F-values were calculated using established methods (Snedecor & Cochran, 1989). F-crit. is the critical F-value (Davis, 1973) based on the degrees of freedom of the dataset; where the F-value is above the F-crit. value, the null hypothesis (that the slope and/or intercept of the respective datasets are identical) can be rejected at the 95% confidence level.

few published xenolith isotope data from the Arabian–Nubian shield. Radiogenic isotope analyses of mantle xenoliths from beneath Arabia (Yemen and Kod Ali Island in the Red Sea; Baker et al., 2002) indicate that the lithosphere has been extensively modified by interaction with the plume and the asthenosphere, and in fact could be modelled by these same three end-members. The rejuvenated Pan-African lithosphere is not conducive to preserving ancient high  ${}^{3}\text{He}/{}^{4}\text{He}$ . Lithosphere with high time-integrated  ${}^{3}\text{He}/(U+Th)$  that would allow preservation of ancient  ${}^{3}\text{He}$  enrichments cannot exist within or below the continents, and is unlikely to exist within the upper mantle, based on  ${}^{3}\text{He}/{}^{4}\text{He}$  in MORB (Day et al., 2005).



**Fig. 5.** (a)  $\varepsilon_{Hf}$  vs  $\varepsilon_{Nd}$  and (b)  $^{206}Pb/^{204}Pb$  vs  $\varepsilon_{Hf}$  for the Quaternary Main Ethiopian Rift sample suite and for the Gulf of Aden basalts [Nd and Pb isotopic data from Schilling et al. (1992)]. Model end-members (see Table 6) are shown as black open stars for the Afar plume and depleted mantle. Pan-African lithosphere plots off the diagram at this scale. The dashed-line field represents the range of values for C (Hanan & Graham, 1996; Hanan et al., 2000). The fields for HIMU, depleted Indian Ocean MORB from the Australian-Antarctic Discordance region of the Southeast Indian Ridge, and depleted Pacific MORB from the East Pacific Rise Garrett Fracture Zone are also shown for reference (data from Chaffey et al., 1989; Salters & White, 1998; Wendt et al., 1999; Hanan et al., 2004). The mantle array in (a) (dashed line) is represented by  $\varepsilon_{Hf} = 1.4 \varepsilon Nd +$ 2.8 (J. Blichert-Toft, unpublished compilation of >3000 oceanic basalts). Symbols are the same as in previous plots. (c)  $\Delta^{208} \text{Pb}/^{204} \text{Pb}$  vs  $\Delta\epsilon_{Hf}$  for the most depleted MORB from the Pacific and Indian Ocean basins, represented by the Garrett Fracture Zone (Wendt et al., 1999) and the Australian-Antarctic Discordance on the

The source origin of the Main Ethiopian Rift basalts cannot be solely within the sub-continental lithosphere. In the <sup>208</sup>Pb/<sup>204</sup>Pb and <sup>207</sup>Pb/<sup>204</sup>Pb vs <sup>206</sup>Pb/<sup>204</sup>Pb isotope diagrams the new combined data do not fall along a single linear correlation, ruling out a simple binary source mixing origin for the lavas. At least three end-member components are required to account for the entire dataset. The three-component mixing model invoked by Schilling et al. (1992) for the mantle source of the lavas from the Gulf of Aden, its extension into the Gulf of Tadjoura, and subaerial basalts from the Ardoukoba Rift in east Afar calls upon interaction between a mantle plume, EM-like Pan-African continental lithosphere and the MORBsource asthenosphere, to explain the Nd-Sr-Pb isotope compositions. In general, these three end-members enclose the entire range of the Nd-Sr-Pb isotope ratios of the new Main Ethiopian Rift data presented here. Schilling et al. (1992) associated the Afar plume with a HIMU radiogenic Pb isotope signature. However, the Butajira, Debre Zeyit, and Wonji Fault Belt trends converge at a lower <sup>206</sup>Pb/<sup>204</sup>Pb (~l9·6) than the characteristic HIMU value (206 Pb/204 Pb >20; Zindler & Hart, 1986), consistent with an Afar plume composition similar to 'C' (Fig. 4), the 'common' isotopic composition observed in oceanic basalts (Hanan & Graham, 1996). This observation is further supported by the  $\varepsilon_{Nd}$  vs  $\varepsilon_{Hf}$  diagram, in which the Wonji Fault Belt, Debre Zevit, and Butajira data trends all fall above and parallel to the mantle array, and do not show mixing relationships toward the HIMU composition, which plots below the mantle array (Fig. 5). In terms of three-component mixing, in addition to the relatively high <sup>206</sup>Pb/<sup>204</sup>Pb plume end-member, the Main Ethiopian Rift data distribution requires the other two end-members to have lower <sup>206</sup>Pb/<sup>204</sup>Pb, with one having relatively higher <sup>207</sup>Pb/<sup>204</sup>Pb than the other. These constraints are consistent with the Pan-African continental lithosphere (high <sup>207</sup>Pb/<sup>204</sup>Pb) and the asthenosphere (MORB-source mantle with low <sup>207</sup>Pb/<sup>204</sup>Pb) end-members of the Schilling et al. (1992) model for the Gulf of Aden, Gulf of Tadjoura, and Afar basalts. Here we retain the isotopic composition of the Pan-African lithosphere and depleted mantle end-members proposed by Schilling et al. (1992). We further extend and modify the three-component

Southeast Indian Ridge (after Hanan *et al.*, 2004). Pacific MORB have  $\Delta^{208} Pb/^{204} Pb < 20$ , whereas Indian Ocean MORB have  $\Delta^{208} Pb/^{204} Pb > 20$ . In terms of  $\Delta \epsilon_{Hf}$ , the Pacific data are confined to the SW quadrant of the diagram with  $\Delta \epsilon_{Hf} < 1$ . Depleted Indian Ocean MORB have positive  $\Delta \epsilon_{Hf}$ , and plot in the NE quadrant with  $\Delta \epsilon_{Hf} > 1$ . The distinction between the Pacific and Indian Ocean MORB derives from pollution of the mantle source of the latter by continental material (Hanan *et al.*, 2004). The data for the Main Ethiopian Rift and the Gulf of Aden show Hf and Pb isotope signatures similar to those of Indian Ocean MORB derived from mantle sources polluted with continental material. Debre Zeyit sample DZ-1003 is labeled.

