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Extreme geomagnetic reversal frequency during the Middle Cambrian as
 revealed by the magnetostratigraphy of the Khorbusuonka section
 (northeastern Siberia)

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#### 11 Abstract

12 We present new magnetostratigraphic results obtained for the Drumian stage (504.5-13 500.5 Ma; Epoch 3/Middle Cambrian) from the Khorbusuonka sedimentary section in 14 northeastern Siberia. They complement previous data that did not allow the 15 determination of a reliable estimate of the geomagnetic reversal frequency during this time. Magnetization of the samples is carried by a mixture of magnetite and hematite 16 17 in various proportions. Thermal demagnetization makes it possible to distinguish two 18 magnetization components. The low unblocking temperature (<350°C) component 19 has a steep inclination and likely originates from remagnetization in a recent field. At 20 higher temperatures, the magnetization isolated possesses the two polarities. Its 21 direction is usually well determined; however, for a noticeable set of samples, a 22 strong overlap between the demagnetization spectra of the two components prevents 23 the determination of reliable directions, although their polarities are well established. 24 The directions from 437 samples define a sequence of 78 magnetic polarity intervals,

25 22 of which are observed in a single sample. Biostratigraphic data available from the 26 Khorbusuonka section indicate that the duration of the studied section is ~3 Myr. A 27 geomagnetic reversal frequency of 26 reversals per Myr is therefore estimated for the 28 Drumian, reduced to 15 reversals per Myr if only the polarity intervals defined by at 29 least two consecutive samples are retained. This is an extreme reversal rate, similar 30 to that obtained for the Late Ediacaran (late Precambrian), ~50 Myr earlier, which has been related to a late nucleation of the inner core. The reversal frequency 31 32 appears to have drastically dropped for ~3-4 Myr from a value probably >20 reversals per Myr during the Drumian to ~1.5 reversals per Myr during the Furongian/Upper 33 Cambrian. Such a sharp decrease is consistent with a transition at a ~1-Myr 34 35 timescale, probably caused by threshold effects in core processes, between two 36 geodynamo modes, one characterized by reversals occurring at frequencies ranging 37 from 1 to 5 reversals per Myr, and the other marked by hyperactivity of the reversing 38 process, with reversal rates >15 reversals per Myr.

39

*Keywords*: Magnetostratigraphy, Drumian, Middle Cambrian, Magnetic reversal
 frequency, extreme reversal rate, geodynamo

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#### 43 **1. Introduction**

The frequency at which the polarity of the geomagnetic field reverses has motivated many studies to focus on the acquisition of new data to scrutinize variations in frequency over time (e.g. Pavlov and Gallet, 2005), on the statistical analysis of polarity reversal sequences (e.g. McFadden, 1984), as well as on the origin of temporal variations in reversal frequency in relation to core and mantle

49 dynamics (e.g. Besse and Courtillot, 1986; Driscoll and Olson, 2011; Biggin et al., 50 2012; Hounslow et al., 2018). Two operational modes have been proposed for the geodynamo (McFadden and Merrill, 1995): (i) a regime without geomagnetic polarity 51 52 reversals over several tens of millions of years, producing events called superchrons; 53 and (ii) a regime marked by polarity reversals occurring at various frequencies over time. For the Phanerozoic Eon, covering the last ~550 Myr, the marine magnetic 54 anomalies available since the Middle Jurassic, together with magnetostratigraphic 55 56 data obtained from sedimentary outcrops, have revealed the existence of three superchrons spaced ~150 Myr apart (e.g. Pavlov and Gallet, 2005; Biggin et al., 57 58 2012). Between these superchrons, the geomagnetic reversal frequency varied from 59 low values of ~1 reversal per Myr to high values of ~5 reversals per Myr (e.g. McFadden and Merrill, 2000). 60

61 The analysis of changes in magnetic reversal frequency has led to many 62 proposals or hypotheses. This has particularly been the case for the last 100 Myr, for 63 which the geomagnetic polarity timescale is likely the best established. Nevertheless, different evolutionary models have been suggested, advocating either a more or less 64 65 complex, non-stationary reversing process since ~83 Ma ago, that is since the end of 66 the normal-polarity Cretaceous superchron (e.g. Mazaud et al., 1983; Gallet and Courtillot, 1995; McFadden and Merrill, 2000 and references therein), a stationary 67 68 process marked by a jump-like evolution of the mean durations of magnetic polarity intervals estimated over tens of Myr-long segments (Lowrie and Kent, 2004), or a 69 70 combination of these two processes (Gallet and Hulot, 1997; Hulot and Gallet, 2003). 71 This indeterminacy in the nature of the modulation of the process that causes 72 geomagnetic reversals is problematic for deciphering the interactions between 73 mantle and core dynamics.

