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### Atypically depleted upper mantle component revealed by Hf isotopes at Lucky Strike segment

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#### Abstract:

The Earth's upper mantle is commonly depicted as a "marble cake" assemblage of a refractory component mingled with varying amounts of geochemically diverse enriched components. This depleted component controls the major element composition of Mid-Oceanic Ridge Basalts (MORBs). It has long been considered a fairly uniform reservoir, depleted by early melting and well homogenized by subsequent convective stirring. We present new Hf, Nd, Pb and Sr isotopes and trace element data for basalts from the center of the Lucky Strike segment of the Mid-Atlantic Ridge (MAR). Despite the limited size of our sampling area (1.6 km along-axis by 5 km across), these data show a large range of isotopic compositions (similar to that of the whole MAR). It has recently been shown that part of the reason for not observing more segments with correlated Hf–Nd isotope systematics may be the lack of fine-scale sampling. Our data confirm this idea and show strong correlations between all isotope ratios. The systematics of Hf–Nd isotopes in Lucky Strike basalts define an atypical correlation, distinct from the global mantle array, with an anomalously high  $\epsilon$ Hf for a given  $\epsilon$ Nd. We illustrate that this atypical trend is the signature of an ancient refractory mantle rather than the product of kinetic processes during the melting event. The existence of this mantle component at Lucky Strike allows us to discuss the structure of the upper mantle around the Azores.

#### Highlights

► High spatial resolution sampling of MAR segment shows small scale mantle heterogeneity. ► Samples define an atypical Hf–Nd isotope correlation, distinct from the mantle array. ► Upper mantle near the Azores hotspot contains different refractory components.

Keywords: Oceanic basalts ; Mid-Atlantic Ridge ; Lucky Strike ; Mantle heterogeneity ; Hf isotopes

## 26 **1. Introduction**

27 It is now well established that the upper mantle is heterogeneous (Hofmann et al., 2003). This geochemical heterogeneity is evidenced by the composition of basalts which erupted along mid-28 29 oceanic ridges, which reveals the existence of broad mantle isotopic domains (e.g., the DUPAL 30 anomaly (Dupré and Allègre, 1983)). Each mantle domain is characterized by its intrinsic lithological 31 and geochemical heterogeneity, suggesting different convective histories (Hamelin et al., 2011; 32 Hanan et al., 2004; Meyzen et al., 2007; Vlastelic et al., 1999). At a local scale, basalt geochemical 33 variations are interpreted toto resultfrom partial melting of a marble-cake upper mantle (Allègre et 34 al., 1984; Morgan and Morgan, 1999; Stracke et al., 2005). This geochemical heterogeneity of ridge 35 basalts has been explained by the addition of various amounts of the FOZO component (Hanan and 36 Graham, 1996; Hart et al., 1992; Stracke et al., 2005) to a uniformly depleted source (Depleted 37 MORB Mantle, DMM). This second asthenospheric component is commonly inferred to have been 38 depleted by previous melting events but subsequently homogenized by convective stirring (Zindler 39 and Hart, 1986). However, recent studies of abyssal peridotites have revealed that distinct ancient 40 depletion signals may still be identified in refractory mantle domains (Harvey et al., 2006; Liu et al., 41 2008; Salters and Dick, 2002; Seyler et al., 2003; Stracke et al., 2011). This observation adds a new 42 level of complexity to the interpretation of MORB data and emphasizes the heterogeneous nature of 43 the depleted mantle.

Lu–Hf and Sm–Nd isotope systems have proven to be useful in tracking the signatures of ancient mantle depletion (e.g., Andres et al., 2004; Graham et al., 2006; Salters et al.; Stracke et al., 2011). Their parent/daughter ratios are higher in the residue than in the melt, thus developing a radiogenic signature in time. The similar behavior of these two radiogenic isotope systems during mantle melting results in strongly correlated <sup>177</sup>Hf/<sup>176</sup>Hf –<sup>143</sup>Nd/<sup>144</sup>Nd in oceanic basalts, the so-called

"mantle array" (Patchett and Tatsumoto, 1980). However, deviations from this basic model have long 49 50 been documented, in particular for mid-oceanic ridge samples (Chauvel and BlichertToft, 2001; 51 Debaille et al., 2006; Patchett and Tatsumoto, 1980; Salters and White, 1998). For example, a lack of 52 correlation between these two systems characterizes MORBs collected between 22°N and 35°N (Debaille et al., 2006). Anomalously high values of <sup>177</sup>Hf/<sup>176</sup>Hf for a given <sup>143</sup>Nd/<sup>144</sup>Nd have also 53 54 been measured along Mohns ridge (Blichert-Toft et al., 2005). These deviations from the mantle array model were initially interpreted as the result of isotopic disequilibrium between melts and their 55 56 mantle source across the garnet-spinel transition beneath the mid-oceanic ridge (Blichert-Toft et al., 57 2005; Debaille et al., 2006). More recently, a systematic study of Hf-Nd isotopes in oceanic basalts 58 has shown an array of parallel trends on a global scale (Salters et al., 2011). These authorsattributed 59 these sub-parallel arrays to the involvement of varying amounts of highly residual peridotites within 60 the upper mantle. According to this model, the non-uniformity of the depleted component together 61 with the lack of fine-scale sampling could explain the lack of correlation between EHf and ENd along 62 the global mid-oceanic ridge system.

63 In the central part of the Lucky Strike ridge segment along the Mid-Atlantic Ridge (MAR), the exceptional sampling resolution gives an opportunity to examine the different hypothesese for Lu-Hf 64 65 and Sm-Nd isotope systematics in the mid-oceanic ridge environment at a local scale. A recently 66 published petrogenetic model for this area has shown the existence of significant geochemical 67 variations in the Lucky Strike segment mantle source (Gale et al., 2011). Based on this study, we have selected a suite of representative samples spanning the whole range of geochemical variation 68 and located within an 8km<sup>2</sup> in the central portion of the segment. The Lucky Strike segment is also 69 proximal to, and influenced by the Azores hotspot (Dosso et al., 1999; Langmuir et al., 1997; 70 71 Schilling, 1975). Therefore, our new sample collection can also provide new constraints on the origin of the geochemical anomaly associated with the Azores hot spot-MAR interaction. 72

## 73 2. Geological setting and small scale sampling

74 The area studied is located along the MAR, South-West of the Azores triple junction. 75 Geochemical evidence together with geomorphological and geophysical characteristics demonstrate a 76 clear plume-ridge interaction in this region (Asimow et al., 2004; Gale et al., 2011; Gente et al., 2003; Schilling, 1975). In this area, the MAR is characterized by a succession of segments bound by 77 78 left-lateral, non-transform offsets. The Lucky Strike segment (fig. 1) is 60 km-long and bound by 79 non-transform discontinuities at 37°00'N and 37°35'N. It is a slow ridge segment characterized by a 80 full spreading rate of 21 mm/yr. This segment, located south west of the Azores hotspot, is 81 significantly shallower and volcanically more robust than other MAR segments (Cannat et al., 1999; 82 Escartín et al., 2001). The axial lithospheric structure and composition are typical of slow spreading 83 ridges, characterized by a well-developed, 11-12km-wide axial valley, and a deepening of 1600m of 84 the valley floor towards the segment ends (Detrick et al., 1995) where serpentinized peridotites 85 outcrop (Gràcia et al., 1997). The prominent seamount at the segment center (the Lucky Strike 86 volcano) is 7km wide and 15km along axis. It is underlain by an axial magma chamber 3.4km below 87 the seafloor (Singh et al., 2006) and hosts an active hydrothermal field at its summit (Langmuir et al., 88 1997). Several volcanic ridges spread north and south from this central volcano suggesting along axis 89 diking, while magnetic data suggest that magmatic accretion in the center of the Lucky Strike 90 segment is presently focused in a narrow graben at the top of the central volcano (Miranda et al., 91 2005). This area has been well studied since the hydrothermal field discovery in 1992 (Langmuir et 92 al., 1997), with a uniquely dense sampling among slow-spreading segments; more than 200 basaltic 93 samples have been collected with dredges, wax cores, submersibles, and other deep-sea vehicles. 94 Sample distribution along the segment is uneven, with the highest density on the central volcano. We 95 report here new Sr, Nd, Pb and Hf isotopic data for 18 samples collected there during the

Graviluck06 and Bathyluck09 cruises (Table 1), using the Nautile submersible (IFREMER) and
Remotely Operated Vehicles (Victor; Figure 1).