Table 6: Isotope composition and elemental abundances for the model end-members

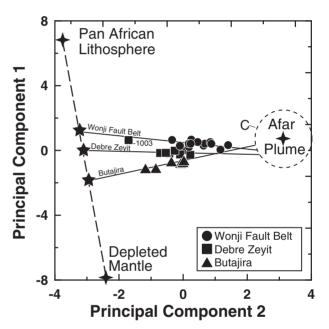
	Sr (ppm)	Nd (ppm)	Pb (ppm)	<sup>87</sup> Sr/ <sup>86</sup> Sr	<sup>143</sup> Nd/ <sup>144</sup> Nd	<sup>206</sup> Pb/ <sup>204</sup> Pb	<sup>207</sup> Pb/ <sup>204</sup> Pb	<sup>208</sup> Pb/ <sup>204</sup> Pb
Afar plume	160	14	29	0.7035	0.512875	19.5	15.6	39-2
DM	75	4	5	0.7022	0.51335	17·5	15-3	36-6
Pan-African	225	22	39	0.7075	0.5121	17.85	15.75	39.75

These values are identical to those reported by Schilling *et al.* (1992) with the exception of the Afar plume end-member. Based on the present study, the Afar plume is now considered to be of a C mantle composition (Hanan & Graham, 1996) and the average isotopic values of C are used here. DM, depleted mantle; Pan African, Pan-African lithosphere. [See the Appendix of this paper and Schilling *et al.* (1992, appendix) for more model details.]

mixing model between the Afar plume, depleted upper mantle, and Pan-African lithosphere proposed by Schilling *et al.* (1992) and apply the model to the new Main Ethiopian Rift data (Tables 1–6).

In summary, the three end-member isotope compositions for the Main Ethiopian Rift ternary mixing model are dictated by the following constraints (see also Schilling et al., 1992 for additional detail): (1) they must span the entire Main Ethiopian Rift dataset in Pb-Nd-Hf-Sr multi-isotope space (i.e. define a triangular space that surrounds the Main Ethiopian Rift dataset); (2) the convergence of the three pseudo-binary mixing trends for the Debre Zeyit, Butajira, and Wonji Fault Belt sub-populations constrains the composition of the radiogenic Pb end-member (the Afar plume); (3) the Afar plume has the isotopic composition of C (Hanan & Graham, 1996; Hanan et al., 2000); (4) the least radiogenic of the two low <sup>206</sup>Pb/<sup>204</sup>Pb end-members is assumed to represent the depleted asthenosphere sampled along normal mid-ocean ridge segments; (5) the low <sup>206</sup>Pb/<sup>204</sup>Pb with a relatively high <sup>207</sup>Pb/<sup>204</sup>Pb lithospheric-like end-member represents Pan-African lithosphere and is EM-like in its isotopic composition. We retain the Nd, Sr, and Pb isotopic compositions and elemental concentrations of the Pan-African lithosphere and depleted asthenosphere components used by Schilling et al. (1992). The Main Ethiopian Rift mixing model differs from the previous model in the choice of the Afar plume isotope composition, as outlined above. The Hf isotope data (Tables 1 and 2; Figs 4 and 5a) for the Gulf of Aden and the Main Ethiopian Rift strongly support the suggestion that the Afar plume composition is C-like rather than of HIMU composition.

The choice of end-members and the validity of the three-component mixing model deduced from the Pb isotope data were tested using principal component analysis (Fig. 6; for an explanation of principal component analysis applied to basalts see, for example, Schilling *et al.*, 1992;



**Fig. 6.** PCl–PC2 projection of a principal component analysis for  $^{206}\text{Pb}/^{204}\text{Pb},~^{207}\text{Pb}/^{204}\text{Pb},~^{208}\text{Pb}/^{204}\text{Pb},~\epsilon_{Nd},~\text{and}~\epsilon_{Sr}~\text{using}$  the Main Ethiopian Rift data and the end-members defined in Table 6. The lines are least-squares regressions through the three sub-populations of data from Debre Zeyit, Butajira, and the Wonji Fault Belt. The labeled point (DZ-1003) was included in the principal component analysis, but omitted from the Debre Zeyit regression because it contains norite xeno-liths and may have been affected by crustal assimilation associated with volcanism at a nearby silicic volcanic center (Rooney *et al.*, 2005). The eigenvectors PCl =70·9% and PC2 = 27·2% account for >98·1% of the total variance. The three-component mixing model is justified because two principal components (and thus, three geochemical end-members) sufficiently account for the Sr–Nd–Pb isotopic variability in the Main Ethiopian Rift lavas, and the three end-members lie within a single plane of mixing. The approximate field of the C-like Afar plume is shown as a dashed-line circle.

Hanan & Schilling, 1997; Blichert-Toft et al., 2003, 2005; Agranier et al., 2005; Debaille et al., 2006). Considering only the Pb isotopes, 99.5% of the population variance is accounted for within the plane of the first two

eigenvectors. Principal component analysis using 206Pb/ <sup>204</sup>Pb-<sup>207</sup>Pb/<sup>204</sup>Pb-<sup>208</sup>Pb/<sup>204</sup>Pb-<sup>87</sup>Sr/<sup>86</sup>Sr-<sup>143</sup>Nd/<sup>144</sup>Nd vields an equivalent result, with the first two eigenvectors accounting for 98·1% of the variance. Hf isotopes were not included given the smaller dataset and the principal component analysis requirement that each sample has the same number of variables. Furthermore, we have not assigned Hf isotope values to the end-members because (1) there are few or no Hf isotope data available to characterize the continental lithosphere in this region, and (2) the sparse Hf isotope data for the Main Ethiopian Rift and East African Rift in general do not currently show enough spread to determine accurate mixing trajectories. The principal component analysis results suggest that hyperbolic mixing of the geochemical end-members is minimal; that is, the elemental ratios of Pb, Sr, and Nd, in the hybrid end-member (comprising asthenospheric upper mantle and a small proportion of continental lithosphere) are approximately equal to those in the Afar plume end-member, resulting in linear mixing trends in multi-isotope ratio plots (e.g. Hanan & Schilling, 1997; Douglass & Schilling, 2000). The linear arrays in multiisotope space define converging pseudo-binary mixing trends (Figs 3, 4 and 6). The converging trends for the Wonji Fault Belt, Debre Zevit, and Butajira lavas apparently result from a special case of ternary mixing (case 3 described by Douglass & Schilling, 2000) in which the mass proportions of two of the end-members (depleted upper mantle and continental lithosphere) co-vary systematically along the rift, maintaining an essentially fixed ratio within each extensional zone. This type of mixing (in which the ratio of asthenosphere to continental lithosphere mass fraction is nearly constant, whereas the ratio of plume to asthenosphere markedly varies) is difficult to attribute to a single-stage mixing event, because in that case all of the components should be equally well mixed (Douglass & Schilling, 2000). An ordered, two-stage mixing process is required, during which the asthenosphere and continental lithosphere interact to form a hybrid, before mixing with the Afar plume.

