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Review of the global models used within phase 1 of the Chemistry–Climate Model Initiative (CCMI)

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Abstract. We present an overview of state-of-the-art chemistry–climate and chemistry transport models that are used within phase 1 of the Chemistry–Climate Model Initiative (CCMI-1). The CCMI aims to conduct a detailed evaluation of participating models using process-oriented diagnostics derived from observations in order to gain confidence in the models’ projections of the stratospheric ozone layer, tropospheric composition, air quality, where applicable global climate change, and the interactions between them. Interpretation of these diagnostics requires detailed knowledge of the radiative, chemical, dynamical, and physical processes incorporated in the models. Also an understanding of the degree to which CCMI-1 recommendations for simulations have been followed is necessary to understand model responses to anthropogenic and natural forcing and also to explain inter-model differences. This becomes even more important given the ongoing development and the ever-growing complexity of these models. This paper also provides an overview of the available CCMI-1 simulations with the aim of informing CCMI data users.

1 Introduction

Climate models have been evolving considerably in recent decades. From relatively simple beginnings, ever more components have been added, until presently the “Earth System Models” (ESMs) define the state of the field. Such models simultaneously serve a variety of purposes including simulating air quality, tropospheric chemistry, stratospheric ozone, and global climate. These applications are strongly coupled; e.g. various air pollutants are climate active, stratospheric and tropospheric ozone are coupled in a variety of ways, and climate change affects atmospheric composition and vice versa. Previous-generation models were usually constructed for just one of these purposes (e.g. Morgenstern et al., 2010; Lamarque et al., 2013). Furthermore, increasingly biogeochemical feedbacks are considered, e.g. in the form of organic aerosol precursors emitted from land and ocean surfaces. However, the additional complexity characterizing ESMs makes these simulations more difficult to evaluate because previously ignored feedbacks now need to be considered.

The purpose of this paper is to document the internal make-up of 20 chemistry–climate models (CCMs) and chemistry-transport models (CTMs) participating in phase 1 of the Chemistry–Climate Model Initiative (CCMI; Eyring et al., 2013a), a combined activity of the International Global Atmospheric Chemistry (IGAC) and Stratosphere–troposphere Processes And their Role in Climate (SPARC) projects. CCMs are a major stepping stone on the way towards ESMs, combining physical climate models with an explicit representation of atmospheric chemistry. CCMI-1 continues the legacy of previous CCM inter-comparisons, par-

ticularly the Chemistry–Climate Model Validation (CCMVal; SPARC, 2010) and the Atmospheric Chemistry and Climate Model Intercomparison Project (ACCMIP; Lamarque et al., 2013). These precursors had more limited aims than CCMI; in the case of CCMVal, the main aim was to inform the World Meteorological Organization’s Scientific Assessments of Ozone Depletion (WMO, 2007, 2011, 2015) with state-of-the-science information about stratospheric ozone and its past and projected future evolution. The main outcome was SPARC (2010), a comprehensive assessment of the performance of stratospheric CCMs. In the case of ACCMIP, the aim was to inform the 5th Assessment Report of the Intergovernmental Panel on Climate Change (IPCC, 2013) about the roles of “near-term climate forcers”, notably tropospheric ozone and aerosols, in historical and future climate change. CCMI-1 builds on CCMVal and ACCMIP by encouraging the participation of coupled atmosphere–ocean–stratosphere–troposphere models that represent the various ways in which stratospheric and tropospheric composition are coupled to each other and to the physical climate more consistently than in their predecessor models. A second phase of CCMI, CCMI-2, is planned, where simulations will be conducted jointly with the Aerosol Comparisons between Observations and Models (AEROCOM) project and will contribute to the 6th Coupled Model Intercomparison Project (CMIP6) under the Aerosols and Chemistry Model Intercomparison Project (AerChemMIP; Collins et al., 2016). These will be performed with next-generation models under development and will not be discussed here (<http://blogs.reading.ac.uk/ccmi/>).

For CCMVal, Morgenstern et al. (2010) described salient model features of CCMVal-2 models, mostly in tabular forms. Their paper builds on numerous publications that describe individual models, or aspects thereof. Here we present an update to the tables in Morgenstern et al. (2010), focussing in the accompanying text on the changes that have occurred in the participating models since CCMVal-2. (In all but three cases, older versions of the CCMI-1 models had participated in CCMVal-2; see below.) This paper is meant to support other publications evaluating the CCMI-1 simulations by providing an overview of the make-up of CCMI-1 models as well as a comprehensive literature list for further reading.

2 Participating models

There are 20 models participating in CCMI-1 (Tables 1 and 2). With three exceptions (CHASER (MIROC-ESM), TOMCAT, MOCAGE), each participating model had a predecessor model in CCMVal-2; hence, the focus here will be on developments since CCMVal-2. MOCAGE participated in ACCMIP (Lamarque et al., 2013). Corresponding to the much broader scope of CCMI, relative to CCMVal-2, salient developments in these models include whole-atmosphere

Table 1. Participating models and contact information. For abbreviations of institution names, see the authors' affiliations.

Model names	Institutions	Principal investigators (PIs)	Email addresses
ACCESS CCM	U. Melbourne, AAD, NIWA	K. Stone, R. Schofield, A. Klekociuk, D. Karoly, O. Morgenstern	stonek@mit.edu robyn.schofield@unimelb.edu.au andrew.klekociuk@aad.edu.au
CCSRNIES MIROC3.2	NIES	H. Akiyoshi, Y. Yamashita	hakiyosi@nies.go.jp yyousuke@jamstec.go.jp
CESM1 CAM4-chem	NCAR	S. Tilmes, J.-F. Lamarque	tilmes@ucar.edu lamar@ucar.edu
CESM1 WACCM	NCAR	D. Kinnison, R. R. Garcia, A. K. Smith, A. Gettelman, D. Marsh, C. Bardeen, M. Mills	dkin@ucar.edu rgarcia@ucar.edu aksmith@ucar.edu andrew@ucar.edu marsh@ucar.edu bardeenc@ucar.edu mmills@ucar.edu
CHASER (MIROC-ESM)	U. Nagoya, JAMSTEC, NIES	K. Sudo, T. Nagashima	kengo@nagoya-u.jp nagashima.tatsuya@nies.go.jp
CMAM	CCCma, Environment and Climate Change Canada	D. Plummer, J. Scinocca	david.plummer@canada.ca john.scinocca@canada.ca
CNRM-CM5-3	CNRM Météo-France CNRS – CERFACS	M. Michou, D. Saint-Martin	martine.michou@meteo.fr david.saint-martin@meteo.fr
EMAC	DLR-IPA, KIT-IMK-ASF, KIT-SCC-SLC, FZJ-IEK-7, FUB, UMZ-IPA, MPIC, CYI	P. Jöckel, H. Tost, A. Pozzer, M. Kunze, O. Kirner, S. Brinkop, D. S. Cai, J. Eckstein, F. Frank, H. Garny, K.-D. Gottschaldt, P. Graf, V. Grewe, A. Kerkweg, B. Kern, S. Matthes, M. Mertens, S. Meul, M. Nützel, S. Oberländer-Hayn, R. Ruhnke, R. Sander	messy_admin@lists.mpic.de patrick.joeckel@dlr.de
GEOSCCM	NASA/GSFC	L. D. Oman, S. E. Strahan	luke.d.oman@nasa.gov susan.e.strahan@nasa.gov
GFDL-AM3/CM3	NOAA GFDL	M. Y. Lin, L. W. Horowitz	meiyun.lin@noaa.gov larry.horowitz@noaa.gov
HadGEM3-ES	MOHC	F. M. O'Connor, N. Butchart, S. C. Hardiman, S. T. Rumbold	fiona.oconnor@metoffice.gov.uk neal.butchart@metoffice.gov.uk
LMDz-REPROBUS	IPSL	S. Bekki, M. Marchand, F. Lott, D. Cugnet, L. Guez, F. Lefevre, S. Szopa	slimane@latmos.ipsl.fr marion.marchand@latmos.ipsl.fr
MOCAGE	Météo France CNRS	B. Josse, V. Marecal	beatrice.josse@meteo.fr virginie.marecal@meteo.fr
MRI-ESM1r1		M. Deushi, T. Y. Tanaka, K. Yoshida	mdeushi@mri-jma.go.jp yatanaka@mri-jma.go.jp kyoshida@mri-jma.go.jp
NIWA-UKCA	NIWA	O. Morgenstern, G. Zeng	olaf.morgenstern@niwa.co.nz guang.zeng@niwa.co.nz
SOCOL	PMOD/WRC, IAC/ETHZ	E. Rozanov, A. Stenke, L. Revell	eugene.rozanov@pmodwrc.ch andrea.stenke@env.ethz.ch laura.revell@env.ethz.ch
TOMCAT	U. Leeds	S. Dhomse, M. P. Chipperfield	m.chipperfield@leeds.ac.uk s.s.dhomse@leeds.ac.uk

Table 1. Continued.

Model names	Institutions	Principal investigators (PIs)	Email addresses
ULAQ-CCM	U. L'Aquila	G. Pitari, E. Mancini, G. Di Genova	gianni.pitari@aquila.infn.it eva.mancini@aquila.infn.it glauco.digenova@aquila.infn.it
UMSLIMCAT	U. Leeds	S. Dhomse, M. P. Chipperfield	m.chipperfield@leeds.ac.uk s.s.dhomse@leeds.ac.uk
UMUKCA-UCAM	U. Cambridge	N. L. Abraham, A. T. Archibald, R. Currie, J. A. Pyle	luke.abraham@atm.ch.cam.ac.uk ata27@cam.ac.uk rc454@cam.ac.uk john.pyle@atm.ch.cam.ac.uk

chemistry (almost all CCMVal-2 models have been further developed to include tropospheric chemistry), coupling (now several of the CCMI-1 models include an interactive ocean), increased resolution for several of them, and progress in various areas of model physics.

In the following, we comment on noteworthy developments relative to the CCMVal-2 models (SPARC, 2010). Apart from some very high-level information (such as model names and contact information), all tabulated information about the models is in the Supplement.

2.1 Family relationships between the models

Like in CCMVal-2, some family relationships are apparent between different models (Table S1). For example, ACCESS-CCM, NIWA-UKCA, UMUKCA-UCAM, and HadGEM3-ES use similar atmosphere, ocean, and sea ice components (ACCESS-CCM and UMUKCA-UCAM are atmosphere-only). GFDL-AM3 is the atmosphere-only equivalent of GFDL-CM3. EMAC and SOCOL are both based on different versions of the ECHAM5 climate model. LMDz-REPROBUS-CM6 can be coupled to a similar version of the Nucleus for a European Model of the Ocean (NEMO; Madec, 2008) as NIWA-UKCA and HadGEM3-ES; however, at the time of writing this paper only the atmosphere-only CM5 version has been used for CCMI-1 simulations. CCSRNIES MIROC 3.2 uses a similar version of the MIROC atmosphere model as CHASER. CESM1 CAM4-chem is the low-top counterpart of CESM1 WACCM; i.e. troposphere–stratosphere aspects of the two models are generally identical.

2.2 Atmosphere grids and resolution

In CCMVal-2, one model (AMTRAC) used a novel grid that was neither latitude–longitude nor spectral, namely the cubed-sphere grid. While several modelling centres presently are working on new-generation models based on this or similar novel grids, in the CCMI-1 ensemble the GFDL successors to AMTRAC continue to use this grid; GEOSCCM has adopted this grid (Table 3). It is anticipated that more models

will adopt novel grids in future model inter-comparisons, to improve scalability and computing efficiency.

The horizontal resolution of most models is unchanged versus CCMVal-2 and in the case of MOCAGE, ACCMIP, respectively (Table 3). ULAQ CCM, HadGEM3-ES, MRI ESM, CMAM, CNRM-CM5-3 (for chemistry), SOCOL, and LMDz-REPROBUS-CM6 have increased their horizontal resolution; now resolution ranges between roughly 5° to less than 2°. In several cases, versus CCMVal-2 the models have increased their vertical resolution, particularly CNRM-CM5-3, LMDz-REPROBUS, MRI ESM, ULAQ CCM, and HadGEM3-ES (Tables 3, S2). Vertical ranges are essentially unchanged versus the models' CCMVal-2 (or ACCMIP, for MOCAGE) counterparts with the uppermost model levels ranging from 35 km (for MOCAGE) to 140 km (for CESM1 WACCM). All but two models (CESM1 CAM4-chem, model top at 200 Pa, MOCAGE – 500 Pa) completely cover the stratosphere (assuming a stratopause pressure of 100 Pa). CESM1 CAM4-chem is the low-top equivalent of CESM1 WACCM, so simulations by these two models can be analysed for the role of the model top height (Table S2).