## 98 **3. Analytical methods**

99 Trace elements were measured using the Brucker 820 ICP-MS quadrupole instrument at the 100 Laboratoire de Planétologie et de Géodynamique in Nantes. Dissolutions were performed on 50 mg 101 of sample at a temperatures of 120°C for 12 hours in 1 mL 8N HNO<sub>3</sub> and 0.5 mL ~23N HF. Samples 102 were evaporated to dryness at 80°C, then re-dissolved in 2mL 8N HNO<sub>3</sub> and re-evaporated. The final 103 dissolution was made in 4 mL 4N HNO<sub>3</sub> before a 1:5000 dilution in 0.3N HNO3. Ge, Rh, In, Tm and 104 Bi were used for internal standard normalization and the following standards were used for 105 calibration curves: BHVO-2, W2, DNC-1, BIR-1 and the LDEO (Columbia University) in-house 106 standard Mid-Atlantic Ridge basalt MAR. The LDEO in-house standard K1919, collected from the 107 same lava flow as BHVO-1, was also analyzed several times within each analytical run to correct for 108 instrumental drift. The continental basalt reference material BR, from the CRPG Nancy was analyzed 109 several times to test our analytical reproducibility. All data are reported in table 1.

110 For isotope measurements, powdered samples (400-700 mg) were dissolved in Savillex vials with 111 concentrated HF-HBr ultrapure acids (3:1 in volume). Pb, Hf, Sr and Nd were separated from the 112 same dissolution using four different columns starting with Pb to minimize contamination. The Pb 113 extraction technique used was from (Manhes et al., 1978), using HBr 0.5M and HCl 6M on AG 1-X8 114 anion resin. Pb blanks measured using this procedure were < 100 pg, and thus are negligible relative 115 to the amount of sample analyzed. The effluent containing Hf, Sr and the rare earths were evaporated 116 and taken up in 6M HCl, evaporated again and taken up in 1 ml of 0.5M HCl/0.15M HF. It was then 117 ready to pass through an AG 50-X8 cation column (microcolumn Savillex, 30 ml, 6.4mm ID x 118 9.6mm OD x 25 cm). The first 6 mls containing the Hf-Ti fraction were eluted in 0.5M HCl/0.15M HF. The rest of the elution was completed using 3M HCl to separate Rb from Sr and 4M HNO<sub>3</sub> to separate Ba from rare earths. Nd was separated from Ce and Sm using 0.2M HCl and 0.35M HCl on a 0.8x4cm Biorad column loaded with LN Eichrom resin. The Hf-Ti fraction was evaporated and taken up in 6M HCl with a few  $\mu$ l of H2O2. It was loaded on a column made of a pipette tip filled with 100 mg of LN Eichrom resin. Ti was first eluted with 10ml 6M HCl with 50 $\mu$ l of H<sub>2</sub>O<sub>2</sub> and Hf was collected in 5 ml of 2M HF.

125 Sr isotope ratio measurements were carried out by thermal ionization mass spectrometry using a Finnigan Mat26x and a Thermo Finnigan Triton. Compositions were normalized for instrumental 126 mass fractionation relative to  ${}^{86}$ Sr/ ${}^{88}$ Sr = 0.1194.  ${}^{87}$ Sr/ ${}^{86}$ Sr of the NBS987 Sr standard during the 127 128 course of the analyses yielded 0.710235  $\pm$ 45 (n=4, 2 $\sigma$ ) for the MAT26x and 0.710255  $\pm$ 20 (n=2) for 129 the Triton. Isotopic compositions of Hf and Pb were measured at Ifremer, Brest, using a MC-ICPMS 130 Neptune. Repeat measurements of the Hf isotope standard JMC 475 (Denver) during the course of the analyses yielded  ${}^{176}$ Hf/ ${}^{177}$ Hf = 0.282145 ±13 (n=8, 2 $\sigma$ ). The results were normalized to the value 131 of 0.282158 using  ${}^{179}$ Hf/ ${}^{177}$ Hf = 0.7325 for mass fractionation correction. The Pb isotope data are 132 133 reported relative to published values for NBS 981 (Catanzaro et al., 1968). The samples were spiked 134 with thallium to correct for mass fractionation. Based on repeated runs of NBS 981, the estimated external precision for Pb analyses is  $\pm 0.02\%$ ,  $2\sigma$  for  $^{206}$ Pb/ $^{204}$ Pb and  $^{207}$ Pb/ $^{204}$ Pb and  $\pm 0.03\%$ ,  $2\sigma$  for 135 <sup>208</sup>Pb/<sup>204</sup>Pb. During the course of the analyses, 10 replicates of the Pb isotope standard NIST981 gave 136 an average of 16.930  $\pm 0.003$  (2 $\sigma$ ) and 15.482  $\pm 0.003$  (2 $\sigma$ ) and 36.668  $\pm 0.008$  (2 $\sigma$ ) for  ${}^{206}\text{Pb}/{}^{204}\text{Pb}$ . 137 <sup>207</sup>Pb/<sup>204</sup>Pb and <sup>208</sup>Pb/<sup>204</sup>Pb respectively. Isotopic compositions of Nd were measured at Institut 138 139 Universitaire Européen de la Mer, Brest, using a Thermo Finnigan, Triton. The measurements were carried out in static mode. All Nd data are fractionation corrected to <sup>146</sup>Nd/<sup>144</sup>Nd=0.7219. During the 140 course of the study, analyses of the La Jolla standard were performed and gave an average of 141  $^{143}$ Nd/ $^{144}$ Nd = 0.511843 ±4 (n=4 2 $\sigma$ ). 142

143 Using the same analytical method, BCR-2 gave values of  ${}^{176}$ Hf/ ${}^{177}$ Hf = 0.282861 ±9,  ${}^{206}$ Pb/ ${}^{204}$ Pb = 144 18.7496 ±7,  ${}^{207}$ Pb/ ${}^{204}$ Pb = 15.6175 ±7,  ${}^{208}$ Pb/ ${}^{204}$ Pb = 38.7291 ±21,  ${}^{87}$ Sr/ ${}^{86}$ Sr = 0.705008 ±8, 145  ${}^{143}$ Nd/ ${}^{144}$ Nd = 0.512629 ±8.

## 146 **4. Results, geochemical variations in Lucky Strike basalts**

147 The Lucky Strike segment contributes to the definition of the general geochemical gradient South 148 of the Azores (Dosso et al., 1999; Gale et al., 2011; Langmuir et al., 1997; Schilling, 1975). Samples 149 from this area have long been characterized as enriched in incompatible elements compared to 150 normal-MORB with La<sub>N</sub>/Sm<sub>N</sub> and K<sub>2</sub>O/TiO<sub>2</sub> ratios higher than 1.25 and 0.11 respectively. Our data 151 are in good agreement with this first order observation. Lucky Strike samples show varying levels of 152 incompatible element enrichment, with a bi-modal distribution of  $K_2O/TiO_2$  ratios (Fig. 2a). We have 153 defined two different groups named "transitional-MORB" (0.11 < K<sub>2</sub>O/TiO<sub>2</sub> < 0.25) and "enriched-154 MORB" ( $K_2O/TiO_2 > 0.25$ ), consistent with the definitions in Gale et al. (2011). These two groups 155 also show different petrographic characteristics: transitional-MORB samples are fresh aphyric 156 basalts, whereas enriched-MORB samples are slightly altered vesicular basalts with plagioclase (0-157 20%) and olivine (0-3%) phenocrysts. Groundmass textures are microcrystalline for E-MORB 158 samples and vary from microcrystalline to glassy for T-MORB samples.