We carried out the three-component Pb, Sr, and Nd isotope mixing model using basic principles described previously (Schilling *et al.*, 1992; Hanan & Schilling, 1997; Douglass & Schilling, 2000; Hanan *et al.*, 2000) (Fig. 7). The modeling indicates that the mixtures are dominated by the asthenosphere (depleted mantle component) (mass fraction = 49–80%), and that the mass fraction of the Afar plume (C-like Afar plume component) (10–43%) is larger than that of the continental lithosphere (Pan-African lithosphere component) (6–17%). The mass fractions of the plume and depleted mantle components co-vary ( $r^2$ =0·92), and the contribution from the Afar plume decreases systematically away from the hypothesized plume center beneath Lake Abhe (Schilling *et al.*, 1992)

(Fig. 7). The <sup>3</sup>He/<sup>4</sup>He ratio (Table 3) broadly co-varies positively with the mass proportion of the Afar plume in the three-component mixture, providing further evidence that the converging pseudo-binary mixing arrays in isotope–isotope space can be used to constrain the Pb–Sr–Nd isotopic composition of the Afar plume component.

Quaternary basalts (MgO >7 wt %) from the Djibouti and Red Sea region [e.g. the Quaternary Moussa Ali volcano and the Maranda-Inkir volcanic centers located NE of the Afar plume center, and from the Sept Frères volcanic islands in the Red Sea (Deniel et al., 1994)] provide additional support for the model of radial dispersion from a plume centered beneath the Lake Abhe region (Fig. 4). The Djibouti and Red Sea data fall along extensions of the converging Pb isotope trends for the Main Ethiopian Rift lavas, within the region defined by the three-component mixing model end-members

# Plume-lithosphere interaction in the East African Rift system

Previous studies have shown that the extent of upper mantle pollution in the vicinity of an upwelling and dispersing plume decreases systematically with radial distance from the plume center (Schilling, 1985; Hanan *et al.*, 1986; Hanan & Schilling, 1989). Afar plume influence has been observed to extend from Afar for ~700 km along the Gulf of Aden, into the Red Sea, and along the Ethiopian sector of the East African Rift (Barrat *et al.*, 1990; Schilling *et al.*, 1992; Marty *et al.*, 1993, 1996; Deniel *et al.*, 1994; Scarsi & Craig, 1996; Moreira *et al.*, 1999; Hopp *et al.*, 2004). The Afar plume head is envisioned to have flattened beneath the lithosphere like a pancake, and now has a radial influence of the order of 700–1400 km (Schilling *et al.*, 1992; Volker *et al.*, 1997).

Based on the laboratory models of Griffiths & Campbell (1990), Schilling et al. (1992) proposed that the flattened Afar plume head has a torus-like structure, containing an outer ring of deeply derived material located about 450 km (± 150 km) from the center of the upwelling, and an inner region of entrained, depleted mantle. In this model the plume tail appears to be tilted toward Afar (Schilling et al., 1992). Along the Sheba Ridge an anomaly in radiogenic Pb, Sr, Nd, and Hf isotopes is located above the proposed torus ring. East of the anomaly (48°E) the MORB are LILE-depleted and show a slight increase in radiogenic Pb and Sr isotope compositions (Schilling et al., 1992). To the west of the anomaly, lavas from the Gulf of Tadjoura and Asal Rift are more similar to the Main Ethiopian Rift lavas compared with the West Sheba Ridge MORB, and have Pb, Sr, Nd, and Hf isotope compositions that indicate a greater involvement of continental lithosphere in their genesis. The western part of the proposed torus would underlie the Main Ethiopian Rift in

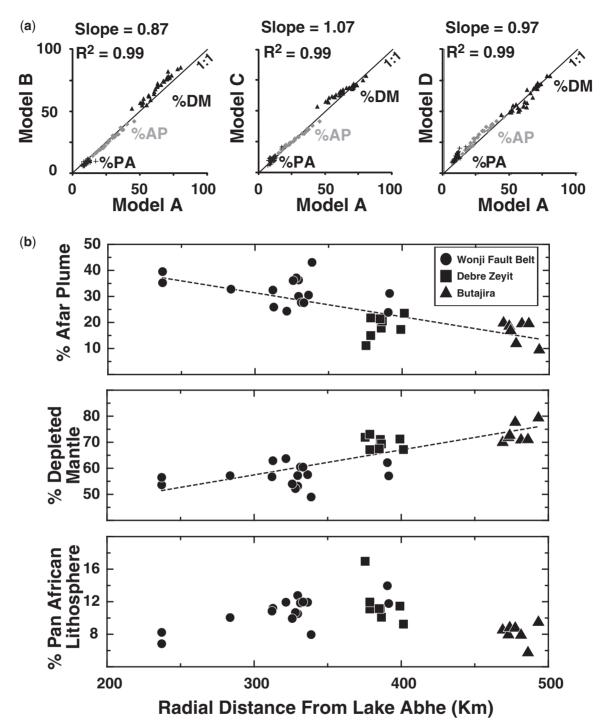


Fig. 7. (a) Correlation diagrams for mixing model solutions. With this model, a complete solution providing the percentage of plume, depleted mantle, and continental lithosphere for a single sample can only be determined using two isotope ratios at a time. Thus each 'model' uses a pair of isotopes from each sample to yield percentages of each component (PA, Pan-African; AP, Afar plume; DM, depleted mantle). We have compared the results of using one pair of isotopes with an alternate pair to test the model's self-consistency. Model  $A = {}^{207}Pb/{}^{204}Pb - {}^{206}Pb/{}^{204}Pb - {}^{206}Pb/{}^{2$ 

