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# Silicate-carbonate liquid immiscibility: insights from the Crevier alkaline intrusion (Quebec)

Groulier P.-A.<sup>1,2</sup>, Turlin F.<sup>1,2</sup>, André-Mayer A.S.<sup>1\*</sup>, Ohnenstetter D.<sup>3</sup>, Crépon A.<sup>1</sup>, Boulvais

P.<sup>4</sup>, Poujol M.<sup>4</sup>, Rollion-Bard C.<sup>5</sup>, Zeh A.<sup>6</sup>, Moukhsil A.<sup>7</sup>, Solgadi F<sup>7</sup>, El Basbas A.<sup>7,8</sup>

<sup>1</sup> GeoRessouces laboratory, Université de Lorraine-CNRS, 54000 Nancy, France

<sup>2</sup> Now at Université du Québec à Montréal, Département des Sciences de la Terre et de l'Atmosphère, 201, ave. du Président-Kennedy, Montréal, QC, Canada H2X 3Y7

<sup>3</sup>CRPG, CNRS, Université de Lorraine, Vandœuvre-lès-Nancy, France

<sup>4</sup> Univ rennes, CNRS, Géosciences Rennes - UMR 6118, F-35000 Rennes

<sup>5</sup> Institut de Physique du Globe de Paris (IPGP), UMR CNRS 7154, Université Paris Diderot, 1 rue Jussieu, F-75238 Paris Cedex 05, France

<sup>6</sup> KIT-Karlsruhe Institute of Technology, Campus South, Institute for Applied Geosciences, Mineralogy and Petrology, 76131, Karlsruhe, Germany

<sup>7</sup> Ministère des Ressources Naturelles du Québec, Direction du Bureau de l'Exploration Géologique du Québec, 400, boulevard Lamaque, Val-d'Or (Québec) J9P 3L4

<sup>8</sup>Now at Faculté des Sciences Université Moulay Ismail BP: 11201 - Zitoune 50060 - Meknès, Maroc

## \*Corresponding author: <u>anne-sylvie.andre@univ-lorraine.fr</u> - +33 (0)3 72 74 55 66

ORCID Numbers: Anne-Sylvie ANDRE-MAYER: 0000-0002-9307-1419 Pierre-Arthur GROULIER: 0000-0003-2386-9270 Marc POUJOL: 0000-0001-8682-2926 François TURLIN: 0000-0002-0378-6972

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#### Abstract

This contribution explores the petrogenetic relationships between silicate and carbonatitic rocks in the Crevier Alkaline Intrusion (CAI, Québec, Canada). The CAI is located in the Proterozoic Grenville Province and is composed of a suite of undersaturated peralkaline rocks from ijolite to nepheline syenite and carbonatites. Petrogenetic relationships between different undersaturated alkaline igneous rocks, carbonate-bearing and carbonate-free nepheline syenite and carbonatites observed in the CAI suggest that (i) carbonate-bearing and carbonate-free silicate rocks are comagmatic with carbonatite, and that (ii) both silicate and carbonatitic liquids are fractionated from an ijolitic parental magma that has undergone liquid immiscibility. One of the observed facies is characterized by spectacular ocelli of carbonate-bearing nepheline syenite in a matrix of carbonatite. The younger nepheline syenite facies can be divided into two groups based on the presence or absence of magmatic carbonates. Both groups are characterized by the presence of pyrochlore-group minerals that carry the Nb-Ta mineralization.

We specifically use accessory minerals such as zircon, pyrochlore and apatite to constrain the temporal and physicochemical parameters of the immiscibility process. By coupling (i) mineral textures, (ii) trace elements, (iii) Ti-in-zircon thermometry, and (iv) oxygen isotope compositions, we have traced the crystallization of zircon before, during and after the immiscibility process. The results allowed us to constrain the minimum temperature of this process at ~815-865°C. In addition, two magmatic populations of pyrochlore are identified through their petrographic and geochemical characteristics within the younger nepheline syenite facies. Pyrochlore from the earlier ocelli facies of carbonate-bearing nepheline syenite follow a Nb-Ta differentiation trend, whereas pyrochlore from the younger carbonate-free nepheline syenite follow a more classical Nb-Ti trend. Following the complete immiscibility between the silicate and carbonatitic liquids, the fractionation between Nb and Ta stopped while a new generation of Nb-rich pyrochlore grew, displaying a more classical Nb-Ti fractionation trend and a more constant Nb/Ta ratio in the nepheline syenite.

**KEYWORDS:** Immiscibility; Nepheline syenite; Carbonatite; Accessory minerals; Crevier alkaline intrusion

#### INTRODUCTION

Carbonatites often occur in close spatial association with undersaturated alkaline igneous rocks with compositions varying between pyroxenites, nephelinites and/or ijolites, and more evolved nepheline syenites (Le Bas, 1987; Gittins, 1989; Bailey, 1993; Harmer, 1999; Winter, 2001; Mitchell, 2005; Millong & Groat, 2013). The petrogenetic link between carbonatite and alkaline silicate rocks, both of mantle origin, is still discussed (Bell & Blenkinsop, 1989; Bell, 1998; Bell et al., 1998; Winter, 2001; Halama et al., 2005, Bell & Simonetti, 2010; Mollex et al., 2018). Some authors argue that carbonatites represent primary melts extracted from a carbonate-rich mantle source (Wyllie & Huang, 1976; Wallace & Green, 1988; Bédard & Chown, 1992; Dalton & Wood, 1993; Dalton & Presnall, 1998; Harmer & Gittins, 1998; Moore & Wood, 1998; Brey et al., 2008; Dasgupta et al., 2009; Hammouda et al., 2009). Others plead for fractional crystallization from a CO<sub>2</sub>-rich parent silicate melt (Watkinson & Wyllie, 1971; Lee & Wyllie, 1994; Korobeinikov et al., 1998; Nielsen & Veksler, 2002; Ulmer & Sweeney, 2002), or for immiscible separation from a carbonate-bearing silicate melt (Koster van Groos & Wyllie, 1973; Freestone & Hamilton, 1980; Kjarsgaard & Hamilton, 1988; Lee & Wyllie, 1997; Brooker, 1998; Kjarsgaard, 1998; Veksler et al., 1998a; Halama et al., 2005; Brooker & Kjarsgaard, 2011; Potter et al., 2017).

Although the immiscibility between silicate and carbonate liquids is poorly documented and only observed *in-situ* in few localities, it represents an attractive model for the genesis of alkaline silicate-carbonatite complexes (Le Bas, 1987). It has been proposed for a number of occurrences, e.g. the Gardiner complex (Greenland, Nielsen, 1980; Veksler *et al.*, 1998b), the Maoniuping complex (China, Xu *et al.*, 2004), Grønnedal-Ìka (Greenland, Halama *et al.*, 2005; Taubald et al., 2004), Kerimasi (Tanzania, Guzmics *et al.*, 2011), and in the Grenville Province at Oka (Quebec, Canada, Treiman & Essene, 1985). Moreover, in the Kenya rift region, several studies support the hypothesis that the large amount of calcio- to natro-carbonatite volcanic rocks (Nasira, Kerimasi, Shombole, Oldoinyo Lengai) originated from liquid immiscibility (Freestone & Hamilton, 1980; Kjarsgaard & Peterson, 1991; Dawson *et al.*, 1994; Kjarsgaard *et al.*, 1995; Church & Jones, 1995; Veksler *et al.*, 1998b; Mitchell, 2009; Mitchell & Dawson, 2012; Sharygin *et al.*, 2012; Sekisova *et al.*, 2015; Potter *et al.*, 2017; Mollex, 2017).

Major minerals such as calcite, pyroxene, feldspar and nepheline, are commonly used to trace silicate-carbonate liquid immiscibility, using *in-situ* mineral chemistry and isotopes (O, Nd) (Bell, 1998; Harmer, 1999; Haynes *et al.*, 2003; Marks *et al.*, 2004a; 2004b; Halama et *al.*, 2005; Guzmics *et al.*, 2008; Guzmics *et al.*, 2012; Morteani *et al.*, 2013). Other studies have used petrographic features coupled to melt inclusions (Nielsen *et al.*, 2017, Mitchell, 2009; Guzmics *et al.*, 2011, 2012; Sekisova *et al.*, 2015; Potter *et al.*, 2017, Baudoin *et al.*, 2018), isotopic constraints (Bell, 1998; Xu *et al.*, 2004; Halama *et al.*, 2005; Li *et al.*, 2016), as well as experimental data (Koster van Groos & Wyllie, 1963; Lee & Wyllie, 1996; Suk, 2001; Brooker & Kjarsgaard, 2011; Veksler *et al.*, 1998a; Veksler *et al.*, 2012) have also been used to trace immiscibility, with variable degrees of success.

The Crevier Alkaline Intrusion (CAI) is located in the central Grenville Province (Quebec, Canada). The intrusion is composed of a suite of undersaturated alkaline rocks that range from ijolite to nepheline syenite and alkaline syenite with comagmatic carbonatites. Therefore, it constitutes a suitable target to investigate silicate-carbonatitic liquid immiscibility. In this study, we present (i) macro- to microscopic textural characterization, (ii) whole-rock geochemistry, (iii) oxygen and carbon stable isotope of carbonates, (iv) major and/or trace element composition of accessory minerals from

silicate (zircon and pyrochlore) and carbonatite facies (apatite), (v) oxygen isotope of zircon, and (vi) U-Pb dating of zircon and apatite from silicate and carbonatite facies, respectively. We argue that the undersaturated peralkaline silicate rocks and carbonatites of the CAI were produced by/generated through multistage carbonate-silicate immiscibility processes.

## THE PROTEROZOIC GRENVILLE PROVINCE

The Grenville Province, located along the southeastern Laurentian margin in Canada, is composed of Archean and Proterozoic terranes reworked during a major continent-continent collision during the late-Mesoproterozoic to the early-Neoproterozoic (Fig. 1a; Rivers, 1997; Rivers *et al.*, 2012; Rivers, 2015).The evolution of the Grenville Province can be separated into two major stages: (i) a long-lived active margin built at the margin of southeastern Laurentia from the Paleoproterozoic to the Mesoproterozoic (ca. 1.71 to 1.18 Ga) where juvenile crust formed at discrete times in arc and back-arc systems (Rivers & Corrigan, 2000; Whitmeyer & Karlstrom, 2007; Rivers *et al.*, 2012; Bickford *et al.*, 2015; Groulier *et al.*, 2018a), and (ii) a major continental collision, the Grenvillian Orogeny (ca. 1090 to 960 Ma) (e.g. Rivers, 1997; Rivers *et al.*, 2012; McLelland *et al.*, 2013; Turlin *et al.*, 2018).

## FIGURE 1

## The Grenvillian Orogeny

The Grenvillian Orogeny is recorded as two distinct phases in two distinct tectonometamorphic domains. The Allochthonous Belt is composed of terranes formed outboard of and accreted to the Laurentian margin during the Mesoproterozoic (Rivers *et* 

*al.*, 2012). This belt records the Ottawan crustal thickening phase (ca. 1090-1020 Ma, e.g. Rivers *et al.*, 2012). The Allochthonous Boundary Thrust (ABT) separates the Allochthonous Belt from the underlying Parautochthonous Belt that corresponds to lithotectonic units of the Superior Province and their cover sequences, and to previously accreted arcs reworked during the Grenvillian Orogeny (Rivers *et al.*, 1989, 2012; Gower & Krogh, 2002; Groulier *et al.*, 2018a). The Parautochthonous Belt records the Rigolet crustal thickening phase (ca. 1005-960 Ma, Rivers, 2009; Rivers *et al.*, 2012). The Grenvillian orogeny is interpreted to represent one of the best examples of a large, hot and long-duration orogen (LHO, Beaumont *et al.*, 2006; Rivers, 2008, 2009).