2.3 Advection

In CCMVal-2, there were some models that used different transport schemes for hydrological versus chemical tracers (Morgenstern et al., 2010). For CCMI-1, all CCMs (which transport both types of tracers) employ the same transport scheme for both (Tables S3, S4). This makes the advection of all tracers physically self-consistent.

2.4 Time stepping and calendars

Atmosphere models use a variety of time steps for the dynamical core, physical processes, radiation, transport, chemistry, and the coupling of chemistry and dynamics for different reasons. Generally the choice of time step is the result of a compromise between the computational cost associated with short time steps and the reduced accuracy associated with larger time steps. There is a considerable amount of diversity in the CCMI-1 ensemble regarding the choices of time steps

Table 2. Model versions and key references. More details on the CCMVal-2 versions are in Morgenstern et al. (2010).

Model	Revision/version	Reference(s)	CCMVal-2 precursor model
ACCESS CCM, NIWA-UKCA	MetUM 7.3	Morgenstern et al. (2009, 2013), Stone et al. (2016)	UMUKCA-UCAM (MetUM 6.1)
CCSRNIES MIROC3.2	3.2	Imai et al. (2013); Akiyoshi et al. (2016)	CCSRNIES
CESM1 CAM4-chem	CCMI_23	Tilmes (2015b)	CAM3.5
CESM1 WACCM	CCMI_30	Solomon et al. (2015); Garcia et al. (2016), Marsh et al. (2013)	WACCM v3.5.48
CHASER (MIROC-ESM)	v4.5	Sudo et al. (2002); Sudo and Akimoto (2007), Watanabe et al. (2011), Sekiya and Sudo (2012, 2014)	N/A
CMAM	v2.1	Jonsson et al. (2004); Scinocca et al. (2008)	CMAM
CNRM-CM	v5-3	Voltaire et al. (2012); Michou et al. (2011), http://www.cnrm-game-meteo.fr/	CNRM-ACM
EMAC	v2.51	Jöckel et al. (2010, 2016)	EMAC
GEOSCCM	v3	Molod et al. (2012, 2015), Oman et al. (2011, 2013)	GEOSCCM
GFDL-AM3	v3	Donner et al. (2011), Lin et al. (2014, 2015a, b)	AMTRAC3
GFDL-CM3	v3 (CMIP5)	Griffies et al. (2011); John et al. (2012), Levy II et al. (2013)	AMTRAC3
HadGEM3-ES	HadGEM3 GA4.0, NEMO 3.4, CICE, UKCA, MetUM 8.2	Walters et al. (2014); Madec (2008), Hunke and Lipscombe (2008); Morgenstern et al. (2009), O'Connor et al. (2014); Hardiman et al. (2016),	UMUKCA-METO (MetUM 6.1)
LMDz-REPROBUS-CM5 & CM6	IPSL-CM5 & CM6	Marchand et al. (2012); Szopa et al. (2012), Dufresne et al. (2013). No reference yet on CM6	LMDZrepro
MRI-ESM1r1	v1.1	Yukimoto et al. (2012, 2011), Deushi and Shibata (2011)	MRI
MOCAGE	v2.15.1	Josse et al. (2004); Guth et al. (2016)	N/A
SOCOL	v3	Revell et al. (2015), Stenke et al. (2013)	SOCOL v2.0 (Schraner et al., 2008)
TOMCAT	v1.8	Chipperfield (1999, 2006)	N/A
ULAQ-CCM	v3, yr 2012	Pitari et al. (2014)	ULAQ
UMSLIMCAT	v1	Tian and Chipperfield (2005)	UMSLIMCAT
UMUKCA-UCAM	MetUM 7.3	Morgenstern et al. (2009), Bednarz et al. (2016)	UMUKCA-UCAM (MetUM 6.1)

for different processes (Table S6). In addition, most models now use the Gregorian or the 365-day calendars (whereas in CCMVal-2, for reasons of easier handling of averages and climatologies, often a 360-day calendar was used). Only the Met Office Unified Model (MetUM)-based models (ACCESS CCM, HadGEM3-ES, NIWA-UKCA, UMSLIMCAT, UMUKCA-UCAM) and ULAQ still use the 360-day calendar.

2.5 Horizontal diffusion

Numerical diffusion relates to the impossibility to represent transport in an exact manner on a discrete grid. Errors occurring in such a process can usually be described as an unphysical, numerical diffusion process. It is an unavoidable aspect of numerical climate models. Transport schemes are generally designed to minimize numerical diffusion. However, in addition, several models require explicit diffusion for stability (Table S7). In CESM1 CAM4-chem

Table 3. Governing equations, horizontal discretization, and vertical grid of the atmosphere component of models. NH is non-hydrostatic; PE is primitive equations; QG is quasi-geostrophic; F[D,V]LL is finite [difference, volume] on lat–long grid; STQ is spectral transform quadratic; STL is spectral transform linear; CP is Charney–Phillips; TA is hybrid terrain-following altitude; TP is hybrid terrain-following pressure; NTP is non-terrain-following pressure; FVCS is finite volume cubed sphere; T21 $\approx 5.6^\circ \times 5.6^\circ$; T42 $\approx 2.8^\circ \times 2.8^\circ$; T47 $\approx 2.5^\circ \times 2.5^\circ$; T63 $\approx 1.9^\circ \times 1.9^\circ$; TL159 $\approx 1.125^\circ \times 1.125^\circ$.

Model name	Gov. eq.	Hor. disc.	Resolution	Vert. grid	Top level	Top of model	Coord. sys.	Comment
ACCESS CCM								
NIWA-UKCA	NH	FDLL	$3.75^\circ \times 2.5^\circ$	CP60	84 km	84 km	TA	Arakawa-C
UMUKCA-UCAM								
CCSRNIES MIROC3.2	PE	STQ	T42	L34	1.2 Pa	1 Pa	TP	
CHASER (MIROC-ESM)	PE	STQ	T42	L57	56 km	56 km	TP	Arakawa-C
CESM1 CAM4-chem	PE	FVLL	$1.9^\circ \times 2.5^\circ$	L26/56	200 Pa	100 Pa	TP	Lin (2004)
CESM1 WACCM	PE	FVLL	$1.9^\circ \times 2.5^\circ$	L66/88	140 km	140 km	TP	Lin (2004)
CMAM	PE	STL	T47	L71	0.08 Pa	0.0575 Pa	TP	
CNRM-CM5-3	PE	STL	T63	L60/89	7/8 Pa	0 Pa	TP	
EMAC	PE	STQ	T42	L47/90	1 Pa	0 Pa	TP	
GEOSCCM	PE	FVCS	$\sim 2^\circ \times 2^\circ$	L72	1.5 Pa	1 Pa	TP	
GFDL-AM3/CM3	NH	FVCS	$\sim 2^\circ \times 2^\circ$	L48	86 km	86 km	TA	Donner et al. (2011)
HadGEM3-ES	NH	FDLL	$1.875^\circ \times 1.25^\circ$	CP85	85 km	85 km	TA	Arakawa-C
MRI-ESM1r1	PE	STL	TL159	L80	1 Pa	0 Pa	TP	
LMDz-REPROBUS	PE	FVLL	$3.75^\circ \times 2.5^\circ$	L39/79	$\sim 70/80$ km	$\sim 70/80$ km	TA	Arakawa-C
MOCAGE	CTM	FDLL	$2^\circ \times 2^\circ$	L47	500 Pa	500 Pa	TP	
SOCOL	PE	STL	T42	L39	1 Pa	0 Pa	TP	
TOMCAT	CTM	FVLL	$2.8^\circ \times 2.8^\circ$	L60	10 Pa	0 Pa	TP	ERA-Interim
ULAQ CCM	QG	STL	T21	CP126	4 Pa	4 Pa	NTP	
UMSLIMCAT	PE	FDLL	$3.75^\circ \times 2.5^\circ$	L64	1 Pa	0.77 Pa	TP	Arakawa-B

and CESM1 WACCM, hyperdiffusion is applied to the smallest scales, and through Fourier transformation and filtering the effective resolution is kept the same at all latitudes. Several models (ACCESS CCM, NIWA-UKCA, UMUKCA-UCAM, GEOSCCM, HadGEM3-ES, LMDz-REPROBUS, MOCAGE) do not contain explicit diffusion in most of their domains. “Sponges” are generally used to prevent reflection of planetary or Rossby waves off the model top, except for CMAM, the MRI-ESM, and HadGEM3-ES. The MetUM family of models also requires diffusion over the poles (ACCESS CCM, NIWA-UKCA, UMUKCA-UCAM, HadGEM3-ES). The need for polar filtering should disappear with the future adoption of “novel” grids that no longer have any singularities at the poles.

2.6 Quasi-biennial oscillation

Essentially, the same CCMs that used nudging for CCMVal-2 continue to use nudging to impose a quasi-biennial oscillation (QBO) in their models. However, nudging is performed in a more sophisticated way, with SOCOL, CCSRNIES MIROC3.2, and EMAC now using smooth transitions at the edges of the nudged region (Table S8). Other models do not impose a QBO (except for the specified-dynamics simulations). This means that the QBO is either occurring spontaneously in these models (validating this and other aspects of model behaviour is beyond the scope of this paper), or it is simply absent.

2.7 Orographic and non-orographic gravity wave drag

Gravity waves are the result of vertical displacements of air in the presence of stratification, which can be due, e.g., to mountains, frontal systems, or tropospheric convective activity. They can either be dissipated if they encounter critical levels (at which the phase speed equals the background winds), or they continue to propagate upwards and increase in amplitude, in accordance with the decreasing air density. Eventually they can break, leading to deceleration of the mean flow. This process contributes to the driving of the stratospheric Brewer–Dobson Circulation, but also affects the temperature structure of the middle atmosphere. Their horizontal scale is mostly below the grid scale, meaning that this process needs to be parameterized. Gravity waves are also poorly observed, contributing to a substantial diversity of approaches to representing this process (Table S9). The paucity of observations leads to gravity wave drag being often used to tune better known model diagnostics such as stratospheric temperatures or age of air.

Gravity wave drag (GWD) is usually divided into two components for modelling: orographic and non-orographic drag. The representation of orographic drag is based on the interaction of flow with topography, a relatively well-known process. Non-orographic gravity waves by contrast are geographically poorly constrained; therefore, often relatively simple approaches, not taking into account any tropospheric meteorology, are used. However, in contrast to

CCMVal-2, several models now link non-orographic drag to tropospheric processes such as convection (CNRM-CM5-3, CESM1 WACCM). This means that in these models, possible changes in the GWD sources associated with climate change are represented.

2.8 Physical parameterizations

References for the descriptions of the models' physical parameterizations such as turbulent vertical fluxes and dry convection, moist convection, cloud microphysics, aerosol microphysics, and cloud cover can be found in Tables S10 and S11. Several models have renewed their physics parameterizations since CCMVal-2, namely ACCESS CCM, NIWA-UKCA, CNRM-CM5-3, GFDL-CM3/AM3, HadGEM3-ES, LMDz-REPROBUS, MRI-ESM, SOCOL, and UМУKCA-UCAM.

2.9 Cloud microphysics

Clouds remain a very substantial source of uncertainty and inter-model differences, e. g. regarding climate sensitivity (for a summary of current understanding see chapter 7 of IPCC, 2013). Small-scale variability, non-equilibrium processes, cloud-aerosol interactions, and other processes all contribute to this. The CCM1-1 model ensemble is characterized by some considerable diversity in approaches to cloud microphysics (Table S12), and most models have implemented changes in the way clouds are represented, relative to CCMVal-2 (where clouds never were a particular focus).

2.10 Tropospheric chemistry

In contrast to CCMVal-2 (which did not focus on tropospheric chemistry), a majority of CCM1-1 models now explicitly represent tropospheric ozone chemistry (Table S13). Six models do not represent any non-methane hydrocarbon (NMHC) chemistry (CCSRNIES MIROC3.2, CMAM, CNRM-CM5-3, LMDz-REPROBUS, TOMCAT, UMSLIMCAT). In LMDz-REPROBUS, a climatological, zonally invariant tropospheric composition is prescribed below 400 hPa. Unlike stratospheric chemistry, tropospheric chemistry is too complex to be incorporated comprehensively in a CCM. The need to include an affordable yet skilled tropospheric chemistry scheme drives some diversity in the chemistry schemes and correspondingly the represented NMHC source gases. Several schemes use lumping, whereby emissions of a non-represented NMHC source gas are implemented as emissions of a represented one. Sometimes this species is denoted as "a lumped species" or "OTHC" (other carbon; Table S15). SOCOL has the simplest organic chemistry scheme in the ensemble (the only organic NMHC source gas, disregarding HCHO, is isoprene, C₅H₈). By contrast, the CTM MOCAGE and several CCMs represent 10 or more NMHC source gases.

