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1 **Extreme geomagnetic reversal frequency during the Middle Cambrian as**  
2 **revealed by the magnetostratigraphy of the Khorbusuonka section**  
3 **(northeastern Siberia)**

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10

11 **Abstract**

12 We present new magnetostratigraphic results obtained for the Drumian stage (504.5-  
13 500.5 Ma; Epoch 3/Middle Cambrian) from the Khorbusuonka sedimentary section in  
14 northeastern Siberia. They complement previous data that did not allow the  
15 determination of a reliable estimate of the geomagnetic reversal frequency during this  
16 time. Magnetization of the samples is carried by a mixture of magnetite and hematite  
17 in various proportions. Thermal demagnetization makes it possible to distinguish two  
18 magnetization components. The low unblocking temperature (<350°C) component  
19 has a steep inclination and likely originates from remagnetization in a recent field. At  
20 higher temperatures, the magnetization isolated possesses the two polarities. Its  
21 direction is usually well determined; however, for a noticeable set of samples, a  
22 strong overlap between the demagnetization spectra of the two components prevents  
23 the determination of reliable directions, although their polarities are well established.  
24 The directions from 437 samples define a sequence of 78 magnetic polarity intervals,

25 22 of which are observed in a single sample. Biostratigraphic data available from the  
26 Khorbusuonka section indicate that the duration of the studied section is ~3 Myr. A  
27 geomagnetic reversal frequency of 26 reversals per Myr is therefore estimated for the  
28 Drumian, reduced to 15 reversals per Myr if only the polarity intervals defined by at  
29 least two consecutive samples are retained. This is an extreme reversal rate, similar  
30 to that obtained for the Late Ediacaran (late Precambrian), ~50 Myr earlier, which  
31 has been related to a late nucleation of the inner core. The reversal frequency  
32 appears to have drastically dropped for ~3-4 Myr from a value probably >20 reversals  
33 per Myr during the Drumian to ~1.5 reversals per Myr during the Furongian/Upper  
34 Cambrian. Such a sharp decrease is consistent with a transition at a ~1-Myr  
35 timescale, probably caused by threshold effects in core processes, between two  
36 geodynamo modes, one characterized by reversals occurring at frequencies ranging  
37 from 1 to 5 reversals per Myr, and the other marked by hyperactivity of the reversing  
38 process, with reversal rates >15 reversals per Myr.

39

40 *Keywords* : Magnetostratigraphy, Drumian, Middle Cambrian, Magnetic reversal  
41 frequency, extreme reversal rate, geodynamo

42

### 43 **1. Introduction**

44 The frequency at which the polarity of the geomagnetic field reverses has  
45 motivated many studies to focus on the acquisition of new data to scrutinize  
46 variations in frequency over time (e.g. Pavlov and Gallet, 2005), on the statistical  
47 analysis of polarity reversal sequences (e.g. McFadden, 1984), as well as on the  
48 origin of temporal variations in reversal frequency in relation to core and mantle

49 dynamics (e.g. Besse and Courtillot, 1986; Driscoll and Olson, 2011; Biggin et al.,  
50 2012; Hounslow et al., 2018). Two operational modes have been proposed for the  
51 geodynamo (McFadden and Merrill, 1995): (i) a regime without geomagnetic polarity  
52 reversals over several tens of millions of years, producing events called superchrons;  
53 and (ii) a regime marked by polarity reversals occurring at various frequencies over  
54 time. For the Phanerozoic Eon, covering the last ~550 Myr, the marine magnetic  
55 anomalies available since the Middle Jurassic, together with magnetostratigraphic  
56 data obtained from sedimentary outcrops, have revealed the existence of three  
57 superchrons spaced ~150 Myr apart (e.g. Pavlov and Gallet, 2005; Biggin et al.,  
58 2012). Between these superchrons, the geomagnetic reversal frequency varied from  
59 low values of ~1 reversal per Myr to high values of ~5 reversals per Myr (e. g.  
60 McFadden and Merrill, 2000).

61         The analysis of changes in magnetic reversal frequency has led to many  
62 proposals or hypotheses. This has particularly been the case for the last 100 Myr, for  
63 which the geomagnetic polarity timescale is likely the best established. Nevertheless,  
64 different evolutionary models have been suggested, advocating either a more or less  
65 complex, non-stationary reversing process since ~83 Ma ago, that is since the end of  
66 the normal-polarity Cretaceous superchron (e.g. Mazaud et al., 1983; Gallet and  
67 Courtillot, 1995; McFadden and Merrill, 2000 and references therein), a stationary  
68 process marked by a jump-like evolution of the mean durations of magnetic polarity  
69 intervals estimated over tens of Myr-long segments (Lowrie and Kent, 2004), or a  
70 combination of these two processes (Gallet and Hulot, 1997; Hulot and Gallet, 2003).  
71 This indeterminacy in the nature of the modulation of the process that causes  
72 geomagnetic reversals is problematic for deciphering the interactions between  
73 mantle and core dynamics.



74 More recently, Gallet and Pavlov (2016) proposed a third geodynamo mode  
75 characterized by a hyper reversal frequency, with frequencies reaching more than  
76 10-15 reversals per Myr (note that this range of values is much higher than the one  
77 previously used by Meert et al. (2016) to define the hyperactive reversal regime, i.e.  
78 > 4-6 reversals per Myr). This regime would be distinct from the one characterized by  
79 frequencies ranging from ~1 to ~5 reversals per Myr, referred to as the 'normal'  
80 reversing mode by Gallet and Pavlov (2016). This idea was based on the observation  
81 of two periods marked by reversal rates that were much higher than those known  
82 from, for instance, the Miocene, being of the order of ~4-5 reversals per Myr, and so  
83 far considered to be high. The first episode of hyper reversal frequency concerns the  
84 Middle Jurassic. This has been proposed from analysis of the oldest marine magnetic  
85 anomalies measured in the western Pacific (e.g. Tivey et al., 2006 and references  
86 therein). According to these authors, the geomagnetic reversal frequency around 160  
87 Ma achieved values as high as ~12-14 reversals per Myr during a period of a few  
88 Myr, although this duration remains poorly established. A second episode was dated  
89 from the end of the Precambrian, around 550-560 Ma ago. Mainly defined from  
90 magnetostratigraphic data obtained from Baltica and its margins, and from Siberia  
91 (Popov et al., 2005; Shatsillo et al., 2015; Bazhenov et al., 2016; see also Halls et al.,  
92 2015 for data obtained from Canada/Laurentia), this latter is characterized by an  
93 extreme reversal frequency of probably >20 reversals per Myr (Bazhenov et al.,  
94 2016). Note that this was based on a rough estimate of sedimentation rates in the  
95 sections studied.

96 The hypothesis of a distinct hyperactivity mode of the reversing process would  
97 make it possible to account, on one hand, for the fact that the reversal frequency  
98 values observed for the two episodes mentioned above are in no way comparable to

99 those of the other periods and, on the other hand, for the probable existence of long  
100 stationary segments in the evolution of the duration of the magnetic polarity intervals  
101 (Gallet and Hulot, 1997; Hulot and Gallet, 2003; Lowrie and Kent, 2004). Gallet and  
102 Pavlov (2016) also suggested that transitions between the three operational  
103 geodynamo modes may have occurred on a ~1-Myr timescale, in accordance with  
104 rapid core dynamics, whereas their occurrence and time distribution would depend  
105 on the thermal conditions prevailing at the core-mantle boundary, driven by mantle  
106 dynamics. In this way, this scenario would be consistent with the widely accepted  
107 idea that the long-term evolution of magnetic reversal frequency has been modulated  
108 by slow (~100-Myr timescale) mantle convection (e.g. McFadden and Merrill, 1984;  
109 Courtillot and Besse, 1987; Courtillot and Olson, 2007; Driscoll and Olson, 2011;  
110 Pétrélis et al., 2011; Biggin and al., 2012; Hounslow et al., 2018).

111         It should be stressed, however, that the data allowing constraint of the highest  
112 -or extreme reversal frequency values remain very few, and with poor age control for  
113 the most remarkable episode at the end of the Precambrian. For this reason, we  
114 resampled the Middle Cambrian Khorbusuonka section located in northeastern  
115 Siberia, which had previously been studied by Gallet et al. (2003). The latter study  
116 revealed quite a large number of polarity reversals, defining a frequency of ~6 to ~8  
117 reversals per Myr; however, given the large number of magnetic polarity intervals  
118 present in this section, the sampling that had been carried out at the time was clearly  
119 insufficient to determine a reliable reversal frequency value, and so only a minimum  
120 estimate was possible. Thanks in particular to recent investigations (Korovnikov and  
121 Tokarev, 2018), this biostratigraphically-precisely-dated section was therefore a  
122 prime target for providing a firm constraint on a very high magnetic reversal  
123 frequency. In fact, below, we demonstrate that it is extreme.