Following a syn-Ottawan peak of metamorphism that reached up to granulite-facies conditions, the (i) intrusion of pegmatitic granite dykes at ca. 1005-1000 Ma under close to solidus conditions (Turlin *et al.*, 2018), (ii) titanite cooling ages at ca. 1006-987 Ma (Dunning & Indares, 2010; Indares & Dunning, 2018; Groulier *et al.*, 2018b), and (iii) modeled closure temperatures (~550-450 °C) of apatite crystals at ca. 960 Ma (Turlin *et al.*, 2018) document a slow cooling of the medium to low pressure belt (aM-L*P*, Fig. 1b) at a rate of ~2-6°C/Ma. This thermal evolution has been attributed to a lithospheric thinning triggered by slab retreat (Turlin *et al.*, 2018) or to a lithospheric foundering (also referred to delamination; Corrigan & Hanmer, 1997; Groulier *et al.*, 2018b).

#### The Grenvillian magmatism

The two orogenic phases of the Grenvillian Orogeny are marked by the emplacement of intrusive rocks of distinct geochemical affinities. The Ottawan phase induced a voluminous and widespread magmatic activity dominated by Anorthosite-Mangerite-Charnockite-Granite (AMCG) and Monzonite-Diorite-Gabbro-Syenite (MDGS) suites (see

Rivers *et al.*, 2012 and references therein), and potassic to ultrapotassic intrusions that are assumed to be formed by the melting of an enriched and metasomatized mantle-source impacted by the long-lived subduction beneath Laurentia (Corriveau & Morin, 2000; Rivers *et al.*, 2012 and references therein; Groulier *et al.*, 2018b). The magmatic activity during the Rigolet phase is composed of either parautochthonous- or allochthonous-derived ca. 1005-983 Ma peraluminous pegmatitic granite dykes, and two-micas garnetiferous leucogranite bodies (Groulier *et al.*, 2018b; Turlin *et al.*, 2017, 2019), all of them hosted in the Allochthonous Belt. From ca. 980 to 950 Ma, the magmatism evolved towards a more alkaline signature consistent with the emplacement of felsic to minor silica-undersatured intrusion emplaced within the Interior Magmatic Belt of the Allochthonous Belt (Bourne, 1991; Gower *et al.*, 1991, 2008; Rivers, 1997; Gower & Krogh, 2002).

#### The Crevier Alkaline Intrusion (CAI)

The CAI was discovered in 1975 following airborne radiometric surveys and was then carefully studied and mapped using different techniques (drilling campaigns, induced polarization, airborne magnetic survey) leading to the discovery of rare metals (Nb, Ta, U) occurrences carried by minerals of the pyrochlore supergroup (Bergeron, 1980; Laplante, 1980; Harrisson, 1981; Bernier & Moorhead, 2000; Goutier, 2006). The CAI hosts the Crevier deposit, a significant resource of Nb and Ta with a total tonnage (measured and indicated) of 25.4 Mt grading at 0.2 wt.% Nb<sub>2</sub>O<sub>5</sub> and 234 ppm Ta, and of 15.4 Mt inferred resources grading at 0.17 wt.% Nb<sub>2</sub>O<sub>5</sub> and 252 ppm Ta (Duplessis & Girard, 2010). The property is owned by a joint venture between Niobay Metals Inc. and Niobec Inc. The CAI is emplaced within charnockitic suites and gneisses of the Allochthonous Belt of the

central Grenville Province (Quebec), north of the Lac Saint-Jean region (Figs. 1 and 2). The CAI outcrops over 25 km<sup>2</sup> and is mainly composed of undersaturated peralkaline silicate and carbonatite rocks (Fig. 2b). Following the discovery of this igneous complex, Bergeron (1980) studied the petrology and geochemistry of the different units while Laplante (1980) worked on the Nb-Ta-U mineralization. In addition, Harrisson (1981) obtained a K-Ar age of 879 Ma in rocks from the CAI. Despite the information acquired by these studies, some numerous unknowns still remained, especially regarding the accurate crystallization age of the complex, the nature of its source and of the magmatic processes that led to the coexistence of silicate and carbonatite rocks. The intrusion is oriented N320 (Fig. 2b), following the general orientation of the Waswanipi-Saguenay lineament (Fig. 1b, Bernier & Moorhead, 2000; Solgadi et al., 2015) that hosts the Paleoproterozoic Montviel (ca. 1894 Ma, David et al., 2006; Goutier, 2006) and the late-Neoproterozoic Saint-Honoré Nb-REE bearing carbonatites (Niobec mine, 582.1±1.8 Ma, in Tremblay et al., 2015). The location of these peralkaline complexes with economic importance along a major crustal lineament highlights the importance of such structure during the emplacement of igneous bodies of similar composition.

## **FIGURE 2**

## SAMPLING AND ANALYTICAL METHODS

## Sampling

Rock material from the CAI sampled for whole-rock geochemistry, U-Pb geochronology, and O and C isotopes come from three outcrops striped for exploration, trenches, and delineation drill holes (from North to South CR09-152, CR09-106 and CR09-125, Fig. 2b) made by the property's previous owner Minière Du Nord (MDN) company. Key areas of the stripped outcrops were mapped (Fig. 3a; for more details, see Groulier *et al.*, 2014). Samples from the drill holes are representative of the whole alkaline suite (Table 1).

## TABLE 1

### Whole-rock geochemistry

Major elements were measured by ICP-AES (Inductively Coupled Plasma - Atomic Emission Spectrometry) and trace elements by ICP-MS (Inductively Coupled Plasma - Mass Spectrometry) by the laboratory Activation Lab (Ancaster, Ontario; sample preparation and digestion methods for geochemistry packages available on http://www.actlabs.com/). Results are reported in Table 2.

### TABLE 2

## Electron probe microanalyser (EPMA)

Major and trace element compositions were measured using a Cameca computer controlled SX-100 equipped with five wavelength dispersive spectrometers (WDS) in the GeoRessouces laboratory (Nancy, France). For all the analyses the beam size was set to 1  $\mu$ m<sup>2</sup>, the background time for each measured element correspond to half the measurement time of the corresponding element, and each X-ray line has been carefully selected in order to prevent peak overlaps and, therefore, interferences.

For quantitative analyses of pyrochlore, major and trace elements (U, Th, Sn, Sc, Y, La, Ce, Mn, Ca, Sr, Ba, Pb, Na, K, W, Nb, Ta, Si, Ti, Zr, Al, Fe, F) were measured using an accelerating voltage of 20 kV and a beam current of 20 nA. Peak counting time was set to 10 s for Na, Al, Si, K, Ca, Ti, Mn, Fe, Sr, Y, Zr, Nb, Sn, Ta, W, Th and 20 s for Ba, La, Ce, Pb, U, F. Standards used for these measurements were uraninite (U), thorianite

(Th), cassiterite (Sn), Sc metal (Sc), Y-monazite (Y), monazite (La, Ce), geikielite (Mn, Ti), andradite (Ca), celestite (Sr), baryte (Ba), galena (Pb), albite (Na, Si), orthoclase (K), W metal (W), LiNbO<sub>3</sub> (Nb), LiTaO<sub>3</sub> (Ta), zircon (Zr), corundum (Al), hematite (Fe), F-apatite (F). The calculation correction used was PAP (Pouchou & Pichoir, 1991) and the pyrochlore composition was calculated following the formula A<sub>2-m</sub>B<sub>2</sub>X<sub>6-w</sub>Y<sub>1-n</sub> from Atencio *et al.* (2010) with m, w and n corresponding to the vacancies on the different sites. The cations U<sup>4+</sup>, Th<sup>4+</sup>, Sn<sup>4+</sup>, Sc<sup>3+</sup>, Y<sup>3+</sup>, La<sup>3+</sup>, Ce<sup>3+</sup>, Mn<sup>2+</sup>, Ca<sup>2+</sup>, Sr<sup>2+</sup>, Ba<sup>2+</sup>, Pb<sup>2+</sup>, Na<sup>+</sup> and K<sup>+</sup> were attributed to the A-site, W<sup>6+</sup>, Nb<sup>5+</sup>, Ta<sup>5+</sup>, Si<sup>4+</sup>, Ti<sup>4+</sup>, Zr<sup>4+</sup>, Al<sup>3+</sup> and Fe<sup>3+</sup> were attributed to the B-site, while F<sup>-</sup> and OH<sup>-</sup> were attributed to the X/Y-sites. The results for pyrochlore are summarized in Table 3 and detailed in Table S1.

#### TABLE 3

For quantitative analyses of apatite, major and trace elements (P, Ca, Cl, F, Na, Si, Mg, Fe, Sr, Y, Eu, Dy, Yb, S, Mn, Zr, La, Ce, Nd, Sm, Gd, Pb, Th, U) were measured using an accelerating voltage of 15 kV and a beam current of 12 nA. Peak counting time was set to 10 s for each element. The standards used for these measurements were apatite (Ca, P), albite (Na, Si), olivine (Mg, Fe), celestite (Sr), geikielite (Mn), monazite (La, Ce, Nd), baryte (S), topaz (F), vanadinite (Cl). Formula units were calculating based on 12.5 oxygens, based on an ideal formula of Ca<sub>5</sub>(PO<sub>4</sub>)<sub>3</sub>(F,Cl,OH). The cations Ca<sup>2+</sup>, Na<sup>+</sup>, Mg<sup>2+</sup>, Fe<sup>3+</sup>, Sr<sup>2+</sup>, Mn<sup>2+</sup>, La<sup>3+</sup>, Ce<sup>3+</sup> and Nd<sup>3+</sup> were attributed to the Ca-site, P<sup>5+</sup>, Si<sup>2+</sup> and S<sup>6+</sup> were attributed to the P-site, while F<sup>-</sup> and Cl<sup>-</sup> were attributed to the F-site. The results for apatite are provided in Table 4.

#### TABLE 4

## Trace element analyses of zircon using laser ablation-inductively coupled plasmasector field-mass spectrometry (LA-ICP-SF-MS)

Trace element analyses were carried out on zircon grains from the carbonate-bearing coarse-grained nepheline syenite (sample 12-PA-011-B11) at the Goethe University of Frankfurt (Germany), using laser ablation-inductively coupled plasma-sector field-mass spectrometry (LA-ICP-SF-MS). The laser spot size and the repetition rates were set to 50 µm and 5.5 Hz, respectively. Analysis time for each spot was overall 90 s, with 30 s washout, 30 s for the background (no ablation) and 30 s for the sample. The intensities of the following isotopes were recorded: <sup>44</sup>Ca, <sup>49</sup>Ti, <sup>57</sup>Fe, <sup>88</sup>Sr, <sup>139</sup>La, <sup>140</sup>Ce, <sup>141</sup>Pr, <sup>146</sup>Nd, <sup>147</sup>Sm, <sup>151</sup>Eu, <sup>157</sup>Gd, <sup>158</sup>Gd, <sup>159</sup>Tb, <sup>161</sup>Dy, <sup>165</sup>Ho, <sup>167</sup>Er, <sup>169</sup>Tm, <sup>172</sup>Yb, <sup>175</sup>Lu, <sup>232</sup>Th, and <sup>238</sup>U. The GJ-1 zircon, NIST SRM-612 and NIST SRM-614 reference materials were monitored as standards during the analytical session. Relative errors obtained for the analyzed trace elements on reference material are between 2 and 8% (1 sigma). Raw data of LA-ICP-SF-MS analyses were processed using the software GLITTER<sup>™</sup> (Griffin *et al.*, 2008). Results are summarized in Table 5 and detailed results are proposed in Table S2.