74 More recently, Gallet and Pavlov (2016) proposed a third geodynamo mode 75 characterized by a hyper reversal frequency, with frequencies reaching more than 10-15 reversals per Myr (note that this range of values is much higher than the one 76 77 previously used by Meert et al. (2016) to define the hyperactive reversal regime, i.e. 78 > 4-6 reversals per Myr). This regime would be distinct from the one characterized by 79 frequencies ranging from ~1 to ~5 reversals per Myr, referred to as the 'normal' reversing mode by Gallet and Pavlov (2016). This idea was based on the observation 80 81 of two periods marked by reversal rates that were much higher than those known 82 from, for instance, the Miocene, being of the order of ~4-5 reversals per Myr, and so 83 far considered to be high. The first episode of hyper reversal frequency concerns the 84 Middle Jurassic. This has been proposed from analysis of the oldest marine magnetic anomalies measured in the western Pacific (e.g. Tivey et al., 2006 and references 85 86 therein). According to these authors, the geomagnetic reversal frequency around 160 87 Ma achieved values as high as ~12-14 reversals per Myr during a period of a few 88 Myr, although this duration remains poorly established. A second episode was dated 89 from the end of the Precambrian, around 550-560 Ma ago. Mainly defined from 90 magnetostratigraphic data obtained from Baltica and its margins, and from Siberia 91 (Popov et al., 2005; Shatsillo et al., 2015; Bazhenov et al., 2016; see also Halls et al., 92 2015 for data obtained from Canada/Laurentia), this latter is characterized by an 93 extreme reversal frequency of probably >20 reversals per Myr (Bazhenov et al., 94 2016). Note that this was based on a rough estimate of sedimentation rates in the sections studied. 95

The hypothesis of a distinct hyperactivity mode of the reversing process would make it possible to account, on one hand, for the fact that the reversal frequency values observed for the two episodes mentioned above are in no way comparable to

99 those of the other periods and, on the other hand, for the probable existence of long 100 stationary segments in the evolution of the duration of the magnetic polarity intervals 101 (Gallet and Hulot, 1997; Hulot and Gallet, 2003; Lowrie and Kent, 2004). Gallet and 102 Pavlov (2016) also suggested that transitions between the three operational 103 geodynamo modes may have occurred on a ~1-Myr timescale, in accordance with 104 rapid core dynamics, whereas their occurrence and time distribution would depend 105 on the thermal conditions prevailing at the core-mantle boundary, driven by mantle 106 dynamics. In this way, this scenario would be consistent with the widely accepted 107 idea that the long-term evolution of magnetic reversal frequency has been modulated 108 by slow (~100-Myr timescale) mantle convection (e.g. McFadden and Merrill, 1984; 109 Courtillot and Besse, 1987; Courtillot and Olson, 2007; Driscoll and Olson, 2011; 110 Pétrélis et al., 2011; Biggin and al., 2012; Hounslow et al., 2018).

111 It should be stressed, however, that the data allowing constraint of the highest 112 -or extreme reversal frequency values remain very few, and with poor age control for 113 the most remarkable episode at the end of the Precambrian. For this reason, we 114 resampled the Middle Cambrian Khorbusuonka section located in northeastern 115 Siberia, which had previously been studied by Gallet et al. (2003). The latter study 116 revealed guite a large number of polarity reversals, defining a frequency of ~6 to ~8 117 reversals per Myr; however, given the large number of magnetic polarity intervals 118 present in this section, the sampling that had been carried out at the time was clearly 119 insufficient to determine a reliable reversal frequency value, and so only a minimum 120 estimate was possible. Thanks in particular to recent investigations (Korovnikov and 121 Tokarev, 2018), this biostratigraphically-precisely-dated section was therefore a 122 prime target for providing a firm constraint on a very high magnetic reversal 123 frequency. In fact, below, we demonstrate that it is extreme.