124 It is also worth pointing out that this new information should be placed in a  
125 more general geophysical perspective, including considering the possible effect of  
126 inner-core nucleation on geodynamo behavior (e.g. Driscoll, 2016; Landeau et al.,  
127 2017; Lhuillier et al, 2019; Bono et al., 2019), bearing in mind that the beginning of  
128 this nucleation could have been as recent as ~600-700 Ma ago (e.g. Olson et al.,  
129 2013; Labrosse, 2015; Driscoll, 2016), or the interactions between Earth's different  
130 envelopes (e.g. Kirschvink et al., 1997; Meert et al., 2016). The Cambrian was  
131 indeed punctuated by several geological events as major as the explosion in faunal  
132 diversity (e.g. Kirschvink et al., 1997), attested to in Canada by the Burgess shales  
133 dated as Middle Cambrian, or the possible occurrence, in the Lower Cambrian, of an  
134 episode of rapid and major true polar wander (Kirschvink et al., 1997; Mitchell et al.,  
135 2010).

136

## 137 **2. Age and new sampling of the Khorbusuonka section**

138 The section, named Khorbusuonka in the present study, is located along the  
139 left bank of the Khorbusuonka River, about 170 km southwest of the city of Tiksi  
140 ( $\lambda=71^{\circ} 28' 47''$  N;  $\varphi=123^{\circ} 49' 57''$  E; northeastern Siberian platform; Fig. 1). Several  
141 Lower and Middle Cambrian sections have been studied along this river, involving  
142 the identification of five successive sedimentary formations, named, from the oldest  
143 to the most recent, Kessyusa (Late Ediacaran-Lower Cambrian), Erkeket (Lower to  
144 Middle Cambrian), and Kuonamka, Yunkyulyabit-Yuryakh and Tyuessala, all Middle  
145 Cambrian (see description in Korovnikov, 2002; Korovnikov and Tokarev, 2018).  
146 Note that Lower and Middle Cambrian refer to Epochs 2 and 3, respectively (e.g.  
147 Ogg et al., 2016). In the present study, the magnetostratigraphic investigations were  
148 focused on the Yunkyulyabit-Yuryakh Formation and on the lower to basal part of the

149 Tyuessala Formation (Fig. 2). Paleomagnetic sampling was carried out on the  
150 section referred to as Section #2 by Korovnikov and Tokarev (2018), being also the  
151 same section as that previously studied by Gallet et al. (2003). This contains a  
152 complete record of the Yunkyulyabit-Yuryakh Formation over a thickness of ~74 m.  
153 The ~85-m-thick part of the section that was sampled consists of alternation of flat-  
154 lying limestones and marls of varying colors -grayish, reddish, sometimes greenish  
155 (Fig. 2).

156         The dating of the Yunkyulyabit-Yuryakh and Tyuessala formations in the  
157 studied section was based on the discovery of trilobites and brachiopods along the  
158 entire sedimentary thickness, as well as in the other sections of the Khorbusuonka  
159 river. A detailed (and recent) description of the findings has been given in Korovnikov  
160 and Tokarev (2018) (see also Gallet et al., 2003 and references therein). Three  
161 regional trilobite zones have been documented -*Tomagnostus fissus* over a thickness  
162 of ~9 m, *Corynexochus perforatus-Anopolenus henrici* over a thickness of ~22 m and  
163 *Anomacarioides limbataeformis* over a thickness of ~54 m. These zones mark the  
164 upper part of the Amgan stage (the first 9 m of the section) and the lower part of the  
165 Mayan stage (the remaining part of the section; Fig. 2) in the Siberian system  
166 (e.g. Krasnov et al., 1983; Korovnikov and Tokarev, 2018). It is worth mentioning that  
167 the thickness of the *Anomacarioides limbataeformis* interval in the section  
168 investigated corresponds to less than one-third of that of the entire trilobite zone in  
169 the area (Korovnikov and Tokarev, 2018).

170         A question has arisen concerning the correlation between the Siberian  
171 biostratigraphic scale and the international Geological Time Scale. The situation  
172 seems to now have been resolved, whereas it was not at the time of Gallet et al.  
173 (2003). The studied section is unambiguously dated as Drumian (ex-Stage 6), as

174 recognized in the international system and is part of Epoch 3 of the Cambrian  
175 (=Middle Cambrian; e.g. Babcock et al., 2007; Howley and Jiang, 2010 and  
176 references therein). The beginning of this stage corresponds to the base of  
177 *Tomagnostus fissus* because these trilobites appear on the Siberian Platform at the  
178 same stratigraphic level as the species called *Ptychagnostus atavus*, which is a key  
179 species for defining the lower limit of the Drumian stage (e.g. Babcock et al., 2007;  
180 Korovnikov and Tokarev, 2018 and references therein). In the Khorbusuonka section,  
181 the base of *Tomagnostus fissus* has been determined in the thin, ~2.5-m-thick  
182 Kuonamka Formation, composed of dark gray to black limestones and marls,  
183 conformably underlying the Yunkyulyabit-Yuryakh Formation (Korovnikov and  
184 Tokarev, 2018). This indicates that the basal part of the Drumian stage is absent  
185 from the section that was studied. Similarly, the upper to terminal part of the Drumian  
186 stage is absent because the *Lejopyge laevigata* trilobite zone is also absent from the  
187 section (the base of this species marks the top of the Drumian).

188         According to the most recent version of the Geological Time Scale (e.g. Ogg  
189 et al., 2016), the Drumian stage lasted 4 Myr, between 504.5 and 500.5 Ma. The  
190 Khorbusuonka section, therefore, has a duration of less than 4 Myr, probably of the  
191 order of ~3 Myr, although it is difficult to estimate precisely the extent of the missing  
192 portions of the *Tomagnostus fissus* and *Anomacarioides limbataeformis* trilobite  
193 zones.

194         During the summer of 2016, we collected a time-series of ~550 hand-samples  
195 through the Yunkyulyabit-Yuryakh and Tyuessala formations, with stratigraphic  
196 intervals of about 10 to 20 cm between the samples. For comparison, the  
197 magnetostratigraphic sequence obtained by Gallet et al. (2003) for the same part of  
198 Khorbusuonka Section #2 relied on 119 samples. It should be noted that to the steep

199 slope (Fig. 1) and the risk of falling rocks made the sampling conditions quite difficult.  
200 As a result, samples from the two field works, conducted 15 years apart, could not be  
201 located with sufficient accuracy relative to each other to allow us to combine the two  
202 data sets into a single one.

203

### 204 **3. Paleomagnetic Results**

205 The analyses were carried out in the paleomagnetism laboratories at the  
206 Institut de Physique du Globe de Paris and the Institute of Physics of the Earth  
207 (Russian Academy of Sciences) in Moscow. In both cases, magnetization  
208 measurements were performed using a 2G cryogenic magnetometer installed in a  
209 magnetically shielded room. Not surprisingly, the paleomagnetic behavior observed  
210 in this study shared all the characteristics previously identified by Gallet et al. (2003)  
211 from a smaller collection of samples.

212 The samples were thermally demagnetized using more than 15 temperature  
213 steps. These demagnetizations allowed the isolation of two magnetization  
214 components (Fig. 3). A first low unblocking temperature component was isolated at  
215 up to 300-350°C. This component shows a steep inclination and clearly has the  
216 direction of the present day field at the site (Table 1; Fig. 3, 4a). The characteristic  
217 magnetization component was then isolated up to the Curie temperature of magnetite  
218 for a set of samples (Fig. 3c,f), and up to the Curie temperature of hematite for the  
219 others (Fig. 3a-b,d-e). The magnetization appeared to be carried by a mixture of  
220 magnetite and hematite in different proportions, with the same magnetic polarity  
221 recorded by the two magnetic phases. This combination of magnetic carriers, evident  
222 from the thermal demagnetization of the natural remanent magnetization (NRM; Fig.

223 3), was confirmed by the thermal demagnetization of three-axis isothermal remanent  
224 magnetization (IRM) acquired in three orthogonal directions (Fig. 5; Lowrie, 1990).  
225 The high unblocking temperature magnetization component possessed the two  
226 magnetic polarities (Fig. 3, 4b). As also shown by Gallet et al. (2003), many samples  
227 showed a strong overlap between the demagnetization spectra of the two  
228 magnetization components between  $\sim 350^{\circ}\text{C}$  and  $\sim 400\text{-}500^{\circ}\text{C}$ , which made it difficult  
229 to accurately estimate the direction of the characteristic component for a significant  
230 number ( $\sim 20\%$ ) of samples (Fig. 6). In these cases, however, it was possible to  
231 determine an approximate direction, which, although biased by the remagnetization  
232 (secondary) component, gave clear information about its magnetic polarity. These  
233 directions were used to establish the magnetostratigraphic sequence (Fig. 7), but not  
234 to estimate a mean direction for each of the two magnetic polarity states (Fig. 4b).  
235 The 317 directions that were best determined had maximum angular deviation (MAD)  
236 most generally  $< 7.5^{\circ}$  ( $\sim 90\%$  in this category) and  $< 5^{\circ}$  for  $\sim 67\%$  of the samples in this  
237 same category. Despite this selection (after the elimination of 18 deviant directions  
238 likely associated with geomagnetic polarity transitions; Fig. S1), the two mean  
239 directions of the different polarities were not antipodal at the 95% confidence level  
240 ( $\gamma = 11.3^{\circ}$  for  $\gamma_c = 4.0^{\circ}$ ; Table 1; McFadden and McElhinny, 1991), likely because of a  
241 residual contamination by the remagnetization component (Fig. 4c). Nevertheless,  
242 the general mean direction obtained for the high-temperature magnetization  
243 component (Table 1) was identical to the mean direction that was previously  
244 estimated by Gallet et al. (2003) from a much smaller dataset ( $\gamma = 2.8^{\circ}$  for  $\gamma_c = 3.4^{\circ}$ ;  
245 McFadden and McElhinny, 1991). It is worth noting that the data from 114 samples  
246 were rejected mainly because only the secondary magnetization component in the  
247 present-day field direction was recorded or their magnetization directions were too

248 scattered upon heating (Fig. S1a-d). This relatively high number, which represents  
249 ~20% of the total collection, is probably related to the very large number of magnetic  
250 reversals observed in the studied section (see below). However, a significant number  
251 of rejected samples were found in the Tyuessala Formation, showing a lithology  
252 generally less favorable for paleomagnetic analyses than that of the Yunkyulyabit-  
253 Yuryakh Formation.