southernmost Ethiopia, where the plume tail is supposedly tilted toward the west. Geophysical studies have not imaged a distinct tail structure, but instead show a broad low-velocity structure tilted westward (Benoit et al., 2006). According to the torus model, our new Main Ethiopian Rift sample locations would correspond to a traverse from inside the plume torus towards and across the edge of the proposed location of the western ring. The observed gradient of decreasing radiogenic Pb, Sr, and Nd away from the plume center and toward the proposed ring location is opposite to what would be predicted by the torus model. Furthermore, melting in the center of the torus plume was suggested to be dominated by lithospheric mantle and depleted asthenosphere entrained by the plume. However, based on the regular spatial variations in the Pb-Sr-Nd-Hf-He isotope ratios along the Main Ethiopian Rift, and the results of our mixing model, it is apparent that all three mantle source components, Pan-African lithosphere, depleted mantle, and the Afar plume, contribute to the petrogenesis of the Main Ethiopian Rift basalts located in the region of the proposed inner ring. The Asal rift and the Gulf of Tadjoura lavas have Pb isotope ratios consistent with radial dispersion of the plume and their distance of about 150-300 km from the plume center. We observe a gradient in plume influence (Figs 2 and 7) along the Main Ethiopian Rift that is expressed by a systematic co-variation between plume and depleted asthenosphere mass fractions (Fig. 7). The near-constant ratio of African continental lithosphere to depleted upper mantle for each of the extensional zones, revealed by the pseudo-binary arrays for the Wonji Fault Belt, Debre Zeyit, and Butajira subgroups, requires two stages of mixing (Douglass & Schilling, 2000). The first stage involves mixing of asthenospheric mantle and continental lithosphere, followed by a second stage during which this hybrid upper mantle mixes with the Afar plume. Detachment and foundering of lithosphere during rifting is one mechanism for contaminating the upper mantle with continental material (Arndt & Goldstein, 1989). This process may occur when the lithosphere becomes ductile, as a result of thermal or flexural weakening during extension (Burov et al., 2007). The hot Afar plume head beneath the rift also contributes to thermal weakening by conductive heating. As the lithosphere is heated and thinned during plume impingement and incipient rifting (e.g. Kent et al., 1992), portions of the uppermost mantle beneath the lithospheric mechanical boundary layer (McKenzie & Bickle, 1988) are admixed with lithospheric material. This upper mantle pollution with continental lithospheric material occurs prior to when the uppermost mantle becomes a sink for the buoyant and dispersing plume (Sleep, 2008). This process is influenced by the underside topography of the continental lithosphere, which may channel plume material into 'thin spots' without involving plume-driven thinning (Thompson & Gibson, 1991; Ebinger & Sleep, 1988; Sleep et al., 2002; Shervais & Hanan, 2008). For example, the isotopic anomaly on the West Sheba Ridge adjacent to the transition between continental and oceanic lithosphere may relate to a step decrease in lithospheric thickness that acts as a sink for plume material (Fig. 2c). The two-stage ordered mixing along the Main Ethiopian Rift suggests that the sub-continental lithosphere acted as a levee against plume flow along the rift until the tectonic forces and geometry of extension associated with the triple junction creation led to formation of the Main Ethiopian Rift arm.

# Implications for the composition of the upper mantle

The sequence of mixing described above has implications for the origin of upper mantle heterogeneity, as well as for the composition and origin of mantle plumes. Asthenospheric mantle that was once associated with continental rifting may retain evidence of continental lithosphere contamination (McKenzie & O'Nions, 1983; Hawkesworth et al., 1986; Arndt & Goldstein, 1989; Hanan & Graham, 1996; Hanan et al., 2004). Pollution of the upper mantle by continental lithosphere (and/or lower crust) has been observed in the Tertiary basalts of Iceland (Hanan & Schilling, 1997; Hanan et al., 2000), along the Mid-Atlantic Ridge north of Iceland (Andres et al., 2004; Blichert-Toft et al., 2005) and along the Gakkel Ridge in the Arctic (Goldstein et al., 2008), in the South Atlantic (Hawkesworth et al., 1986; Carlson et al., 1996; Kamenetsky et al., 2001; Andres et al., 2002) and along the Southwest (Escrig et al., 2005; Janney et al., 2005; Meyzen et al., 2007) and Southeast Indian Ridges (Hanan et al., 2004; Graham et al., 2006). Our observations and model for the Ethiopian sector of the East African Rift indicate that mantle plume material interacts with continental lithosphere and upper mantle only after thermomechanical thinning of the lithosphere has progressed sufficiently to form a sink for the rising and dispersing plume material. Lithospheric detachment and mixing into the asthenosphere during continental rift evolution may, therefore, account for much of the range of ambient upper mantle compositions sampled by mid-ocean ridge volcanism away from island hotspots. A corollary is that the isotopic composition of plume-related lavas, such as continental flood basalts and some oceanic islands, may not always represent the isotopic composition of the deep mantle plume source alone, but instead carry an integrated record of plume interaction and mixing with upper mantle that has been variably polluted with continental material such as the Yellowstone Plume (Hanan et al., 2008) and eastern Atlantic cases (Geldmacher et al., 2011).

#### CONCLUSIONS

The isotopic variations of basalts erupted along the Main Ethiopian Rift at Debre Zeyit, Butajira, and the Wonji Fault Belt result from mixing between the Afar plume, the asthenospheric upper mantle and the African continental lithosphere. The isotopic composition of the Afar plume resembles the 'C' component on the basis of converging arrays in multi-isotope space. The influence of this plume signature systematically decreases along the Main Ethiopian Rift away from Afar into southern Ethiopia. Mixing between the plume, lithosphere, and ambient upper mantle follows an ordered sequence of events. Within each volcanic belt the Afar plume mixes with a hybrid reservoir that is composed of different proportions of lithosphere and depleted mantle. Within each volcanic belt, the mass proportions of continental lithosphere and upper mantle are relatively constant, requiring that the mixing event that generated these hybrid reservoirs occurred prior to their mixing with the Afar plume.

Detachment and dispersal of continental lithosphere into the upper mantle during plume—rift interaction follows an ordered sequence of mixing events. This mixing produces large-scale pollution of the upper mantle by continental lithosphere that is recorded in the 'background' geochemical signatures of MORB, even those erupting far away from mantle hotspots or plumes. Furthermore, the distinct isotopic compositions of plume-related lavas may not necessarily represent intrinsic mantle source signatures, but rather a complex history of mixing involving components derived from both the upper and lower mantle.