124 It is also worth pointing out that this new information should be placed in a 125 more general geophysical perspective, including considering the possible effect of 126 inner-core nucleation on geodynamo behavior (e.g. Driscoll, 2016; Landeau et al., 127 2017; Lhuillier et al, 2019; Bono et al., 2019), bearing in mind that the beginning of 128 this nucleation could have been as recent as ~600-700 Ma ago (e.g. Olson et al., 129 2013: Labrosse, 2015: Driscoll, 2016), or the interactions between Earth's different 130 envelopes (e.g. Kirschvink et al., 1997; Meert et al., 2016). The Cambrian was 131 indeed punctuated by several geological events as major as the explosion in faunal diversity (e.g. Kirschvink et al., 1997), attested to in Canada by the Burgess shales 132 133 dated as Middle Cambrian, or the possible occurrence, in the Lower Cambrian, of an 134 episode of rapid and major true polar wander (Kirschvink et al., 1997; Mitchell et al., 135 2010).

136

#### 137 **2. Age and new sampling of the Khorbusuonka section**

138 The section, named Khorbusuonka in the present study, is located along the 139 left bank of the Khorbusuonka River, about 170 km southwest of the city of Tiksi 140 ( $\lambda$ =71° 28' 47" N;  $\varphi$ =123° 49' 57" E; northeastern Siberian platform; Fig. 1). Several Lower and Middle Cambrian sections have been studied along this river, involving 141 142 the identification of five successive sedimentary formations, named, from the oldest 143 to the most recent, Kessyusa (Late Ediacaran-Lower Cambrian), Erkeket (Lower to 144 Middle Cambrian), and Kuonamka, Yunkyulyabit-Yuryakh and Tyuessala, all Middle 145 Cambrian (see description in Korovnikov, 2002; Korovnikov and Tokarev, 2018). 146 Note that Lower and Middle Cambrian refer to Epochs 2 and 3, respectively (e.g. 147 Ogg et al., 2016). In the present study, the magnetostratigraphic investigations were 148 focused on the Yunkyulyabit-Yuryakh Formation and on the lower to basal part of the

Tyuessala Formation (Fig. 2). Paleomagnetic sampling was carried out on the section referred to as Section #2 by Korovnikov and Tokarev (2018), being also the same section as that previously studied by Gallet et al. (2003). This contains a complete record of the Yunkyulyabit-Yuryakh Formation over a thickness of ~74 m. The ~85-m-thick part of the section that was sampled consists of alternation of flatlying limestones and marls of varying colors -grayish, reddish, sometimes greenish (Fig. 2).

156 The dating of the Yunkyulyabit-Yuryakh and Tyuessala formations in the 157 studied section was based on the discovery of trilobites and brachiopods along the 158 entire sedimentary thickness, as well as in the other sections of the Khorbusuonka 159 river. A detailed (and recent) description of the findings has been given in Korovnikov and Tokarev (2018) (see also Gallet et al., 2003 and references therein). Three 160 161 regional trilobite zones have been documented - Tomagnostus fissus over a thickness 162 of ~9 m, Corynexochus perforatus-Anopolenus henrici over a thickness of ~22 m and 163 Anomacarioides limbataeformis over a thickness of ~54 m. These zones mark the 164 upper part of the Amgan stage (the first 9 m of the section) and the lower part of the 165 Mayan stage (the remaining part of the section; Fig. 2) in the Siberian system 166 (e.g. Krasnov et al., 1983; Korovnikov and Tokarev, 2018). It is worth mentioning that 167 the thickness of the Anomacarioides limbataeformis interval in the section 168 investigated corresponds to less than one-third of that of the entire trilobite zone in the area (Korovnikov and Tokarev, 2018). 169

A question has arisen concerning the correlation between the Siberian biostratigraphic scale and the international Geological Time Scale. The situation seems to now have been resolved, whereas it was not at the time of Gallet et al. (2003). The studied section is unambiguously dated as Drumian (ex-Stage 6), as

174 recognized in the international system and is part of Epoch 3 of the Cambrian (=Middle Cambrian; e.g. Babcock et al., 2007; Howley and Jiang, 2010 and 175 176 references therein). The beginning of this stage corresponds to the base of 177 *Tomagnostus fissus* because these trilobites appear on the Siberian Platform at the 178 same stratigraphic level as the species called *Ptychagnostus atavus*, which is a key 179 species for defining the lower limit of the Drumian stage (e.g. Babcock et al., 2007; 180 Korovnikov and Tokarev, 2018 and references therein). In the Khorbusuonka section, 181 the base of *Tomagnostus fissus* has been determined in the thin, ~2.5-m-thick 182 Kuonamka Formation, composed of dark gray to black limestones and marls, 183 conformably underlying the Yunkyulyabit-Yuryakh Formation (Korovnikov and 184 Tokarev, 2018). This indicates that the basal part of the Drumian stage is absent 185 from the section that was studied. Similarly, the upper to terminal part of the Drumian 186 stage is absent because the *Lejopyge laevigata* trilobite zone is also absent from the 187 section (the base of this species marks the top of the Drumian).