254         The two elements that concern, on one hand, the choice of the magnetic  
255 polarity of the directions obtained, and on the other hand, the primary nature of the  
256 characteristic magnetization isolated at Khorbusuonka are worth discussing (see also  
257 discussion in Gallet et al., 2003). Numerous paleomagnetic results have shown that  
258 the Siberian Platform was located in the Southern Hemisphere during the Cambrian,  
259 rotated by 180° from its current orientation; it then crossed the Equator towards the  
260 end of the Middle Ordovician (e.g. Khramov, 1987; Gallet and Pavlov, 1996; Pavlov  
261 and Gallet, 2005; Torsvik et al., 2012). As a consequence, paleomagnetic directions  
262 with a negative (resp. positive) inclination and a declination of ~160° (resp. ~340°)  
263 have a normal (resp. reverse) polarity (Fig. 4b).

264         The primary nature of the characteristic magnetization was deduced from the  
265 following evidence: (i) whether predominantly carried by magnetite or hematite, the  
266 directions of the magnetization component are similar and the geomagnetic polarity  
267 does not depend on the predominant magnetic phase in the samples. This showed  
268 that the hematite likely has an early diagenetic (or detritic?) origin; (ii) most of the 18  
269 abnormal (or deviating) directions isolated through the section are observed at the  
270 transition between two magnetic polarity intervals (Fig. S1e-f). These most likely  
271 reflect a transitional configuration of the geomagnetic field during a polarity reversal;  
272 (iii) the mean direction estimated for the Khorbusuonka section (note that the



273 averaging tended to reduce the bias due to the recent magnetic field) is almost  
274 identical at the 95% confidence level to the tilt-corrected mean direction (and after its  
275 transfer to the location of the Khorbusuonka section) previously obtained by Pavlov  
276 and Gallet (2001) for the Middle Cambrian part of the Kulumbe section, located  
277 nearly 1500 km away ( $\gamma=3.7^\circ$ ;  $\gamma_c=3.5^\circ$ ; McFadden and McElhinny, 1991). In contrast  
278 the two mean directions are significantly different before bedding correction ( $\gamma=16.0^\circ$ ;  
279  $\gamma_c=3.6^\circ$ ; McFadden and McElhinny, 1991), which gives a positive, albeit partial, fold  
280 test. It is also worth mentioning that the Kulumbe section is slightly younger than the  
281 Khorbusuonka section, since the lower part of the Kulumbe section has been dated  
282 as being at the top of the Drumian *Anomacarioides limbataeformis* trilobite zone (e.g.  
283 Pavlov and Gallet, 2005; Kouchinsky et al., 2008); and (iv) the paleomagnetic pole  
284 defined from the characteristic component isolated at Khorbusuonka is different from  
285 all other paleomagnetic poles of younger age known for the Siberian Platform (e.g.  
286 Khramov, 1987; Torsvik et al., 2012).

287

#### 288 **4. Discussion and conclusions**

289 The results described above indicate that the Khorbusunka section contains a  
290 record of the geomagnetic field that prevailed during the Drumian stage (Epoch  
291 3/Middle Cambrian). The new data offer an average resolution gain of 3.7 compared  
292 to the results of Gallet et al. (2003) (i.e. the ratio between the 437 new samples and  
293 the 119 samples used by Gallet et al., 2003). This gain varied according to the  
294 trilobite zones, from  $\sim 3.4$  for the *Tomagnostus fissus* zone, to  $\sim 2.9$  for the  
295 *Corynexochus perforatus-Anopolenus henrici* and  $\sim 4.1$  for the *Anomacarioides*  
296 *limbataeformis* zone. They show the occurrence of 78 magnetic polarity intervals  
297 along the 85 m of the studied section, of which 22 intervals were defined by only one

298 sample (Fig. 7). An extreme, but probably simplistic, approach would be to eliminate  
299 the latter intervals by considering them too uncertain, which would then imply a  
300 sequence of 46 magnetic polarity intervals; however, there is no a priori reason why  
301 the directions of the samples concerned should be considered to be less well defined  
302 than the others. The most likely reason for the intervals defined by a single sample is  
303 that they were very short and the sampling rate was still too low compared to the  
304 prevailing geomagnetic reversal frequency. For the same stratigraphic sequence,  
305 Gallet et al. (2003) found 42 magnetic polarity intervals, with 18 intervals defined by a  
306 single sample, or a series of 18 intervals after the exclusion of the latter. An important  
307 observation is that, despite the very significant difference in the number of polarity  
308 intervals found between these two studies, a fairly convincing correlation can be  
309 observed between the two magnetostratigraphic sequences (Fig. 8).

310 A comparison is reported in Fig. 8 between the Khorbusuonka results and  
311 those found in a section sampled along the Kulumbe river (Pavlov and Gallet, 2001).  
312 This section was dated by the presence of the *Anomacarioides limbataeformis* and  
313 *Lejopyge laevigata* trilobite zones: that is, it covers the top of the Drumian and the  
314 lower part of the Guzhangian stage (ex-stage 7; e.g. Ogg et al., 2016). In the  
315 Kulumbe section, there was a strong contrast between the *Anomacarioides*  
316 *limbataeformis* zone, marked by numerous geomagnetic reversals, and the *Lejopyge*  
317 *laevigata* zone, with only a few geomagnetic reversals detected. It was not possible  
318 to achieve a convincing correlation between the Khorbusuonka and Kulumbe  
319 sequences (Fig. 8). A first option would be that this difficulty primarily results from the  
320 large number of polarity reversals observed in both Drumian sections, a difficulty  
321 further amplified by possible sampling rate effects and/or due to differences in  
322 sedimentation rate between the two sections. The second would be that the Kulumbe

323 section is slightly younger than the Khorbusuonka section (see also Pavlov and  
324 Gallet, 2005). The most important point, however, is that, whatever the option  
325 retained, the data from Kulumbe confirm the presence of numerous magnetic polarity  
326 reversals during the *Anomacarioides limbataeformis* zone of the Drumian stage.

327         Based on the chronostratigraphic timescale considered for the Cambrian (e.g.  
328 Ogg et al., 2016), and on the biostratigraphic data available for the Middle Cambrian  
329 deposits found along the Khorbusuonka river, the duration of the Khorbusuonka  
330 section analyzed in the present study would have a duration of ~3 Myr. This relatively  
331 short duration implies a geomagnetic reversal frequency as high as ~26 reversals per  
332 Myr, if all polarity intervals are considered, or 15 reversals per Myr if the intervals  
333 defined by a single sample are excluded. Even considering a misinterpretation of  
334 some of the magnetic polarity intervals defined by a single sample, it follows that the  
335 geomagnetic reversal frequency during the Drumian stage was likely >20 reversals  
336 per Myr. This is obviously an extreme reversal rate, if we refer, for example, to the  
337 evolution of the geomagnetic reversal frequencies over the last 100 million years. In  
338 addition, the new paleomagnetic data show that, during this period of hyperactivity of  
339 the reversing process, the geometry of the geomagnetic field probably remained  
340 essentially dipolar, with the slight non-antipodality of the two directions of opposite  
341 polarity observed at Khorbusuonka being likely due to a residual contamination by  
342 the recent field (Fig. 4c).