## TABLE 5

## U-Pb dating on zircon and apatite using LA-ICP-(SF-)MS

U-Pb analyses of zircon grains from the carbonate-bearing coarse-grained nepheline syenite (sample 12-PA-011-B11) were carried out by LA-ICP-SF-MS at the Goethe University of Frankfurt (Germany). Methods and instruments were described by Gerdes & Zeh (2006, 2009), with modifications explained in Zeh & Gerdes (2012). Results are summarized in Table 5 and detailed results for unknowns and standard measurements are provided in Table S3.

U-Pb analyses of apatite grains from carbonatites were carried out by LA-ICP-MS directly on thin sections at Géosciences Rennes (France) using an ESI NWR193UC excimer 193 nm laser coupled to a quadripole Agilent 7700x ICP-MS. Methods and instruments were described by Pochon *et al.* (2016), the operating conditions are given in Table S4. Results are listed in Table 6.

## TABLE 6

## *In-situ* ion probe oxygen isotope of zircon

Ion probe oxygen isotope analyses were performed at the Centre de Recherches Pétrographiques et Géochimiques (CRPG, Nancy, France) using the Cameca IMS 1270. Secondary Ion Mass Spectrometry (SIMS) oxygen isotope analyses of zircon were carried out using a Cs<sup>+</sup> primary ion beam accelerated at 10 kV with an intensity of 10 nA. The <sup>16</sup>O and <sup>18</sup>O ion intensities were measured simultaneously in multi-collection mode using two off-axis Faraday cups, as described in detail in Nemchin *et al.* (2006). Oxygen isotope are reported in  $\delta^{18}$ O notation as per mil deviation from the <sup>18</sup>O/<sup>16</sup>O of V-SMOW. The analyses were corrected for the instrumental mass fractionation using the 91500-zircon standard with a  $\delta^{18}$ O value of 9.86‰ (Wiedenbeck *et al.*, 2004) to which the data for unknowns was normalized. Results are summarized in Table 5 and detailed results are listed in Table S5.

## Oxygen and Carbon isotope analyses of carbonates

Coarse-grained calcite texture from the carbonatite facies was crushed with pestle and mortar to obtain powders. A mass of 11 to 12 mg of calcite powder was digested in anhydrous phosphoric acid ( $H_3PO_4$ ) in a double bottom reactor at a temperature of 50°C

during a few hours in order to complete the reaction. Analyses on the recovered  $CO_2$  gas were realized on a VG Optima triple collector mass spectrometer at Géosciences Rennes (France). The long term analysis of in-house standard Prolabo Rennes and international standard NBS 18 (Fen carbonatite) allows estimating the overall uncertainty at ±0.20‰ for O and ±0.10‰ for C. The results are listed in Table 7.

## TABLE 7

## **FIELD RELATIONSHIPS**

The previous work of Bergeron (1980) identified a complex assemblage of various types of silica-undersaturated rocks in the Crevier Alkaline Intrusion (CAI). An updated characterization of the main lithologies/facies constitutive of the intrusion was based on airborne aeromagnetic data, detailed mapping, drill-core logging, exploration trenches and stripped outcrops (Figs. 2b, 3a-c). They are described below following their relative timing of emplacement.

## Carbonate-bearing fine to coarse-grained nepheline syenite

The outer nepheline syenite (Fig. 2b) corresponds to a polyphase intrusive body with different facies from fine- to dominant coarse-grained texture (Fig. 3a). The coarse-grained type is carbonate-bearing and the dominant feldspathoid is nepheline but some grains of cancrinite and sodalite have been observed in drill cores. It forms most of the intrusion and is crosscut by fine- to medium-grained nepheline syenite dykes which are randomly injected (Fig. 3b). The rocks do not exhibit any evidence of solid-state deformation.

## **Banded core**

The banded core is oriented N320 (Fig. 2b) and interpreted to represent parallel dykes by Bergeron (1980). The banded core is composed of four main facies, namely (i) ijolite-urtite (ijolite consists of pyroxene and 30 to 70% modal nepheline while urtite are leuco-ijolite with more than 70% modal nepheline; Le Maitre, 2002), (ii) silicocarbonatite, (iii) nepheline syenite, and (iv) calcio- and ferrocarbonatite.

## Silicocarbonatites

Outside of the banded core, the silicocarbonatites appear as dykes that crosscut the outer nepheline syenite (Fig. 3a, c). They are crosscut by the pegmatitic nepheline syenite dyke swarms (Fig. 3a, c) and by calcio- and ferro-carbonatites (Fig. 3h).

## Pegmatitic nepheline syenite dyke swarms

Both the banded core and the surrounding nepheline syenite are crosscut in their center by a dyke swarm of pegmatitic nepheline syenite (Figs. 2b, 3a-d, 3f) and carbonatite (Figs. 3a, 3c, 3e-h). The dyke swarms of pegmatitic nepheline syenite are oriented N320, i.e. parallel to the banded core and to the general orientation of the CAI, with an average dip of 80° toward the North-East. They correspond to the main economic target due to their high Nb and Ta contents. Detailed mapping showed that these dykes were emplaced *en echelon* with an apparent dextral shear-movement (Fig. 3a). The dykes are characterized by margins of megacrystic nepheline and albite (up to 50 cm long) together with dark mica crystallizing inward from the dyke walls (Fig. 3d). The center of the dykes (feeder conduits) is composed of fine-grained nepheline syenite which corresponds to the liquid from which the pegmatitic nepheline syenite crystallized. In the dyke swarms, two facies of nepheline syenite can be differentiated based on their mineralogy and field relationships. The (1) most primitive (oldest) facies corresponds to carbonate-bearing nepheline syenite characterized by the presence of magmatic carbonates. One of the feeder conduits display a spectacular ocelli texture with spheres of carbonate-bearing nepheline syenite rimmed by dark mica in a carbonatite and apatite matrix (Fig. 4a-c). The (2) youngest facies correspond to a carbonate-free nepheline syenite which seems to be contemporaneous with the carbonatites. Both the carbonate-bearing and carbonate-free nepheline syenites are enriched in Nb and Ta as shown by the presence of pyrochlore as an accessory phase.

## Calcio- and ferro-carbonatites

In the southern part of the CAI the carbonatites appear as dykes (Fig. 3a, 3c) or veins closely associated with the dyke swarms of pegmatitic nepheline syenite (Fig. 3f). They are more abundant in the central and northern parts (banded core) of the CAI where they can form homogeneous bodies and dykes. The dominant calciocarbonatite and the more evolved ferrocarbonatite represent the youngest types of carbonatites within the complex (Fig. 3e, g).

## Nepheline-bearing alkali syenite and syenite

In the southern part of the complex a small stock of nepheline-bearing alkali syenite and syenite crosscut a fine to coarse-grained nepheline syenite that will not be further discussed in the manuscript (Fig. 2b).

## **FIGURE 3**

## PETROGRAPHY

## ljolite-urtite

The ijolite-urtite series of the banded core of the CAI is a fine-grained rock with a porphyritic texture marked by the presence of euhedral nepheline and perthitic alkali-feldspar phenocrysts with Carlsbad twins (Fig. 4d). The groundmass is composed of an assemblage of nepheline-alkali feldspar-phlogopite-aegirine with primary calcite. Accessory minerals are represented by euhedral titanite and interstitial Fe-Ti oxides. As the aegirine represents around 10% of the whole assemblage, this rock is designated as leuco-ijolite based on the classification of Mitchell (1996), and as urtite or ijolite based on the percentage of nepheline (30-70% in this case) according to the classification of Platt (1996).

## **FIGURE 4**

## Nepheline syenite

The dominant lithology found in the CAI is a holocrystalline and equigranular coarsegrained nepheline syenite that corresponds to the outer rim of the complex (Figs. 2b, 4e). Alkali feldspar is albite dominant, but perthitic orthoclase are commonly observed together with minor microcline. The grains commonly display undulose extinction, twin deformation and grain boundary migration. Nepheline is commonly altered into cancrinite, some grains of sodalite were observed in drill cores. In addition, in some facies, abundant primary carbonates (calcite) are found as interstitial phases either at the edges of nepheline and feldspar crystals or as inclusions in the perthitic feldspars, making them carbonate-bearing nepheline syenites (Fig. 4e-f, Table 2). The coarse-grained nepheline syenite is crosscut by fine- to medium-grained nepheline syenite injections, the mineralogy of which is similar to the coarse-grained nepheline syenite. Some dykes display medium-grained centers and fine-grained margins interpreted as chilled margins suggesting a difference in temperature between the cooler host rock and the intruding magma.

### Pegmatitic nepheline syenite and their fine-grained feeder conduits

The pegmatitic part is characterized by centimeter-size crystals of albite, alkali feldspar, nepheline and dark mica with minor proportions of sodalite and carbonates. Nepheline is often altered in cancrinite and the observed accessory minerals are zircon, pyrochlore, pyrite, pyrrhotite and Fe-Ti oxides. The center of the dykes is composed of fine-grained nepheline syenite with variable amounts of carbonate and mica. A weak magmatic foliation underlined by the albite lattes gives to the rock a fluidal, almost trachytic, texture. The feeder conduits are mostly composed of alkali feldspar, albite and nepheline and they are highly enriched in pyrochlore (up to 6%, Fig. 4g). Parts of the dyke swarms display evidence of magmatic-hydrothermal alteration such as crosscutting sodalite veins, nepheline with rims of cancrinite, hematization and carbonation. These zones are characterized by a modal enrichment in pyrochlore and zircon. The presence of carbonates makes the distinction between the older carbonate-bearing and the younger carbonate-free groups.

## Ocelli facies

This facies was observed in only one location. It represents a feeder conduit of a pegmatitic nepheline syenite dyke within one of the dyke swarms. This particular texture

is characterized by spheres of fine-grained and dark mica- and carbonate-bearing nepheline syenite rimmed by mica in a carbonatite and apatite matrix (Figs. 3a, 4a-c). In this facies, pyrochlore grains are observed only within the spheres of nepheline syenite.

## Silicocarbonatite

The silicocarbonatites are melanocratic due to a high modal proportion of dark mica (>40%) and are characterized by magmatic carbonate (>10%), plagioclase, aegirine with varying proportions in Fe-Ti oxides and accessory minerals (zircon, apatite; Fig. 3c, h). Two types of silicocarbonatite are observed in the intrusion: one is enriched in Fe-Ti oxides, and the other one is enriched in aegirine (Fig. 4h). The silicocarbonatites show a strong magmatic foliation underlined by micas {001} cleavage planes and they can be enriched in apatite. These silicocarbonatites form layers in the banded-core and dykes in the outer rim of nepheline syenite of the intrusion (Fig. 3a, c).