For methane (CH₄) the recommendation is to use a single prescribed time-evolving volume-mixing ratio (as defined by the Representative Concentration Pathway (RCP 6.0); Meinshausen et al., 2011) as the global lower boundary condition. This is followed by almost all models. CHASER has an interactive methane scheme, and the EMAC and ULAQ models prescribe CH₄ mixing ratios at the surface under consideration of a hemispheric asymmetry (i.e. there is about 5 % less CH₄ in the Southern than in the Northern Hemisphere). EMAC also prescribes a seasonal cycle for CH₄.

2.11 Stratospheric chemistry

Stratospheric gas-phase chemistry is well-enough understood and sufficiently simple so that it can be treated mostly explicitly, by adopting all relevant reactions for which rate coefficients have been published. Most models follow the Sander et al. (2011b) recommended rates. There is some diversity as to which halogen source gases are considered (Table S13). Six models represented here also participated in the re-assessment of lifetimes of long-lived species (SPARC, 2013). In this context, UМУKCA-UCAM and CESM1 WACCM have expanded their range of halogen source gases, relative to CCMVal-2. Most MetUM-based participants (ACCESS CCM, HadGEM3-ES, NIWA-UKCA, UMSLIMCAT) continue to lump chlorine source gases, which are dominated by chlorofluorocarbons (CFCs), into only two representatives (CFC-11, CFC-12; Morgenstern et al., 2009). SOCOL and the CESM model family, at 14 and 12 species including Halon-1211, respectively, have the largest number of chlorine source gases. For bromine, the recommendation was to include the short-lived constituents di-bromomethane (CH₂Br₂) and bromoform (CHBr₃; Eyring et al., 2013a); about half of the models follow this recommendation. All models represent CH₃Br (the most abundant bromine source gas); several also include Halon-1211, Halon-1301, and/or Halon-2402. CH₃Br in some cases is lumped with other bromine source gases not represented (ACCESS CCM, HadGEM3-ES, NIWA-UKCA, UMSLIMCAT). EMAC also has a representation of a sea salt aerosol source of gas-phase halogen, which may be of importance to the tropospheric oxidizing capacity (Allen et al., 2007), and a larger range of very short-lived organic bromine compounds, which likely influence tropospheric and stratospheric ozone chemistry. There do not appear to be fundamental differences with respect to how stratospheric chemistry was treated for CCMVal-2. For example, all models have an explicit representation of methane oxidation to produce stratospheric water vapour that is based on similar reactions and rates. CESM1 WACCM, which covers the upper atmosphere, explicitly treats ion/neutral chemistry important in that region. Other than that, there appears to be no significant characteristic of the formulation of stratospheric chemistry that would, e.g., distinguish low-top from high-top models. Morgenstern et al. (2010) provided a more exhaustive dis-

discussion of stratospheric chemistry in CCMVal-2 models that is still generally relevant.

2.12 Stratospheric and tropospheric heterogeneous chemistry

Heterogeneous chemistry (i.e. reactions that require a solid or liquid surface as a catalyst) is crucial to several aspects of atmospheric chemistry, notably the ozone hole and the tropospheric nitrogen cycle. Most of the reactions in Table S17 are chlorine and/or bromine activation reactions, e.g. they turn chlorine from its unreactive forms (HCl, ClONO₂) into reactive forms (that photolyse readily in sunlight). The implementation of heterogeneous chemistry is subject to considerable inter-model differences regarding represented reactions and their associated heterogeneous surface types. Seven models (CCSRNIES MIROC3.2, CESM1 CAM4-chem, CESM1 WACCM, CHASER, CMAM, GFDL CM3/AM3, and MOCAGE) explicitly consider supercooled ternary solutions (STS; mixtures of HNO₃, H₂SO₄, and H₂O) which impact stratospheric chemistry through swelling of droplets and associated denitrification and heterogeneous chemistry. Nitric acid formation (and subsequent nitrogen removal) partly occurs on/in cloud droplets, ice crystals, and on aerosol surfaces. Most models have nitric acid formation occurring on nitric acid trihydrate (NAT) and ice surfaces and also on sulfate aerosol (the details of which depend on how this aerosol is represented). In EMAC and MRI ESM1r1, this process also occurs on sea salt aerosol. Also SO₂ oxidation to form SO₃ (which then further reacts to form sulfate aerosol; the intermediates are often omitted in model formulations) partly occurs in the aqueous cloud phase. Most models include this heterogeneous reaction (Table S18). In several models (ACCESS CCM, CESM1 CAM4-chem, CESM1 WACCM, CHASER (MIROC ESM), GFDL-AM3/CM3, HadGEM3-ES, MOCAGE, NIWA-UKCA, ULAQ, UМУKCA-UCAM) this process involves a gas-phase reaction with OH and aqueous-phase reactions with O₃ and H₂O₂ (Feichter et al., 1996; Kreidenweis et al., 2003; Tie et al., 2005). EMAC treats SO₂ oxidation as part of a complex gas- and aqueous-chemistry mechanism detailed by Jöckel et al. (2016). ULAQ also has SO₂+H₂O₂→SO₃ occurring on upper-tropospheric ice particles (Clegg and Abbatt, 2001). In MRI-ESM1r1, SO₂ reacts with gas-phase OH, O₃, and O(³P) to form SO₃, with rates following Sander et al. (2011b). In all cases, the oxidants are calculated interactively.

2.13 Polar stratospheric clouds (PSCs)

Like for CCMVal-2, the models divide into two groups: those that assume thermodynamical equilibrium for PSCs (i.e. the gas-phase constituents are reduced to their saturation abundances and excess matter is condensed into PSCs), and others (CESM1, EMAC, GEOSCCM, ULAQ) that account for deviations from thermodynamic equilibrium (Ta-

ble S19). Usually, at least two PSC types (type 1: nitric acid trihydrate; type 2: ice) as well as ubiquitous sulfate aerosol are assumed (Table S17). Several models also account for STS (i.e. HNO₃ + H₂SO₄ + H₂O mixtures). With the exception of CMAM, all models account for PSC sedimentation (which leads to denitrification) but assumptions around this vary considerably. Several models impose fixed sedimentation velocities for the different PSC types; in others, these velocities are a function of particle size. In most models, the approach to handling PSCs appears to be unchanged versus SPARC (2010).

2.14 Tropospheric aerosol

The additional focus, relative to CCMVal-2, on tropospheric climate-composition linkages has led to most models including an explicit treatment of tropospheric aerosol, except for CCSRNIES MIROC3.2, EMAC, CNRM-CM5-3, LMDz-REPROBUS, and SOCOL (Table S20). CMAM uses prescribed sulfate aerosol surface area densities in the troposphere for heterogeneous chemistry calculations. In CCMVal-2, most models did not have any representation of tropospheric aerosols.

Most aerosol schemes are “bulk”, i.e. only total mass of an aerosol type is predicted. In bulk schemes, derived quantities such as particle number require assumptions about particle sizes to be made. The ULAQ CCM and MRI-ESM1r1 use sectional approaches, which represent aerosols of different size classes in discrete bins, thus avoiding a priori assumptions on particle size. The ULAQ CCM represents nitrate aerosol in a modal way; i.e. the aerosol size distribution is assumed to be described by one or more log-normal distributions. Modal schemes are computationally more efficient than sectional schemes while also predicting both aerosol size and number. Several models (CNRM-CM5-3, EMAC, SOCOL) use offline representations of aerosol. Types of aerosol included in the models comprise dust, sea salt, organic carbon, black carbon, sulfate, and nitrate (in the case of CHASER and ULAQ). These provide surfaces for heterogeneous chemistry (Table S18) but there appears to be little consistency regarding how this aspect is treated.

2.15 Volcanic effects

There has been considerable progress regarding the physical consistency of volcanic effects in CCMs. Whereas in CCMVal-2, surface area densities and aerosol-induced heating rates were prescribed; now 13 of the CCMI-1 models treat radiative effects online, i.e. calculate or assume an aerosol size distribution for volcanic aerosol in the stratosphere and derive radiative heating rates from this (CCSRNIES MIROC3.2, CESM1 (both versions), CHASER, CMAM, CNRM-CM5-3, EMAC, GEOSCCM, HadGEM3-ES, MRI ESM1r1, SOCOL, ULAQ CCM, UМУKCA-UCAM; Table S21). GFDL-AM3/CM3 prescribes heat-

ing rates associated with the presence of aerosol. The remaining models (ACCESS CCM, NIWA-UKCA, LMDz-REPROBUS, UM-SLIMCAT) do not consider stratospheric volcanic aerosol in their radiation schemes or do not have a radiation scheme (MOCAGE, TOMCAT). Surface area density used in heterogeneous chemistry calculations, with only two exceptions (GFDL-AM3/CM3, UMUKCA-UCAM), follows the CCM1-1 recommendation (Sect. 4).

2.16 Photolysis

In CCMVal-2 and ACCMIP, in both cases half the models used tabulated photolysis rates and interpolation to calculate photolysis rates (Morgenstern et al., 2010; Lamarque et al., 2013). This approach is problematic in the troposphere because of complicating effects of clouds, aerosols, surface albedo, and other factors that are not considered in the pre-calculated tables. It is, however, computationally more efficient. With the new focus, relative to CCMVal-2, on tropospheric chemistry, all models that have explicit tropospheric chemistry also take explicit account of the presence of clouds (Table S22). GFDL-AM3 and the MetUM family (ACCESS CCM, HadGEM3-ES, NIWA-UKCA, UMUKCA-UCAM) have adopted the FAST-JX online formulation of photolysis in the domain (below 60 km, in the case of the MetUM family), and a group of other models continue to use look-up tables but apply corrections accounting for the presence of clouds (CESM1 CAM4-chem, CESM1 WACCM, CMAM, MOCAGE, MRI-ESM1r1, SOCOL). Also in many cases photolysis cross sections have been updated, relative to CCMVal-2.

2.17 Shortwave radiation

In most cases, models are using the same basic schemes as documented in SPARC (2010, Table S23). (The CTMs do not have any explicit treatment of radiation.) However, ULAQ, MRI ESM1r1, the 79-level version of LMDz-REPROBUS, the GFDL models, EMAC, CNRM-CM5-3, CCSRNIES MIROC3.2, and SOCOL have all increased their spectral resolution versus their CCMVal-2 predecessors.

2.18 Longwave radiation

Longwave radiation is treated largely in the same way as documented in SPARC (2010). (Again, the CTMs do not represent this process.) However, again a few models (CNRM-CM5-3, MRI-ESM1r1, SOCOL, ULAQ) have increased their spectral resolution versus their CCMVal-2 predecessors (Table S24).

2.19 Solar forcing

Interactions between the atmosphere and the Sun are considered in an increasingly consistent manner in CCM1-1 models (Tables S25 and S26). All models consider spectrally re-

solved irradiance. In six models (CCSRNIES MIROC3.2, EMAC, HadGEM3-ES, MRI ESM1r1, SOCOL, and ULAQ CCM) photolysis and shortwave radiation are handled consistently; in particular, the solar cycle is handled consistently in both. In the remaining models, the shortwave radiation and photolysis schemes have not specifically been made consistent. SOCOL and the MRI ESM1r1 also consider proton ionization by solar particles. With the exceptions of ACCESS CCM, NIWA-UKCA, and GEOSCCM, all models listed in Table S21 consider solar variability.

2.20 Ocean surface forcing

For the atmosphere-only reference (REF-C1 and REF-C1SD) and sensitivity (SEN-C1) simulations (Sect. 3), ocean surface forcing (sea surface temperatures, sea ice) need to be imposed (Eyring et al., 2013a). Most modelling groups used the Hadley Centre Ice and Sea Surface Temperature (HadISST) dataset (Rayner et al., 2003), as recommended. The LMDz-REPROBUS model uses the Atmosphere Model Intercomparison Project (AMIP) II dataset (http://www-pcmdi.llnl.gov/projects/amip/AMIP2EXPDSN/BCS_OBS/amip2_bcs.htm; Table S27). For the atmosphere-ocean coupled reference (REF-C2) and sensitivity (SEN-C2) simulations (Sect. 3), those models that do not couple to an interactive ocean/sea ice module require climate model fields to be imposed. A substantial variety of different climate model datasets were used for this purpose. In the ULAQ CCM simulations, an ocean surface dataset was used that was derived from a climate model simulation, with mean biases relative to HadISST removed.