343         The extreme geomagnetic reversal frequency observed at Khorbusuonka  
344 appears very similar to that found by Bazhenov et al. (2016) for the Zigan Formation  
345 in the southern Urals (Russia), dated to the late Precambrian (Late Ediacaran), some  
346 50-60 Myr before the Drumian. Furthermore, as in Khorbusuonka, the Late Ediacaran  
347 data also showed that a dipolar field likely prevailed at that time. Beyond these

348 exceptional reversal rates, there is the question of their evolution throughout the  
349 Cambrian and the geophysical implications that could result from it.  
350 Magnetostratigraphic data obtained from Siberia, Australia and China have provided  
351 constraints on that evolution after the Drumian; more precisely, during the terminal  
352 part of the Middle Cambrian (Guzhangian stage), the Furongian Epoch/Upper  
353 Cambrian and during the Ordovician (see synthesis in Pavlov and Gallet, 2005 and  
354 Hounslow, 2016). The Kulumbe sequence showed about 30 geomagnetic polarity  
355 intervals during the Guzhangian, but it should be noted that this sequence is very  
356 preliminary due to the poor quality of the paleomagnetic signal observed in the rocks  
357 of the Orakta Formation (Kouchinsky et al., 2008). Considering a duration of ~3.5  
358 Myr for the Guzhangian stage (Ogg et al., 2016), the reversal frequency would have  
359 reached a value of ~10 reversals per Myr at the end of the Middle Cambrian. In  
360 contrast, the reversal frequency would have been much lower during the Furongian  
361 Epoch/Late Cambrian (~1.5 reversal per Myr), with ~15 to 20 magnetic polarity  
362 intervals over a duration of ~12 Myr (Pavlov and Gallet, 2005; Kouchinsky et al.,  
363 2008; Ogg et al., 2016; Hounslow, 2016). A similar low reversal rate would also have  
364 prevailed at the beginning of the Lower Ordovician, just before the onset of the  
365 reverse-polarity Moyero Superchron (Fig. 9; Gallet and Pavlov, 1996; Pavlov and  
366 Gallet, 1998; 2005; Pavlov et al., 2017). Therefore, the presently available  
367 magnetostratigraphic data show that the magnetic reversal frequency dramatically  
368 dropped from values >20 reversals per Myr to ~1.5 reversal per Myr in ~3-4 Myr,  
369 which is the duration of the Guzhangian stage (with an 'intermediate', albeit high,  
370 value of ~10 reversals per Myr during the Guzhangian).

371         The new magnetostratigraphic data, and those previously obtained by  
372 Bazhenov et al. (2016), next raise the question of the evolution in reversal frequency

373 during the Lower Cambrian. Was the period between ~550 Ma and ~500 Ma  
374 characterized by a single or more episodes of hyperactivity in the reversing process?  
375 Note further that the onset of the hyperactivity episode in the Late Ediacaran has not  
376 yet been documented. To date, the magnetostratigraphic data available for the Lower  
377 Cambrian are neither numerous nor constrained enough to answer this question with  
378 certainty. Even though the data mainly obtained from Siberia (Kirschvink and  
379 Rozanov, 1984; Kirschvink et al., 1991; Gallet et al., 2003) and China (Duan et al.,  
380 2018 and references therein) clearly showed the occurrence of many geomagnetic  
381 polarity reversals during Epoch 2/the Lower Cambrian, it remains difficult to precisely  
382 determine the reversal rates. Duan et al. (2018) recently proposed a frequency of ~7  
383 reversals per Myr between ~524 Ma and 514 Ma, but the stratigraphic sequence  
384 from which the data were obtained was fragmentary; it therefore seems possible (but  
385 not proven) that the reversal frequency was actually higher during this period. The  
386 Lower Cambrian paleomagnetic data obtained at Khorbusuonka by Gallet et al.  
387 (2003) also attest to a very high geomagnetic reversal frequency during the  
388 Tommotian and Atdabanian (Gallet et al., 2003 tentatively proposed a reversal  
389 frequency even higher than that of the Drumian), in addition to the occurrence of a  
390 major true polar wander event near the end of the Lower Cambrian, in agreement  
391 with Kirschvink et al. (1997). The data, however, should be confirmed by further  
392 studies on other sedimentary sections in Siberia or elsewhere. At this stage, it is only  
393 possible to affirm that the reversal frequency was high (>5 reversals per Myr) during  
394 the Lower Cambrian, but the hyperactivity character (>15 reversals per Myr) cannot  
395 be established (Fig. 9).

396 Comparing dipole field behavior in numerical simulations with varying inner-  
397 core size and assuming inner-core nucleation began during the Neoproterozoic,

398 Lhuillier et al. (2019) recently suggested that the hyperactivity of the reversing  
399 process during the latest Precambrian and the Cambrian could be linked to a critical  
400 size of the inner core, around an aspect ratio of  $\sim 0.2$ . During this critical period,  
401 corresponding to significant changes in the flow pattern in the tangent cylinder, the  
402 reversal frequency would have been very sensitive to variations in the vigor of the  
403 convection in the outer core. If these simulations do indeed provide relevant  
404 information for the geodynamo, this sensitivity could account for long duration of the  
405 hyperactivity phenomenon, if it persisted for about 50 Myr, or recurrence of such  
406 events, if several hyperactivity episodes were present, a very different situation from  
407 that observed during the Middle Jurassic (Lhuillier et al., 2019). It should be noted,  
408 however, that just changing the thermal conditions at the core-mantle boundary can  
409 alone account for the full spectrum of temporal variations in geomagnetic reversal  
410 rate (e.g. Driscoll and Olson, 2011; Olson et al., 2013). On another hand, Bono et al.  
411 (2019) showed that the hyperactivity episode at the end of the Precambrian was also  
412 accompanied by very low mean paleointensity. This observation would be consistent  
413 with a weak-field geodynamo state before inner-core nucleation as predicted by  
414 numerical models (Driscoll, 2016; Landeau et al., 2017; see also Bono et al., 2019).  
415 Accurate paleointensity data are not yet available to determine whether this state still  
416 prevailed in the Middle Cambrian. An important point of note, however, is that both  
417 the data from Ediacaran (Bazhenov et al., 2016) and the Middle Cambrian seem to  
418 argue for a dipole-dominated field geometry at the beginning of the inner-core  
419 growth.

420         The Cambrian thus appears to be a period marked by remarkable and  
421 fascinating phenomena affecting all terrestrial envelopes. Linking them together  
422 exceeds the objectives of our study (see for example Kirschvink et al., 1997; Meert et

423 al., 2016). Here, we have emphasized the exceptional nature of the reversing  
424 process during the Middle Cambrian, but the sharp decrease in reversal rate at the  
425 end of the Middle Cambrian provides further important information on the geodynamo  
426 behavior. It argues for threshold effects in core processes (see also Courtillot and  
427 Olson, 2007), which would be consistent with a ~1-Myr-scale transition between two  
428 reversing modes; that is, between the normal and hyperactive reversing modes  
429 defined by Gallet and Pavlov (2016) (Fig. 9).

430

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440

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579

580

581 **Table and Figures captions**

582

583 **Table 1.** Paleomagnetic results obtained from the Drumian part of the Khorbusuonka  
584 section. LTC and HTC refer to the magnetization components isolated in the low and  
585 high unblocking temperature ranges.

586

587 **Fig. 1.** General geology of northeastern Siberia (a) and of the area in the vicinity of  
588 the Khorbusuonka section (b). The red square in panel (a) and the red star in panel  
589 (b) indicate the location of the studied Khorbusuonka section.

590

591 **Fig. 2.** View of the Khorbusuonka section analyzed in the present study (@Yves  
592 Gallet). The simplified lithology, and the dating of the studied sequence, are from  
593 Korovnikov and Tokarev (2018).

594

595 **Fig. 3.** Representative examples of the thermal demagnetization of the samples from  
596 the Khorbusuonka section. Solid (resp. open) circles are in the horizontal (resp.  
597 vertical) plane. The corresponding magnetization moment versus temperature curves  
598 are also reported (right panels).

599

600 **Fig. 4.** Stereographic projections of the directions obtained in the Khorbusuonka  
601 section. Directions at the sample level were isolated for the low (a) and high (b)  
602 unblocking temperature components (LTC and HTC, respectively). (c) Mean

603 directions with opposite polarities. The closed (open) symbols refer to directions in  
604 the lower (upper) hemisphere. All directions were obtained using the PaleoMac  
605 software (Cogné, 2003).

606

607 **Fig. 5.** Thermal demagnetization of three-axis isothermal remanent magnetization  
608 components acquired in fields of 1.9 T (dots), 0.3 T (triangles) and 0.15 T (squares).  
609 These examples show that the magnetization of the samples was likely carried by  
610 various mixtures of magnetite and hematite.

611

612 **Fig. 6.** Examples of thermal demagnetization of samples from the Khorbusuonka  
613 section showing a strong overlap between the low and high unblocking temperature  
614 magnetization components. In these cases, a direction could not be reliably  
615 determined for the high unblocking temperature component (see text). Same  
616 conventions as in Fig. 3-4.

617

618 **Fig. 7.** Magnetostratigraphy of the Drumian sequence of the Khorbusuonka section.  
619 The dark green dots show the directions used to estimate a mean high unblocking  
620 temperature magnetization direction, whereas those in light green color refer to  
621 biased directions, allowing constraint of only the associated magnetic polarity (see  
622 text).

623

624 **Fig. 8.** Comparison between the magnetostratigraphic sequences of Middle  
625 Cambrian age (Epoch 3) obtained from the Khorbusuonka (Gallet et al., 2003; this  
626 study) and Kulumbe (Pavlov and Gallet, 2001) sections.

627

628 **Fig. 9.** Synthesized sketch of the evolution in geomagnetic reversal frequency since  
629 the end of the Precambrian (modified from Gallet et Pavlov, 2016).