#### Carbonatite

On the basis of their texture and chemical composition (mainly Fe), the carbonatites are divided in two types: (i) a fine- to coarse-grained calciocarbonatite dominated by calcite (sövite composition; Figs. 3g, 4i), and (ii) a coarse-grained ferrocarbonatite composed of an assemblage of ankerite, calcite, hematite, and fluorapatite. Interstitial silicate phases such as phlogopite develop at the walls of the carbonatite veins/dykes when crosscutting the nepheline syenite. The two facies are pyrochlore-poor but contain fluorapatite (Fig. 4i), strontianite, REE-carbonates and monazite. Some calciocarbonatites are enriched in sulfides (pyrrhotite, chalcopyrite, molybdenite and argentite) and tellurides.

An envelope of fenitization is marked by the metasomatic crystallization of albite and arfverdsonite in the surrounding rocks (Fig. 2b). This process of fenitization is not further considered in this contribution.

## WHOLE-ROCK GEOCHEMISTRY

In the total alkali versus silica (TAS) diagram (Fig. 5a; Middlemost, 1994), the ijolite-urtite has an intermediate composition between the nepheline syenite and silicocarbonatite lithologies with an alkali content of 11.8 wt.% and a silica content of 43.8 wt.% (Fig. 5a, Table 2). The nepheline syenite defines a geochemical trend with increasing alkali (Na<sub>2</sub>O + K<sub>2</sub>O from 11.5 to 16.3 wt.%) and silica (SiO<sub>2</sub> from 46.4 to 59.8 wt.%) contents. The carbonate-bearing nepheline syenite are generally less evolved and fall in the nepheline monzosyenite fields with lower alkali and silica contents. The silicocarbonatite has lower alkali (from 7.7 to 9.6 wt.%) and silica (SiO<sub>2</sub> from 32.5 to 42.4 wt.%) contents, with higher Fe<sub>2</sub>O<sub>3(t)</sub> (from 4.53 to 16.4 wt.%) and CaO (from 6.6 to 12.7 wt.%, Fig. 5, Table 2) contents. The carbonatites (CaO from 49.8 to 52.4 wt.%) are not represented on this TAS diagram due to their very low SiO<sub>2</sub> content (ranging from 0.2 to 3.2 wt.%, Table 2). Their Mg number (MgO/[Fe<sub>2</sub>O<sub>3(t)</sub>+MgO]\*100) ranges from 30 to 45 (Table 2).

In the A vs B chemical-mineralogical diagram of Debon & Lefort (1988), the evolution from the ijolite towards the nepheline syenite is mineralogically expressed as an enrichment in nepheline, alkali feldspar and plagioclase and a decrease in clinopyroxene (Fig. 5b). Conversely, the ferro- and silico-carbonatite is associated with an enrichment in dark micas, and the calciocarbonatite with an enrichment in calcite (Fig. 5b).

#### FIGURE 5

The different lithologies display different features in Harker diagrams (Figs. 6a-d, see also Table 2). The nepheline syenite is characterized by high values in Na<sub>2</sub>O (from 6.4 to 12.6 wt.%) and Al<sub>2</sub>O<sub>3</sub> (from 16 to 25 wt.%), and low contents of MgO (usually <1 wt.%) and CaO (from 0.3 to 10.4 wt.%). The silicocarbonatite is enriched in MgO (from 1.5 to 9.4 wt.%), TiO<sub>2</sub> (from 0.7 to 4.9 wt.%), and CaO (from 6.6 to 12.7 wt.%), and has lower SiO<sub>2</sub> (between 32 and 43 wt.%) and Al<sub>2</sub>O<sub>3</sub> (between 10 and 16 wt.%) contents. The carbonatites are enriched in CaO (>28 wt.%), Sr (up to 9343 ppm) and REE ( $\Sigma$ REE up to 1414 ppm) with low SiO<sub>2</sub> (<4 wt.%), Na<sub>2</sub>O (<1 wt.%), and Al<sub>2</sub>O<sub>3</sub> (<2 wt.%) contents (Table 2). The ijolite facies displays an intermediate composition between the nepheline syenite and the silicocarbonatite, with SiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, MgO, CaO, Na<sub>2</sub>O and TiO<sub>2</sub> contents of 43.8 wt.%, 17.4 wt.%, 1.0 wt.%, 11.1 wt.%, 7.0 wt.%, and 0.7 wt.% (Fig. 6 and Table 2), respectively. The carbonate-bearing nepheline syenites show compositions that are closer to the ijolite than to the carbonate-free ones.

#### FIGURE 6

The ijolite contains a  $\Sigma$ REE content of 62 ppm and shows a chondrite-normalized REE pattern marked by an enrichment in the LREE relative to the HREE ([La/Yb]<sub>N</sub> = 12, Fig. 7, Table 2). The carbonate-free nepheline syenites contain  $\Sigma$ REE contents between 27 and 133 ppm (Table 2). They are variably enriched in LREE relative to chondrite ([La/Yb]<sub>N</sub> = 7.6-65, Fig. 7, Table 2) and they tend to have flat to slightly fractionated HREE chondrite-normalized patterns ([Gd/Yb]<sub>N</sub> = 0.8-6.3, Fig. 7, Table 2). The carbonate-bearing nepheline syenites are overall more enriched in REE than the carbonate-free nepheline syenites with  $\Sigma$ REE contents between 121 and 319 ppm (Table 2). The silicocarbonatite and carbonatite are more enriched in REE, the former showing  $\Sigma$ REE contents between

274 and 543 ppm, and the latter showing  $\Sigma$ REE contents between 1189 and 1414 ppm (Table 2). Both show a relative enrichment of the LREE compared to HREE in their chondrite-normalized patterns as shown by their [La/Yb]<sub>N</sub> of 19-37 and 19-32, respectively (Fig. 7, Table 2).

#### FIGURE 7

## TEXTURAL AND GEOCHEMICAL CHARACTERISTICS OF ACCESSORY MINERALS Apatite

Apatite is present in both the silicocarbonatite and carbonatite facies. Grains are (sub-)euhedra and range in size from ca 0.2 to 2.0 mm, but some display slight dissolution textures on the rims (Fig. 8a-b). They are rather homogeneous as shown by their CaO and  $P_2O_5$  contents of 54.90±0.13 wt.% and of 43.03±0.15 wt.% (Table 4), respectively, and are fluorapatite with an homogeneous F content of 3.16±0.04 wt.% (Table 4). They plot in the carbonatitic field defined by Piccoli *et al.* (2002) (Fig. 8c-e), and are thus considered of magmatic origin.

#### FIGURE 8

#### Pyrochlore

Previous studies have distinguished two types of pyrochlore, namely (i) primary magmatic grains with complex zoning (Fig. 9a) and (ii) secondary grown or extensively altered grains during the circulation of a magmatic-hydrothermal fluid with dissolution/reprecipitation zones (Groulier et al., 2014). For the purpose of this study we only focus on the magmatic pyrochlore in the studied rock types.

Magmatic pyrochlore is dominantly found in the late nepheline-bearing facies, and particularly within the dyke swarms of pegmatitic nepheline syenite. Even though silicocarbonatite contains some pyrochlore, it is very rare in the carbonatites. The few observed are small in size ( $\leq 10 \ \mu$ m) and highly metamict due to high U content, so that any EPMA analysis is difficult to undertake. Primary magmatic pyrochlore from the pegmatitic dyke swarms of nepheline syenites is euhedral and generally larger than 100  $\mu$ m (Fig. 9a, average of 500  $\mu$ m and up to ca. 2 mm). In the pegmatitic nepheline syenites, the pyrochlore grains are located along the edges of the megacrystic nepheline and albite where it is generally in spatial association with fluorapatite and ilmenite. They appear as disseminated grains within the fine-grained feeder conduits. The pyrochlore grains observed in the ocelli facies are found within the spheres of carbonate-bearing nepheline syenite where they are associated with carbonates and fluorapatite. BSE images reveal that the majority of the primary magmatic pyrochlore grains from all the different facies exhibit magmatic inclusions with complex zoning (Fig. 9a).

All the investigated magmatic grains plot in the pyrochlore field with Nb<sup>5+</sup>+ Ta<sup>5+</sup> > 2Ti<sup>4+</sup> and Nb<sup>5+</sup>> Ta<sup>5+</sup> (Hogarth, 1977; Fig. 9b). Two trends are noticeable: (i) a Nb-Ti trend with relatively constant Ta content in pyrochlore from the carbonate-poor to free pegmatitic nepheline syenite and their feeder conduits, and (ii) a Nb-Ta trend with constant Ti concentrations in the ocelli facies (Fig. 9b). Pyrochlore from the carbonate-poor to free nepheline syenite facies has Nb<sub>2</sub>O<sub>5</sub> contents ranging from 52.50 to 67.24 wt.% (mean of 61.49 wt.%), Ta<sub>2</sub>O<sub>5</sub> from 0.83 to 9.54 wt.% (mean of 6.23 wt.%), TiO<sub>2</sub> from 0.83 to 5.39 wt.% (mean of 2.77 wt.%) and UO<sub>2</sub> below 0.16 wt.% (mean of 0.01 wt.%). The pyrochlore from the carbonate-bearing ocelli facies has Nb<sub>2</sub>O<sub>5</sub> contents ranging from 53.22 to 64.48 wt.% (mean of 57.92 wt.%),  $Ta_2O_5$  from 4.77 to 16.06 wt.% (mean of 11.87 wt.%),  $TiO_2$  from 1.21 to 2.66 (mean of 1.79 wt.%), and  $UO_2$  below 1.94 wt.% (mean of 0.52 wt.%) (Tables 3 and S1).

In the Ta<sup>5+</sup> vs Ti<sup>4+</sup> diagram (Fig. 9c), the pyrochlore from the ocelli facies define a trend marked by an increase of the Ta content with minor variation in Ti implying a dominant  $Nb^{5+} = Ta^{5+}$  (1) isomorphic substitution due to the solid solution between pyrochlore [CaNaNb<sub>2</sub>O<sub>6</sub>(OH,F)] and microlite [CaNaTa<sub>2</sub>O<sub>6</sub>(OH,F)] (Hogarth, 1977; Nasraoui & Bilal, 2000; Atencio, 2010). This isomorphic substitution is responsible for the negative correlation between Nb and Ta (Fig. 9d), and for the variable Nb contents with constant Ti (Fig. 9e).

In the Ta vs Ti diagram (Fig. 9c), the pyrochlore from the carbonate-poor to free nepheline syenite follow a second trend characterized by an increasing Ti content with constant Ta implying a substitution between Nb and Ti. In the Ta vs Nb diagram (Fig. 9d), the pyrochlore from the carbonate-poor to free nepheline syenites show variations in the Nb content with no to few variations in the Ta content. They also display a linear negative correlation between Nb and Ti (Fig. 9e). The evolution of the Ta, Nb and Ti contents marks the replacement of Nb<sup>5+</sup> by Ti<sup>4+</sup> and the involvement of Na<sup>+</sup> and Ca<sup>2+</sup> following the reaction  $Ca^{2+} + Ti^{4+} = Na^+ + Nb^{5+}$  (2). This substitution scheme is dominant for the pegmatitic carbonate-poor to free nepheline syenite facies. It illustrates the solid solution between betafite [CaNaTi<sub>2</sub>O<sub>6</sub>(OH,F)] and pyrochlore [CaNaNb<sub>2</sub>O<sub>6</sub>(OH,F)] (Hogarth, 1977; Nasraoui & Bilal, 2000; Atencio, 2010).