According to the most recent version of the Geological Time Scale (e.g. Ogg et al., 2016), the Drumian stage lasted 4 Myr, between 504.5 and 500.5 Ma. The Khorbusuonka section, therefore, has a duration of less than 4 Myr, probably of the order of ~3 Myr, although it is difficult to estimate precisely the extent of the missing portions of the *Tomagnostus fissus* and *Anomacarioides limbataeformis* trilobite zones.

During the summer of 2016, we collected a time-series of ~550 hand-samples through the Yunkyulyabit-Yuryakh and Tyuessala formations, with stratigraphic intervals of about 10 to 20 cm between the samples. For comparison, the magnetostratigraphic sequence obtained by Gallet et al. (2003) for the same part of Khorbusuonka Section #2 relied on 119 samples. It should be noted that to the steep

slope (Fig. 1) and the risk of falling rocks made the sampling conditions quite difficult.
As a result, samples from the two field works, conducted 15 years apart, could not be
located with sufficient accuracy relative to each other to allow us to combine the two
data sets into a single one.

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#### 204 **3. Paleomagnetic Results**

The analyses were carried out in the paleomagnetism laboratories at the Institut de Physique du Globe de Paris and the Institute of Physics of the Earth (Russian Academy of Sciences) in Moscow. In both cases, magnetization measurements were performed using a 2G cryogenic magnetometer installed in a magnetically shielded room. Not surprisingly, the paleomagnetic behavior observed in this study shared all the characteristics previously identified by Gallet et al. (2003) from a smaller collection of samples.

212 The samples were thermally demagnetized using more than 15 temperature 213 steps. These demagnetizations allowed the isolation of two magnetization 214 components (Fig. 3). A first low unblocking temperature component was isolated at 215 up to 300-350°C. This component shows a steep inclination and clearly has the 216 direction of the present day field at the site (Table 1; Fig. 3, 4a). The characteristic 217 magnetization component was then isolated up to the Curie temperature of magnetite 218 for a set of samples (Fig. 3c,f), and up to the Curie temperature of hematite for the 219 others (Fig. 3a-b,d-e). The magnetization appeared to be carried by a mixture of 220 magnetite and hematite in different proportions, with the same magnetic polarity 221 recorded by the two magnetic phases. This combination of magnetic carriers, evident 222 from the thermal demagnetization of the natural remanent magnetization (NRM; Fig.

3), was confirmed by the thermal demagnetization of three-axis isothermal remanent
magnetization (IRM) acquired in three orthogonal directions (Fig. 5; Lowrie, 1990).

225 The high unblocking temperature magnetization component possessed the two 226 magnetic polarities (Fig. 3, 4b). As also shown by Gallet et al. (2003), many samples 227 showed a strong overlap between the demagnetization spectra of the two 228 magnetization components between ~350°C and ~400-500°C, which made it difficult to accurately estimate the direction of the characteristic component for a significant 229 230 number (~20%) of samples (Fig. 6). In these cases, however, it was possible to 231 determine an approximate direction, which, although biased by the remagnetization 232 (secondary) component, gave clear information about its magnetic polarity. These 233 directions were used to establish the magnetostratigraphic sequence (Fig. 7), but not 234 to estimate a mean direction for each of the two magnetic polarity states (Fig. 4b). 235 The 317 directions that were best determined had maximum angular deviation (MAD) 236 most generally <7.5° (~90% in this category) and <5° for ~67% of the samples in this 237 same category. Despite this selection (after the elimination of 18 deviant directions 238 likely associated with geomagnetic polarity transitions; Fig. S1), the two mean 239 directions of the different polarities were not antipodal at the 95% confidence level 240 ( $\gamma$ =11.3° for  $\gamma$ c=4.0°; Table 1; McFadden and McElhinny, 1991), likely because of a 241 residual contamination by the remagnetization component (Fig. 4c). Nevertheless, 242 the general mean direction obtained for the high-temperature magnetization 243 component (Table 1) was identical to the mean direction that was previously 244 estimated by Gallet et al. (2003) from a much smaller dataset ( $\gamma$ =2.8° for  $\gamma$ c=3.4°; 245 McFadden and McElhinny, 1991). It is worth noting that the data from 114 samples 246 were rejected mainly because only the secondary magnetization component in the 247 present-day field direction was recorded or their magnetization directions were too

scattered upon heating (Fig. S1a-d). This relatively high number, which represents ~20% of the total collection, is probably related to the very large number of magnetic reversals observed in the studied section (see below). However, a significant number of rejected samples were found in the Tyuessala Formation, showing a lithology generally less favorable for paleomagnetic analyses than that of the Yunkyulyabit-Yuryakh Formation.