630

### 631 **Supplementary material**

632 **Fig. S1.** (a-d) Thermal demagnetization of rejected samples. In each case, the  
633 directional data are reported in orthogonal and stereographic projections, together  
634 with the evolution of magnetization as a function of temperature (see text). Same  
635 conventions as in Fig. 3 and 4. (e-f) Two examples of demagnetization behavior for  
636 samples located at a geomagnetic polarity transition.

637



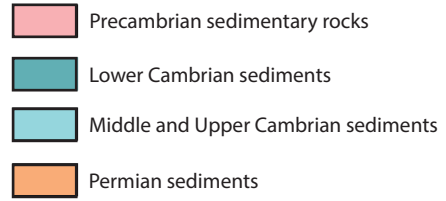
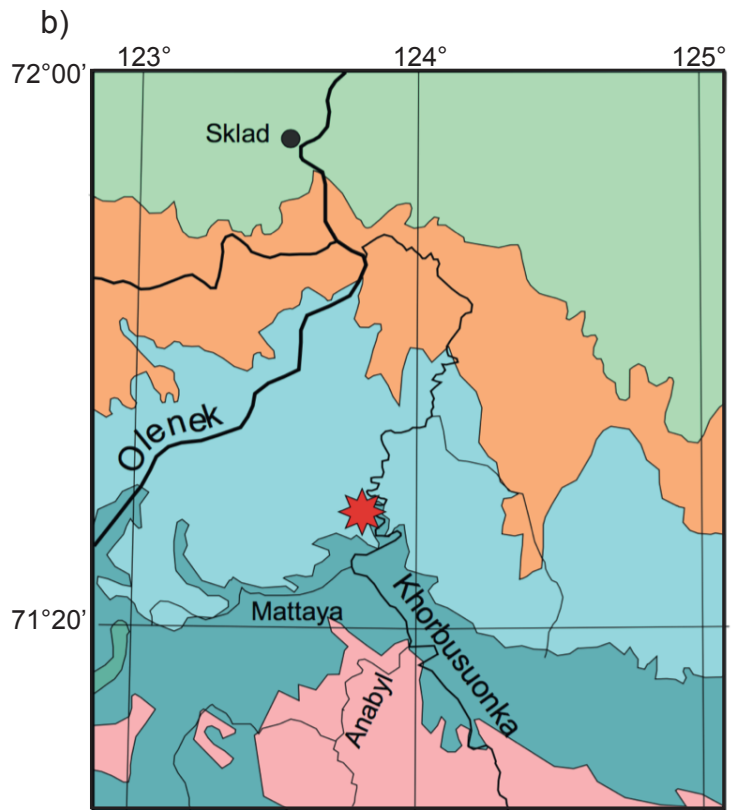
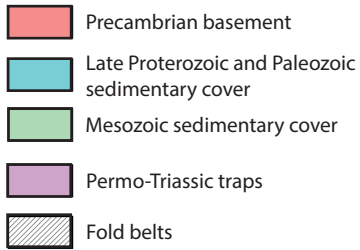
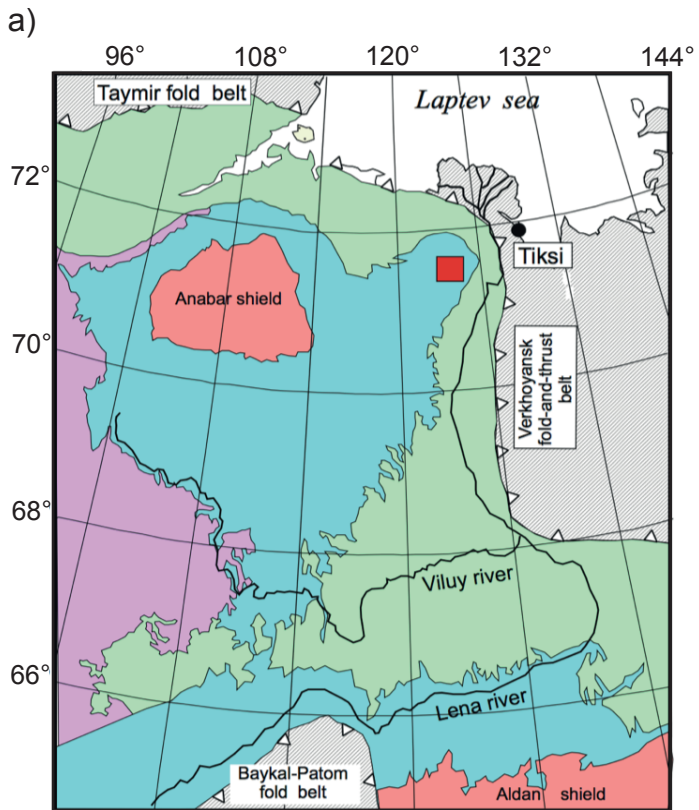


Figure 1

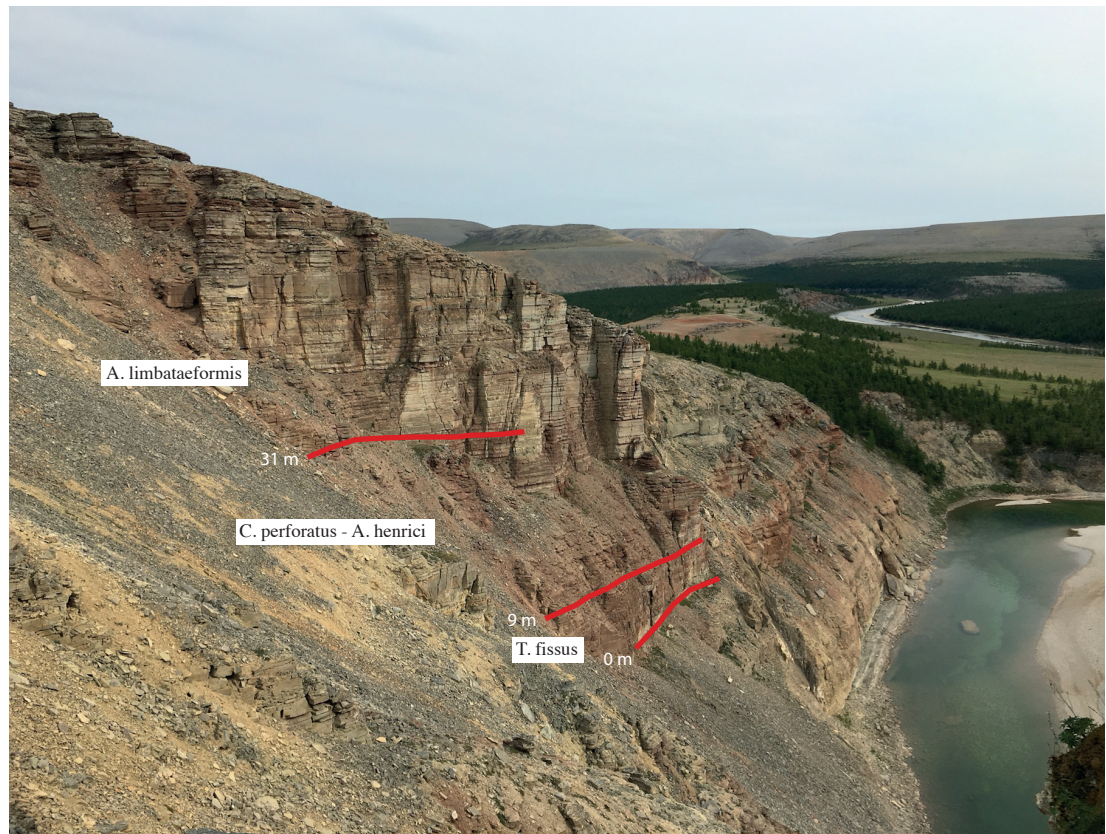
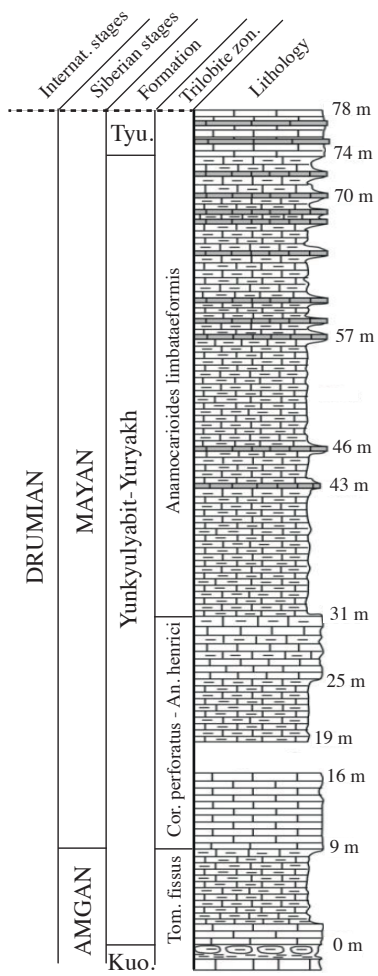