2.21 Ocean coupling

Nine of the CCM1-1 models are coupled, at least for some simulations, to an interactive ocean module (Tables S1 and S28), namely CESM1 CAM4-chem, CESM1 WACCM, CHASER, EMAC, GFDL-CM3, HadGEM3-ES, LMDz-REPROBUS-CM6, MRI-ESM 1r1, and NIWA-UKCA. This is a substantial increase from CCMVal-2 and ACCMIP, when only one model each (CMAM and GISS-E2-R) was coupled to an ocean/sea ice model (Morgenstern et al., 2010; Lamarque et al., 2013). These models therefore self-consistently represent climate change throughout the atmosphere and ocean domains, in contrast to atmosphere-only models where oceanic feedbacks are not considered. Three of the models (HadGEM3-ES, LMDz-REPROBUS-CM6, NIWA-UKCA) use versions of the NEMO ocean model (Madec, 2008). The other five use independent ocean models. For sea ice, apart from HadGEM3-ES and NIWA-UKCA, which use versions of CICE, all of the models use different sea ice modules. The coupling between the atmosphere, ocean, and sea ice modules involves the passing of several physical fields that define the interactions between these modules, which essentially consist of transfers of momentum, heat, moisture, and

salinity. The coupling frequency also varies a lot, from daily to hourly.

2.22 Land surface, soil, and the planetary boundary layer

Land surface properties, such as vegetation and soil type, but also soil moisture, snow, and groundwater have significant climate effects (e.g. influence the severity of droughts and floods), mediated through surface albedo and evapotranspiration (IPCC, 2013). Hence, their representation in climate models is essential especially for regional climate simulations. CCMI-1 models, like IPCC-type climate models, generally have land surface schemes. Like for cloud microphysics, there is some considerable diversity of approaches in treating these three aspects of the models (Table S29).

3 CCMI-1 simulations

In this section, we briefly describe the motivation and some technical details regarding the experiments conducted for CCMI-1. Eyring et al. (2013a) have given more details. The specific forcings imposed are discussed briefly in Sect. 4.

- REF-C1: this experiment is analogous to the REF-B1 experiment of CCMVal-2. Using state-of-knowledge historic forcings and observed sea surface conditions, the models simulate the recent past (1960–2010). The models are free-running.
- REF-C1SD: this is similar to REF-C1 but the models are nudged towards reanalysis datasets, and correspondingly the simulations only cover 1980–2010. (“SD” stands for specified dynamics.) Through a comparison to the REF-C1 simulations, the influence on composition of dynamical biases and differences in variability between the reanalysis and the models can be assessed. This type of experiment had not been conducted for CCMVal-2. Table S27 has details on how nudging is implemented in those models that have conducted the specified-dynamics simulations.
- REF-C2: this experiment is a set of seamless simulations spanning the period 1960–2100, similar to the REF-B2 experiment for CCMVal-2. The experiments follow the WMO (2011) A1 scenario for ozone-depleting substances and the RCP 6.0 (Meinshausen et al., 2011) for other greenhouse gases, tropospheric ozone (O_3) precursors, and aerosol and aerosol precursor emissions. Ocean conditions can either be taken from a separate climate model simulation, or the models can be coupled interactively to ocean and sea ice modules.

In addition to these reference simulations, a variety of sensitivity simulations have been asked for, which are variants

on the reference simulations, typically with just one aspect changed.

- SEN-C2-fODS/SEN-C2-fODS2000: this is the same as REF-C2 but with ozone-depleting (halogenated) substances (ODSs) fixed at their 1960 or 2000 levels, respectively. The SEN-C2-fODS2000 simulations start in 2000, with ODS surface-mixing ratios fixed at their year-2000 values.
- SEN-C2-fGHG: this is similar to REF-C2 but with greenhouse gasses (GHGs) fixed at their 1960 levels, and sea surface and sea ice conditions prescribed as the 1955–1964 average (where these conditions are imposed).
- SEN-C2-fCH4: this experiment is identical to REF-C2 but the methane surface-mixing ratio is fixed to its 1960 value (Hegglin et al., 2016).
- SEN-C2-CH4rcp85: this experiment is identical to REF-C2 but the methane surface-mixing ratio follows the RCP 8.5 scenario, which anticipates a much larger increase in CH_4 than RCP 6.0, approaching 3.8 ppmv in 2100. All other GHGs and forcings follow RCP 6.0. The simulation covers 2005–2100.
- SEN-C2-fN2O: this experiment is identical to REF-C2 but the nitrous oxide surface-mixing ratio is fixed to its 1960 value (Hegglin et al., 2016).
- SEN-C1-fEmis/SEN-C1SD-fEmis: in these experiments, for species with explicit surface emissions such as nitrogen oxides (NO_x), carbon monoxide (CO), non-methane volatile organic compounds (NMVOCs), and aerosol precursors, 1960 emissions are prescribed throughout, allowing the role of meteorological variability in influencing tropospheric composition to be established.
- SEN-C2-fEmis: this is similar to REF-C2 but with surface and aircraft emissions fixed to their respective 1960 levels.
- SEN-C1-Emis/SEN-C1SD-Emis: in these experiments the recommended emission dataset is replaced with an emission dataset of the modellers’ choice, to assess the impact of alternative emissions on tropospheric composition.
- SEN-C2-RCP: this is the same as REF-C2, but with the GHG scenario changed to either RCP 2.6, 4.5, or RCP 8.5. The simulations start in 2000. This means these scenarios differ in their assumptions on GHGs and surface emissions (but not regarding the ODSs). They also require adequate sea surface and sea ice conditions corresponding to the variant climate scenarios, for the atmosphere-only models.

- SEN-C2-GeoMIP: these simulations link CCMI-1 with the Geoengineering Model Intercomparison Project (GeoMIP; Tilmes, 2015a). They are designed to test the impact of proposed efforts to actively manage the Earth's radiation budget to offset the impact of increasing GHGs using sulfur injections.
- SEN-C1-SSI: this is the same as REF-C1 but using a solar forcing dataset with increased UV intensity. (Krivova et al., 2006). (SSI stands for “spectral solar irradiance”.)
- SEN-C2-SolarTrend: this experiment will assess the impact of a possible reduction of solar activity akin to the Maunder Minimum of the 17th and 18th centuries. It is anticipated that the Sun will move out of the recent grand maximum; this would perhaps counteract some anticipated global warming. More details on this experiment are at http://sparcsolaris.gfz-potsdam.de/input_data.php.

4 Forcings used in the reference simulations

Eyring et al. (2013a) and Hegglin et al. (2016) provide full details of the forcings to be used in the above listed CCMI-1 simulations. Here we only comment on selected aspects.

4.1 Greenhouse gases

Most simulations use historical and/or RCP 6.0 mixing ratios for GHGs (Fig. 1a). These are characterized by continuing increases of carbon dioxide (CO_2), which more than doubles between 1960 and 2100. However, the rate of increase reduces at the end of this period. The nitrous oxide (N_2O) surface volume-mixing ratio (VMR) also increases continuously from around 290 ppbv in 1960 to over 400 ppbv in 2100. The methane (CH_4) VMR increases during the 20th century, plateaus between around 2000 and 2030, and subsequently shows a renewed increase, a maximum in around 2070, and then a decrease in the last few decades of the 21st century. In comparison to the REF-B2 simulations of CCMVal-2, CO_2 and N_2O follow similar projections, but CH_4 has a considerably reduced maximum, which also occurs later in the 21st century (SPARC, 2010, Fig. 2.3).

4.2 Ozone-depleting substances (ODSs)

ODSs develop according to the A1 scenario of WMO (2011) (Fig. 1b). There are no major differences with respect to the scenario used by CCMVal-2 (SPARC, 2010, Fig. 2.3). The A1 scenario does not take into account the revised lifetimes of ODSs as documented in SPARC (2013). Test simulations with a scenario based on these revised lifetimes indicate that there would be no significant impact on ozone (WMO, 2015); hence, the recommendation for ODSs has

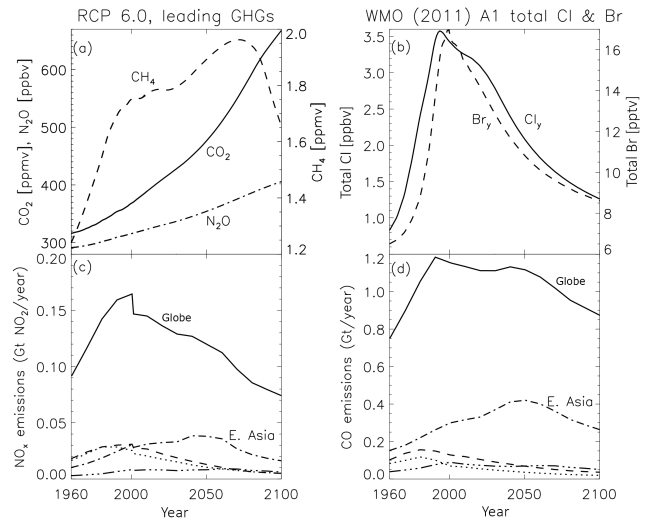


Figure 1. Selected forcings used in the REF-C2 simulations. (a) Carbon dioxide (CO_2 ; solid), methane (CH_4 ; dashed), and nitrous oxide (N_2O ; dash-dotted) surface mass-mixing ratios, following RCP 6.0 (Meinshausen et al., 2011). (b) Total chlorine (Cl; solid) and total bromine (Br) excluding the mass-mixing ratios of the VSLs (dashed; scenario A1 of WMO, 2011). (c) Total nitrogen oxide (NO_x) emissions. Solid: global; dashed: Europe; dotted: North America; dash-dotted: East Asia; dash-dot-dot-dotted: South Asia. (d) Same but for carbon monoxide (CO).

remained unchanged. In addition to long-lived ODSs, modellers are recommended to include CH_2Br_2 and CHBr_3 as bromine source gases. Both are classified as very short-lived species (VSLs). Surface-mixing ratios for both are fixed at 1.2 pptv (giving a total of 6 pptv of bromine). Considering losses of both species in the troposphere, they are meant to deliver the ~ 5 pptv of inorganic bromine to the stratosphere that is thought to originate as VSLs (WMO, 2015).

4.3 Anthropogenic tropospheric ozone and aerosol precursors

For the REF-C2 scenario, for anthropogenic emissions the recommendation is to use MACCity (Granier et al., 2011) until 2000, followed by RCP 6.0 emissions. Figure 1c, d show globally and regionally integrated emissions of nitrogen oxides (NO_x) and carbon monoxide (CO). Globally, efforts to improve air quality, introduced during the late 20th century, cause past and projected future NO_x and CO emissions to peak and then decline. This is clearly seen in European and North American emissions (Fig. 2), but East and South Asian emissions are anticipated to continue to increase – East Asian emissions of NO_x only peak in around 2050 and remain substantial, compared to year-2000 emissions, until the end of this century. Notably, for NO_x there is a discontinuity in 2000 caused by differences in the assumptions on ship emissions between MACCity and RCP 6.0 (Fig. 2). For REF-C1, the MACCity emissions are used throughout the whole 1960–