254 The two elements that concern, on one hand, the choice of the magnetic polarity of the directions obtained, and on the other hand, the primary nature of the 255 256 characteristic magnetization isolated at Khorbusuonka are worth discussing (see also 257 discussion in Gallet et al., 2003). Numerous paleomagnetic results have shown that 258 the Siberian Platform was located in the Southern Hemisphere during the Cambrian, 259 rotated by 180° from its current orientation; it then crossed the Equator towards the 260 end of the Middle Ordovician (e.g. Khramov, 1987; Gallet and Pavlov, 1996; Pavlov 261 and Gallet, 2005; Torsvik et al., 2012). As a consequence, paleomagnetic directions 262 with a negative (resp. positive) inclination and a declination of ~160° (resp. ~340°) 263 have a normal (resp. reverse) polarity (Fig. 4b).

264 The primary nature of the characteristic magnetization was deduced from the 265 following evidence: (i) whether predominantly carried by magnetite or hematite, the 266 directions of the magnetization component are similar and the geomagnetic polarity 267 does not depend on the predominant magnetic phase in the samples. This showed 268 that the hematite likely has an early diagenetic (or detritic?) origin; (ii) most of the 18 269 abnormal (or deviating) directions isolated through the section are observed at the 270 transition between two magnetic polarity intervals (Fig. S1e-f). These most likely 271 reflect a transitional configuration of the geomagnetic field during a polarity reversal; 272 (iii) the mean direction estimated for the Khorbusuonka section (note that the

273 averaging tended to reduce the bias due to the recent magnetic field) is almost 274 identical at the 95% confidence level to the tilt-corrected mean direction (and after its 275 transfer to the location of the Khorbusuonka section) previously obtained by Pavlov 276 and Gallet (2001) for the Middle Cambrian part of the Kulumbe section, located nearly 1500 km away ( $\gamma$ =3.7°;  $\gamma$ c=3.5°; McFadden and McElhinny, 1991). In contrast 277 278 the two mean directions are significantly different before bedding correction ( $\gamma$ =16.0°; 279  $\gamma_{c}=3.6^{\circ}$ ; McFadden and McElhinny, 1991), which gives a positive, albeit partial, fold 280 test. It is also worth mentioning that the Kulumbe section is slightly younger than the 281 Khorbusuonka section, since the lower part of the Kulumbe section has been dated 282 as being at the top of the Drumian Anomacarioides limbataeformis trilobite zone (e.g. 283 Pavlov and Gallet, 2005; Kouchinsky et al., 2008); and (iv) the paleomagnetic pole 284 defined from the characteristic component isolated at Khorbusuonka is different from 285 all other paleomagnetic poles of younger age known for the Siberian Platform (e.g. 286 Khramov, 1987; Torsvik et al., 2012).

287

288 **4. Discussion and conclusions** 

289 The results described above indicate that the Khorbusunka section contains a 290 record of the geomagnetic field that prevailed during the Drumian stage (Epoch 291 3/Middle Cambrian). The new data offer an average resolution gain of 3.7 compared 292 to the results of Gallet et al. (2003) (i.e. the ratio between the 437 new samples and 293 the 119 samples used by Gallet et al., 2003). This gain varied according to the 294 trilobite zones, from ~3.4 for the Tomagnostus fissus zone, to ~2.9 for the 295 Corynexochus perforatus-Anopolenus henrici and ~4.1 for the Anomacarioides 296 *limbataeformis* zone. They show the occurrence of 78 magnetic polarity intervals 297 along the 85 m of the studied section, of which 22 intervals were defined by only one

298 sample (Fig. 7). An extreme, but probably simplistic, approach would be to eliminate 299 the latter intervals by considering them too uncertain, which would then imply a 300 sequence of 46 magnetic polarity intervals; however, there is no a priori reason why 301 the directions of the samples concerned should be considered to be less well defined 302 than the others. The most likely reason for the intervals defined by a single sample is 303 that they were very short and the sampling rate was still too low compared to the 304 prevailing geomagnetic reversal frequency. For the same stratigraphic sequence, 305 Gallet et al. (2003) found 42 magnetic polarity intervals, with 18 intervals defined by a 306 single sample, or a series of 18 intervals after the exclusion of the latter. An important 307 observation is that, despite the very significant difference in the number of polarity 308 intervals found between these two studies, a fairly convincing correlation can be 309 observed between the two magnetostratigraphic sequences (Fig. 8).