Figure 2

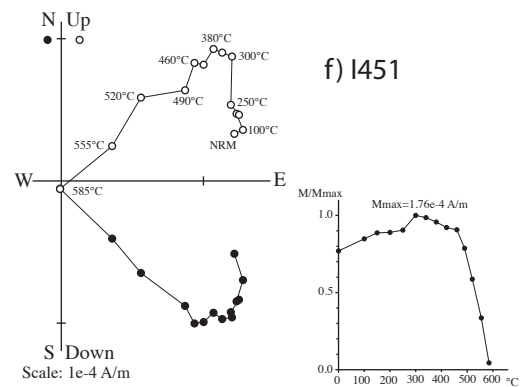
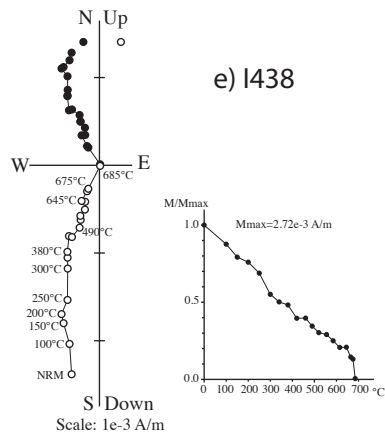
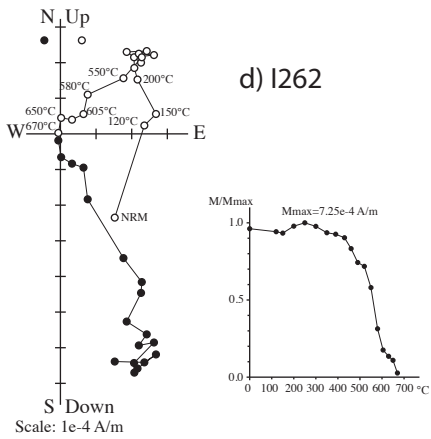
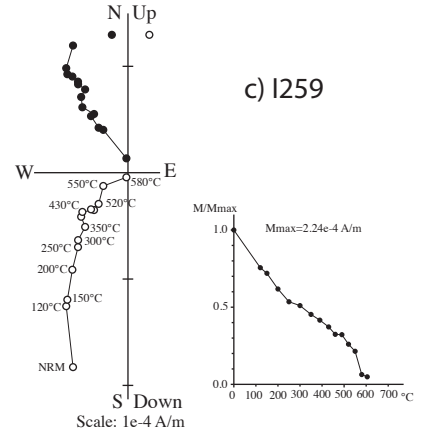
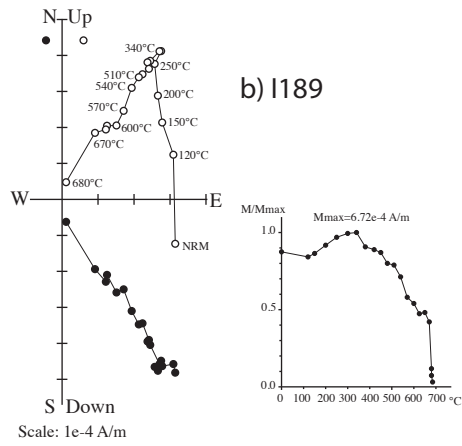
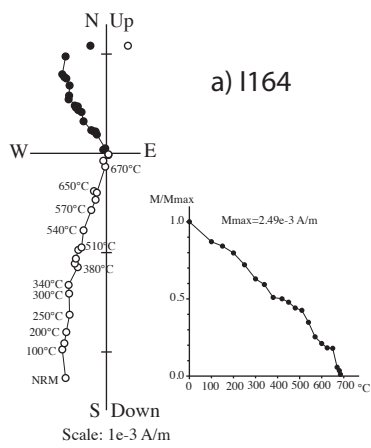