# **ACKNOWLEDGEMENTS**

We thank Kassahun Ejeta, Roeland Doust, Gezahegn Yirgu, and Dereje Ayalew for field assistance and discussion in Ethiopia, Margaret Nitz and Leigh Patterson for help with sample preparation, and Joan Willis and Jasper Konter for their assistance in the clean room. We thank John Lupton for access to the helium isotope laboratory in Newport, OR, which is supported by the NOAA Vents Program. Comments from Nick Rogers helped to clarify early versions of the paper. Helpful comments were provided by Andrew Kerr, two anonymous reviewers, and editor Dominique Weis. Finally, we would like to thank the EAGLE group and in particular Ian Bastow for stimulating discussions.

#### **FUNDING**

This work was supported by National Science Foundation grants to B. Hanan, D. Graham, and T. Furman, and a George H. Deike Jr grant to T. Furman. J. Blichert-Toft received financial support from the French Institut National des Sciences de l'Univers.

#### SUPPLEMENTARY DATA

Supplementary data for this paper are available at *Journal* of *Petrology* online.

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### **APPENDIX**

### TERNARY MIXING MODEL

The isotope composition of a multi-component mixture is a weighted average of the components contributing to the mixture:

$$R_{\rm m} = \sum c_i w_i R_i / \sum c_i w_i \tag{1}$$

where R is the isotope ratio (of Pb, Nd or Sr), subscript i refers to component a, b or c (here a indicates the Afar plume, b depleted mantle, c Pan-African lithosphere), m is the mixture (sample),  $c_i$  is the Pb, Nd, or Sr concentration in component i, and  $w_i$  is the mass fraction contributed by each component i to the mixture.

We also have the closure condition

$$\sum w_i = w_a + w_b + w_c = 1.$$
 (2)

Therefore, when all end-member  $R_i$  and  $c_i$  values are known, the measurement of two isotope ratios (such as  $^{206}\text{Pb}/^{204}\text{Pb}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$ ) for any sample leads to three equations (one for each  $R_{\rm m}$ , plus the closure equation) and three unknowns ( $w_i$  values), and the system is fully solvable.

The chosen elemental concentrations and isotopic end-members are identical to those of Schilling *et al.* (1992), except the plume component, which was given a value corresponding to the C mantle reservoir (Hanan & Graham, 1996). A rationale for the end-member isotopic compositions has been given in the text and by Schilling *et al.* (1992). The derivation of the mixing model, rationale, and method for determining the end-member concentrations have been given in detail by Schilling *et al.* (1992, appendix).

Table A1: Major and trace element data for samples from the Wonji Fault Belt following the methods outlined by Furman et al. (2006)

Location	Lat. (°N)		-		* SiC	) <sub>2</sub> N	lgO	CaO	Mn	0 -	ΓiO <sub>2</sub>	Fe <sub>2</sub> O <sub>3</sub>	Na <sub>2</sub> C	) A	l <sub>2</sub> O <sub>3</sub>	K <sub>2</sub> O	P <sub>2</sub> O <sub>5</sub>	Rb	Sr	Ba	Zr	Υ				
WFB-N-14 WFB-1026				-		01 6· 16 5·		11·34 9·70				11·39 14·28				0·43 0·65		7·1 14·1	514 395		87 159		-			
	Cs	Hf	U	Th	Nb	Та	Sc	Со	Ni	Cr	٧	La	Се	Pr	Nd	Sm	Eu	Gd	Tb	Dγ	<b>/</b> Н	ło	Er	Yb	Lu	Pb
WFB-N-14 WFB-1026			0·26 0·47					9 46 7 42			306 305						9 1·42 4 2·32								0·21 0·41	

<sup>\*</sup>Radial distance from Lake Abhe, Afar.

Table A2: Model mass proportions of the end-members for the Main Ethiopian Rift data from Table 1

ample	% A	far plun	ne using	model:		% De	epleted n	nantle usir	ng model:		% Pan-African lithosphere using model:						
	Α	В	С	D	Av.	А	В	С	D	Av.	А	В	С	D	Av.		
VFB-1023	33	31	31	37	33	57	61	61	49	57	10	8	8	14	10		
VFB-1027	36	36	33	36	35	55	54	61	55	57	9	9	6	9	8		
VFB-1026	41	39	37	41	40	52	54	57	51	54	8	7	5	8	7		
VFB-E99-2	26	24	25	28	26	62	67	66	57	63	12	9	10	14	11		
VFB-N-18	34	31	31	34	32	53	60	60	54	57	13	9	9	12	11		
/FB-1037	24	22	24	27	24	63	70	65	56	64	12	8	11	16	12		

(continued)