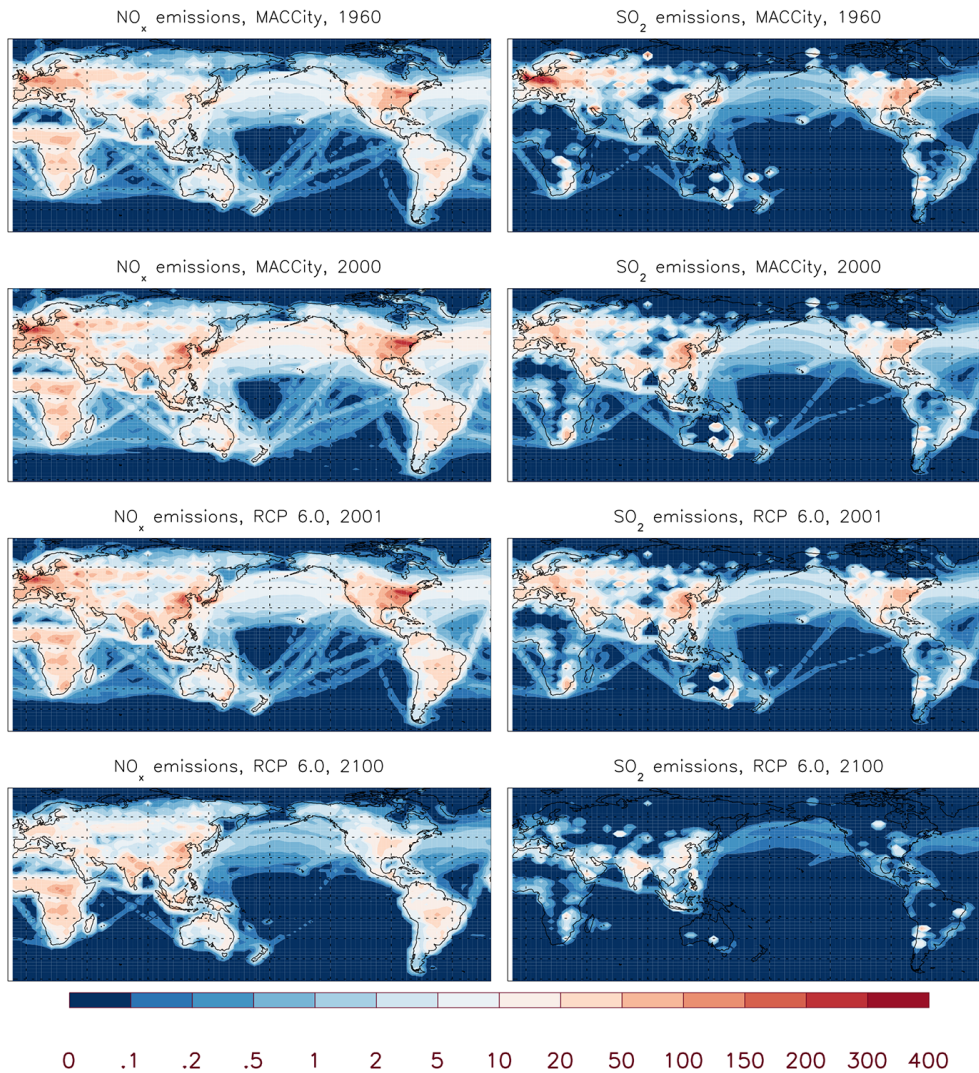


Figure 2. (left) Annual-mean NO_x surface emissions ($10^{-12} \text{ kg (NO}_2\text{) m}^{-2} \text{ s}^{-1}$) in 1960 and 2000, taken from the MACCity emissions database, and in 2001 and 2100, taken from RCP 6.0. (right) Same, but for SO_2 in ($10^{-12} \text{ kg (S) m}^{-2} \text{ s}^{-1}$). Displayed here are the combined surface and stack-height emissions, excluding volcanic emissions.

2010 period covered. For SO_2 , between 1960 and 2000 European (and to a lesser extent North American) emissions drop considerably. This reflects efforts to improve air quality. This trend is anticipated to continue throughout the 21st century. By contrast, East Asian SO_2 emissions increase somewhat during the 20th century. Asia dominates global industrial SO_2 in 2100, in the RCP 6.0 scenario. There is no discontinuity between MACCity and RCP 6.0 for SO_2 emissions.

Whether or not non-represented NMVOCs are lumped with represented ones can result in differences in the actual amount of NMVOCs entering the troposphere (Sect. 2.10).

4.4 Biogenic emissions

For natural (biogenic) emissions, the CCMI-1 recommendation is to use interactive emissions, where available. The

extent to which interactive schemes are used, however, is very species and model dependent, resulting in some diversity of choices regarding biogenic emissions. In the case of soil nitrogen oxide (NO_x) emissions, the majority of models uses prescribed emissions, with the exceptions of EMAC and GEOSCCM which both use the Yienger and Levy (1995) emissions scheme (Table S22). For oceanic dimethyl sulfide (DMS) emissions, most models use interactive emissions schemes with some commonality in the choice of scheme (e.g. Wanninkhof, 1992; Chin et al., 2002), particularly within model families, although a small number of models also use prescribed emissions for oceanic DMS. For biogenic acetone ($(\text{CH}_3)_2\text{CO}$) emissions, all but the CESM1 models either exclude $(\text{CH}_3)_2\text{CO}$ or use prescribed emissions. CESM1 CAM4-chem and CESM1 WACCM, however, use the Model for Emissions of Gases and Aerosols from

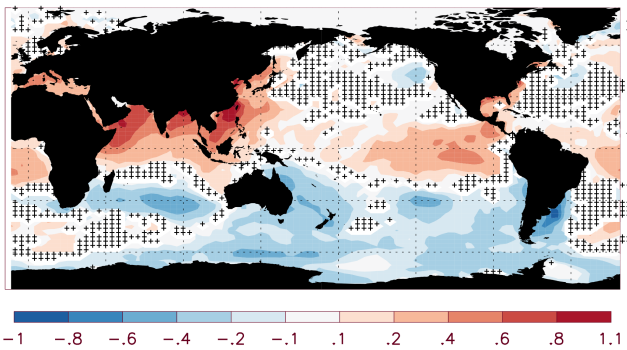


Figure 3. Trend in the annual-mean sea surface temperatures (SSTs), in HadISST for 1960–2010 (K century^{-1}). Trends insignificant at 95 % confidence are stippled.

Nature (MEGAN) 2.1 interactive scheme (<http://lar.wsu.edu/megan>). For ethane (C_2H_6), the CESM1 models also use MEGAN2.1, whereas all other models either exclude C_2H_6 altogether or prescribe emissions. For isoprene (C_5H_8) emissions, about half of the models prescribe emissions and for those that use interactive terrestrial emissions, MEGAN is the predominant emissions scheme of choice. A small number of models include interactive oceanic C_5H_8 emissions. For species whose emissions are not modelled interactively, a variety of different assumptions have been made.

4.5 Sea surface temperature and sea ice

For the REF-C1 and SEN-C1 experiments, sea surface conditions need to be prescribed. For this, as for CCMVal-2, the HadISST climatology (Rayner et al., 2003) is recommended (Sect. 2.20). Variance correction for this monthly mean climatology is recommended following the AMIP II method (<http://www-pcmdi.llnl.gov/projects/amip/AMIP2EXPDSN/BCS/bcsintro.php>). Between 1960 and 2010 there is some warming in the HadISST dataset over various areas of the ocean, but also widespread cooling of the Southern Ocean (Fig. 3). Arctic annual minimum sea ice extent reduces considerably in this period, whereas Antarctic sea ice expands slightly (Fig. 4). HadISST is heavily based on satellite observations which are non-existent for the early part of the record, meaning that trends derived over this period have to be viewed with caution. For the REF-C2 and SEN-C2 experiments, either an interactive ocean/sea ice sub-model is used or pre-calculated sea surface conditions derived from a variety of different climate model simulations as detailed in the previous paragraphs (Table S27).

4.6 Stratospheric aerosol loading

Figure 5 shows the aerosol surface area density at 22 km as recommended by CCMI-1 and imposed by most models (Arfeuille et al., 2013). In comparison to the dataset used for CCMVal-2 (Morgenstern et al., 2010), the most signif-

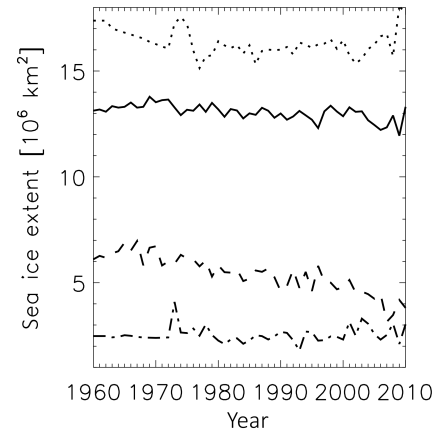


Figure 4. Maximum and minimum monthly mean sea ice extent in the HadISST climatology. Dotted: Antarctic maximum; solid: Arctic maximum; dashed: Arctic minimum; dot-dashed: Antarctic minimum.

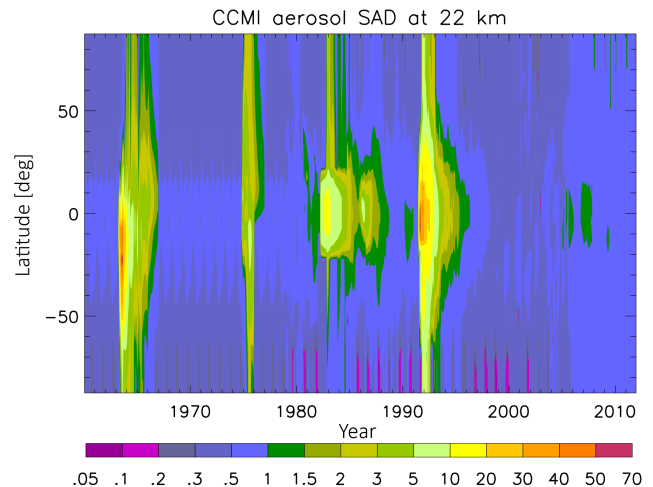


Figure 5. Zonal-mean aerosol surface area density ($\mu\text{m}^2 \text{cm}^{-3}$) at 22 km. The discrete events are due to volcanic eruptions, superimposed on a much smaller non-volcanic background.

icant difference is the insertion of a major volcanic injection of aerosol into the stratosphere in 1974/1975, due to the Fuego (Guatemala) eruption. This had been ignored before. Furthermore, note the increase in aerosol density during the last decade attributed to a series of small volcanic eruptions (Vernier et al., 2011).

4.7 Solar forcing

The recommended solar forcing dataset contains daily solar irradiance, ionization rates by solar protons, and the geomagnetic activity index A_p (<http://solarisheppa.geomar.de/ccmi>). Spectrally resolved solar irradiances for 1960–2010 were calculated with the empirical NRLSSI model (Lean et al., 2005). The spectral grid width (1 nm bins from 0 to 750 nm,

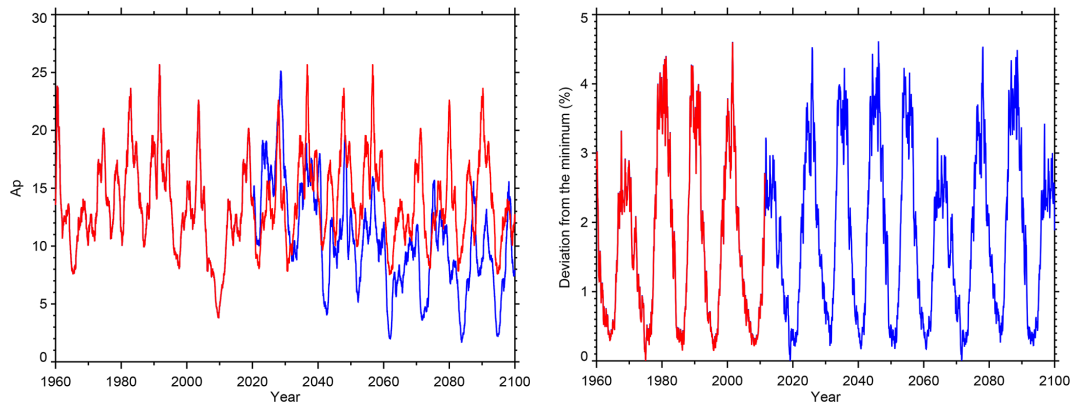


Figure 6. (left) Recommended A_p index time series for SEN-C2_SolarTrend (blue) and the other reference and sensitivity simulations (red). The data are smoothed with 360-day-wide window. (right) Deviation (%) of the recommended monthly mean solar irradiance for the 175–250 nm spectral band from its minimum value, to be used in all simulations. Red: prior to and including 2010, the data are based on observations. Blue: projected solar irradiance post-2010.