Basalts from the transitional group are found throughout the segment whereas enriched basalts are found almost exclusively in the center. Our samples were recovered solely in the central part of the segment within an area of less than 8 km<sup>2</sup>. They display  $1.25 < La_N/Sm_N < 2$  typical of transitional MORB (9 samples) and  $La_N/Sm_N > 2$  typical of enriched MORB (9 samples) showing therefore the entire range in incompatible element enrichment previously reported for the whole Lucky Strike segment (Gale et al., 2011) (Fig. 2a). Figure 2b presents the REE patterns normalized to chondritic CI values for both groups of samples. Transitional samples are characterized by relatively flat REE 166 patterns with slight enrichments in light REE, whereas enriched samples have marked enrichments in 167 light REE (60-70x chondrites) and show slight fractionations between middle and heavy REE. It is 168 worth noting that samples from the Eastern slopes of the central volcano and from the Eastern axial 169 valley wall (named GRA-Nxx) show the same range of variation as that from samples collected 170 exclusively on the summit of the Lucky Strike volcano (named B09-ROCxx). This indicates that the 171 patterns described here do not reflect a temporal variability of basalt compositions at time scales of 172 about 500 ka (spreading age of the sample furthest from the axis) or less. Sr, Nd and Pb isotopes 173 show variations that are in good agreement with previous data from this ridge segment (Gale et al., 174 2011). The most enriched samples exhibit coherent, more radiogenic Pb, Sr and less radiogenic Nd 175 than the intermediate group (fig. 4). In binary isotope diagrams, Lucky Strike samples plot along 176 mixing lines between a depleted mantle component and an enriched component attributed to the 177 Azores hotspot. In these isotopic dimensions, even the less enriched samples from Lucky Strike are 178 significantly more enriched than the DMM, which agrees with their major and trace element 179 compositions. A simple mixing model between the DMM (Salters and Stracke, 2004) and the Azores mantle (Beier et al., 2007) reproduces the variations in  ${}^{207}$ Pb/ ${}^{204}$ Pb,  ${}^{206}$ Pb/ ${}^{204}$ Pb and  $\epsilon$ Nd,  ${}^{87}$ Sr/ ${}^{86}$ Sr. 180

181 Our new dataset shows a range of values from  $\varepsilon$ Hf = +14.3 to +22.1, comparable to values 182 previously reported in ridge segments immediately South and North of Lucky Strike (Agranier et al., 183 2005; Chauvel and BlichertToft, 2001). Despite the limited spatial extent of our sampling area, the 184 new data show a range of variation that is of the same order of magnitude as that reported for the 185 entire MAR (Agranier et al., 2005; Andres et al., 2004; Debaille et al., 2006). It is worth noting that 186 the transitional group extends towards a more radiogenic EHf end-member in the Hf-Nd isotopic 187 diagram and therefore plots above the mantle array. In contrast, in the same diagram, the enriched 188 group extends toward a lower EHf end-member and plots near the Azores isotopic field, below the 189 mantle array (fig. 5). The positive correlation between EHf and ENd agrees with the expected 190 coherence of Sm-Nd and Lu-Hf isotopic systems during magmatic processes, albeit with an unusual 191 slope compared to that of the mantle array (Fig. 5). Compared to other isotopic systems, the 192 systematics of Hf isotopes in this region of the MAR are still poorly known (fig. 3), and the Hf 193 isotopic ratios reported here are the first for the Lucky Strike segment. Basalts collected along the 194 MAR between 46° to 35°N reveal a comparable slope, however, (Agranier et al., 2005; Chauvel and 195 BlichertToft, 2001; Dosso et al., 1999; Salters et al., 2011; Yu et al., 1997) which is strikingly 196 different from the Nd-Hf correlation observed along other regions of the MAR (Fig. 5). The Hf 197 versus Sr isotope diagram shows a similar abnormal correlation compared to the rest of the MAR. 198 We additionally observe that the Lucky Strike transitional end-member seems relatively radiogenic in Pb isotopes (<sup>206</sup>Pb/<sup>204</sup>Pb > 19.35) compared to the classical depleted mantle (Salters and Stracke, 199 2004) ( $^{206}$ Pb/ $^{204}$ Pb  $\approx$  18.20) (Fig. 5a). 200

201 This unusual isotopic signature is however not unique in MORB, having been detected, for 202 example, in basalts from Mohns Ridge (fig. 6). The most depleted samples from this northern 203 Atlantic ridge show values of EHf for a given ENd to be very similar to those of Lucky Strike samples 204 (Blichert-Toft et al., 2005) (Fig. 6). However, important differences exist between the two zones: (i) 205 Mohns Ridge samples define a hyperbolic EHf- ENd array instead of the linear Lucky Strike trend 206 (Blichert-Toft et al., 2005), (ii) values of Pb isotopic ratios in depleted Mohns Ridge samples are unradiogenic ( $^{206}$ Pb/ $^{204}$ Pb  $\approx$  17.95) and (iii) a geographical trend is present along Mohns Ridge, 207 208 whereas the Lucky Strike variability is observed within a small ( $< 8 \text{ km}^2$ ) area on axis, and 209 reproduced in basalts from other segments in the 45-35°N MAR region near the Azores.

## 210 **5. Discussion**

### **5.1 Significant geochemical variations in basalts at local scale.**

### 212 <u>5.1.1. Small scale mantle heterogeneities or disequilibrium melting?</u>

213 At this fine sampling scale, it is important to investigate whether the geochemical variations 214 observed in basalts are the result of mantle heterogeneity or if they are produced by kinetic processes during melting. Early studies have commonly observed a wide range of EHf in MORB for a given 215 216  $\varepsilon$ Nd, leading to the hypothesis that melts may not be isotopically equilibrated with their peridotitic 217 sources (Blichert-Toft et al., 2005; Debaille et al., 2006). According to this hypothesis, a highly 218 variable Hf isotope dataset may result from the disequilibrium melting of a garnet-bearing peridotite 219 in which the individual mineral phases are not in isotopic equilibrium (Blichert-Toft et al., 2005; 220 Debaille et al., 2006). Because garnet incorporates Lu preferentially over Hf, it could develop a time-221 integrated signature with elevated EHf values. Preferential melting of garnet would therefore likely produce highly radiogenic Hf values in basalts. Hf<sup>4+</sup> is expected to possess a lower diffusion 222 coefficient than  $Pb^{2+}$  and  $Nd^{3+}$ , which could explain why this element would be more sensitive to this 223 224 kinetic process. If the isotopic composition of a particular mineral phase dominates the isotopic 225 composition of the generated melt, variable EHf compositions in MORB may be generated from a 226 single mantle domain during the current melting event below the present-day ridge. In this 227 hypothesis, the local depleted mantle identified on EHf diagrams (fig. 5) may not correspond to a real 228 end-member, but may instead result from a deviation toward radiogenic EHf values due to kinetic 229 processes during mantle melting. This disequilibrium melting hypothesis has been proposed to 230 explain the high EHf component found along Mohns and Knipovich ridges (Blichert-Toft et al., 2005) 231 (fig. 6), and to explain the lack of correlation between EHf and other isotopes systems along the MAR 232 between 22 and 35°N (Debaille et al., 2006).

233 Several arguments lead us to rule out this kinetic hypothesis as the cause of observed EHf 234 variations at Lucky Strike. A primary line of evidence stems from the good correlation between EHf 235 and other isotopic systems, which is not expected if kinetic processes influenced this isotopic ratio. 236 An assumption of the disequilibrium melting hypothesis is that partial melting takes place in presence 237 of garnet, which is irreconcilable with geochemical characteristics of the Lucky Strike transitional 238 basalts that do not show any fractionation between medium REE and heavy REE as would be 239 expected if melting occurs in the garnet stability field (Fig. 2b). This conclusion is in good agreement 240 with a recent petrogenetic model of melt supply to the Lucky Strike segment based on basalt 241 geochemistry, indicating that melting is controlled by spinel peridotite for the transitional group 242 (Gale et al., 2011). As recently demonstrated by Salters et al., (2011), a second line of argument 243 against disequilibrium melting is based on experimental studies of diffusion coefficients in garnet. In 244 this particular mineral, Dy diffusively equilibrates over 1 mm in 100,000 years at 1400°C (Van Orman et al., 2002). Even if we consider a diffusion coefficient 10 times slower for Hf<sup>4+</sup>.centimeter-245 246 scale isotopic equilibrium at mantle temperatures is expected to take place over one or two million 247 years. This equilibration time is orders of magnitude shorter than the time needed to build up a 248 radiogenic Hf signature. Disequilibrium melting is thus not a plausible hypothesis to explain the 249 unusually radiogenic Hf characteristic of the most depleted component along Lucky Strike.