#### **FIGURE 9**

Zircon

Zircon grains from the carbonate-bearing coarse-grained nepheline syenite (sample 12-PA-011-B1, outer nepheline syenite from Fig. 2b) can be separated in two main populations based on their textures and chemical compositions. The dominant population corresponds to (sub-)euhedral zircon grains with lengths ranging from 100 to 950 µm and aspect ratios of 1:1 to 1:2 (Figs. 10a and 11a). The CL (CathodoLuminescence) images reveal oscillatory zoned cores truncated by CL-bright rims of irregular shape perhaps formed by coupled dissolution-reprecipitation (Figs. 10a and 11a). The second population corresponds to anhedral zircon grains with lengths ranging from ca. 250 to 700 µm and aspect ratios of 1:1 to 1:2 (Figs. 10c and 11c). The CL images reveal complex oscillatory zoning and local dissolution-reprecipitation rims (Figs. 10c and 11c). No mineral inclusions have been observed in all zircon populations even using the highest SEM resolution in the range of ca. 50 nm.

The (sub-)euhedral and anhedral zircon grains show contrasting geochemical features. The (sub-)euhedral zircons show higher Th/U ratios (generally greater than 1 and up to 6.8), Ca + Fe contents (up to 28005 ppm),  $\Sigma$ LREE (up to 301 ppm), Sr (up to 2460 ppm) and Ti (up to 1536 ppm) contents, compared to the anhedral zircons with lower Th/U ratios (generally below 1 and down to 0.12) and Ca + Fe contents (down to 77 ppm),  $\Sigma$ LREE (between 0.5 and 16 ppm), Sr (between 0.1 and 25 ppm), and Ti (between 1.6 and 14 ppm) (Figure 12a-e, Table 5).

The REE chondrite-normalized patterns of anhedral zircon grains are characterized by (i) significantly higher HREE relative to LREE ( $[Sm/Yb]_N$  from 0.01 to 0.22) independent from the Ca and Fe contents (Fig. 12e, Table 5), (ii) a positive Ce anomaly (Ce/Ce\* from 1.31 to 25.83), and (iii) Eu (Eu/Eu\* from 0.54 to 1.05) anomalies (Fig. 12f, Table 5). In the (sub-)euhedral zircon grains the REE patterns are highly variable and become flatter with

increasing Ca+Fe and  $\Sigma$ LREE contents, as indicated by [Sm/Yb]<sub>N</sub> between 0.01 and 0.20, Ce/Ce\* between 1.28 and 14.04, and Eu/Eu\* between 0.29 and 1.04(Fig. 12e-f, Table 5). It is pertinent to note that trace element analyses obtained from cores and rims of the two (sub-)euhedral zircon grains cover nearly the entire compositional range estimated on all anhedral and (sub)euhedral zircon grains, whereby the cores commonly show higher contents in zircon-incompatible elements (Ca, Fe, LREE, Sr) compared to the related rims (Fig. 12).

## FIGURE 10 FIGURE 11 FIGURE 12

## Zircon crystallization temperatures

Temperatures calculated by applying the Ti-in-zircon geothermometer calibrated by Ferry and Watson (2007) vary between 584 and 1628°C, whereby the derived temperatures obviously increase systematically with the amount of zircon-incompatible elements (Fig. 12d). Zircon core and rim analyses of the anhedral zircon grains, all with relatively low contents in Ca, Fe, Sr and LREE, yield temperatures between 584°C and 781°C ( $2\sigma$  below 16°C, Ferry & Watson, 2007), whereas the (sub-)euhedral grains, having all much higher contents in Ca, Fe, Sr and LREE, yield temperatures between 724 and 1628°C ( $2\sigma$  below 16°C, Ferry & Watson, 2007) (Fig. 12d, Table 5). The highest temperatures (corresponding to the highest Ti contents) were calculated from domains having the highest Ca+Fe contents (as well as Sr and LREE), comprising both core and rim domains (for explanation see next chapter). Temperatures estimated from cores and rims of the

(sub-)euhedral two zircon grains cover nearly the entire temperature range obtained from the anhedral and (sub-)euhedral grains. Here, we note that most of the calculated temperatures are probably meaningless in a geological context for three main reasons. First, the investigated rocks contain neither rutile nor guartz, which are both required by the calibration of Ferry and Watson (2007). A reduced silica activity (aSiO<sub>2</sub>) will result in a temperature underestimate, and a reduced rutile activity (aTiO<sub>2</sub>) to temperature overestimate. Second, and even more important, correlation between Ti-in-zircon temperature and Ca+Fe suggest, that the analyzed Ti content was not only controlled by the zircon lattice (and associated thermal conditions), but very likely also by interstitial domains (see next chapter), causing an overestimation of the zircon crystallization temperature. Third, the composition of most rims, but also of core domains was modified by coupled dissolution-re-precipitation processes, which perhaps also affected the Ti contents. Taking these three processes into account, it seems that Ti-in-zircon temperatures obtained from the anhedral grains most likely represent minimum temperatures, and those obtained from the (sub-)euhedral grains maximum temperatures.

#### STABLE ISOTOPE DATA

Carbonates from the carbonatite facies of the CAI have  $\delta^{13}$ C values from -6.1‰ to -6.8‰ and  $\delta^{18}$ O values from +6.7 to +7.3‰ (Fig. 13, Table 7). In the  $\delta^{13}$ C <sub>v-PDB</sub> vs <sup>18</sup>O<sub>V-SMOW</sub> diagram (Fig. 13), data plot in the primary carbonatite field defined by Taylor *et al.* (1967). Both the (sub-)euhedral and anhedral zircon grains show a wide range of  $\delta^{18}$ O<sub>V-SMOW</sub> values spreading from the supra-mantle to sub-mantle field (mantle signature of +5.3±0.3‰ as defined by Valley *et al.*, 2005). The (sub-)euhedral and anhedral grains have  $\delta^{18}O_{V-SMOW}$  spreading between +2.9 and +7.8‰, and between +1.3 and +5.7‰ (Fig. 14, Table 6), respectively. The  $\delta^{18}O_{V-SMOW}$  values both populations are positively correlated to the Ca and Fe contents (Fig. 14, Table 6).

### FIGURE 13

#### FIGURE 14

## **U-Pb DATING OF ZIRCON AND APATITE**

Twenty-five zircon grains from the carbonate-bearing coarse-grained nepheline syenite have been investigated through 41 spot analyses. Plotted in a concordia diagram, 17 of these analyses above 99% Pb-Pb conc. define a concordant date of 957.5 ± 2.9 Ma (MSWD = 0.74, Fig. 15a, Table 5). The degree of concordance of the <sup>207</sup>Pb/<sup>206</sup>Pb apparent ages varies between the two dominant zircon populations as (sub-)euhedral grains show a strong decrease of the degree of concordance with increasing Ca and Fe contents (Fig. 15b), and the anhedral grains all show high degrees of concordance with variations that are not linked to the Ca and Fe contents (Fig. 15b, Table 5). Both populations show a dominant mode of distribution of their <sup>207</sup>Pb/<sup>206</sup>Pb apparent ages at ca. 957 Ma and a younger one at ca. 850 Ma and ca. 920 Ma for the (sub-)euhedral and anhedral zircon grains (Fig. 15c), respectively.

Fifteen apatite grains from the carbonatite have been investigated through 26 analyses. Due to the generally low radiogenic-Pb/common-Pb ratio, data were plotted in a Tera-Wasserburg diagram (Tera & Wasserburg, 1972). Data define a discordia yielding a lower intercept date of 934  $\pm$  25 Ma and an initial <sup>206</sup>Pb/<sup>207</sup>Pb ratio of 0.945 consistent with the model evolution of Stacey & Kramers (1975) (Fig. 15d).

## DISCUSSION

## Evidence for a silicate-carbonate liquid-immiscibility magmatic process

## The ocelli facies

Ocelli textures, like those described here (Fig. 4a-c), are rare in igneous rocks. Different magmatic processes have been proposed to explain the diversity of ocelli settings and habits (Leveson et al., 1966). Occurrences of ocelli/spherical assemblages that we refer here to as 'ocelli' have been described in alkaline/carbonatitic environment, including carbonatite (Lapin & Vartiainen, 1983), carbonatitic kimberlite (Haggerty & Fung, 2006), or ijolite (Zurevinski & Mitchell, 2015). At Shombole (Kenya-Tanzania), many nephelinite lavas contain globular carbonate-rich segregations, interpreted by Peterson (1989) and further by Kjarsgaard & Peterson (1991) as the result of a silicate-carbonate liquid immiscibility. Potter et al. (2017) described silicate spheroids within a carbonatite groundmass in the Oldoinyo Lengai 1933 lava, and these authors plead again for liquid immiscibility. This process has also been studied using melt inclusions entrapped in cocrystallising minerals such as nepheline, perovskite, pyroxene and magnetite (Kjarsgaard et al., 1995; Nielsen et al., 1997; Xu et al., 2004; Panina, 2005; Thompson et al., 2007; Guzmics et al., 2011, 2012; Sharygin et al., 2012; Sekisova et al., 2015; Potter et al., 2017; Mollex, 2017; Baudoin et al., 2018). These immiscible melts are proposed to be the parental magmas of nephelinitic silicate rocks, calciocarbonatite (sövite) (Guzmics et al., 2012), and in some examples sulfides (Guzmics et al., 2012; Demaiffe et al., 2013), either formed in plutonic or volcanic environments (Le Bas et al., 1981).

When plotted in the Freestone & Hamilton projection (1980), the whole rock composition of the nepheline syenite and silicocarbonatite lithologies of the CAI follow the nepheline-carbonatite melt miscibility gap proposed by Kjarsgaard & Peterson (1991; solvus at 0.5 GPa, 1050-900°C, Fig. 16). The silica-alkali poor composition of the carbonatite does not fit with this experimental miscibility gap. This may be due to the modification of the initial magmatic composition by late-stage magmatic or secondary processes leading to silica-alkali-depleted to -free carbonatite bulk composition. This latter composition, when compared to the geochemistry of carbonatitic melt inclusions and carbonatitic rocks in the same environment, has been interpreted by Guzmics et al. (2012) to be not representative of the initial magmatic composition. To summarize, we propose that the spectacular ocelli facies in the silicocarbonatite likely evidence a silicatecarbonate immiscibility process in the CAI. This would have led to coeval silicate-bearing and carbonate-bearing magmas. The silicate-bearing magmas are represented by preimmiscibility carbonate-bearing coarse-grained nepheline syenite and post-immiscibility carbonate-free pegmatitic facies, whereas the carbonate-bearing magmas are represented by syn-immiscibility ocelli facies and silicocarbonatites, and by postimmiscibility carbonatites (see below).

### Field relationships and whole-rock geochemistry

The field relationships and whole-rock geochemistry from rocks of the CAI are consistent with an (i) evolution from a common carbonate-bearing magma of ijolitic composition, the (ii) fractionation of the ijolitic magma into pre-immiscibility carbonate-bearing nepheline syenite and silicocarbonatites, the (iii) formation of the syn-immiscibility ocelli facies that records the segregation of silicate and carbonatitic liquids, and the (iv) post-immiscibility emplacement of the carbonate-poor to free dyke swarms of nepheline syenite (pegmatitic together with their feeder conduits) and the calcio- and ferro-carbonatites. Field observations show clear separation of the different rock units with no to few evidence of magma mixing or mingling as shown by the lack of enclaves and crosscutting relationships. This is consistent with the whole-rock geochemistry that highlights two distinct trends starting from an ijolite composition. Together they support the model of immiscibility rather than mixing between a silicate and a carbonatitic liquid.