Figure 3

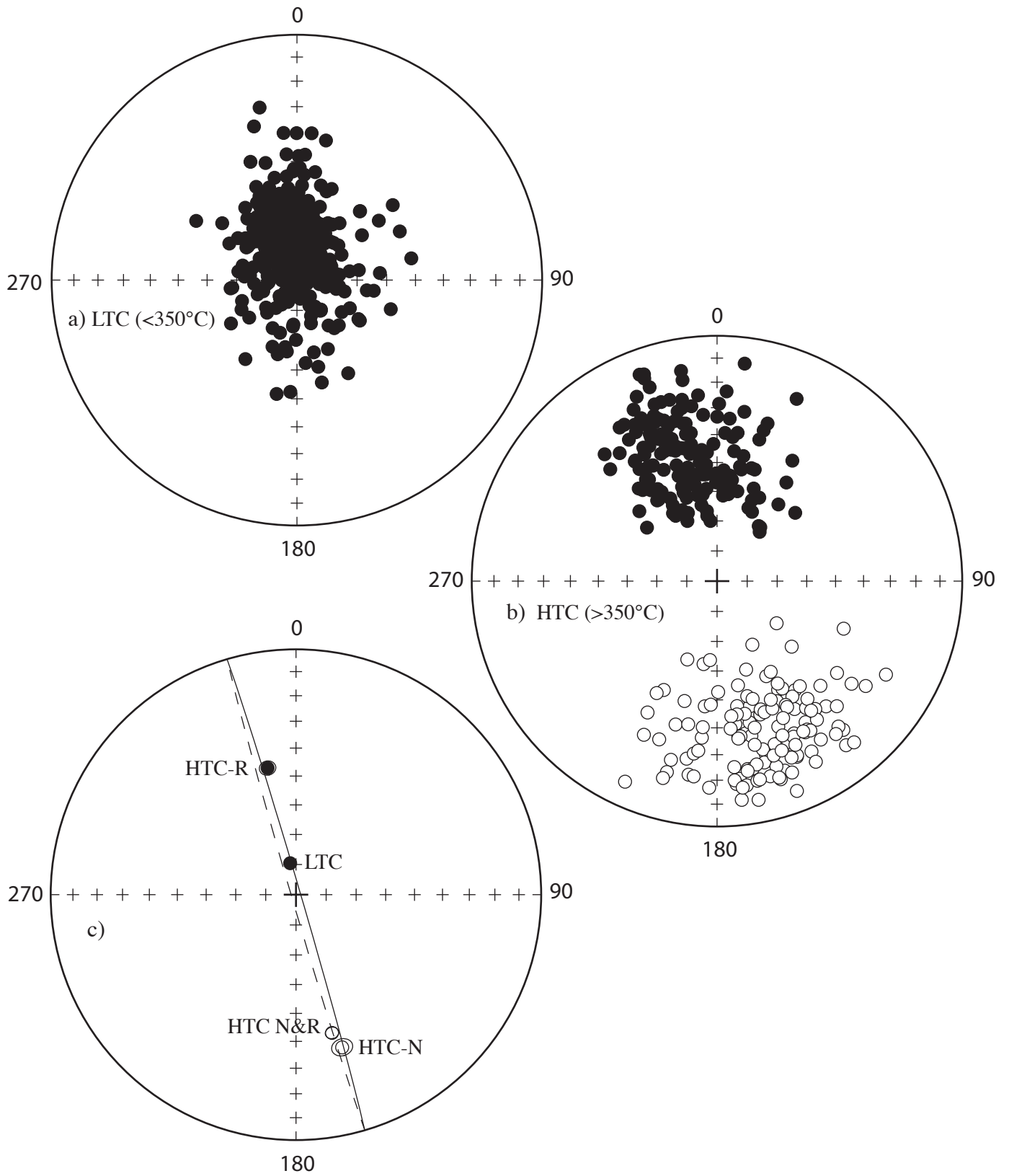


Figure 4

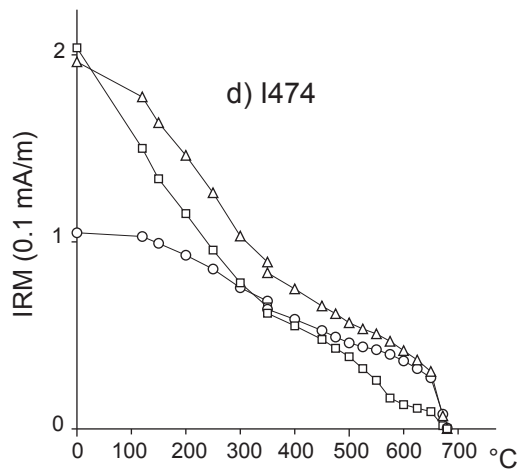
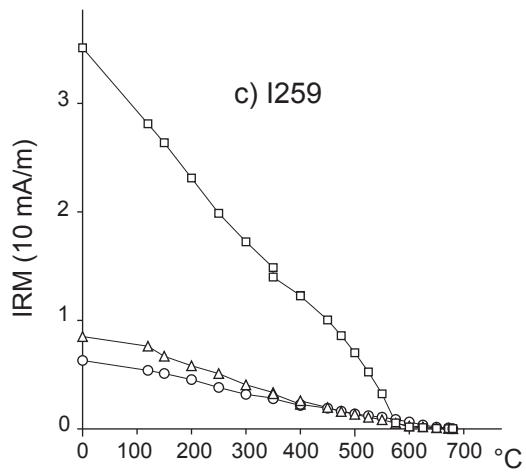
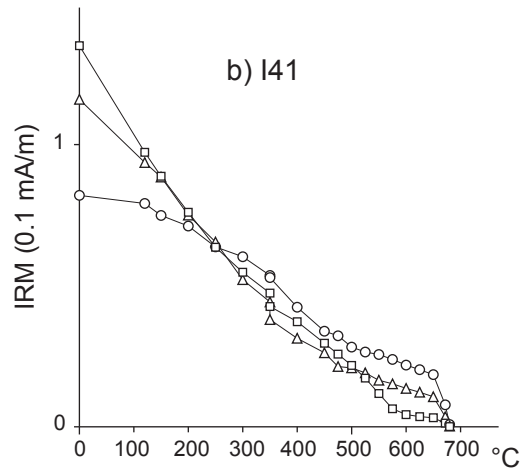
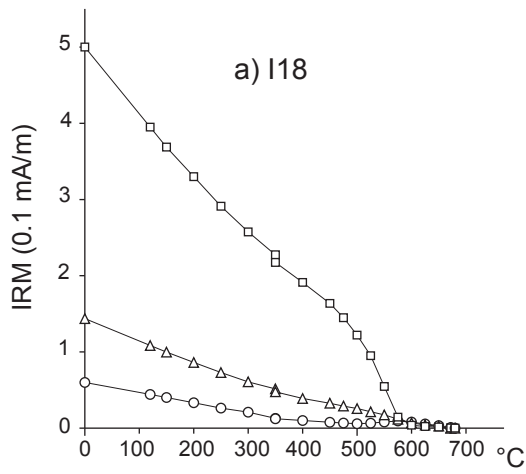
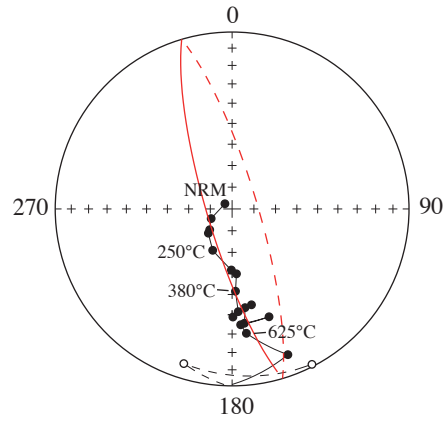
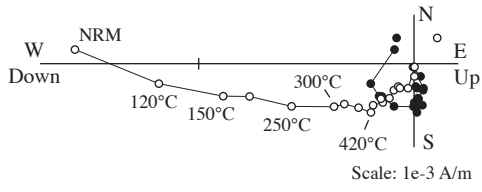
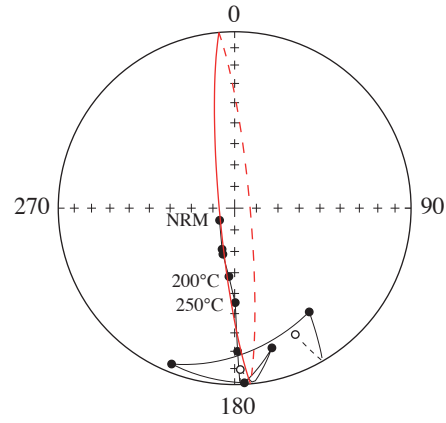
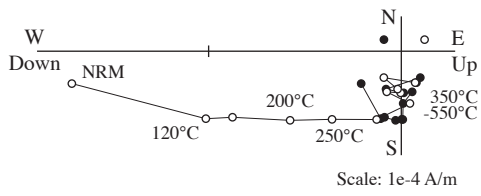


Figure 5

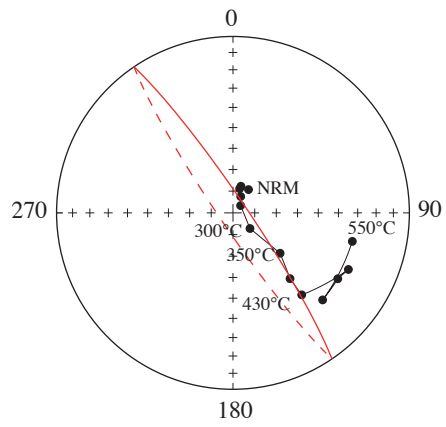
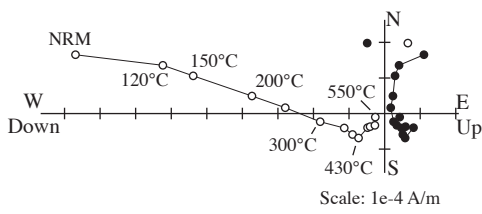
a) I113



b) I213



c) I252



d) I416

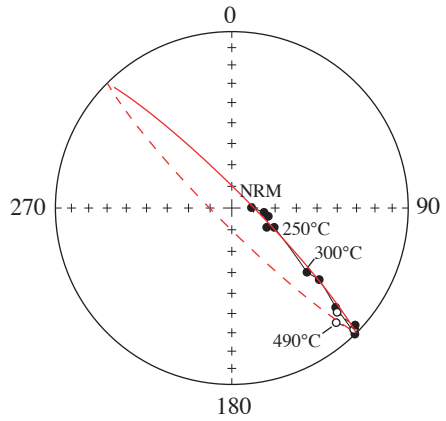
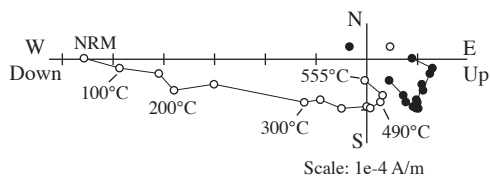