If melting takes place at isotopic equilibrium, the Hf isotope composition of the melt is controlled by the whole rock composition of the peridotite. The presence or absence of garnet in the residual assemblage of the current melting event is therefore irrelevant and the source itself must have an anomalous isotopic signature.

254 <u>5.1.2. Small scale mantle heterogeneity and efficiency of magma mixing.</u>

255 Our study of a high spatial resolution dataset (<5 km) reveals the existence of incompletely mixed 256 coexisting primary magmas in the erupted material, despite the melt lens imaged at 3.4 km below the 257 seafloor (Singh et al., 2006). At first glance, the large variations measured in lithophile element 258 isotopic systems are in conflict with the lack of variation in He and Ne isotopes measured on samples 259 from the same cruises (Moreira et al., 2011). Because noble gas analyses require perfectly fresh glass, 260 He and Ne measurements have been conducted on a slightly different subset of Graviluck06 and 261 Bathyluck09 sampling, excluding samples from the more enriched group. This sampling bias can 262 partly explain the observed discrepancy. However, even within the transitional group of Lucky Strike 263 samples, Sr, Nd, Pb and Hf isotopes illustrate small variations that are not seen in He isotopes. We 264 propose that this difference is explained by the difference of diffusivity in melt between heavy 265 radiogenic isotopes and noble gases. The small-scale geochemical variations of trace elements and 266 Sr, Nd, Pb and Hf isotopes indicate that neither the magma chamber nor the magmatic plumbing 267 system totally homogenize the melts, therefore preserving, at least partially, the local mantle 268 heterogeneity signal. These compositionally distinct liquids can be produced simultaneously if the 269 scales of geochemical heterogeneities in a marble-cake assemblage are smaller than the melting zone. 270 Beneath the Lucky Strike central volcano, we know that the melting column extends 80 km below the 271 seafloor. In addition, geophysical evidencehas shown that the focusing of melts toward the centers of 272 segments is common along slow-spreading segments (Detrick et al., 1995), and thus sampling at the 273 segment center can reflect mantle composition at a larger scale. Depending on the efficiency of 274 along-axis melt migration, the volume of mantle tapped by the magmatic body could extend up to the 275 whole ridge segment (70 km at Lucky Strike). This length-scale represents, therefore, an upper limit 276 for the size of geochemical heterogeneities in the marble-cake assemblage. Within the compositional 277 heterogeneity of the Lucky Strike dataset, the atypical Hf isotopic signature is only shown by the 278 transitional group. This observation suggests that this mantle signature is carried by the local, more 279 refractory, mantle component. This high EHf signature is absent in Azores' islands basalts, most 280 likely due to the difference of lithospheric thickness between the two geological settings (> 50 km 281 beneath São Miguel and  $\approx 10$  km beneath the ridge); continuation of melting to shallower levels 282 beneath a thin lithosphere allows for a higher degree of melting and thus allows for a greater 283 proportion of melts to derive from the more refractory mantle component. Our study also implies that 284 interpretation of large- scale sampling of mantle domains using inadequate sample spacing should be 285 approached with caution. As shown recently by Salters et al. (2011), the long believed lack of 286 correlation between EHf and other isotope systems along mid-oceanic ridges is likely the result of a 287 sampling bias and scarcity of data at a fine scale.

## 288 **5.2 Origin of the mantle component with anomalous high εHf.**

According to a recent petrogenetic model based on major and trace element compositions and Sr, Nd and Pb isotopes, the source of Lucky Strike transitional basalts (called "local DM" in figures 5, 6 and 7) results from a mixing between a DMM-like end member and a metasomatized mantle with an isotopic composition similar to Azores mantle (Gale et al., 2011). However, since the Azores lavas have a  $\varepsilon$ Hf- $\varepsilon$ Nd composition below the mantle array, simple mixing cannot account for the high  $\varepsilon$ Hf values (+24) and relatively low  $\varepsilon$ Nd values (+10) measured in the transitional group (fig. 5).

### 295 <u>5.2.1 Ancient, residual mantle component.</u>

It has long been suggested that Hf-Nd isotope decoupling in the mantle source can be produced by an ancient melting event involving garnet fractionation, and subsequent melting of the resulting mantle residue (Blichert-Toft and Albarede, 1997; Chauvel and BlichertToft, 2001; Salters and Hart, 1989; Salters and Zindler, 1995). Because the ratio of Lu and Hf partition coefficients between garnet and a basaltic liquid ( $K_D_{Grt-liq}^{Lu}/K_D_{Grt-liq}^{Hf}$ ) is very high (> 30), mantle residues produced by melt extraction at great depth would likely be HREE-enriched relative to Hf. An ancient partial melting event in the garnet stability field should yield a high <sup>177</sup>Hf/<sup>176</sup>Hf signal in the residue, and the 303 amplitude of this signal should increase with the time elapsed since the melting event. In contrast, Sm 304 and Nd have similar partition coefficients in garnet and spinel peridotite and ENd is therefore 305 expected to remain almost unaffected by the two different types of melting events. Given a sufficient 306 length of time (1-2 Ga), the residual garnet peridotite should therefore develop an isotopic signature 307 characterized by a steeper slope in *E*Hf versus *E*Nd diagram. This interpretation is consistent with the 308 mixtures of variously-depleted mantle model proposed by Salters et al. (2011) to explain the array of 309 sub-parallel trends of Hf-Nd isotope ratios in oceanic basalts on a global scale. According to this 310 model, ancient residual oceanic lithosphere (ReLish, Salters et al., 2011), produced by the extraction 311 of the crust, is recycled and mingled into the MORB and/or OIB mantle sources. Is the ReLish a 312 possible component of the Lucky Strike mantle assemblage?

313 As mentioned earlier, the recently published petrogenetic model for Lucky Strike basalts (Gale et 314 al., 2011) does not fit the unusual EHf-ENd correlation obtained for our samples. This model can 315 briefly be summarized by a 2-step geochemical history: i) mixing between the DMM (90%) and the 316 Azores mantle (10%), and ii) metasomatism by low-degree melts produced in the garnet stability 317 field in the Azores mantle. The first step defines the regional gradient along the MAR South of the 318 Azores, whereas the second step generates the local variability shown by transitional and enriched 319 groups. We attempted to adjust this petrogenetic model and we replaced the DMM component with a 320 mixture between DMM and ReLish. This latter component is highly depleted and almost completely 321 exhausted of incompatible elements, and its signatures can easily be overpowered by a small addition 322 of enriched components. This could explain why the influence of the ReLish component is not 323 observed on highly incompatible isotopes system (Rb-Sr, U,Th-Pb). We have modeled the trace 324 elements (REE+Hf) together with the Sr-Nd-Pb-Hf isotope compositions of Lucky Strike basalts. The 325 details of the melting models are in the supplementary material and results can be seen on figure 5, 6 and 7. In this model, the ReLish component was produced 1.2 Ga ago, by a 6% melting event in the 326

327 garnet stability field, from a MORB source that has been isolated from the bulk silicate Earth 328 reservoir for 2 Ga. We are able to reproduce the trace element and isotopic systematics of Lucky 329 Strike basalts by a two-steps model: i) mixing of 20% of the ReLish component to a mantle 330 assemblage of DMM (70%) and Azores mantle (10%), and ii) a metasomatism of this local Lucky 331 Strike mantle by low-F melts produced in presence of garnet from the Azores mantle. This solution is 332 not unique and it is possible to fit the data with a different set of parameters. For example, the mass 333 fraction of the ReLish component is inversely proportional to its formation age: a more recent ReLish 334 would simply imply a greater proportion in the mantle assemblage. Our goal is therefore not to 335 quantify these melting events accurately, but rather to show that the presence of a ReLish component 336 in the source of Lucky Strike basalts is able to produce the atypical high EHf end-member (Fig. 6). 337 The presence of such refractory component could also explain the relatively low degree of melting 338 (7%) calculated by Gale et al. (2011) for the present day melting beneath the ridge. Although residual 339 oceanic lithosphere material is probably a ubiquitous mantle component, due to its refractory 340 properties, it is expected to contribute little to the observed global MORB variations.