5 nm bins from 750 nm to 5 μm , 10 nm bins from 5 to 10 μm , 50 nm bins from 10 to 100 μm) allows for easy calculations of the spectral SSI for any specific model spectral grids, which should be applied to calculate the shortwave heating rates in the radiation module and the photolysis in the chemistry scheme. The ionization rates caused by solar protons for the same time period are calculated using the Jackman et al. (2009) approach based on the proton flux measurements by several instruments onboard the GOES satellites. The recommended coefficients for the conversion of the ionization rates to in situ HO_x and NO_x production intensity are also given by Jackman et al. (2009). For the models extending only to the mesopause, a time-varying geomagnetic activity index A_p is provided as a proxy for the thermospheric NO_x influx, which is used to include indirect energetic particle effects using an approach similar to that defined by Baumgaertner et al. (2009). These datasets are recommended to be applied in REF-C1 simulations covering 1960–2010. For SEN-C1-SSI simulations, the SATIRE SSI dataset (Krivova et al., 2006) should be used instead of the NRLSSI data described above. This SSI dataset exhibits larger UV variability, which can have consequences not only for atmospheric heating but also for ozone chemistry (Ermolli et al., 2013). For REF-C2 simulations covering 2010–2100, it is recommended to repeat the SSI, solar proton event, and A_p sequences of the last four solar activity cycles (i.e. cycles 20–23). For the sensitivity experiment SEN-C2_SolarTrend covering 1960–2100, it is advised to introduce a declining trend in the solar activity, reflecting a widely discussed possible decline of solar activity in the future. The proposed trend is based on past solar activity cycles repeated in reverse order. Starting from 2011 it is recommended to apply daily SSI and particle output for the cycles 20, 19, 18, 17, 16, 15, 14, 13, and 12. The programme to build daily future solar forcing for different experiments is available from <http://solarisheppa.geomar.de/ccmi>. Figure 6 illustrates the A_p index evolution for standard and sensitiv-

ity scenarios, showing a decline of the geomagnetic activity in the future, and the recommended solar irradiance for the 175–250 nm spectral band from its minimum value.

5 Availability of simulations

Tables 4–6 summarize the available simulations at the time of writing this paper. As recommended, a large majority of models have performed the reference simulations. A subset has produced REF-C1SD, reflecting that not all models have the capability to be nudged to meteorological fields. The sensitivity simulations, both SEN-C1 and SEN-C2, are less consistently covered, ranging from 1 to 15 simulations.

Most of the model output can be accessed via the British Atmospheric Data Center (BADC), which hosts the CCMI-1 data archive. Some model institutions provide their data however directly via their local Earth System Grid Federation nodes (see the list provided on http://www.met.reading.ac.uk/ccmi/?page_id=251). CESM1 CAM4-chem and CESM1 WACCM data are provided via <https://www.earthsystemgrid.org/search.html?Project=CCMI1>. In some cases, the simulations are complete but have not or not fully been uploaded for public access. In these cases, readers are advised to contact the corresponding model PIs. In particular, for GFDL AM3/CM3 simulations, please contact Meiyun Lin (meiyun.lin@noaa.gov).

6 Conclusions

The purpose of this paper has been to provide some overview information on the internal make-up of CCMI-1 models, broadly characterize the forcings, and give an overview of available simulations under CCMI-1, mainly to inform authors of other papers focussing on scientific results of CCMI-1. We have not assessed model performance, but it

Table 4. Numbers of reference simulations, by model. Numbers in brackets denote simulations that are incomplete at the time of publication. “L39” stands for the 39-level version of LMDz–REPROBUS.

Model name	REF-C1 (1960–2010)	REF-C2 (1960–2100)	REF-C1SD (1980–2010)
ACCESS CCM	1	2	
CCSRNIES MIROC3.2	3	1	1
CESM1 CAM4-chem	3	3	1 (NASA MERRA)
CESM1 WACCM	5	3	1 (NASA MERRA)
CHASER (MIROC-ESM)	1	1	
CMAM	3	1	1
CNRM-CM5-3	4	2	2
EMAC	2	3	4
GEOSCCM	1	1	1
GFDL-AM3			1 (Lin et al., 2014)
GFDL-CM3		5	
HadGEM3-ES	1	1 (+2)	(2)
LMDz–REPROBUS	1 (L39)	1 (L39)	1 (L39)
MRI-ESM1r1	1	1	1
MOCAGE	1	(1)	1
NIWA-UKCA	3	5	
SOCOL	4	1	
TOMCAT			1
ULAQ CCM	3	3	
UMSLIMCAT	1	1	
UMUKCA-UCAM	1	2	1
Total	39	38	19

Table 5. SEN-C1 sensitivity simulations, by model.

Model name	SEN-C1-Emis	SEN-C1SD-Emis	SEN-C1-fEmis	SEN-C1SD-fEmis	SEN-C1-SSI
CESM1 CAM4-chem				3	
CHASER (MIROC-ESM)				1	
GFDL-AM3	5 (Lin et al., 2014)	1		1	
MOCAGE				1	
NIWA-UKCA			1		
TOMCAT					1
ULAQ CCM			1		3
UMSLIMCAT	1		1		1
Total	6	1	3	6	5

Table 6. SEN-C2 sensitivity simulations, by model.

Model name	RCP2.6	RCP4.5	RCP8.5	fODS	fODS2000	fGHG	fEmis	GeoMIP	SolarTrend	fCH4	fN2O	CH4rcp85
ACCESS CCM				2		1						
CCSRNIES MIROC3.2	1	1	1	1	1	1			1			
CESM1 CAM4-chem								3 (1°)				
CESM1 WACCM		1	3	3	3	3	(1)			(1)	(1)	(1)
CHASER (MIROC-ESM)	1	1	1	1		1	1	(1)				
CMAM	1	1	1	1		1				1	1	(1)
GFDL-CM3	1	3	1									
HadGEM3-ES					(3)	(3)						
LMDz–REPROBUS	1 (L39)	1 (L39)	1 (L39)	1 (L39)	1 (L39)	1 (L39)						
NIWA-UKCA				2		3				(1)	(1)	
SOCOL	1	1	1				1			1	1	1
ULAQ CCM	1	1	1	1	1	1	2	2		(1)	(1)	
Total	7	10	10	12	9	15	5	6	4	5	5	3

is clear from this paper that in the years since CCMVal-2 and ACCMIP, considerable progress has been made to improve the models' internal consistency, make them more physically based, and more comprehensive, as well as improving their resolutions. While these developments have to be welcomed, experience shows that simulations with a more physically consistent and comprehensive model, which is less constrained by external forcings, may not compare more favourably against observations than those produced by a more constrained model (e.g. Eyring et al., 2013b). This is particularly the case as Earth System Models increasingly cover aspects of the climate system that are challenging to capture numerically, such as atmospheric chemistry or biogeochemistry of the ocean. This complicates measuring progress in climate modelling and contributes to the perceived "failure" of the climate modelling community to narrow the range of climate futures produced in multi-model inter-comparisons such as the 5th Coupled Model Intercomparison Project (CMIP5) or CCMI. Understanding how this diversity is linked to differences in model formulation can help explain such findings. The purpose of this paper, and a major motivation for CCMI, is to drive progress in this regard.

7 Code and data availability

Readers should contact the model PIs to enquire about conditions of code availability for the 20 models documented in this paper (Table 1).

No model output was used in this paper. For CCMI-1 model data, see Sect. 5. Forcing data used in this paper are described and can be downloaded at <http://blogs.reading.ac.uk/ccmi/reference-simulations-and-forcings>.

Appendix A: Individual model descriptions

A1 ACCESS-CCM and NIWA-UKCA

NIWA-UKCA is a coupled atmosphere–ocean CCM, based on the HadGEM3-AO model (revision 2) coupled to the NIWA-UKCA gas-phase chemistry scheme. It is identical to ACCESS-CCM, except that ACCESS-CCM uses prescribed sea surface conditions in all simulations. Relative to the UMUKCA models used for CCMVal-2, both models now feature a medium-complexity tropospheric hydrocarbon oxidation scheme, including the Mainz Isoprene Mechanism (Pöschl et al., 2000) and the FAST-JX online photolysis scheme (Telford et al., 2013). NIWA-UKCA uses an interactive ocean and sea ice module (Hewitt et al., 2011). In transitioning to HadGEM3, atmospheric physics was updated; in particular, the models now use the PC2 cloud scheme (Wilson et al., 2008). The models are run at a resolution of N48L60 ($3.75^\circ \times 2.5^\circ$) in the atmosphere and (for NIWA-UKCA) $\sim 2^\circ$ and 31 levels in the ocean.

A2 CCSRNIES MIROC3.2

CCSRNIES MIROC3.2 CCM was constructed on the basis of the MIROC3.2 general circulation model (GCM), which was used for future climate projection in the 4th and 5th Assessment Reports of the Intergovernmental Panel on Climate Change (IPCC, 2007, 2013). The updated CCM introduces the stratospheric chemistry module of CCSRNIES CCM that was used for CCMVal-1 and CCMVal-2. CCSRNIES MIROC3.2 CCM has a new higher resolution radiation scheme for the spectral bins (32 bins) than that of CCSRNIES CCM (18 bins). The new CCM uses a semi-Lagrangian scheme for tracer transport, whilst CCSRNIES CCM used a spectral transport scheme. The new CCM is not coupled to the ocean; sea surface temperature (SST) and sea ice are prescribed in the simulations.

A3 CESM1 CAM4-chem and CESM1 WACCM

The Community Earth System Model, version 1 (CESM1) is a coupled climate model for simulating the Earth's climate system. The atmospheric component is the Community Atmosphere Model, version 4 (CAM4) (Neale et al., 2013), which uses a finite volume dynamical core (Lin, 2004) for the tracer advection. Two versions of CAM4 participated in CCMI-1: (1) a lower lid model reaching up to about 40 km altitude (CESM1 CAM4-chem); (2) and a high-top model that extends to approximately 140 km altitude (Whole Atmosphere Community Climate Model Version 4, CESM1 WACCM4). The horizontal resolution used for all CCMI-1 simulations is $1.9^\circ \times 2.5^\circ$ (latitude \times longitude). Both model versions include detailed and identical representation of tropospheric and stratospheric (TS) chemistry and interactive tropospheric aerosols (Tilmes et al., 2016). The polar heterogeneous chemistry was recently updated (Wegner et

al., 2013) and further evaluated by Solomon et al. (2015). CESM1 WACCM also includes a representation of physics and chemistry of the mesosphere–lower thermosphere (MLT) region (Marsh et al., 2013). The TS (CESM1 CAM4-chem) and TSMLT (CESM1 WACCM4) chemical mechanisms include 171 and 183 species, respectively, contained within the O_x , NO_x , HO_x , ClO_x , BrO_x , and FO_x chemical families, along with CH_4 and its degradation products. In addition, 17 primary non-methane hydrocarbons and related oxygenated organic compounds are included. All CCMI-1 scenarios use the same TS and TSMLT chemical mechanisms. The previous CCMVal-2 version of CESM1-WACCM simulated Southern Hemisphere winter and spring temperatures that were too cold compared with observations. Among other consequences, with the recent updates to the heterogeneous chemistry module, this “cold pole bias” leads to unrealistically low ozone column amounts in Antarctic spring. In all CCMI-1 simulations, the cold pole problem is addressed by introducing additional mechanical forcing of the circulation via parameterized gravity waves (Garcia et al., 2016).

A4 CHASER (MIROC-ESM)

The CHASER model (Sudo et al., 2002; Sudo and Akimoto, 2007), developed mainly at Nagoya University and the Japan Agency for Marine–Earth Science and Technology (JAMSTEC), is a coupled CCM, simulating atmospheric chemistry and aerosols. Aerosols are handled by the SPRINTARS module (Takemura et al., 2005). It has been developed also in the framework of the MIROC Earth System Model, MIROC-ESM-CHEM (Watanabe et al., 2010). CHASER simulates detailed chemistry in the troposphere and stratosphere with an online aerosol simulation including production of particulate nitrate and secondary organic aerosols. For this study, the model's horizontal resolution is selected to be T42 ($2.8^\circ \times 2.8^\circ$) with 57 layers in the vertical extending from the surface up to about 55 km altitude. As for the overall model structure, CHASER is fully coupled with the climate model core MIROC, permitting atmospheric constituents (both gases and aerosols) to interact radiatively and hydrologically with meteorological fields in the model. The chemistry component of CHASER considers the O_x – NO_x – HO_x – CH_4 – CO chemical system with oxidation of NMVOCs, halogen chemistry, and the NH_x – SO_x – NO_3 system. In total 96 chemical species and 287 chemical reactions are considered. In the model, primary NMVOCs include C_2H_6 , C_2H_4 , C_3H_8 , C_3H_6 , C_4H_{10} , acetone, methanol, and biogenic NMVOCs (isoprene, terpenes).