#### Pyrochlore geochemistry as a tracer of immiscibility

Accessory minerals such as pyrochlore or REE-bearing minerals have been used to decipher carbonatite features, e.g. chemical evolution from plutonic to volcanic carbonatites, as well as magmatic to hydrothermal processes (Bambi *et al.*, 2012; McCreath *et al.*, 2012; Torro *et al.*, 2012). However, only few studies used these accessory minerals to document petrogenetic processes such as liquid immiscibility (Taubald *et al.*, 2004; Lee *et al.*, 2006; Morteani *et al.*, 2013; Walter *et al.*, 2018; Giebel *et al.*, 2019).

In the CAI, pyrochlore grains are predominantly associated with nepheline syenite, they display evidence of polyphase growth and they are associated with numerous episodes of dissolution and reprecipitation in the magmatic stage (Fig. 9a). We interpret some of these episodes to be linked to disequilibrium processes between the early crystallized generation of magmatic pyrochlore grains and the evolving magma, leading to the formation of several generations of primary magmatic pyrochlore grains.

It has been suggested that the pyrochlore composition can be linked to the evolution of the chemistry of the melt (Hogarth *et al.*, 2000; Jung Lee *et al.*, 2006; Zaitsev *et al.*, 2012),

therefore pyrochlore minerals can theoretically be used as a proxy for the geochemical evolution of the melt from which they crystallized. When compared to pyrochlore grains from other complexes, the pyrochlore from the CAI show some distinctive features. Indeed, the two trends observed in Fig. 9b are unusual at the scale of an alkaline complex. Pyrochlore generally follow one trend, the most common one in alkaline and peralkaline complexes being a dominant trend between Nb and Ti following the solid solution substitution scheme between pyrochlore and betafite (substitution (2), Chakhmouradian & Mitchell, 2002; Shrinivas & Bismayer, 2010; Cordeiro et al., 2011; McCreath et al., 2012). In highly evolved peraluminous rocks, pyrochlore mainly follow a Ta-Nb or a Ta-Ti trend (Ohnenstetter & Piantone, 1992). The pyrochlore from the ocelli facies of the CAI follow a Nb-Ta differentiation trend, whereas the grains from the younger pegmatitic nepheline syenite dykes follow a Nb-Ti trend (Fig. 9c-d). The ocelli facies (Fig. 4a-c) is characterized by spheres of carbonate-bearing nepheline syenite material separated by a matrix of fluorapatite-bearing carbonatite. The younger pegmatitic nepheline syenite dykes are characterized by a lack of primary magmatic carbonates, less fluorapatite and the pyrochlore grains they contain follow a more "classical" Nb-Ti trend (Fig. 9b). The greater Ta content of pyrochlore from the ocelli facies can be explained by the Nb/Ta partitioning behavior in the silicate/carbonate melt, the evolution of the fluorine fugacity (F fugacity) and the metal content of the melt. The K<sub>D</sub> Nb/Ta carb-sil at 0.8-0.9 kbar and temperature ranging from 965 to 1015°C is ~5.1 (Veksler et al., 1998a), indicating that these two elements can be strongly fractionated by an immiscibility process between a silicate and a carbonatitic liquid. Tantalum is likely to be incorporated into the silicate liquid, whereas Nb will more strongly partition into the carbonate liquid. Therefore,

immiscibility between carbonate and silicate liquids would lead to the formation of a Ta-"supersaturated" silicate melt.

Zheng *et al.* (2014) proposed that the concentration of Ta in pyrochlore could also be controlled by the F fugacity of the melt. The pyrochlore grains are closely associated with fluorapatite, and F and Ta contents are negatively correlated within pyrochlore grains (Table 3). The ocelli facies of the CAI is characterized by the presence of fluorapatite aggregates within the carbonatitic matrix. The crystallization of fluorapatite could lead to a decrease of the F fugacity of the melt. Accordingly, the combined action of segregation of a carbonate liquid from a silicate parental melt, and the decrease of the F fugacity would lead to the crystallization of an early generation of Ta-bearing pyrochlore that would subsequently result in a decrease of the available Ta content in the melt. After complete separation of the silicate and carbonatitic liquids, the fractionation between Nb and Ta most probably stopped, thus leading to the crystallization of Nb-rich pyrochlore in the pegmatitic nepheline syenite dykes with more constant Nb/Ta ratios and a more classic Nb-Ti solid solution trend (Fig. 9e).

To sum up, we propose that the two magmatic populations of pyrochlore observed within the nepheline-bearing rocks of the CAI result from the immiscibility between a carbonate and a silicate melt.

## Petrogenetic model for syenite-carbonatite immiscibility

Using the petrological, geochemical, and geochronological constraints obtained on the CAI (field and petrographic data, whole rock geochemical data, characteristics of accessory minerals that are zircon, pyrochlore and apatite), we propose a petrogenetic model of magmatic evolution for the CAI (Fig. 17).

### Stage 1. Emplacement of a mantle-derived ijolitic/urtitic magma

We propose that the ijolite/urtite magma emplaced within a medium to low pressure allochthonous crustal segment of the Grenville Province represents the initial melt of the CAI. Indeed, in spider, Harker and A-B diagrams, the ijolite/urtite composition always plots at the intersection between the silicate and carbonatitic trends (Figs. 5-7, Table 2). This rock is characterized by a nepheline + alkali feldspar + phlogopite + aegirine + apatite + calcite (± Fe-Ti oxides) assemblage (Fig. 17).

Stage 2. Fractionation of the initial melt and formation of nepheline syenite ± carbonates The ijolitic/urtitic magma experienced fractionation that led to the formation of the coarsegrained carbonate-bearing nepheline syenite that constitutes the main unit of the CAI. The observed crosscutting field relationships between the different units (Figs. 2-3) indicate that the coarse-grained carbonate-bearing nepheline syenite was emplaced before the pegmatitic nepheline syenite, the ocelli facies, the silicocarbonatites and the other carbonatites. In the carbonate-bearing nepheline syenite, the (sub-)euhedral zircon population is composed of zircon grains with high and positively correlated Ca, Fe, LREE (mainly La), Sr and Ti contents (Fig. 12b-d, Table 5). Such large ionic radii/low ionic charge trace element does not fit into the zircon lattice; therefore, their high contents are usually related to alteration processes (e.g. Geisler et al., 2001, 2007; Hanchar et al., 2001; Hoskin & Schaltegger, 2003; Zeh et al., 2014). However, it could also result from the entrapment of co-precipitated nano-inclusions into the zircon lattice during zircon growth as proposed by Turlin et al. (2019). This process might have been similar to one recently described by Zeh et al. (2018) to explain relatively high Th and REE contents in authigenic rutile. The striking correlation between Ti and Ca+Fe (REE, Sr) might point to co-precipitation of phases like perovskite (CaTiO<sub>3</sub>), even though pure perovskite can be excluded, as the amount of Ca+Fe commonly exceed the Ti amount by a factor 10 or more. Co-precipitation of nano-phases can also account for the observation that high amounts of Ca, Fe, LREE, Sr, and Ti were not only analyzed on zircon rims of the (sub)euhedral zircon grains, modified by a coupled dissolution-re-precipitation processes, but also on zircon cores still showing an oscillatory zoning (see Fig. 10). On the other hand, low contents in Ca, Fe, LREE, Sr of the anhedral zircon grains suggest zircon growth without co-precipitation and/or entrapment of any Ca-Fe-Ti-rich nanophase.

The finding of both zircon generations, with and without high Ca-Fe-LREE-Sr contents, as well as of intermediate zircon (see Fig. 12) within the same rock sample suggest that the melt environment changed significantly during zircon growth, allowing the entrapment of Ca-Fe-Ti-rich nanophases during an early stage of zircon growth, but not during a later stage. We interpret the (sub-)euhedral population to be the first zircon generation to crystallize in a carbonate-saturated liquid, followed by the intermediate zircon generation. This interpretation is in agreement with the commonly supra-mantle  $\delta^{18}O_{V-SMOW}$  values obtained from pristine cores of the (sub-)euhedral population, ranging between the mantle value of +5.3±0.3‰ (Valley, 2005) and +7.8‰, in agreement with syenite-carbonatite melt extraction from an enriched mantle source (Marks & Markl, 2017 and references therein). The sub-mantle values down to +2.9‰ (Fig. 15), might result from alteration of the primary zircon grains by a coupled dissolution-re-precipitation process after growth of the (sub)euhedral grains, in order to minimize the lattice stain caused by the incorporation of the abundant zircon-incompatible trace elements.
The anhedral zircon grains perhaps formed later in the syenite magma, after abundant extraction of carbonate-rich fluid from the initially mixed melt. This unmixing would not only prevent the co-precipitation of Ca-Fe-Ti-rich phases during zircon growth, but would also cause the fractionation of  $\delta^{18}O_{V-SMOW}$  between the carbonate-rich and silica-rich melt phase, responsible for the generally lower  $\delta^{18}O_{V-SMOW}$  of the anhedral zircon grains (+1.3 to +5.7‰) compared to the (sub)euhedral grains (+2.9 to +7.8‰; Fig. 14, Table 6); although post-growth alteration of the anhedral zircon grains cannot be excluded.

## Stage 3. Immiscibility between carbonatitic and silicate liquids

This stage corresponds to the silicate-carbonate liquid immiscibility process, i.e. to the physical separation of both liquids. The main arguments in favor of this specific immiscibility stage are: (i) the presence of coalescent nepheline syenite droplets/spheres in a carbonatitic matrix (ocelli facies; Fig. 4a-c), (ii) emplacement of silicocarbonatites and the ocelli facies with Ta-enriched pyrochlore grains, (iii) different  $\delta^{18}O_{V-SMOW}$  of (sub)euhedral and anhedral zircon, and finally (iv) dissolution/reprecipitation rim textures mainly around the (sub-)euhedral zircons (Figs. 10a, 11a, 17).

#### Stage 4. Post-immiscibility conditions

Following their complete separation, the silicate and carbonatitic liquids continued to evolve individually during a post-immiscibility stage. Pyrochlore grains associated with this stage within the pegmatitic nepheline syenite follow a Nb-Ti solid solution and are characterized by higher Nb/Ta ratios (Tables 3 and S1). The anhedral zircon grains of the coarse-grained nepheline syenite show low Ca, Fe, LREE, Sr and Ti contents, similar to

the ones of the rims in the intermediate zircon grains (Fig. 12b-d, Table 5). These features support crystallization of anhedral zircon in a carbonate-poor silicate liquid, i.e. post-immiscibility. In addition, except for one analysis that falls within the mantle domain ( $\delta^{18}O_{V-SMOW}$  values of +5.8±0.6‰), all  $\delta^{18}O_{V-SMOW}$  values of these zircon grains are in the submantle domain, with values between +1.3±0.4 and +5.1±0.5‰ (Fig. 14, Table 5). They indicate a crystallization in equilibrium with the carbonatites. Indeed, at 500-800°C, using the isotopic fractionation factor of Zheng (1993), zircon in equilibrium with carbonatites would have  $\delta^{18}O_{V-SMOW}$  values lower by 2.5 to 4.5‰ than the  $\delta^{18}O_{V-SMOW}$  values of the carbonatites that is of +6.7 to +7.3‰ (Fig. 4, Tables 5 and 7), namely values between +2.2‰ and +4.8‰, close to the range actually measured.

*In-situ* LA-ICP-MS U-Pb dating of zircon and apatite constrain the timing of emplacement of the nepheline syenite at 957.5  $\pm$  2.9 Ma (Fig. 15a) and of the carbonatite at 934  $\pm$  25 Ma (15d), respectively.