310 A comparison is reported in Fig. 8 between the Khorbusuonka results and 311 those found in a section sampled along the Kulumbe river (Pavlov and Gallet, 2001). 312 This section was dated by the presence of the Anomacarioides limbataeformis and 313 Lejopyge laevigata trilobite zones: that is, it covers the top of the Drumian and the 314 lower part of the Guzhangian stage (ex-stage 7; e.g. Ogg et al., 2016). In the 315 Kulumbe section, there was a strong contrast between the Anomacarioides 316 *limbataeformis* zone, marked by numerous geomagnetic reversals, and the *Lejopyge* 317 laevigata zone, with only a few geomagnetic reversals detected. It was not possible 318 to achieve a convincing correlation between the Khorbusuonka and Kulumbe 319 sequences (Fig. 8). A first option would be that this difficulty primarily results from the 320 large number of polarity reversals observed in both Drumian sections, a difficulty 321 further amplified by possible sampling rate effects and/or due to differences in 322 sedimentation rate between the two sections. The second would be that the Kulumbe

section is slightly younger than the Khorbusuonka section (see also Pavlov and
Gallet, 2005). The most important point, however, is that, whatever the option
retained, the data from Kulumbe confirm the presence of numerous magnetic polarity
reversals during the *Anomacarioides limbataeformis* zone of the Drumian stage.

327 Based on the chronostratigraphic timescale considered for the Cambrian (e.g. 328 Ogg et al., 2016), and on the biostratigraphic data available for the Middle Cambrian 329 deposits found along the Khorbusuonka river, the duration of the Khorbusuonka 330 section analyzed in the present study would have a duration of ~3 Myr. This relatively 331 short duration implies a geomagnetic reversal frequency as high as ~26 reversals per 332 Myr, if all polarity intervals are considered, or 15 reversals per Myr if the intervals 333 defined by a single sample are excluded. Even considering a misinterpretation of 334 some of the magnetic polarity intervals defined by a single sample, it follows that the 335 geomagnetic reversal frequency during the Drumian stage was likely >20 reversals 336 per Myr. This is obviously an extreme reversal rate, if we refer, for example, to the 337 evolution of the geomagnetic reversal frequencies over the last 100 million years. In 338 addition, the new paleomagnetic data show that, during this period of hyperactivity of 339 the reversing process, the geometry of the geomagnetic field probably remained 340 essentially dipolar, with the slight non-antipodality of the two directions of opposite 341 polarity observed at Khorbusuonka being likely due to a residual contamination by the recent field (Fig. 4c). 342

The extreme geomagnetic reversal frequency observed at Khorbusuonka appears very similar to that found by Bazhenov et al. (2016) for the Zigan Formation in the southern Urals (Russia), dated to the late Precambrian (Late Ediacaran), some 50-60 Myr before the Drumian. Furthermore, as in Khorbusuonka, the Late Ediacaran data also showed that a dipolar field likely prevailed at that time. Beyond these

348 exceptional reversal rates, there is the question of their evolution throughout the 349 Cambrian and the geophysical implications that could result from it. 350 Magnetostratigraphic data obtained from Siberia, Australia and China have provided 351 constraints on that evolution after the Drumian; more precisely, during the terminal 352 part of the Middle Cambrian (Guzhangian stage), the Furongian Epoch/Upper 353 Cambrian and during the Ordovician (see synthesis in Pavlov and Gallet, 2005 and 354 Hounslow, 2016). The Kulumbe sequence showed about 30 geomagnetic polarity 355 intervals during the Guzhangian, but it should be noted that this sequence is very 356 preliminary due to the poor quality of the paleomagnetic signal observed in the rocks 357 of the Orakta Formation (Kouchinsky et al., 2008). Considering a duration of ~3.5 358 Myr for the Guzhangian stage (Ogg et al., 2016), the reversal frequency would have 359 reached a value of ~10 reversals per Myr at the end of the Middle Cambrian. In 360 contrast, the reversal frequency would have been much lower during the Furongian 361 Epoch/Late Cambrian (~1.5 reversal per Myr), with ~15 to 20 magnetic polarity 362 intervals over a duration of ~12 Myr (Pavlov and Gallet, 2005; Kouchinsky et al., 363 2008; Ogg et al., 2016; Hounslow, 2016). A similar low reversal rate would also have 364 prevailed at the beginning of the Lower Ordovician, just before the onset of the 365 reverse-polarity Moyero Superchron (Fig. 9; Gallet and Pavlov, 1996; Pavlov and 366 Gallet, 1998; 2005; Pavlov et al., 2017). Therefore, the presently available 367 magnetostratigraphic data show that the magnetic reversal frequency dramatically 368 dropped from values >20 reversals per Myr to ~1.5 reversal per Myr in ~3-4 Myr. 369 which is the duration of the Guzhangian stage (with an 'intermediate', albeit high, 370 value of ~10 reversals per Myr during the Guzhangian).