Table A2: Continued

WFB-KO-12 30 28 29 34 30 57 61 60 50 57 13 11 11 11 17 WFB-KO-4B 38 35 34 39 36 51 56 57 49 53 12 9 9 13 WFB-N-20 37 35 n.a. 39 37 52 56 n.a. 49 52 11 9 n.a. 12 WFB-1035 27 26 27 31 28 61 65 62 54 61 12 9 11 15 WFB-N-15 30 30 28 34 31 58 59 63 51 58 12 11 9 15 WFB-N-15 30 30 28 34 31 58 59 63 51 58 12 11 9 15 WFB-N-15 30 30 28 34 31 58 59 63 51 58 12 11 9 9 12 WFB-N-16 37 35 35 38 36 53 57 56 50 54 11 9 9 12 WFB-N-21 27 26 26 31 28 61 63 64 53 60 12 10 10 16 WFB-P-P-20 31 29 n.a. 33 31 57 62 n.a. 52 57 12 9 n.a. 14 WFB-N-01 23 21 23 27 24 63 69 64 53 62 13 10 13 20 DZ-1003 11 10 12 12 11 71 81 68 67 72 18 9 20 21 DZ-1004 23 20 22 22 22 22 63 72 67 67 67 67 13 8 11 11 DZ-1005 15 14 15 15 15 15 71 77 71 72 73 13 9 13 13 DZ-1013 18 16 18 18 18 17 70 76 70 69 71 12 8 12 12 13 DZ-1009 21 19 21 24 21 68 74 68 61 67 11 7 11 15 DZ-1007 18 16 18 18 18 18 70 75 70 69 71 12 8 12 8 12 12 DZ-1008 21 19 21 21 21 21 68 73 68 68 69 11 8 11 11	Sample	% A	far plun	ne using r	model:		% De	epleted n	nantle using	g model:		% Pa	n-African	lithosphere ι	ising mode	el:
WFB-KO-48         38         35         34         39         36         51         56         57         49         53         12         9         9         13           WFB-N-20         37         35         n.a.         39         37         52         56         n.a.         49         52         11         9         n.a.         12           WFB-1035         27         26         27         31         28         61         65         62         54         61         12         9         11         15           WFB-N-22         46         41         41         44         43         44         52         53         47         49         10         7         6         9           WFB-N-15         30         30         28         34         31         58         59         63         51         58         12         111         9         15           WFB-N-15         30         30         28         34         31         58         59         63         51         58         12         111         9         16           WFB-N-12         27         26         26		Α	В	С	D	Av.	Α	В	С	D	Av.	А	В	С	D	Av.
WFB-N-20   37   35   n.a.   39   37   52   56   n.a.   49   52   11   9   n.a.   12   NFB-N-20   37   35   n.a.   39   37   52   56   n.a.   49   52   11   9   n.a.   12   NFB-N-22   46   41   41   44   43   44   52   53   47   49   10   7   6   9   NFB-N-15   30   30   28   34   31   58   59   63   51   58   12   11   9   9   15   NFB-N-21   27   26   26   31   28   61   63   64   53   60   12   10   10   16   NFB-E-99-20   31   29   n.a.   33   31   57   62   n.a.   52   57   12   9   n.a.   14   NFB-N-01   23   21   23   27   24   63   69   64   53   62   13   10   13   20   02-1003   11   10   12   12   11   71   81   68   67   72   18   9   20   21   02-1004   23   20   22   22   22   23   25   24   66   71   77   71   72   73   13   9   13   13   02-1013   18   16   18   18   17   70   76   70   69   71   12   8   12   13   02-1014   24   22   23   25   24   66   71   68   61   67   11   7   7   7   11   15   02-1006   21   19   21   21   21   21   68   73   68   68   69   70   72   8   5   9   9   9   9   9   9   9   9   9	WFB-KO-12	30	28	29	34	30	57	61	60	50	57	13	11	11	17	13
WFB-1035	WFB-KO-4B	38	35	34	39	36	51	56	57	49	53	12	9	9	13	11
WFB-N-22         46         41         41         44         43         44         52         53         47         49         10         7         6         9           WFB-N-15         30         30         28         34         31         58         59         63         51         58         12         11         9         15           WFB-1036         37         35         35         38         36         53         57         56         50         54         11         9         9         12           WFB-N-21         27         26         26         31         28         61         63         64         53         60         12         10         10         16           WFB-N-21         27         26         26         31         28         61         63         64         53         60         12         10         10         16           WFB-N-01         23         21         23         27         24         63         69         64         53         62         13         10         13         20           27-1003         11         10         12	WFB-N-20	37	35	n.a.	39	37	52	56	n.a.	49	52	11	9	n.a.	12	11
WFB-N-15 30 30 28 34 31 58 59 63 51 58 12 11 9 15 WFB-N36 37 35 35 38 36 53 57 56 50 54 11 9 9 12 WFB-N-21 27 26 26 26 31 28 61 63 64 53 60 12 10 10 16 WFB-E99-20 31 29 n.a. 33 31 57 62 n.a. 52 57 12 9 n.a. 14 WFB-N-01 23 21 23 27 24 63 69 64 53 62 13 10 13 20 D2-1003 11 10 12 12 11 71 81 68 67 72 18 9 20 21 D2-1004 23 20 22 22 22 63 72 67 67 67 67 13 8 11 11 D2-1005 15 14 15 15 15 15 71 77 71 72 73 13 9 13 13 D2-1013 18 16 18 18 17 70 76 70 69 71 12 8 12 12 13 D2-1014 24 22 23 25 24 66 71 68 64 67 10 7 9 11 D2-1009 21 19 21 24 21 68 74 68 61 67 11 7 12 8 12 D2-1007 18 16 18 18 18 17 70 75 70 69 71 12 8 12 D2-1008 21 19 21 21 21 68 73 68 68 69 11 8 11 11 B3-1044 20 19 21 21 21 20 71 76 69 70 72 8 5 9 9 BJ-1048 10 9 11 11 10 80 84 78 78 80 9 6 11 12 BJ-1049 13 11 14 13 13 78 83 74 78 78 9 5 12 9 BJ-1051 19 18 20 20 19 72 77 70 71 73 8 6 9 9 9 BJ-1051 19 18 20 20 19 72 77 70 71 73 8 6 9 9 9 BJ-1051 19 18 20 20 19 72 77 70 71 73 8 6 9 9 9 BJ-1051 19 18 20 20 19 72 77 70 71 73 8 6 9 9 9 BJ-1051 19 18 20 20 19 72 77 70 71 73 8 6 9 9 9 BJ-1051 19 18 20 20 19 72 77 70 71 73 8 6 9 9 9 BJ-1051 19 18 20 20 19 72 77 70 71 73 8 6 9 9 9	WFB-1035	27	26	27	31	28	61	65	62	54	61	12	9	11	15	12
NFB-1036 37 35 35 38 36 53 57 56 50 54 11 9 9 9 12  NFB-1036 37 35 35 38 36 53 57 56 50 54 11 9 9 9 12  NFB-N-21 27 26 26 26 31 28 61 63 64 53 60 12 10 10 10 16  NFB-E99-20 31 29 n.a. 33 31 57 62 n.a. 52 57 12 9 n.a. 14  NFB-N-01 23 21 23 27 24 63 69 64 53 62 13 10 13 20  DZ-1003 11 10 12 12 11 71 81 68 67 72 18 9 20 21  DZ-1004 23 20 22 22 22 63 72 67 67 67 67 13 8 11 11  DZ-1005 15 14 15 15 15 15 71 77 71 72 73 13 9 13 13  DZ-1013 18 16 18 18 17 70 76 70 69 71 12 8 12 13  DZ-1009 21 19 21 24 21 68 74 68 61 67 11 7 11 15  DZ-1009 21 19 21 24 21 68 74 68 61 67 11 7 11 15  DZ-1007 18 16 18 18 18 18 70 75 70 69 71 12 8 12 12 12  DZ-1008 21 19 21 21 21 21 68 73 68 68 69 11 8 11 11  DZ-1008 21 19 21 21 21 20 71 76 69 70 72 8 5 9 9  SJ-1044 20 19 11 11 10 80 84 78 78 78 80 9 6 11 12  SJ-1049 13 11 14 13 13 78 83 74 78 78 78 78 9 5 12 9  SJ-1045 21 19 18 20 20 19 72 77 70 71 73 8 6 6 9 9	NFB-N-22	46	41	41	44	43	44	52	53	47	49	10	7	6	9	8