A5 CMAM

Compared with the model version used for CCMVal-2, the CMAM used for the CCMI-1 simulations calculates chemistry throughout the troposphere, though the only hydrocarbon considered is methane. While CMAM was interactively coupled to an ocean model for CCMVal-2, specified SSTs

and sea ice fields were used for all CCMI-1 simulations. The horizontal resolution has increased from T31 to T47 and while spectral advection is still used for all chemical tracers, for HNO_3 and NO_x a logarithmic transformation of the mixing ratio (Scinocca et al., 2008) is advected to better preserve the strong horizontal gradients in the troposphere. For CCMVal-2 a constant dry deposition velocity was used to provide a tropospheric sink for selected species; here wet deposition is calculated interactively with the stratiform/deep convection parameterizations and dry deposition uses a “big-leaf” approach that is tied to the model land surface scheme. The look-up table for photolysis rates has been expanded to take account of surface albedo, and a correction to the clear-sky rates is made for clouds following the approach of Chang et al. (1987). Hydrolysis of N_2O_5 in the troposphere has been included, using a monthly varying climatology of sulfate aerosols from a more recent version of the Canadian climate model (von Salzen et al., 2013) and reaction probabilities of Davis et al. (2008) assuming ammonium sulfate.

A6 CNRM-CM5-3

The CNRM-CM5-3 CCM is based on the CNRM-CM5-3 AOGCM of CNRM/CERFACS, whose version 5.1 has been used in CMIP5 simulations and is described by Voltaire et al. (2012). The CCM includes some fundamental changes from the previous version (CNRM-ACM), which was extensively evaluated in the context of the CCMVal-2 validation activity. The most notable changes concern the radiation code of the GCM (Morcrette, 1990, 1991; Morcrette et al., 2001), the parameterization of non-orographic gravity waves, stochastic parameterization triggered by convection as described by Lott and Guez (2013), and the inclusion of the detailed stratospheric chemistry online within the GCM (Michou et al., 2011). To clarify, CCMI-1 simulations have been performed in an AMIP-type mode, the atmospheric GCM (v6.03) being forced by SSTs and sea ice, without the use of the SURFEX external surface scheme.

A7 EMAC

The Modular Earth Submodel System (MESSy; Jöckel et al., 2005, 2006, 2010) is a software package providing a framework for a standardized, bottom-up implementation of Earth System Models with flexible complexity. “Bottom-up” means, the MESSy software provides an infrastructure with generalized interfaces for the standardized control and interconnection (coupling) of ESM components (dynamic cores, physical parameterizations, chemistry packages, diagnostics, etc.), which are called submodels. MESSy comprises currently about 60 submodels (coded according to the MESSy standards) in different categories: infrastructure (i.e. framework) submodels, atmospheric chemistry related submodels, physics related submodels, and diagnostic submodels. The ECHAM/MESSy Atmospheric Chemistry (EMAC) model

uses the MESSy to link multi-institutional computer codes to the core atmospheric model, i.e. the 5th generation European Centre Hamburg general circulation model (Roeckner et al., 2003, 2006). Updates used for CCMI-1 (EMAC version 2.51) are documented in detail by Jöckel et al. (2016).

A8 GEOSCCM

The Goddard Earth Observing System Chemistry-Climate Model (GEOSCCM) is based on the GEOS-5 GCM (Molod et al., 2012, 2015) coupled to the stratospheric and tropospheric (StratTrop) Global Modeling Initiative (GMI) chemical mechanism (Strahan et al., 2007; Duncan et al., 2007). This version uses a C48 cubed-sphere grid, which has been regridded to 2.5° longitude $\times 2^\circ$ latitude horizontal resolution and 72 vertical layers up to 80 km. The response of tropospheric ozone to variations in the El Niño–Southern Oscillation (ENSO) compared to observations were described by Oman et al. (2011, 2013). An earlier version of the model contributed to the ACCMIP activity (Lamarque et al., 2013).

A9 GFDL-AM3 AND GFDL-CM3

AM3 is the atmospheric component of the Geophysical Fluid Dynamic Laboratory (GFDL) global coupled atmosphere–ocean–land–sea ice model (CM3), which includes interactive stratosphere–troposphere chemistry and aerosols at C48 cubed-sphere horizontal resolution (approximately $2^\circ \times 2^\circ$) (Donner et al., 2011; Austin et al., 2013; Naik et al., 2013). In support of CCMI-1, we conduct a suite of multi-decadal hindcast simulations (1979–2014) designed to isolate the response of atmospheric constituents to historical changes in human-induced emissions, methane, wildfires, and meteorology. Details of these simulations are described by Lin et al. (2014, 2015a, b). We implement a height-dependent nudging technique, relaxing the model to NCEP u and v with a timescale of 6 h in the surface level, but weakening the nudging strength linearly with decreasing pressure (e.g. relaxing with a timescale of 60 h by 100 hPa and 600 h by 10 hPa) (Lin et al., 2012a). To quantify stratospheric influence on tropospheric ozone, we define a stratospheric ozone tracer relative to a dynamically varying tropopause and subjecting it to chemical and depositional loss in the same manner as odd oxygen in the troposphere (Lin et al., 2012b, 2015a). These AM3 simulations have been evaluated against a broad suite of observations. Analysis of satellite measurements, daily ozonesondes, and multi-decadal in situ observation records indicates that the nudged GFDL-AM3 model captures many salient features of observed ozone over the North Pacific and North America, including the influences from Asian pollution events (Lin et al., 2012a), deep tropopause folds (Lin et al., 2012b), as well as their variability on interannual to decadal timescales (Lin et al., 2014, 2015a) and long-term trends (Lin et al., 2015b). The model also captures interannual variability of ozone in the lower stratosphere and its re-

sponses to ENSO events and volcanic aerosols as measured by ozonesondes (Lin et al., 2015a).

A10 HadGEM3-ES

The Met Office model (HadGEM3-ES, formerly UMUKCA-METO) has changed significantly since CCMVal-2. The underlying atmosphere model is now HadGEM3 (Walters et al., 2014), with horizontal resolution increased from 3.75° longitude \times 2.5° latitude to 1.875° longitude \times 1.25° latitude, and the number of levels spanning the model domain 0–85 km increased from 60 to 85. The move to the HadGEM3 model has significantly reduced two critical biases seen in UMUKCA-METO simulations in which stratospheric air was too old and the tropical tropopause too warm (Morgenstern et al., 2009). As a consequence of the improvements in tropical tropopause temperatures, water vapour concentrations entering the stratosphere are no longer prescribed and are now interactively determined by the model. For the scenario simulations coupled ocean (NEMO vn3.4; Madec, 2008) and sea ice (CICE vn4.1; Hunke and Lipscombe, 2008) modules are now included. Significant developments to the UKCA chemistry component (Morgenstern et al., 2009; O'Connor et al., 2014) include the replacement of the stratosphere-only scheme used in UMUKCA-METO with a combined stratosphere–troposphere chemistry scheme, with increased numbers of tracers, chemical species and reactions, the Mainz Isoprene Scheme (MIM, Pöschl et al., 2000), interactive lightning emissions (O'Connor et al., 2014), interactive photolysis rates (FAST-JX; Telford et al., 2013), the CLASSIC aerosol scheme (Bellouin et al., 2011), and a resistance-type approach to dry deposition (Wesely, 1989; O'Connor et al., 2014).

The model, in the HadGEM2 predecessor version, participated in ACCMIP. Relative to this configuration, changes include improved vertical resolution and range, a whole-atmosphere chemistry scheme with expanded NMVOC chemistry, and online photolysis, as detailed above.

A11 LMDz–REPROBUS

LMDz–REPROBUS is a coupled CCM, formed by the coupling of the LMD GCM and the REPROBUS atmospheric chemistry module. When linked to the NEMO ocean model, the configuration is identical to the IPSL atmosphere–ocean climate model but with atmospheric chemistry. The LMDZrepro model used for CCMVal-2 had 50 levels and a resolution of 2.5° latitude \times 3.75° longitude with a top at about 65 km. The CCM1-1 simulations already completed have been performed with the CMIP5 version (LMDz–REPROBUS-CM5) that have 39 levels and a resolution of 2.5° latitude \times 3.75° longitude with a top at about 70 km (Dufresne et al., 2013). We plan to rerun the same CCM1-1 simulations with the CM6 version that has 79 levels and

a resolution of 1.25° latitude \times 2.5° longitude with a top at about 80 km.

A12 MOCAGE

MOCAGE (Modèle de Chimie Atmosphérique de Grande Echelle) is a Météo-France's CTM. MOCAGE combines the RACM (Stockwell et al., 1997) tropospheric and the REPROBUS (Lefèvre et al., 1994) stratospheric chemistry schemes, consistently applied from the surface to the model top. It simulates 109 gaseous species (there are no aerosols in these CCM1-1 runs) that are grouped in families, with 91 being transported. In the stratosphere, nine heterogeneous reactions are described, using the parameterization of Carslaw et al. (1995a). Moreover, 52 photolysis and 312 thermal reactions are considered. The photolysis rates follow look-up tables and are modified to account for cloudiness, following Chang et al. (1987). The model includes a reaction pathway for $\text{HO}_2 + \text{NO}$ to yield HNO_3 (Butkovskaya et al., 2007).

The resolution of the model is $2^\circ \times 2^\circ$ on a latitude–longitude grid, with 47 levels to 5 hPa. For the REF-C1SD and SEN-C1SD-fEmis experiments, we use ERA-Interim forcings. For the REF-C1 and REF-C2 experiments, the meteorological forcing is taken from an update of the CNRM-CM model, which was used for CMIP5 simulations (Voldoire et al., 2012). However, convective transport of species is recomputed following the parameterization by Bechtold et al. (2001). Convective in-cloud scavenging is determined in the updraft (Mari et al., 2000), whereas wet deposition due to stratiform precipitations follow Giorgi and Chameides (1986). A 1-year simulation has been performed to compute dry deposition velocities following the Wesely (1989) approach. Values have been averaged to get monthly diurnal profiles. The same values have been used for all simulations.

Except for lightning NO_x , natural emissions are monthly mean distributions taken from Global Emissions Initiative (GEIA) inventories. Lightning NO_x is parameterized in the convection scheme following Price et al. (1997) and is hence climate-sensitive. Methane concentrations were prescribed at the surface following a monthly zonal climatology taking the evolution of the global value as a function of RCPs into account.

A13 MRI-ESM1r1

MRI-ESM1r1 is an updated version of the Earth System Model MRI-ESM1, which was used for future climate projection in the 5th Assessment Report of the Intergovernmental Panel on Climate Change (IPCC, 2013). The vertical resolution of MRI-ESM1r1 (L80) is improved compared to MRI-ESM1 (L48). The SCUP coupler (Yoshimura and Yukimoto, 2008) is used to couple the atmosphere, ocean, aerosol, and (gas-phase) chemistry modules, which make up MRI-ESM1r1. The chemistry module is MRI-CCM2 (Deushi and Shibata, 2011), which is an updated version of MRI-CCM1

used for CCMVal-2. In MRI-CCM2, a tropospheric hydrocarbon oxidation scheme of medium complexity is newly added.

A14 SOCOLv3

Since CCMVal-2, the SOCOL model (Solar Climate Ozone Links; Stenke et al., 2013) has significantly changed. SOCOLv2, which participated in CCMVal-2, was a combination of the GCM MA-ECHAM4 (Manzini et al., 1997) and the CTM MEZON (Rozanov et al., 1999; Egorova et al., 2003), while the third and current version, SOCOLv3, is based on MA-ECHAM5 (Roeckner et al., 2006; Manzini et al., 2006). The advection of chemical trace species is now calculated by the flux-form semi-Lagrangian scheme of Lin and Rood (1996) instead of the previously applied hybrid advection scheme. This change made the mass correction applied to certain tracers in SOCOLv2 obsolete. Furthermore, the unsatisfying separation of tropospheric and stratospheric water vapour fields in SOCOLv2 has also become obsolete. SOCOLv3 considers only one water vapour field, i.e. the ECHAM5 water vapour. Advection, convection, and the tropospheric hydrological cycle are calculated by the GCM, while chemical water vapour production/destruction as well as PSC formation are calculated by the chemistry module.

For CCMI-1 SOCOL was run with T42 horizontal resolution, which corresponds approximately to $2.8^\circ \times 2.8^\circ$, and with 39 vertical levels between the Earth's surface and 0.01 hPa (~ 80 km). Further important modifications for the CCMI-1 set-up include an isoprene oxidation mechanism (Pöschl et al., 2000), the online calculation of lightning NO_x emissions (Price and Rind, 1992), treatment of the effects produced by different energetic particles (Rozanov et al., 2012), updated reaction rates and absorption cross sections (Sander et al., 2011b), improved solar heating rates (Sukhodolov et al., 2014), as well as a parameterization of cloud effects on photolysis rates (Chang et al., 1987). Furthermore, the ODS species are no longer transported as families, but as separate tracers.