The new magnetostratigraphic data, and those previously obtained by Bazhenov et al. (2016), next raise the question of the evolution in reversal frequency

373 during the Lower Cambrian. Was the period between ~550 Ma and ~500 Ma 374 characterized by a single or more episodes of hyperactivity in the reversing process? 375 Note further that the onset of the hyperactivity episode in the Late Ediacaran has not 376 yet been documented. To date, the magnetostratigraphic data available for the Lower 377 Cambrian are neither numerous nor constrained enough to answer this question with certainty. Even though the data mainly obtained from Siberia (Kirschvink and 378 379 Rozanov, 1984; Kirschvink et al., 1991; Gallet et al., 2003) and China (Duan et al., 380 2018 and references therein) clearly showed the occurrence of many geomagnetic 381 polarity reversals during Epoch 2/the Lower Cambrian, it remains difficult to precisely 382 determine the reversal rates. Duan et al. (2018) recently proposed a frequency of ~7 383 reversals per Myr between ~524 Ma and 514 Ma, but the stratigraphic sequence 384 from which the data were obtained was fragmentary; it therefore seems possible (but 385 not proven) that the reversal frequency was actually higher during this period. The 386 Lower Cambrian paleomagnetic data obtained at Khorbusuonka by Gallet et al. 387 (2003) also attest to a very high geomagnetic reversal frequency during the 388 Tommotian and Atdabanian (Gallet et al., 2003 tentatively proposed a reversal 389 frequency even higher than that of the Drumian), in addition to the occurrence of a 390 major true polar wander event near the end of the Lower Cambrian, in agreement 391 with Kirschvink et al. (1997). The data, however, should be confirmed by further 392 studies on other sedimentary sections in Siberia or elsewhere. At this stage, it is only 393 possible to affirm that the reversal frequency was high (>5 reversals per Myr) during 394 the Lower Cambrian, but the hyperactivity character (>15 reversals per Myr) cannot 395 be established (Fig. 9).

396 Comparing dipole field behavior in numerical simulations with varying inner-397 core size and assuming inner-core nucleation began during the Neoproterozoic,

398 Lhuillier et al. (2019) recently suggested that the hyperactivity of the reversing 399 process during the latest Precambrian and the Cambrian could be linked to a critical 400 size of the inner core, around an aspect ratio of ~0.2. During this critical period, 401 corresponding to significant changes in the flow pattern in the tangent cylinder, the 402 reversal frequency would have been very sensitive to variations in the vigor of the 403 convection in the outer core. If these simulations do indeed provide relevant 404 information for the geodynamo, this sensitivity could account for long duration of the 405 hyperactivity phenomenon, if it persisted for about 50 Myr, or recurrence of such 406 events, if several hyperactivity episodes were present, a very different situation from 407 that observed during the Middle Jurassic (Lhuillier et al., 2019). It should be noted, 408 however, that just changing the thermal conditions at the core-mantle boundary can 409 alone account for the full spectrum of temporal variations in geomagnetic reversal 410 rate (e.g. Driscoll and Olson, 2011; Olson et al., 2013). On another hand, Bono et al. 411 (2019) showed that the hyperactivity episode at the end of the Precambrian was also 412 accompanied by very low mean paleointensity. This observation would be consistent 413 with a weak-field geodynamo state before inner-core nucleation as predicted by 414 numerical models (Driscoll, 2016; Landeau et al., 2017; see also Bono et al., 2019). 415 Accurate paleointensity data are not yet available to determine whether this state still 416 prevailed in the Middle Cambrian. An important point of note, however, is that both 417 the data from Ediacaran (Bazhenov et al., 2016) and the Middle Cambrian seem to 418 argue for a dipole-dominated field geometry at the beginning of the inner-core growth. 419

The Cambrian thus appears to be a period marked by remarkable and fascinating phenomena affecting all terrestrial envelopes. Linking them together exceeds the objectives of our study (see for example Kirschvink et al., 1997; Meert et

423 al., 2016). Here, we have emphasized the exceptional nature of the reversing 424 process during the Middle Cambrian, but the sharp decrease in reversal rate at the 425 end of the Middle Cambrian provides further important information on the geodynamo 426 behavior. It argues for threshold effects in core processes (see also Courtillot and 427 Olson, 2007), which would be consistent with a ~1-Myr-scale transition between two 428 reversing modes; that is, between the normal and hyperactive reversing modes 429 defined by Gallet and Pavlov (2016) (Fig. 9).