#### 341

#### 5.2.2 Sub-continental lithospheric mantle.

342 A second potential candidate for the atypical source signature in Lucky Strike transitional basalts 343 is the Sub-Continental Lithospheric Mantle (SCLM). Compositions of the SCLM are primarily based 344 on xenolith and xenocryst samples. These samples show a large range of Hf isotopic compositions up 345 to very radiogenic values (Simon et al., 2007). This variability in Hf isotopes is associated with less 346 variable and less radiogenic Nd isotopic compositions.composition The SCLM consists mainly of 347 ultramafic rocks, ranging from lherzolites to dunites. Despite the observed petrological diversity of 348 this reservoir, a recent tomographic study has suggested that most Archean sub-continental 349 lithospheric mantle originally consisted of depleted dunites and harzburgites (Griffin et al., 2009). In 350 intraplate volcanic areas, these refractory peridotites are subsequently metasomatized and refertilized. In xenolith datasets, the decoupling between Hf and Nd isotopes in this subcontinental reservoir has been interpreted as a result of depletion events followed by carbonatite-type metasomatism (Simon et al., 2007). This process is expected to link very radiogenic ɛHf values with relatively less radiogenic Nd and enrichment in highly incompatible trace elements, consistent with the geochemical characteristics of the transitional Lucky Strike basalts (Fig. 5).

356 Several studies have argued for contamination of the MORB source by either the crustal or the 357 mantle part of the continental lithosphere (e.g. Hanan et al., 2004; Janney et al., 2005). Because the 358 Rb/Sr ratio is very high in the continental crust, addition of this material to the depleted mantle would dramatically increase <sup>87</sup>Sr/<sup>86</sup>Sr values. In contrast, the mantle portion of the continental lithosphere is 359 360 not expected to have an exceptionally high Rb/Sr ratio and is therefore a more likely component. The 361 presence of relics of SCLM preserved in the upper mantle in the North-central Atlantic area has been 362 a matter of debate for the last 15 years. For example, SCLM is suspected to explain the distinctive 363 metasomatic parageneses and garnet signature in spinel-peridotite xenoliths from Cape Verde 364 Archipelago (Bonadiman et al., 2005). Similarly, the enriched component in the lavas from São 365 Miguel has been inferred to contain a contribution from localized pieces of SCLM material (Madureira et al., 2011; Moreira et al., 1999; Widom et al., 1997). According to these studies, this 366 367 sub-continental lithospheric mantle originally resided beneath North-western Africa or Iberia and was 368 delaminated during rifting upon the opening of the Atlantic Ocean basin. However, based on the low 369 EHf measured in lavas from Azorean islands, other studies have argued against this hypothesis (Beier 370 et al., 2007; Elliott et al., 2007). In contrast to these studies, we propose that SCLM material is a 371 potential candidate for the origin of the transitional group basalts, and should be seen as a possible 372 refractory component of the mantle rather than the source of the enriched basalts. As previously 373 discussed in §5.1.2, this is in good agreement with the fact that this signature is exclusively observed at the ridge axis, where the shallower melting compared to Azores islands allows more refractorymaterial to melt.

It is important to note that the geochemical composition of the SCLM remains poorly constrained. It is therefore difficult to produce an incontrovertible mixing model involving this reservoir. SCLM samples show highly heterogeneous isotopic signatures, which seems to be inconsistent with the tightly defined nature of the more depleted Lucky Strike end-member. However, based on our present day knowledge of the Hf-Nd isotope characteristics of SCLM xenoliths, the addition of this type of material to the Lucky Strike upper mantle could account for the abnormal isotopic signature of the local depleted end-member.

383 Distinguishing whether the abnormal Hf-Nd isotopic signature is the result of the presence of 384 residual oceanic lithosphere previously melted in the garnet stability field (*i.e.* ReLish) (§5.2.1) or the 385 contribution of sub-continental lithospheric mantle (§5.2.2), is not straightforward. Very ancient 386 depleted signatures in Os isotopes have long been suggested as a characteristic of subcontinental lithospheric mantle. But recent studies of abyssal peridotites have shown a wide range of <sup>187</sup>Os/<sup>188</sup>Os 387 388 ratios at slow mid-oceanic ridge settings extending to very unradiogenic values (Liu et al., 2008). If 389 highly depleted Os isotope signatures are not unique to the SCLM, there are no isotopic or chemical 390 signatures that unambiguously distinguish the two proposed hypotheses.

### **5.3. Geodynamical implications.**

We have shown that the unusual Hf isotope signature in basalts from Lucky Strike is best explained by the melting of a refractory component in the mantle near the Azores rather than by a kinetic effect during the current melting event. Even if we cannot specify whether this refractory component is a ReLish or a SCLM, we have shown that this component is mixed with the depleted mantle and the Azores mantle prior to the addition of the metasomatic agent (low-F melts derived from the Azores; Gale et al., 2011). The existence of this component and its mixing relationship raisethe question of the structure of the upper mantle around the Azores.

399 Since only limited data are available for Hf isotopes compared to other isotope systems, it is not 400 possible to map precisely the extent of the particular Hf signature recognized in our local scale study. 401 However, the Nd vs. Hf isotopes of segments North and South of the Lucky Strike segment show a 402 similar correlation (fig. 5). The latitudinal extension of this atypical correlation is therefore at least 403 between 46°N to 35°N, and centered at the latitude of the Azores hotspot. This along-axis extension 404 is significantly shorter than the regional gradients defined by Sr isotopes or La/Sm ratios (Dosso et 405 al., 1999; Gale et al., 2011; Langmuir et al., 1997; Schilling, 1975). By analogy to observations along 406 the MAR near Ascension Island (Paulick et al., 2010), the hypothesis involving mixing between one 407 enriched component and two different depleted components (DMM and ReLish/SCLM) is expected 408 to produce a cloud or two different trends in Nd-Hf isotope diagrams. Since only one trend is seen 409 from 46°N to 35°N (Agranier et al., 2005; Chauvel and BlichertToft, 2001; Dosso et al., 1999; Yu et 410 al., 1997) (fig. 4), we can infer that the depleted components in the system are well homogenized.

A first possible model for the upper mantle structure in the Lucky Strike and Azores domain is that the refractory mantle component is a primary constituent of the Azores hotspot, mixing with the regional mantle (fig. 8A). In this hypothesis, this refractory component has been recycled in the asthenosphere and brought back to the surface by the hotspot. The initial depletion event could either be the result of deep ancient melting in an oceanic domain (ReLish) or of some mantle material from the base of the continental lithosphere (SCLM) which was recycled deep in the mantle.

417 An alternative model is an upper mantle constituted of different coexisting refractory materials 418 prior to the arrival of the Azores plume (fig. 8B). In this hypothesis, the atypical depleted component 419 could be a ubiquitous residual mantle component (*i.e.* ReLish) or relics of the SCLM left behind by 420 the North-East drifting of the African continent (fig. 8B). The presence of the particular isotopic 421 signature in the transitional Lucky Strike samples could be the result of a change of the melting 422 regime near the Azores (Asimow et al., 2004; Schilling, 1975). Between 46°N and 35°N, the thin 423 lithosphere and the local increase of mantle potential temperature (Asimow et al., 2004; Schilling, 424 1975) (+ 35°C) associated with the thermal influence of the Azores enhances mantle melting. These 425 pressure and temperature conditions would allow melting of the more refractory mantle identified 426 around the Azores region. Future Hf-Nd studies along segments near Lucky Strike will provide 427 constraints on the geographical extension of this refractory component and allow study of the along-428 axis evolution of its mixing proportions with the ambient mantle.