A15 TOMCAT

TOMCAT is a global 3-D offline chemical transport model (Chipperfield, 2006). The model is usually forced by ECMWF meteorological (re)analyses, although GCM output can also be used. When using ECMWF fields, as in the CCMI-1 experiments, the model reads in the 6-hourly fields of temperature, humidity, vorticity, divergence, and surface pressure. The resolved vertical motion is calculated online from the vorticity. The model has parameterizations for sub-grid-scale tracer transport by convection (Stockwell and Chipperfield, 1999; Feng et al., 2011) and boundary layer mixing (Holtslag and Boville, 1993). Tracer advection is performed using the conservation of second order moments scheme by Prather (1986). The CTM can be used with a

variety of chemistry and aerosol schemes including stratospheric chemistry (Chipperfield et al., 2015), tropospheric chemistry (e.g. Monks et al., 2012) and idealized tracers. For the CCMI-1 experiments, the model was run at horizontal resolution of $2.8^\circ \times 2.8^\circ$ with 60 levels from the surface to ~ 60 km. Experiments with stratospheric chemistry and idealized tracers were performed.

A16 ULAQ-CCM

The ULAQ-CCM is a climate–chemistry coupled model with an interactive aerosol module (a compact description was given by Morgenstern et al., 2010, for CCMVal-2). Since then, the following updates have been made to the model (Pitari et al., 2014): (a) increase in horizontal and vertical resolution; (b) inclusion of a numerical code for the formation of upper tropospheric cirrus cloud ice particles (Kärcher and Lohmann, 2002; Pitari et al., 2015a); (c) upgrade of the radiative transfer code for calculations of photolysis, heating rates, and radiative forcing. This is a two-stream δ -Eddington approximation operating online in the ULAQ-CCM, used for photolysis rate calculation at UV–visible (Vis) wavelengths, solar heating rates, and radiative forcing at UV–Vis–NIR (near-infrared) bands (Randles et al., 2013; Pitari et al., 2015b). In addition, a companion broadband, k -distribution longwave radiation module is used to compute radiative transfer and heating rates in the planetary infrared spectrum (Chou et al., 2001; Pitari et al., 2015c). Calculations of photolysis rates and radiative fluxes have been evaluated in the framework of CCMVal (SPARC, 2010) and AeroCom inter-comparison campaigns (Randles et al., 2013). The chemistry–aerosol module is organized with all medium and short-lived species grouped in families. It includes the major components of tropospheric aerosols (sulfate, carbonaceous aerosol, soil dust, sea salt), with calculation at each size bin of surface fluxes, removal and transport terms, in external mixing conditions. A modal approximation is used for nitrate aerosols. Wet and dry deposition is treated following Müller and Brasseur (1995), using a climatological cloud distribution. Lower-stratospheric denitrification and dehydration are calculated using the predicted size distribution of PSC particles.

A17 UMSLIMCAT

The UMSLIMCAT has only undergone minor changes since CCMVal-2. The model is based on a old version of the MetUM and, although it performs well in terms of stratospheric chemistry and dynamics, the model is not actively developed. Core UMSLIMCAT simulations are performed in order to increase the range of simulations available and to provide some continuity with previous CCM studies. Compared to CCMVal-2, the minor model updates are (i) updated photolysis scheme with an improved treatment of ozone profiles in the online look-up table, (ii) the use of the CCMI-1

aerosol surface area density (SAD), and (iii) an updated solar flux representation. Dhomse et al. (2011, 2013, 2015) describe the implementation of this representation and present an analysis of solar flux variability and volcanic aerosol in the model.

A18 UMUKCA-UCAM

UMUKCA-UCAM is an atmosphere-only CCM, based on the HadGEM3 model (revision 2). The chemistry in UMUKCA-UCAM is based on a similar scheme as was used in the UMUKCA models in CCMVal-2 (focusing on the chemistry of stratosphere; Bednarz et al., 2016), but with an explicit treatment of halogen source gases, i.e. no lumping. Since CCMVal-2, significant improvements to the model physics have been made, and although the model resolution is degraded to run at N48L60 ($3.75^\circ \times 2.5^\circ$) in the atmosphere, the model physics is identical to HadGEM3. Relative to the UMUKCA models used for CCMVal-2, the FAST-JX online photolysis scheme (Telford et al., 2013) is now included, as are interactive lightning emissions, the CLASSIC aerosol scheme (Bellouin et al., 2011), and a resistance-type approach to dry deposition (Wesely, 1989).

Appendix B: Deviations from CCMI-1 recommendations

We list here the ways in which simulations and model setups deviate from Eyring et al. (2013a). Furthermore, simulations submitted to the archive that are additional to those solicited by Eyring et al. (2013a) and Hegglin et al. (2016) are described here. Errors with CCMI-1 models or simulations that come to light after publication of this paper will be documented at <https://blogs.reading.ac.uk/ccmi/badc-data-access/data-errata-and-notes/>.

B1 ACCESS CCM and NIWA-UKCA

For some simulations, anthropogenic NMVOC emissions were held at their 1960 levels for 1960–2000 in about half of the NIWA-UKCA simulations. This error was picked up and corrected for the later simulations but remains in earlier simulations. Simulations affected by this problem include: REF-C1 (r2, r3), REF-C2 (r1, r2, r3, r4), SEN-C2-fODS (r1), and SEN-C2-fGHG (r1). Not affected are REF-C1 (r1), SEN-C1-fEMIS (r1), REF-C2 (r5), SEN-C2-fODS (r2), SEN-C2-fGHG (r2, r3), SEN-C2-fCH4 (r1), and SEN-C2-fN2O (r1).

As noted before, ACCESS CCM and NIWA-UKCA do not consider the radiative impacts of stratospheric aerosol. Also there is no variance correction applied to sea surface temperatures in the simulations without interactive ocean.

B2 CCSRNIES MIROC3.2

HadISST1 data were used for REF-C1 and REF-C1SD simulations. Chemical reactions important in the troposphere are not included, but the stratospheric chemistry scheme is just used in the troposphere. Solar radiation at wavelengths shorter than 177.5 nm is not considered except for Lyman- α . Atmospheric ionization by solar protons is not included.

B3 CMAM

The ACCMIP historical database of emissions (Lamarque et al., 2010) was used for the REF-C1 and REF-C1SD simulations up to the year 2000, with the RCP8.5 emissions used for the following years. It was also used up to 2000 for the REF-C2 and associated scenario simulations. Emissions at intermediate years were linearly interpolated from the years given in the database. An additional emission of CO of 250 Tg(CO) year⁻¹ was included to account for CO from isoprene oxidation, with the emissions distributed following the monthly emissions of isoprene from Guenther et al. (1995). No variance correction was applied to the specified SSTs.

B4 EMAC

Due to a unit conversion error at data import, the extinction of stratospheric aerosols was too low, by a factor of approximately 500. The effect of stratospheric background aerosol on radiative heating rates has been tested by sensitivity simulations and estimated to be smaller than the interannual standard deviation. However, the dynamical effects of large volcanic eruptions (Mt. Pinatubo in 1991, El Chichón in 1982, etc.) are essentially not represented in the simulations, except for the contribution to the tropospheric temperature signal induced by the prescribed SSTs. The chemical effects (through heterogeneous chemistry), however, are included, since the prescribed aerosol surface areas were treated correctly.

Next, due to an error in the model set-up, the timing of the road traffic emissions was unfortunately wrong; instead of updating the monthly input fields every month, they have been updated only every year, and thus in 1950 emissions of January 1950 have been used, in 1951 the emissions of February 1950, etc.

And last, but not least, some of the diagnostic tracers have been treated differently, as detailed by Jöckel et al. (2016).

More details on the deviations from the CCMI-1 recommendations are documented by Jöckel et al. (2016, see their Sect. 3.12 and Table A1).

B5 GFDL-AM3

The AM3_BASE simulation (i.e. REF-C1SD) applies interannually varying emissions of aerosol and ozone precursors from human activity, based on Lamarque et al. (2010) for 1980–2000 and RCP 8.5 projections (Riahi et al., 2011)

Table B1. Summary of forcings and emissions data used in the GFDL-AM3 hindcast simulations, with italics indicating where the data differ from the CCMI-1 recommendations. L2014 = Lin et al. (2014).

Experiment	L2014 name	Meteorology	Period	RF	CH ₄	Anth. emissions	Fire emissions
REF-C1SD	BASE	NCEP <i>u</i> & <i>v</i>	1980–2010	REF-C1	REF-C1	REF-C1 (Except SO ₂ , BC, and OC after 1996)	REF C1 (RETRO before 1996, GFEDv3 for 1997–2010)
SEN-C1SD-fEmis	FIXEMIS	as REF-C1SD	1980-2010	REF-C1	2000*	1970–2010 climatology*	1970–2010 climatology*
SEN-C1SD-Emis	IAVFIRE	as REF-C1SD	1980-2010	CCMI-1	2000*	1970–2010 climatology*	REF-C1
SEN-C1-Emis	AMIP	N/A	1960–2010	REF-C1	2000*	<i>O₃precursors: FIXEMIS;</i> <i>aerosol precursors: REF-C1</i>	1970–2010 climatology

beyond 2005, linearly interpolated for intermediate years. The AM3_FIXEMIS simulation (i.e. SEN-C1SD-fEmis), with anthropogenic and biomass burning emissions set to the 1970–2010 climatology and methane held constant at 2000 levels, is designed to isolate the role of meteorology. The IAVFIRE simulation (i.e. SEN-C1SD-Emis) applies interannual-varying monthly mean emissions from biomass burning based on Schultz et al. (2008) for 1970–1996 and GFEDv3 (Van der Werf et al., 2010) for 1997–2010. Otherwise, all forcings are the same as in FIXEMIS. The BASE, FIXEMIS, and IAVFIRE simulations with modified emissions are nudged to NCEP reanalysis winds over 1980 to 2010. We also conduct four ensemble simulations without nudging, driven by prescribed sea surface temperatures (SSTs) and atmospheric radiative forcing agents over 1960 to 2010 (SEN-C1-Emis; Table B1). In SEN-C1SD-fEmis, emissions of ozone and aerosol precursors are fixed to the 1970–2010 climatology, instead of the 1980 levels recommended by CCMI-1. Due to an error in data processing, anthropogenic emissions of aerosol precursors (SO₂, BC, and OC) after 1996 in the REF-C1SD simulation do not follow the CCMI-1 recommendation. Otherwise denoted in Table B1, all forcings follow the CCMI-1 recommendations.

B6 HadGEM3-ES

The specified-dynamics simulation (REF-C1SD) uses an anomaly correction to the ERA-I forcing data, as outlined in McLandress et al. (2014). Two REF-C2 simulations only start in 2000. The three SEN-C2-fGHG simulations are forced with fixed year-2000 GHG-mixing ratios not 1960 ones, and also only start in 2000.

B7 MRI-ESM1r1

The molecular weight of sulfate aerosols (SO₄) due to volcanic eruptions was inappropriately set to that of sulfur atom (S) in our REF-C1 and REF-C1SD simulations. As a result, the amount of volcanic aerosol in these simulations was one-third of its correct amount. Molecular weights of other aerosols (anthropogenic, biogenic, dust, etc.) are appropriately treated.

B8 SOCOLv3

Sea surface temperatures for REF-C2 and all sensitivity simulations based on REF-C2 were taken from the CESM1-CAM5 model.

B9 ULAQ CCM

All CCMI-1 experiments have been conducted following the CCMI-1 recommendations. For the sensitivity cases SEN-C2-fGHG, SEN-C2-fODS, and SEN-C2-fODS2000, the following procedure has been used for CH₄, N₂O, and CFCs, which are both GHGs and ODSs. These species were fixed in the radiation–dynamics–climate modules in the fGHG experiment, leaving them to evolve in time for chemistry. The opposite choice was made for the two fODS experiments (1960, 2000), i.e. fixing these species in chemistry and letting them evolve in the radiation–dynamics–climate modules.

B10 UMUKCA-UCAM

The stratospheric aerosol climatology used is SPARC (2006), and is included in the chemistry, photolysis, and radiation schemes. Surface emissions (of NO_x, CO, and HCHO) and the NO_x aircraft emissions are the same as used in the CCM-Val2 REF-B2 simulation.

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