Figure 6

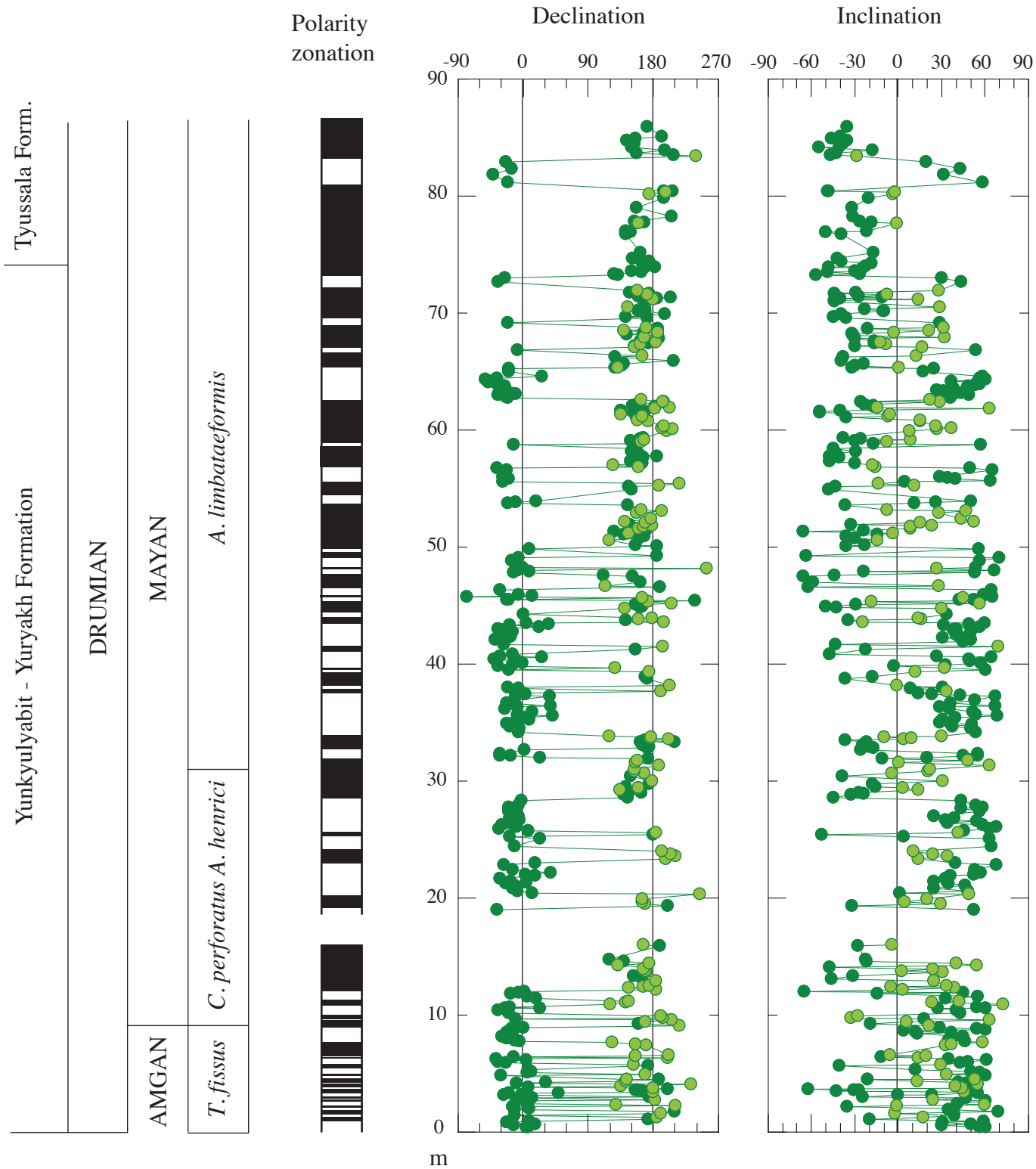


Figure 7



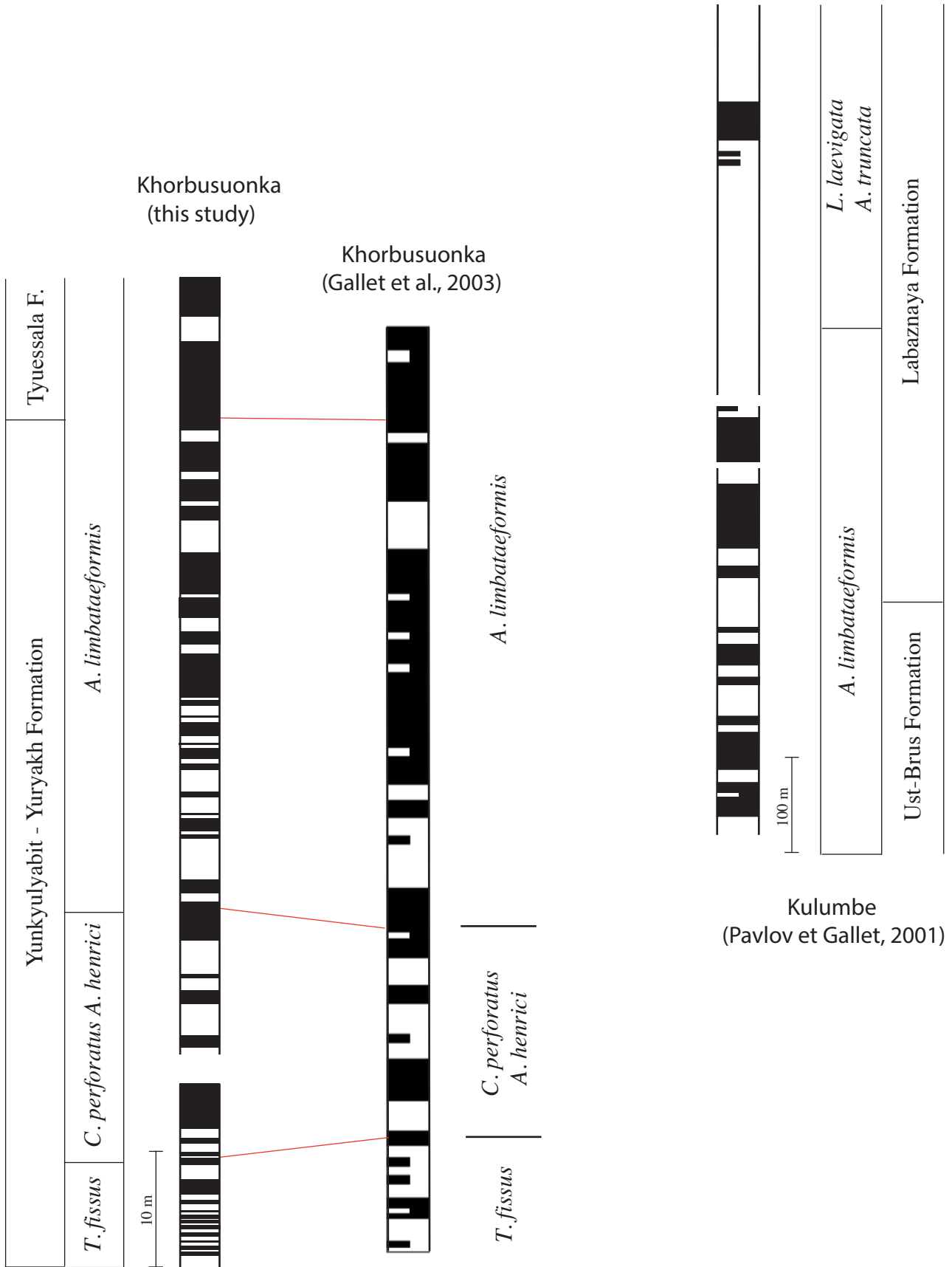


Figure 8



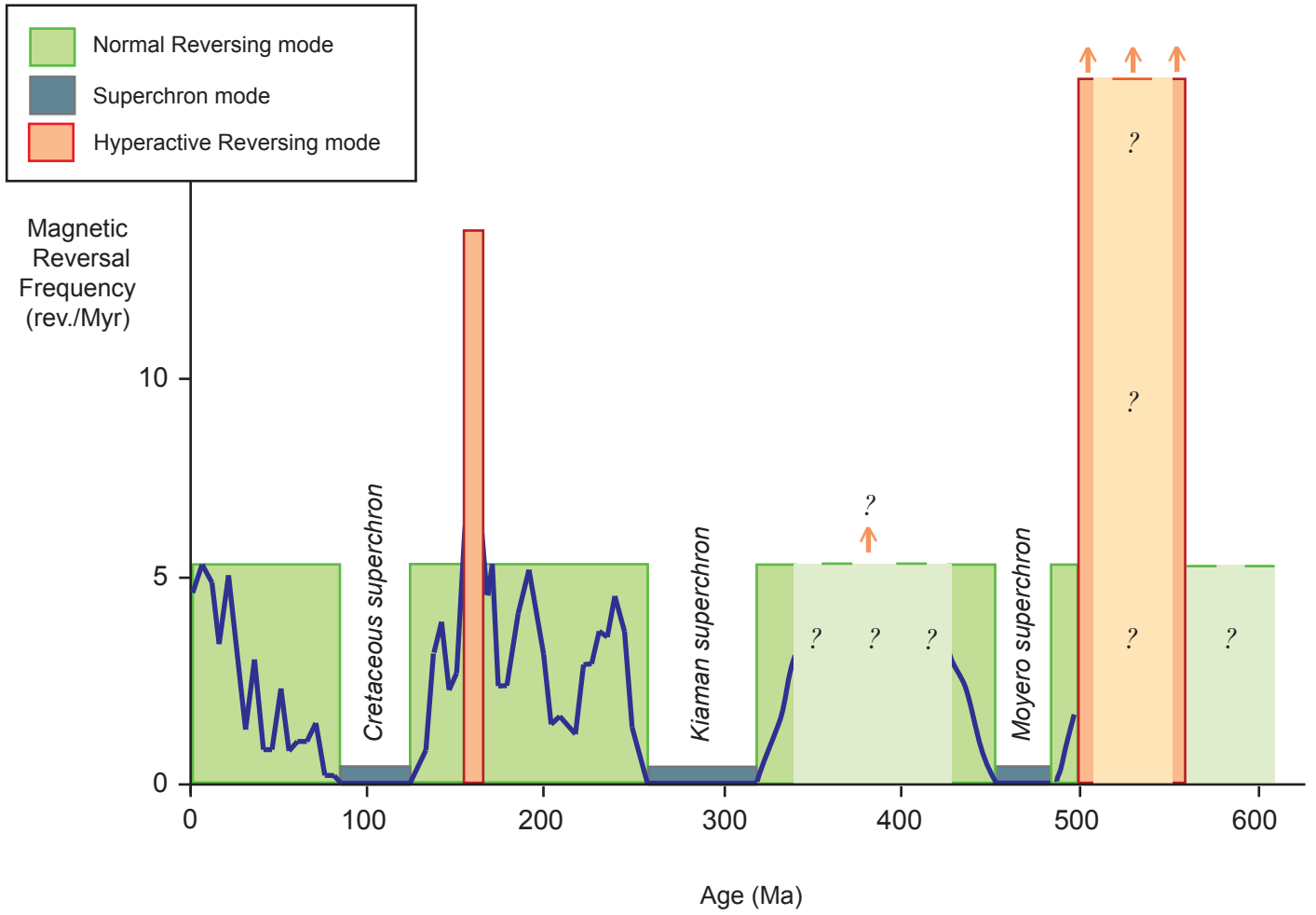


Figure 9

| Khorbusuonka section<br>$\lambda=71^{\circ}28'47''\text{N}$<br>$\Phi=123^{\circ}49'57''\text{E}$ | N   | In situ |       |      |               | Tilt corrected |       |      |               | Paleomagnetic pole |       |         |
|--|-----|---------|-------|------|---------------|----------------|-------|------|---------------|--------------------|-------|---------|
|  |     | Dec°    | Inc°  | K    | $\alpha_{95}$ | Dec°           | Inc°  | K    | $\alpha_{95}$ | Lat°               | Long° | dp°/dm  |
| LTC  | 515 | 349.5   | 79.4  | 27.5 | 1.2°          | 349.6          | 79.4  | 27.5 | 1.2           |                    |       |         |
| HTC-Normal   | 131 | 163.1   | -35.2 | 16.4 | 3.1           | 163.1          | -35.3 | 16.4 | 3.1           |                    |       |         |
| HTC-Reverse  | 168 | 347.4   | 46.1  | 19.3 | 2.5           | 347.4          | 46.1  | 19.3 | 2.5           |                    |       |         |
| HTC N+R  | 299 | 165.3   | -41.4 | 16.6 | 2.1           | 165.4          | -41.4 | 16.6 | 2.1           | -41.6              | 141.8 | 1.6/2.6 |

Table 1