## 429 **6.** Conclusion

430 We report a strong correlation between Hf and Nd isotopic systems from MORBs within the 431 Lucky Strike segment along the MAR, which is contrary to the commonly accepted decoupling of 432 these systems. In good agreement with recent results by Salters et al. (2011), our results suggest that 433 the lack of correlation between *E*Hf and other isotope systems along the global mid-oceanic ridge 434 system is the result of sampling bias. Part of the reason for not observing more segments with 435 correlated Hf-Nd isotope systematics may be the lack of fine-scale, dense sampling. In order to 436 discuss the origin and significance of large-scale mantle composition, alternative solutions are (i) 437 high spatial resolution sampling of the ridge or (ii) statistical techniques such as spectral analyses 438 (Agranier et al., 2005), where the dataset interpretation is weighted by the sampling density.

The Hf-Nd isotope correlation seen in Lucky Strike samples is significantly different from the global mantle array defined by MORB and OIB, and is characterized by an unusually high εHf values for a given εNd. In every binary isotopic plot, Lucky Strike basalts define a trend from this atypical radiogenic Hf mantle component toward the Azores enriched mantle. We have shown that kinetic 443 effects cannot account for the observed Hf-Nd systematics in Lucky Strike basalts and that this trend 444 is best explained by mixing enriched and depleted components in the local mantle. If the origin of the 445 enriched component is clearly related to the Azores hotspot, the origin of the depleted mantle source 446 is more controversial. We propose that this unusual end-member could either indicate the presence of 447 some residual oceanic lithosphere or some sub-continental lithospheric mantle left behind by the 448 drifting of the African continent. At Lucky Strike, enhanced melting conditions associated with the 449 Azores plume allow us to detect this refractory component within basalts erupted on-axis. This 450 observation points towards the existence in the mantle of domains of refractory material that do not 451 melt in normal mid-oceanic ridge conditions, thus contributing little to the global MORB 452 compositions. This study emphasizes the heterogeneous nature of the depleted mantle long 453 considered the archetype of a homogeneous mantle domain.

454

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## 463 **Figure captions**

464

465 Figure 1: Bathymetric maps of (A) the Mid-Atlantic Ridge, (B) Lucky Strike segment and (C) Arctic 466 ridges. Locations of our new samples (square symbols) are shown together with published samples: 467 of the entire MAR (gray circles), the Azores geochemical anomaly (yellow circles) and Mohns ridge 468 (green triangles). Symbols are the same in all the figures, published data are from (Agranier et al., 469 2005), (Andres et al., 2004; Andres et al., 2002), (Blichert-Toft et al., 2005), (Castillo and Batiza, 470 1989), (Chauvel and BlichertToft, 2001), (Debaille et al., 2006), (Dosso et al., 1993; Dosso et al., 471 1999; Dosso et al., 1991), (Douglass et al., 1999), (Frey et al., 1993), (Fontignie and Schilling, 1996), 472 (Hamelin et al., 1984), (Hanan et al., 1986), (Ito et al., 1987), (Machado et al., 1982), (Nowell et al., 1998), (Paulick et al., 2010), (Salters, 1996), (Schiano et al., 1997), (Schilling et al., 1994), (Shirey et 473 474 al., 1987), (Yu et al., 1997).

Figure 2: A) histogram showing the bimodal distribution of Lucky Strike samples. Two distinct peaks define the two petrological groups described in this study: "Transitional" if  $K_2O/TiO_2 < 0.25$ and "enriched" if  $K_2O/TiO_2 > 0.25$ . Data are from this study together with unpublished data (Bezos personal communication), Gale et al. (2011) and Ferreira (2006). B) Diagram showing normalized REE concentration of the two groups.

480 Figure 3: Geochemical gradient south of the Azores. The center of Lucky Strike segment shows a
481 large range of geochemical variation.

**Figure 4:** Binary isotopic plots showing Lucky Strike data compared to previously published MAR and Azores data. Mixing models between a depleted (DMM, (Salters and Stracke, 2004)) and an enriched (Azores, Gale et al., 2011) ) components are shown. The chosen Azores mantle composition assumed is that of the least enriched end-member (Sete Cidades samples, Beier et al. (2007)). Figure 5: Hf isotopes versus other isotopic systems. The model modified from (Gale et al., 2011) is shown as black lines. Small dark squares along these lines represent percentage of low-F melts additions (1% step) to the local Depleted Mantle (Local DM, see text for further explanations). Compare to the mantle array, new data from Lucky Strike segment define a trend with very radiogenic Hf composition for a given εNd. The same atypical correlation can be observed along the MAR in samples collected between 46 and 35°N (yellow circles).

Figure 6: Comparison between Lucky Strike data and Mohns ridge sample, previously described as
geochemically abnormal with εHf for a given εNd. The same model shown in figure 5 is reported
here, fractions of low-F melts added to the regional mantle are shown in percent.

495 Figure 7: Hf isotopes versus La/Hf, and Nd isotopes versus La/Nd. Metasomatism of the regional 496 mantle by low-F melts allow to reproduce the geochemical heterogeneity measured at Lucky Strike.

497 Figure 8: Two different models for the structure of the upper mantle near the Azores. In the left 498 column, the atypical depleted component is a primary constituent of the Azores hotspot. In the right 499 column, this component resides in the upper mantle prior to the plume arrival.

500 **Supplementary figure:** Lu/Hf vs Sm/Nd diagram. Green and orange lines represent melting models 501 (batch and fractional) with residual spinel and residual garnet, respectively. The black line represents 502 low-F melts metasomatism model (see text for further explanations). The Lucky Strike abnormally 503 high Lu/Hf ratio for a given Sm/Nd ratio is well explained by the influence of this recent 504 metasomatism event.

505

Table 1: Major element, trace element and isotopes data for samples collected at Lucky Strike
segment center. AII127-D15G has previously been published in Dosso et al. (1999), except new Pb

- 508 and Hf isotope measurements done on a MC-ICPMS at SDSU in collaboration with Barry Hanan.<sup>b</sup>
- Sr isotope measured using a Thermo Finnigan Triton. <sup>c</sup> Sr isotope measured using a Finnigan
  MAT26x.
- 511