AVFB-N-21         27         26         26         31         28         61         63         64         53         60         12         10         10         16           AVFB-P99-20         31         29         n.a.         33         31         57         62         n.a.         52         57         12         9         n.a.         14           AVFB-N-01         23         21         23         27         24         63         69         64         53         62         13         10         13         20           DZ-1003         11         10         12         12         11         71         81         68         67         72         18         9         20         21           DZ-1004         23         20         22         22         22         63         72         67         67         67         13         8         11	NFB-N-15	30	30	28	34	31	58	59	63	51	58	12	11	9	15	12
NFB-E99-20 31 29 n.a. 33 31 57 62 n.a. 52 57 12 9 n.a. 14 NFB-NOT 23 21 23 27 24 63 69 64 53 62 13 10 13 20 NFB-NOT 23 21 23 27 24 63 69 64 53 62 13 10 13 20 NFB-NOT 23 20 21 12 11 71 81 68 67 72 18 9 20 21 NFB-NOT 23 20 22 22 22 63 72 67 67 67 13 8 11 11 NFB-NOT 105 15 14 15 15 15 15 71 77 71 72 73 13 9 13 13 NFB-NOT 105 15 14 15 15 15 15 71 77 71 72 73 13 9 13 13 NFB-NOT 105 15 14 15 15 15 15 71 77 71 72 73 13 9 13 13 NFB-NOT 105 15 14 15 15 15 15 71 77 71 72 73 13 9 13 13 NFB-NOT 105 15 14 15 15 15 15 71 77 71 72 73 13 13 9 13 13 NFB-NOT 105 15 14 15 15 15 15 71 77 71 72 73 13 13 9 13 13 NFB-NOT 105 15 14 15 15 15 15 71 77 71 71 72 73 13 13 9 13 13 NFB-NOT 105 15 14 15 15 15 15 71 77 71 71 72 73 13 13 9 13 13 NFB-NOT 105 15 14 15 15 15 15 71 77 71 71 72 73 13 13 9 13 13 NFB-NOT 105 15 14 15 15 15 15 15 71 77 71 71 72 73 13 13 9 13 13 NFB-NOT 105 15 14 15 15 15 15 15 71 77 71 71 72 73 13 13 9 13 13 NFB-NOT 105 15 14 15 15 15 15 15 15 71 77 71 71 72 73 13 13 9 13 13 NFB-NOT 105 15 14 15 15 15 15 15 15 71 77 71 70 71 12 8 12 12 12 12 12 12 12 12 12 12 12 12 12	WFB-1036	37	35	35	38	36	53	57	56	50	54	11	9	9	12	10
NYFB-N-01         23         21         23         27         24         63         69         64         53         62         13         10         13         20           DZ-1003         11         10         12         12         11         71         81         68         67         72         18         9         20         21           DZ-1004         23         20         22         22         22         63         72         67         67         67         13         8         11         11           DZ-1005         15         14         15         15         15         71         77         71         72         73         13         9         13         13           DZ-1013         18         16         18         18         17         70         76         70         69         71         12         8         12         13           DZ-1014         24         22         23         25         24         66         71         68         64         67         10         7         9         11           DZ-1009         21         19         21 <t< td=""><td>WFB-N-21</td><td>27</td><td>26</td><td>26</td><td>31</td><td>28</td><td>61</td><td>63</td><td>64</td><td>53</td><td>60</td><td>12</td><td>10</td><td>10</td><td>16</td><td>12</td></t<>	WFB-N-21	27	26	26	31	28	61	63	64	53	60	12	10	10	16	12
DZ-1003       11       10       12       12       11       71       81       68       67       72       18       9       20       21         DZ-1004       23       20       22       22       22       63       72       67       67       67       13       8       11       11         DZ-1005       15       14       15       15       15       71       77       71       72       73       13       9       13       13         DZ-1013       18       16       18       18       17       70       76       70       69       71       12       8       12       13         DZ-1014       24       22       23       25       24       66       71       68       64       67       10       7       9       11         DZ-1009       21       19       21       24       21       68       74       68       61       67       11       7       11       15         DZ-1007       18       16       18       18       18       70       75       70       69       71       12       8       12       12	VFB-E99-20	31	29	n.a.	33	31	57	62	n.a.	52	57	12	9	n.a.	14	12
DZ-1004       23       20       22       22       22       63       72       67       67       67       13       8       11       11         DZ-1005       15       14       15       15       15       71       77       71       72       73       13       9       13       13         DZ-1013       18       16       18       18       17       70       76       70       69       71       12       8       12       13         DZ-1014       24       22       23       25       24       66       71       68       64       67       10       7       9       11         DZ-1009       21       19       21       24       21       68       74       68       61       67       11       7       11       15         DZ-1009       21       19       21       24       21       68       74       68       61       67       11       7       11       15         DZ-1007       18       16       18       18       18       70       75       70       69       71       12       8       12       12	VFB-N-01	23	21	23	27	24	63	69	64	53	62	13	10	13	20	14
DZ-1005       15       14       15       15       15       17       77       71       72       73       13       9       13       13         DZ-1013       18       16       18       18       17       70       76       70       69       71       12       8       12       13         DZ-1014       24       22       23       25       24       66       71       68       64       67       10       7       9       11         DZ-1009       21       19       21       24       21       68       74       68       61       67       11       7       11       15         DZ-1007       18       16       18       18       18       70       75       70       69       71       12       8       12       12         DZ-1008       21       19       21       21       21       68       73       68       68       69       11       8       11       11         3J-1044       20       19       21       21       20       71       76       69       70       72       8       5       9       9 <td>DZ-1003</td> <td>11</td> <td>10</td> <td>12</td> <td>12</td> <td>11</td> <td>71</td> <td>81</td> <td>68</td> <td>67</td> <td>72</td> <td>18</td> <td>9</td> <td>20</td> <td>21</td> <td>17</td>	DZ-1003	11	10	12	12	11	71	81	68	67	72	18	9	20	21	17