# The Crevier alkaline intrusion in the late-Grenvillian tectonic frame: implications for the formation of rare metal deposits in late- to post-collisional settings

The late stages of the Grenvillian Orogeny (ca. 980-950 Ma) are marked by the thermal relaxation of the continental crust following gravitational spreading and weakening of the convergence that led to orogenic collapse and thinning of the crust associated with low cooling rates of ca. 2-6°C/Ma (Turlin *et al.*, 2018). This is witnessed in the medium to low pressure crustal segment of the Allochthonous Belt (aM-L*P*, Fig. 1b) by the apatite diffusion modeling that yielded a closure temperature of apatite crystals of 550-450°C at ca. 960 Ma (Turlin *et al.*, 2018). Two mechanisms have been proposed to account for this

thermal evolution, a slab retreat (Turlin *et al.*, 2018) or a lithospheric foundering (also referred to delamination; Corrigan & Hanmer, 1997; Groulier *et al.*, 2018b) of the subcontinental lithospheric mantle. This late evolution of the Grenvillian Belt is associated with small volumes of late- to post-tectonic mantle-derived within-plate silica-saturated (sometimes fluorine-bearing) or silica-undersaturated alkaline magmas.

In the central Grenville Province, this event is characterized by the intrusion of syenitic bodies of the undersaturated CAI at ca. 960-955 Ma (U-Pb on zircon, this contribution). It is composed of nepheline syenite and carbonatitic rocks consistent with low-degree of partial melting of a volatile-rich mantle source (Le Bas, 1987; Gittins, 1989; Bailey, 1993; Winter, 2001; Mitchell, 2005; Millong & Groat, 2013). Such mantle enrichment is expected after the extensive metasomatism of the Grenvillian subcontinental lithospheric mantle during the Proterozoic as the southeastern Laurentian margin was an active margin over a period of 500 Ma (Rivers & Corrigan, 2000; Corriveau & Morin, 2000; Whitmeyer & Karlstrom, 2007; Rivers et al., 2012; Groulier et al., 2018b). The shape of the CAI indicates a strong structural control following the general orientation of a major crustal lineament known as the Waswanipi-Saguenay corridor (Bernier & Moorhead, 2000). This major crustal lineament is characterized by the presence of the Paleoproterozoic Nb-REEbearing Montviel carbonatitic complex (David et al., 2006) and by the Neoproterozoic Nbbearing Saint-Honoré carbonatite (Tremblay et al., 2015). The presence of three major Nb(±Ta in the CAI) deposits along this lineament highlights the importance of two crucial factors for the formation of such deposits: (i) a fertile and enriched subcontinental lithospheric mantle source able to produce carbonatitic and undersaturated magma, and (ii) a crustal weakness to facilitate the ascent and emplacement of the Nb(±Ta)-bearing

magma within the crust during extensional tectonic setting (late/post-tectonic to anorogenic).

Accordingly, the ca. 980-950 Ma late- to post-Grenvillian period is prospective for the formation of  $Nb(\pm Ta)$ -bearing magmas as it is characterized by discrete and low-degree partial melting of the mantle in an overall extensional tectonic setting.

#### CONCLUSION

Petrogenetic relationships between different undersaturated alkaline igneous rocks, carbonate-bearing and carbonate-free nepheline syenite and carbonatites observed in the Crevier Alkaline Intrusion (CAI) are proposed to result from the immiscibility between a carbonate and a silicate melt. The spectacular ocelli facies observed in the pegmatitic nepheline syenite dyke swarms is considered to evidence the syn-immiscibility stage, whereas the carbonate-free nepheline syenite and the calcio- and ferro-carbonatites are considered to be post-immiscibility. Accessory minerals (pyrochlore, zircon, and apatite) record geochemical and isotopic features that (i) argue for this immiscibility process, and (ii) allow to constrain the thermal conditions of immiscibility at a minimum temperature of ~815-865°C and its timing (U-Pb on zircon of the nepheline syenite at 957.5  $\pm$  2.9 Ma, and on apatite of the carbonatite at  $934 \pm 25$  Ma). Two generations of pyrochlore have been described: (i) a syn-immiscibility Ta-U-Ca-enriched type, and (ii) a post-immiscibility Nb-Ti-Th-Na-enriched type. We propose that the behavior of Nb and Ta is strongly connected to the immiscibility process and to the F-fugacity of the melt. This ca. 980-950 Ma late- to post-Grenvillian period is prospective for the formation of Nb(±Ta)-bearing magmas as it is characterized by discrete and low-degree partial melting of the mantle in an overall extensional tectonic setting.

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#### **Figure captions**

Figure 1. (a) Position of the Grenville orogen in Laurentia: Paleozoic and younger cover are omitted (modified from Hoffman, 1989; Rivers, 2008; Rivers et al., 2012). The northern dashed line represents the boundary between Internal and External Paleoproterozoic Laurentia and the southern dashed line represents the Grenville Front (GF). Abbreviations: 1 = exposed Grenville Province, light grey represents the inferred extension of subsurface allochthonous Grenville Province; 2 = Granite-Rhyolite Igneous Province, ca. 1.50-1.34 Ga and reworked equivalents in the Grenville Province; 3 = Mid-Continental Rift system; 4 = Paleoproterozoic orogens, ca. 1.9-1.8 Ga, ca. 1.65 Ga and reworked equivalents in the Grenville Province; 5 = Archean cratons; GF = Grenville Front. (b) Simplified tectonic map of the Grenville Province (modified after Rivers, 2008; Rivers et al., 2012). Letters in stars represent the localization of the main alkaline suites intruding the Allochthonous Belt of the Grenville Province from the early-Neoproterozoic to Mesozoic and their association to post-Grenvillian grabens: 1 = Crevier alkaline suite  $(958.7 \pm 3.7 \text{ Ma})$ ; 2 = Manitou and Iron Islands alkaline complexes (ca. 577 Ma); 3 = St-Honoré (ca. 582.1±1.8 Ma); 4 = Oka (ca. 124 Ma). Abbreviation: ABT = Allochthon Boundary Thrust.

**Figure 2.** (a) Simplified regional geology of the Crevier alkaline suite. (b) Detailed mapping of the Crevier alkaline intrusion (modified after Bergeron, 1980; Groulier *et al.*, 2014). Numbers refer to the drill core samples. Abbreviation: NTS = National Topographic System.

**Figure 3.** (a) Detailed mapping of the exploration stripped outcrops located in the central south-eastern part of the CAI highlighting the different crosscutting relations. (b) Detailed mapping of the different facies of nepheline syenite in the main unit. (c) Field photograph showing the relation between the pegmatitic nepheline syenite dyke swarm, the silicocarbonatite and the fine- to coarse-grained nepheline syenite (nepheline syenite on the figure). (d) Dykes of pegmatitic nepheline syenite together with their finer-grained feeder conduits and characterized by walls of megacrystic nepheline and albite (up to 50 cm long) together with dark mica crystallizing inward from the dyke walls. (e) Ferrocarbonatite with iron carbonate and calcite. (f) Pegmatitic nepheline syenite crosscuts by veins of carbonatite with biotite rims. (g) Fragment of coarse-grained carbonatite within fine-grained carbonatite. (h) Hand specimen of silicocarbonatite.

**Figure 4.** Microphotographs of the lithologies from the Crevier alkaline intrusion. (a) to (c) Macrophotographs of the spectacular ocelli-textured/oecelli facies, with spheres of carbonate-bearing nepheline syenite rimmed by mica in a carbonatite and apatite-rich matrix. (d) ljolite characterized by porphyritic texture and marked by euhedral nepheline, clinopyroxene of aegirine composition, perthitic alkali-feldspar phenocrysts with Carlsbad twins, and titanite as accessory mineral. (e) and (f) Nepheline syenite with albite, perthitic alkali feldspar and nepheline locally altered in cancrinite. (g) holocrystalline and equigranular coarse-grained nepheline syenite, with perthitic alkali feldspars, albite, nepheline, commonly altered in cancrinite, and pyrochlore and Fe-oxides as accessory minerals. (h) aegirine-mica-albite-carbonate-bearing silicocarbonatite enriched in Fe-Ti oxides. (i) Coarse-grained carbonatite with fluoroapatite grains. Abbreviations: Ab = albite;

Aeg = aegirine; Ap = apatite; Cal = calcite; Carb = carbonates; Ccn = cancrinite; Coarseg. = coarse-grained; Ne/Neph. = nepheline; Ox = oxide; Pcl = pyrochlore; Phl = phlogopite.

**Figure 5.** (a) Total alkali versus silica (TAS) diagram (Middlemost, 1994) showing the medium composition of ijolite (black star) between the carbonate-bearing and carbonate-free nepheline syenite (dark grey symbols) and the silicocarbonatite (white symbols). (b) Peraluminous index A (Al-(K+Na+2Ca)) versus differentiation index B (Fe+Mg+Ti) diagram (Debon & Le Fort, 1988) placing the location of the mineralogical poles (muscovite, biotite, quartz, alkali feldspar, pyroxene, aegirine, nepheline and calcite).

**Figure 6.** Harker diagrams showing the relation between SiO<sub>2</sub> (wt.%) and Na<sub>2</sub>O (wt.%) (a), Al<sub>2</sub>O<sub>3</sub> (wt.%) (b), CaO (wt.%) (c), and TiO<sub>2</sub> (wt.%) (d) contents. Two geochemical trends are identified from an initial ijolitic composition, one with the nepheline syenite lithologies, the other one with the silicocarbonaite lithologies, same legend than in Fig. 5, white triangle represents the carbonatite.

**Figure 7.** REE spectra of the different Crevier alkaline intrusion facies. All facies display fractionation between LREE and HREE. The ijolite is among the least enriched rocks of the CAI, its REE pattern is marked by little fractionation between LREE and HREE. The fine- to coarse-grained nepheline syenite facies are variably enriched in LREE relative to chondrite as it tends to have flat HREE patterns. <sup>1</sup>: Normalization after McDonough and Sun (1995).

**Figure 8.** (a) and (b) SEM images of apatite from the carbonatite. (c) to (e) Mn-Na-Ca, Mn-Fe-Ca and Mn-Fe-Ce ternary plots of apatite compositions. Shaded areas represent apatite formed from carbonatitic melts (Piccoli & Candella, 2002).

Figure 9. Petrographic and geochemical features of pyrochlore observed in the nepheline syenite, pegmatitic nepheline syenite and carbonate-bearing ocelli-textured/orbicular facies. (a) The BSE images reveal that a majority of primary magmatic pyrochlore grains exhibit complex zoning and intern patterns. (b) Nb-Ti-Ta ternary diagram showing the fields of composition of pyrochlore, betafite and microlite. The primary grains analyzed in the CAI are all falling in the pyrochlore field (Nb-dominant). Pyrochlore grains from the peqmatitic nepheline svenite (silicate-bearing) and carbonate-bearing ocellitextured/orbicular nepheline svenite show different trend of evolution. (c) Ta<sup>5+</sup> vs Ti<sup>4+</sup> diagram (values in a.p.f.u.) displaying the different substitution trends for pyrochlore from silicate-bearing and carbonate-bearing facies. Pyrochlore grains from carbonate-bearing facies define a trend marked by an increase of the Ta content with small variations in Ti, implying that the isomorphic substitution Nb<sup>5+</sup> = Ta<sup>5+</sup> is dominant (d) Diagram Nb<sup>5+</sup> vs Ta<sup>5+</sup> (values in a.p.f.u.), the arrows highlight the substitution trends for pyrochlore from NS and CS facies. (e) Ti<sup>4+</sup> vs Nb<sup>5+</sup> diagram (values in a.p.f.u.) highlighting the isomorphic substitution of Nb<sup>5+</sup> by Ta<sup>5+</sup> with the pyrochlore from the carbonate-bearing facies that form a linear trend. (f) Ta<sup>5+</sup> vs Nb/Ta plot showing the distribution of values for pyrochlore grains from silicate- and carbonate-bearing facies. Abbreviations: Ab = albite; Cal = calcite; Pcl = pyrochlore; Phl = phlogopite.