## 513 Table 1

Long. (°)         -32.283         -32.282         -32.282         -32.287         -32.276           Lat. (°)         37.291         37.291         37.291         37.297         37.299           Type         T-MORB         E-MORB         T-MORB         E-MORB         E-MORB           SiO <sub>2</sub> 5.1.4         49.3         5.1.4         51.5         48.8         50.1           TiO <sub>2</sub> 1.07         1.08         1.05         1.05         1.28         1.20           Al <sub>2</sub> O <sub>3</sub> 14.7         17.0         14.8         15.1         16.4         16.7           FeO         9.8         6.4         9.9         9.5         7.5         6.8           MnO         0.18         0.12         0.18         0.19         0.15         0.13           MgO         8.1         9.4         8.0         8.2         8.2         8.5           CaO         12.1         13.9         12.2         12.4         14.1         14.3           Na <sub>2</sub> O         2.0         0.46         0.16         0.18         0.60         0.70           P <sub>2</sub> O <sub>5</sub> 0.14         0.19         0.12         0.13         0.27         0.23	Sample	AII127- D15G	B09-ROC01	B09-ROC07	B09-ROC08	B09-ROC11	B09-ROC13
Lat. (°)         37.291         37.284         37.291         37.291         37.297         37.299           Type         T-MORB         E-MORB         T-MORB         T-MORB         E-MORB         E-MORB         E-MORB           SiO2         51.4         49.3         51.4         51.5         48.8         50.1           TiO2         1.07         1.08         1.05         1.05         1.28         1.20           Al <sub>2</sub> O <sub>3</sub> 14.7         17.0         14.8         15.1         16.4         16.7           FeO         9.8         6.4         9.9         9.5         7.5         6.8           MnO         0.18         0.12         0.18         0.19         0.15         0.13           MgO         8.1         9.4         8.0         8.2         8.2         8.5           CaO         12.1         13.9         12.2         12.4         14.1         14.3           NayO         2.2         2.3         2.3         2.2         2.4         2.3           K2O         0.00         0.46         0.16         0.18         0.60         0.70           Sum         99.9         101.0         100.7	Long. (°)	-32.283	-32.282	-32.282	-32.282	-32.277	-32.276
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	Lat. (°)	37.291	37.284	37.291	37.291	37.297	37.299
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	Type	T-MORB	E-MORB	T-MORB	T-MORB	E-MORB	E-MORB
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	SiO <sub>2</sub>	51.4	49.3	51.4	51.5	48.8	50.1
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$TiO_2$	1.07	1.08	1.05	1.05	1.28	1.20
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$Al_2O_3$	14.7	17.0	14.8	15.1	16.4	16.7
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	FeO	9.8	6.4	9.9	9.5	7.5	6.8
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	MnO	0.18	0.12	0.18	0.19	0.15	0.13
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	MgO	8.1	9.4	8.0	8.2	8.2	8.5
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	CaO	12.1	13.9	12.2	12.4	14.1	14.3
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	Na <sub>2</sub> O	2.2	2.3	2.3	2.2	2.4	2.3
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$K_2O$	0.20	0.46	0.16	0.18	0.60	0.70
L.O.I. Sum $0.25$ 99.9 $-0.63$ 101.0 $-0.51$ 	$P_2O_5$	0.14	0.19	0.12	0.13	0.27	0.23
Sum         99.9         101.0         100.7         101.0         101.5         102.0           La         12.45         12.45         5.74         5.74         16.59         15.58           Ce         24.49         24.49         12.32         12.32         31.36         29.74           Pr         2.94         2.94         1.71         1.71         3.65         3.49           Nd         12.35         12.35         8.04         8.04         14.99         14.15           Sm         2.75         2.75         2.39         2.39         3.19         3.11           Eu         0.98         0.98         0.89         0.89         1.11         1.07           Gd         3.19         3.19         3.27         3.27         3.70         3.55           Dy         3.07         3.07         3.88         3.88         3.56         3.41           Ho         0.63         0.63         0.85         0.85         0.75         0.72           Er         1.74         1.74         2.39         2.39         2.08         2.00           Yb         1.62         1.62         2.34         2.34         1.97	L.O.I.		0.25	-0.63	-0.51	0.97	0.07
La12.4512.455.745.7416.5915.58Ce24.4924.4912.3212.3231.3629.74Pr2.942.941.711.713.653.49Nd12.3512.358.048.0414.9914.15Sm2.752.752.392.393.193.11Eu0.980.980.890.891.111.07Gd3.193.193.273.273.703.55Tb0.500.500.580.580.580.55Dy3.073.073.883.883.563.41Ho0.630.630.850.850.750.72Er1.741.742.392.392.082.00Yb1.621.622.342.341.971.91Lu0.250.250.360.360.300.29Hf1.91.871.621.622.051.84 $^{87}Sr/^{86}Sr$ 0.7029450.703118 <sup>b</sup> 0.702957 <sup>b</sup> 0.702933 <sup>c</sup> 0.703109 <sup>b</sup> 0.703049 <sup>c</sup> $^{80}Pb/^{204}Pb$ 18.82219.33718.88618.9419.15919.219 $^{206}Pb/^{204}Pb$ 15.53815.59215.54415.54615.57915.582 $^{208}Pb/^{204}Pb$ 38.42138.89838.51138.54738.77338.801 $\epsilon Hf$ 21.38-21.6520.6914.3614.76 <td>Sum</td> <td>99.9</td> <td>101.0</td> <td>100.7</td> <td>101.0</td> <td>101.5</td> <td>102.0</td>	Sum	99.9	101.0	100.7	101.0	101.5	102.0
La12.4512.455.745.7416.5915.58Ce24.4924.4912.3212.3231.3629.74Pr2.942.941.711.713.653.49Nd12.3512.358.048.0414.9914.15Sm2.752.752.392.393.193.11Eu0.980.980.890.891.111.07Gd3.193.193.273.273.703.55Tb0.500.500.580.580.580.55Dy3.073.073.883.883.563.41Ho0.630.630.850.850.750.72Er1.741.742.392.392.082.00Yb1.621.622.342.341.971.91Lu0.250.250.360.360.300.29Hf1.91.871.621.622.051.8487 Sr/ <sup>86</sup> Sr0.7029450.703118 <sup>b</sup> 0.702957 <sup>b</sup> 0.702933 <sup>c</sup> 0.703109 <sup>b</sup> 0.703049 <sup>c</sup> $\epsilon Nd$ 8.977.049.159.127.187.28 $t^{208} Pb/^{204} Pb$ 15.53815.59215.54415.54615.57915.582 $t^{208} Pb/^{204} Pb$ 38.42138.89838.51138.54738.77338.801 $\epsilon Hf$ 21.38-21.6520.6914.3614.76							
Ce24.4924.4912.3212.3231.3629.74Pr2.942.941.711.713.653.49Nd12.3512.358.048.0414.9914.15Sm2.752.752.392.393.193.11Eu0.980.980.890.891.111.07Gd3.193.193.273.273.703.55Tb0.500.500.580.580.580.55Dy3.073.073.883.883.563.41Ho0.630.630.850.850.750.72Er1.741.742.392.392.082.00Yb1.621.622.342.341.971.91Lu0.250.250.360.360.300.29Hf1.91.871.621.622.051.84 $^{87}Sr/^{86}Sr$ 0.7029450.703118 <sup>b</sup> 0.702957 <sup>b</sup> 0.702933 <sup>c</sup> 0.703109 <sup>b</sup> 0.703049 <sup>c</sup> $^{206}Pb/^{204}Pb$ 18.82219.33718.88618.9419.15919.219 $^{207}Pb/^{204}Pb$ 15.53815.59215.54415.54615.57915.582 $^{208}Pb/^{204}Pb$ 38.42138.89838.51138.54738.77338.801 $\epsilon Hf$ 21.38-21.6520.6914.3614.76	La	12.45	12.45	5.74	5.74	16.59	15.58
Pr2.942.941.711.713.653.49Nd12.3512.358.048.0414.9914.15Sm2.752.752.392.393.193.11Eu0.980.980.890.891.111.07Gd3.193.193.273.273.703.55Tb0.500.500.580.580.580.55Dy3.073.073.883.883.563.41Ho0.630.630.850.850.750.72Er1.741.742.392.392.082.00Yb1.621.622.342.341.971.91Lu0.250.250.360.360.300.29Hf1.91.871.621.622.051.84**	Ce	24.49	24.49	12.32	12.32	31.36	29.74
Nd12.3512.358.048.0414.9914.15Sm2.752.752.392.393.193.11Eu0.980.980.890.891.111.07Gd3.193.193.273.273.703.55Tb0.500.500.580.580.580.55Dy3.073.073.883.883.563.41Ho0.630.630.850.850.750.72Er1.741.742.392.392.082.00Yb1.621.622.342.341.971.91Lu0.250.250.360.360.300.29Hf1.91.871.621.622.051.84*** <td>Pr</td> <td>2.94</td> <td>2.94</td> <td>1.71</td> <td>1.71</td> <td>3.65</td> <td>3.49</td>	Pr	2.94	2.94	1.71	1.71	3.65	3.49
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Nd	12.35	12.35	8.04	8.04	14.99	14.15
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	Sm	2.75	2.75	2.39	2.39	3.19	3.11
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Eu	0.98	0.98	0.89	0.89	1.11	1.07
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Gd	3.19	3.19	3.27	3.27	3.70	3.55
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Tb	0.50	0.50	0.58	0.58	0.58	0.55
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Dy	3.07	3.07	3.88	3.88	3.56	3.41
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Но	0.63	0.63	0.85	0.85	0.75	0.72
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Er	1.74	1.74	2.39	2.39	2.08	2.00
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	Yb	1.62	1.62	2.34	2.34	1.97	1.91
Hf1.91.871.621.622.051.84 $^{87}$ Sr/ $^{86}$ Sr0.7029450.703118 <sup>b</sup> 0.702957 <sup>b</sup> 0.702933 <sup>c</sup> 0.703109 <sup>b</sup> 0.703049 <sup>c</sup> $\epsilon$ Nd8.977.049.159.127.187.28 $^{206}$ Pb/ $^{204}$ Pb18.82219.33718.88618.9419.15919.219 $^{207}$ Pb/ $^{204}$ Pb15.53815.59215.54415.54615.57915.582 $^{208}$ Pb/ $^{204}$ Pb38.42138.89838.51138.54738.77338.801 $\epsilon$ Hf21.38-21.6520.6914.3614.76	Lu	0.25	0.25	0.36	0.36	0.30	0.29
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Hf	1.9	1.87	1.62	1.62	2.05	1.84
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	<sup>87</sup> Sr/ <sup>86</sup> Sr	0.702945	0.703118 <sup>b</sup>	0.702957 <sup>b</sup>	0.702933 <sup>c</sup>	0.703109 <sup>b</sup>	0.703049 <sup>c</sup>
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	εNd	8.97	7.04	9.15	9.12	7.18	7.28
$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	<sup>206</sup> Pb/ <sup>204</sup> Pb	18.822	19.337	18.886	18.94	19.159	19.219
208 Pb/204 εHf38.42138.89838.51138.54738.77338.801εHf21.38-21.6520.6914.3614.76	<sup>207</sup> Pb/ <sup>204</sup> Pb	15.538	15.592	15.544	15.546	15.579	15.582
εHf 21.38 - 21.65 20.69 14.36 14.76	<sup>208</sup> Pb/ <sup>204</sup> Pb	38.421	38.898	38.511	38.547	38.773	38.801
	εHf	21.38	-	21.65	20.69	14.36	14.76