DZ-1013       18       16       18       18       17       70       76       70       69       71       12       8       12       13         DZ-1014       24       22       23       25       24       66       71       68       64       67       10       7       9       11         DZ-1009       21       19       21       24       21       68       74       68       61       67       11       7       11       15         DZ-1007       18       16       18       18       18       70       75       70       69       71       12       8       12       12         DZ-1008       21       19       21       21       21       68       73       68       68       69       11       8       11       11         3J-1044       20       19       21       21       20       71       76       69       70       72       8       5       9       9         3J-1048       10       9       11       11       10       80       84       78       78       80       9       6       11       12	DZ-1004	23	20	22	22	22	63	72	67	67	67	13	8	11	11	11
02-1014       24       22       23       25       24       66       71       68       64       67       10       7       9       11         02-1009       21       19       21       24       21       68       74       68       61       67       11       7       11       15         02-1007       18       16       18       18       18       75       70       69       71       12       8       12       12         02-1008       21       19       21       21       21       68       73       68       68       69       11       8       11       11         3J-1044       20       19       21       21       20       71       76       69       70       72       8       5       9       9         3J-1048       10       9       11       11       10       80       84       78       78       80       9       6       11       12         3J-1049       13       11       14       13       13       78       83       74       78       78       9       5       12       9	Z-1005	15	14	15	15	15	71	77	71	72	73	13	9	13	13	12
DZ-1009       21       19       21       24       21       68       74       68       61       67       11       7       11       15         DZ-1007       18       16       18       18       18       18       75       70       69       71       12       8       12       12         DZ-1008       21       19       21       21       21       68       73       68       68       69       11       8       11       11         3J-1044       20       19       21       21       20       71       76       69       70       72       8       5       9       9         3J-1048       10       9       11       11       10       80       84       78       78       80       9       6       11       12         3J-1049       13       11       14       13       13       78       83       74       78       78       9       5       12       9         3J-1051       19       18       20       20       19       72       77       70       71       73       8       6       9       9	DZ-1013	18	16	18	18	17	70	76	70	69	71	12	8	12	13	11
DZ-1007       18       16       18       18       18       70       75       70       69       71       12       8       12       12         DZ-1008       21       19       21       21       21       68       73       68       68       69       11       8       11       11         BJ-1044       20       19       21       21       20       71       76       69       70       72       8       5       9       9         BJ-1048       10       9       11       11       10       80       84       78       78       80       9       6       11       12         BJ-1049       13       11       14       13       13       78       83       74       78       78       9       5       12       9         BJ-1051       19       18       20       20       19       72       77       70       71       73       8       6       9       9         BJ-1045       21       19       22       21       21       70       75       68       70       71       9       6       10       9     <	DZ-1014	24	22	23	25	24	66	71	68	64	67	10	7	9	11	9
OZ-1008       21       19       21       21       21       68       73       68       68       69       11       8       11       11         3J-1044       20       19       21       21       20       71       76       69       70       72       8       5       9       9         3J-1048       10       9       11       11       10       80       84       78       78       80       9       6       11       12         3J-1049       13       11       14       13       13       78       83       74       78       78       9       5       12       9         3J-1051       19       18       20       20       19       72       77       70       71       73       8       6       9       9         3J-1045       21       19       22       21       21       70       75       68       70       71       9       6       10       9	DZ-1009	21	19	21	24	21	68	74	68	61	67	11	7	11	15	11
3J-1044 20 19 21 21 20 71 76 69 70 72 8 5 9 9 8 3J-1048 10 9 11 11 10 80 84 78 78 80 9 6 11 12 3J-1049 13 11 14 13 13 78 83 74 78 78 9 5 12 9 3J-1051 19 18 20 20 19 72 77 70 71 73 8 6 9 9 3J-1045 21 19 22 21 21 70 75 68 70 71 9 6 10 9	DZ-1007	18	16	18	18	18	70	75	70	69	71	12	8	12	12	11
3J-1048 10 9 11 11 10 80 84 78 78 80 9 6 11 12 3J-1049 13 11 14 13 13 78 83 74 78 78 9 5 12 9 3J-1051 19 18 20 20 19 72 77 70 71 73 8 6 9 9 3J-1045 21 19 22 21 21 70 75 68 70 71 9 6 10 9	DZ-1008	21	19	21	21	21	68	73	68	68	69	11	8	11	11	10
3J-1049     13     11     14     13     13     78     83     74     78     78     9     5     12     9       3J-1051     19     18     20     20     19     72     77     70     71     73     8     6     9     9       3J-1045     21     19     22     21     21     70     75     68     70     71     9     6     10     9	3J-1044	20	19	21	21	20	71	76	69	70	72	8	5	9	9	8
3J-1051 19 18 20 20 19 72 77 70 71 73 8 6 9 9 3J-1045 21 19 22 21 21 70 75 68 70 71 9 6 10 9	3J-1048	10	9	11	11	10	80	84	78	78	80	9	6	11	12	10
3J-1045 21 19 22 21 21 70 75 68 70 71 9 6 10 9	3J-1049	13	11	14	13	13	78	83	74	78	78	9	5	12	9	9
	3J-1051	19	18	20	20	19	72	77	70	71	73	8	6	9	9	8
3J-1053 21 19 22 21 21 70 76 67 70 71 9 6 11 9	3J-1045	21	19			21	70		68	70	71	9	6	10	9	9
3J-1043 17 16 19 18 18 74 78 70 71 73 9 6 11 10	3J-1053	21					70		67	70		9	6	11		9

n.a., no data exist for one or more of the isotopes necessary to complete this system. Av. is the average of the four models and is used to plot Fig. 7. Model  $A=^{207}Pb/^{204}Pb-^{206}Pb/^{204}Pb$ , Model  $B=^{208}Pb/^{204}Pb-^{206}Pb/^{204}Pb$ , Model  $C=^{143}Nd/^{144}Nd-^{206}Pb/^{204}Pb$ , Model  $D=^{87}Sr/^{86}Sr-^{206}Pb/^{204}Pb$ .