430

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#### 581 **Table and Figures captions**

582

**Table 1.** Paleomagnetic results obtained from the Drumian part of the Khorbusuonka
section. LTC and HTC refer to the magnetization components isolated in the low and
high unblocking temperature ranges.

586

Fig. 1. General geology of northeastern Siberia (a) and of the area in the vicinity of
the Khorbusuonka section (b). The red square in panel (a) and the red star in panel
(b) indicate the location of the studied Khorbusuonka section.

590

Fig. 2. View of the Khorbusuonka section analyzed in the present study (@Yves
Gallet). The simplified lithology, and the dating of the studied sequence, are from
Korovnikov and Tokarev (2018).

594

Fig. 3. Representative examples of the thermal demagnetization of the samples from
the Khorbusuonka section. Solid (resp. open) circles are in the horizontal (resp.
vertical) plane. The corresponding magnetization moment versus temperature curves
are also reported (right panels).

599

Fig. 4. Stereographic projections of the directions obtained in the Khorbusuonka section. Directions at the sample level were isolated for the low (a) and high (b) unblocking temperature components (LTC and HTC, respectively). (c) Mean

directions with opposite polarities. The closed (open) symbols refer to directions in
the lower (upper) hemisphere. All directions were obtained using the PaleoMac
software (Cogné, 2003).

606

Fig. 5. Thermal demagnetization of three-axis isothermal remanent magnetization
components acquired in fields of 1.9 T (dots), 0.3 T (triangles) and 0.15 T (squares).
These examples show that the magnetization of the samples was likely carried by
various mixtures of magnetite and hematite.

611

**Fig. 6.** Examples of thermal demagnetization of samples from the Khorbusuonka section showing a strong overlap between the low and high unblocking temperature magnetization components. In these cases, a direction could not be reliably determined for the high unblocking temperature component (see text). Same conventions as in Fig. 3-4.

617

**Fig. 7.** Magnetostratigraphy of the Drumian sequence of the Khorbusuonka section. The dark green dots show the directions used to estimate a mean high unblocking temperature magnetization direction, whereas those in light green color refer to biased directions, allowing constraint of only the associated magnetic polarity (see text).

623

Fig. 8. Comparison between the magnetostratigraphic sequences of Middle Cambrian age (Epoch 3) obtained from the Khorbusuonka (Gallet et al., 2003; this study) and Kulumbe (Pavlov and Gallet, 2001) sections.

- Fig. 9. Synthesized sketch of the evolution in geomagnetic reversal frequency since
  the end of the Precambrian (modified from Gallet et Pavlov, 2016).
- 630

#### 631 Supplementary material

**Fig. S1.** (a-d) Thermal demagnetization of rejected samples. In each case, the directional data are reported in orthogonal and stereographic projections, together with the evolution of magnetization as a function of temperature (see text). Same conventions as in Fig. 3 and 4. (e-f) Two examples of demagnetization behavior for samples located at a geomagnetic polarity transition.











Figure 5



Figure 6







Kulumbe (Pavlov et Gallet, 2001)



Figure 9

Khorbusuonka section $\lambda = 71^{\circ}28'47''$ N	Ν	In situ	Tilt corrected	Paleomagnetic pole
Φ=123°49'57"E		Dec° Inc° K $\alpha_{95}$	$Dec^\circ$ Inc $^\circ$ K $lpha_{95}$	Lat° Long° dp°/dm
LTC	515	349.5 79.4 27.5 1.2°	349.6 79.4 27.5 1.2	
HTC-Normal	131	163.1 -35.2 16.4 3.1	163.1 -35.3 16.4 3.1	
HTC-Reverse	168	347.4 46.1 19.3 2.5	347.4 46.1 19.3 2.5	
HTC N+R	299	165.3 -41.4 16.6 2.1	165.4 -41.4 16.6 2.1	-41.6 141.8 1.6/2.6

Table 1