Sample B09-ROC14 B09-ROC15 B09-ROC20 B09-ROC21 B09-ROC22	2 B09-ROC23
Long. (°) -32.276 -32.275 -32.280 -32.280 -32.280	-32.278
Lat. (°) 37.300 37.301 37.291 37.291 37.291	37.290
Type E-MORB E-MORB T-MORB T-MORB T-MORB	T-MORB
SiO <sub>2</sub> - 48.7 51.6 51.8 51.7	51.6
TiO <sub>2</sub> - 1.01 1.06 1.06 1.05	1.05
Al <sub>2</sub> O <sub>3</sub> - 17.9 14.3 15.0 15.0	15.1
FeO - 6.7 10.3 10.0 9.9	9.5
MnO - 0.14 0.19 0.19 0.18	0.18
MgO - 8.7 8.2 8.1 8.1	8.2
CaO - 14.5 12.3 12.3 12.2	12.5
Na <sub>2</sub> O - 2.2 2.3 2.3 2.3	2.2
K <sub>2</sub> O - 0.41 0.16 0.16 0.16	0.18
$P_2O_5$ - 0.18 0.13 0.13 0.12	0.13
L.O.I 0.50 -0.78 -0.77 -0.77	-0.68
Sum - 101.5 100.8 101.3 101.1	101.0
La 14.03 14.03 4.98 4.98 5.01	5.54
Ce 26.41 26.41 11.25 11.26 11.26	12.15
Pr 3.17 3.17 1.59 1.58 1.59	1.68
Nd 12.85 12.85 7.55 7.57 7.65	7.92
Sm 2.85 2.85 2.34 2.34 2.34	2.37
Eu 1.00 1.00 0.88 0.88 0.88	0.89
Gd 3.29 3.29 3.25 3.26 3.27	3.25
Tb 0.53 0.53 0.59 0.59 0.59	0.58
Dy 3.26 3.26 3.96 3.95 3.93	3.84
Ho 0.68 0.68 0.87 0.87 0.87	0.84
Er 1.90 1.90 2.46 2.47 2.45	2.37
Yb 1.83 1.83 2.45 2.46 2.45	2.34
Lu 0.28 0.28 0.38 0.38 0.38	0.36
Hf 1.69 1.69 1.59 1.59 1.58	1.61
<sup>87</sup> Sr/ <sup>86</sup> Sr - $0.70304^{\circ}$ $0.702915^{\circ}$ $0.702926^{\circ}$ $0.702921^{\circ}$	0.702962 <sup>c</sup>
εNd 6.10 7.00 9.25 9.40 9.21	9.15
206 Pb/204 Pb 19.355 - 18.888 18.884 18.89	18.932
$^{207}$ Pb/ $^{204}$ Pb 15.591 - 15.543 15.545 15.544	15.547
$^{208}$ Pb/ $^{204}$ Pb 38.93 - 38.507 38.513 38.511	38.542
εHf 14.32 14.98 21.97 22.12 21.76	-

516 Table 1

Sample	B09-ROC24	GRA N05-1	GRA N06-1	GRA N10-2	GRA N10-3	GRA N12-1
Long. (°)	-32.278	-32.220	-32.275	-32.237	-32.257	-32.235
Lat. (°)	37.290	37.285	37.291	37.280	37.285	37.280
Туре	T-MORB	E-MORB	E-MORB	T-MORB	E-MORB	E-MORB
SiO <sub>2</sub>	52.1	50.4	48.1	51.7	51.0	50.3
$TiO_2$	1.06	1.1	1.12	1.09	0.97	1.01
$Al_2O_3$	15.2	18.7	17.5	15.0	16.6	16.2
FeO	9.6	5.9	6.8	9.4	7.5	7.8
MnO	0.18	0.1	0.14	0.17	0.14	0.16
MgO	8.2	6.8	8.2	8.0	8.3	8.8
CaO	12.6	14.4	15.1	12.4	13.4	13.3
Na <sub>2</sub> O	2.3	2.4	2.0	2.2	2.1	2.1
$K_2O$	0.18	0.3	0.46	0.18	0.41	0.30
$P_2O_5$	0.13	0.2	0.24	0.14	0.18	0.20
L.O.I.	-0.53	1.2	1.01	-0.26	0.21	0.53
Sum	102.1	102.1	101.5	101.2	101.6	101.5
La	5.60	8.51	14.01	6.08	11.65	8.68
Ce	12.36	16.93	27.49	13.41	22.36	17.54
Pr	1.70	2.19	3.36	1.83	2.68	2.22
Nd	8.03	9.69	13.90	8.39	11.08	9.63
Sm	2.40	2.50	3.15	2.45	2.58	2.38
Eu	0.90	1.02	1.12	0.92	0.92	0.91
Gd	3.27	3.16	3.57	3.29	3.13	2.96
Tb	0.59	0.53	0.57	0.58	0.51	0.50
Dy	3.87	3.38	3.50	3.78	3.23	3.15
Но	0.85	0.72	0.73	0.82	0.69	0.67
Er	2.38	1.98	2.02	2.32	1.93	1.88
Yb	2.36	1.91	1.92	2.27	1.86	1.85
Lu	0.37	0.30	0.29	0.35	0.29	0.29
Hf	1.62	2.01	2.09	1.66	1.66	1.71
<sup>87</sup> Sr/ <sup>86</sup> Sr	0.702981 <sup>b</sup>	-	-	-	-	0.702978 <sup>b</sup>
εNd	8.97	7.85	-	-	8.07	8.39
<sup>206</sup> Pb/ <sup>204</sup> Pb	18.944	_	19.229	19.013	19.238	19.003
$^{207}$ Pb/ $^{204}$ Pb	15.548	-	15.59	15.56	15.583	15.56
$^{208}$ Ph/ $^{204}$ Ph	38 557	_	38 833	38 624	38 807	38 62
	21.04	17.04	14.22	10.024	17 46	20.02
tH3	21.04	17.84	14.33	19.89	17.46	20.05

#### Table 1

## 523 **References**

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Figure 2













Figure 7