**Figure 10**. Cathodoluminescence images of the zircon grains of the coarse-grained nepheline syenite. The indicated zircon number refers to Table 5, where all analyses are summarized. Trace elements (blue) and  $\delta^{18}$ O (red) results are plotted on each zircon. (a) (Sub)-euhedral zircon grains showing oscillatory zoned cores and sharp dissolution-reprecipitation rims, (b) Zircon grains n°15 and n°19 that distinguished themselves from the populations (a) and (b), as they present significant geochemical variations from core to rim (see Fig. 10 and Table 5), and (c) Anhedral zircons grains, with complex oscillatory zoning and local dissolution-reprecipitation rims.

**Figure 11.** Cathodoluminescence images of the zircon grains of the coarse-grained nepheline syenite. The indicated zircon number refers to Table 5, where all analyses are summarized. U-Pb LA-ICP-MS results are plotted on each zircon, with indication of degree of discordance. (a) (Sub)-euhedral zircon grains showing oscillatory zoned cores and sharp dissolution-reprecipitation rims, (b) Zircon grains n°15 and n°19 that distinguished themselves from the populations (a) and (b), as they present significant geochemical variations from core to rim (see Fig. 10 and Table 5), and (c) Anhedral zircons grains, with complex oscillatory zoning and local dissolution-reprecipitation rims.

**Figure 12.** Trace element geochemistry of the zircons of the coarse-grained nepheline syenite (a) Ca + Fe (ppm) vs Th/U diagram. (b) Ca + Fe (ppm) vs  $\Sigma$ LREE (ppm) diagram. (c) Ca + Fe (ppm) vs Sr (ppm) diagram. (d) Ca + Fe (ppm) vs Temperature (°C) calculated with Ti-in-zircon thermometer ± 60 to 70 °C (Ferry & Watson, 2007). Note that the Ti-in zircon thermometer yields consistent temperatures for Ca-Fe-poor grains whereas

calculated temperatures for Ca-Fe-rich grains are unrealistic. (e) Ca + Fe (ppm) vs  $Sm_N/Yb_N$  diagram. (f) REE patterns normalized to chondrite.

**Figure 13.**  $\delta^{13}$ C (‰ PDB) vs  $\delta^{18}$ O (‰ SMOW) diagram of carbonates from carbonatites. The grey area defines the primary carbonatite field (Taylor *et al.*, 1967). Arrows indicate the main processes responsible for isotopic composition variations (Deines, 1989; Démeny *et al.*, 1998). Abbreviations: High T = high temperature; Low T = low temperature.

**Figure 14.** (a) Discrimination of the (Sub)-euhedral and anhedral zircon population regarding their δ<sup>18</sup>O composition. (Sub)-euhedral zircon are interpreted as witness of the pre- to syn-immiscibility with carbonate signature, whereas anhedral zircon may be more characteristic of post-immiscibility with no carbonate contribution in their  $\delta^{18}O_{V-SMOW}$  signature. Light grey area represents the mantle zircon value, as the dark grey area the carbonatite  $\delta^{18}O_{V-SMOW}$  composition. (b)  $\delta^{18}O_{V-SMOW}$  vs Ca + Fe (ppm) diagram of zircon showing the positive correlation between the  $\delta^{18}O_{V-SMOW}$  signatures and the Ca + Fe contents of zircon grains. Note that anhedral Ca-Fe-poor zircon have  $\delta^{18}O_{V-SMOW}$  signatures that range from 1.9 and 3.15‰ whereas (sub)-euhedral Ca-Fe-rich zircon have  $\delta^{18}O_{V-SMOW}$  signatures that trend towards the carbonates' signatures. A temporal path is proposed for zircon crystallization from mantle zircon value to carbonatite composition before or during the immiscibility process, ending with low values following immiscibility as observed in zircon from a carbonate-free syenite.

**Figure 15.** (a) LA-ICP-MS U-Pb ages of zircon from the coarse-grained nepheline syenite facies (SNg) (n = 17); (b)  $^{206}$ Pb/ $^{207}$ Pb age degree of concordance (%) vs the Ca+Fe content in ppm diagram of zircon. It expresses the correlation between high level of concordance and the low content of Ca+Fe. (c) Probability diagram of U-Pb ages on zircon; (d) LA-ICP-MS U-Pb ages of apatite from the carbonatite facies (n = 24).

**Figure 16.** Plot of the different lithologies of the CAI in the Freestone and Hamilton projection (1980). The whole rock composition of the nepheline syenite and silicocarbonatite lithologies follow the nepheline-carbonatite melt miscibility gap proposed by Kjarsgaard and Peterson (1991; solvus at 0.5 GPa, 1050-900°C). The carbonatite composition does not fit with this experimental miscibility gap, with a silica-alkali-poor composition. This may be link to the modification of the initial magmatic composition by late-stage magmatic or secondary processes leading to silica-alkali-depleted to -free carbonatite bulk compositions.

**Figure 17.** Synthesis of the results presented in this study allowing to decipher the evolution of the Crevier alkaline intrusion through the immiscibility process between a syenitic and a carbonatitic melts. Abbreviations: Allocht. Crust = Allochthonous crust; Cal = calcite; Ccn = cancrinite; Ne/Neph. = nepheline; PhI = phlogopite; Temp. = temperature; Zrc = zircon.

#### **Table caption**

**Table 1.** Present coordinates and results presented in this study of the various facies of the Crevier alkaline intrusion from the central Grenville Province sampled for this study, ordered by outcrop. Abbreviation: Neph. = nepheline; Cb-bearing NS = Carbonate-bearing nepheline syenite.

**Table 2.** Whole-rock geochemistry of the plutonic rocks from the various facies identified in the Crevier alkaline intrusion from the central Grenville Province, ordered by rock type. Abbreviations: C = (Fe,Ca)-carbonatite; Cb-bearing NS = carbonate-bearing nepheline syenite; LOI = loss on ignition; Mg# = (Mg)/(Mg+Fe); NS = nepheline syenite; NSg = coarse-grained nepheline syenite; NSp = pegmatitic nepheline syenite; SC = Silicocarbonatite.

**Table 3.** Representative magmatic pyrochlore compositions and structural formula from different syenite facies of the Crevier alkaline intrusion from the central Grenville Province obtained using EPMA. Abbreviations: apfu = atoms per formula unit; Neph. Sye. = nepheline syenite. General formula:  $A_{2-m}B_2X_{6-w}Y_{1-n}$  (Atencio et al., 2010) with m, w and n corresponding to the vacancies of the different sites.

**Table 4.** Apatite compositions from carbonatite facies (sample 12-PA-007-A6) of the Crevier alkaline intrusion from the central Grenville Province obtained using EPMA. They are reported with confidence interval of 95% in brackets (n = number of analyses).

**Table 5.** Summary of the results obtained on zircon grains from the carbonate-bearing nepheline syenite (sample 12-PA-011-B11) of the Crevier alkaline intrusion from the central Grenville Province. The zircon grains are grouped according to the classification of Figs. 10 and 11. Abbreviation: Temp. = temperature.

**Table 6.** U-Pb dating of apatite from carbonatite facies (sample 12-PA-007-A6) of the Crevier alkaline intrusion from the central Grenville Province.

**Table 7.**  $\delta^{18}O_{V-SMOW}$  and  $\delta^{13}C_{V-PDB}$  composition of calcite from carbonatite facies (samples 12-PA-019-A, 12-PA-007-A6 and 12-PA-009-A5) of the Crevier alkaline intrusion from the central Grenville Province.

**Table S1.** Magmatic pyrochlore compositions and structural formula from different syenite facies of the Crevier alkaline intrusion from the central Grenville Province obtained using EPMA. apfu = atoms per formula unit; Neph. Sye. = nepheline syenite. General formula:  $A_{2-m}B_2X_{6-w}Y_{1-n}$  (Atencio et al., 2010) with m, w and n corresponding to the occupancy of the different sites.

**Table S2.** Results of trace elements LA-ICP-MS measurements of zircon grains from the carbonate-rich coarse-grained nepheline syenite (sample 12-PA-032) of the Crevier alkaline intrusion from the central Grenville Province. Abbreviation: PGD = pegmatitic granite dyke.

**Table S3.** Results of LA-ICP-MS U-Pb dating of zircon grains from the carbonate-rich coarse-grained nepheline syenite (sample 12-PA-032) of the Crevier alkaline intrusion from the central Grenville Province. Abbreviation: conc. = degree of concordance.

**Table S4.** Operating conditions for the LA-ICP-MS equipment dedicated to U-Pb dating of apatite grains.

**Table S5.** Results of SIMS O isotope analyses of zircon grains from the carbonate-bearing coarse-grained nepheline syenite (sample 12-PA-032) of the Crevier alkaline intrusion from the central Grenville Province.



## Figure 1 (black and white; 2 columns fitting)





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Manuscript submitted to Journal of Petrology Figure 3 (a) to (h) (color; 2 columns fitting)



# Figure 4 (color; 2 columns fitting)






## Figure 6 (black and white; 2 columns fitting)



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## Figure 8 (black and white; 2 columns fitting)

Compositional field of apatite formed from carbonatitic melts (Piccoli & Candella, 2002)



Figure 9 (black and white; 2 columns fitting)





- 245 ← Ca+Fe (ppm)
- 3.05 ← δ<sup>18</sup>O (‰ V-SMOW) value



Legend: ○ U-Pb analysis 961 ← <sup>206</sup>Pb/<sup>207</sup>Pb age (Ma) (100) ← Degree of concordance (%)



## Figure 12 (black and white; 2 columns fitting)







Figure 15 (black and white; 1 column fitting)



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Figure 16 (black and white; 2 columns fitting)



Figure Man (color, white age, lands caperolientation)



**Table 1.** Present coordinates and results presented in this study of the various facies of the Crevier alkaline intrusion from the central Grenville Province sampled for this study, ordered by outcrop. Abbreviation: Cb = carbonate; Neph. = nepheline.

					Pyrochlore	Apat	tite		Zircon		
Sample <sup>1</sup>	Easting <sup>2</sup>	Northing <sup>2</sup>	Facies	Whole-rock geochemistry <sup>3</sup>	Mineral composition <sup>4</sup>	Mineral composition <sup>5</sup>	U-Pb dating <sup>6</sup>	Trace elements <sup>7</sup>	U-Pb dating <sup>8</sup>	O isotopes <sup>9</sup>	O, C isotopes <sup>10</sup>
002-G1	661630	5481331	Neph. syenite		X						
004-C1	661307	5481637	Neph. syenite	х							
004-C2			Neph. syenite	х	х						
006-C2	661278	5481665	Neph. syenite		х						
006-C3			Neph. syenite	х	х						
007-A6	661230	5481797	Carbonatite			Х	х	х			x
007-B1			Neph. syenite	х							
007-D1			Ijolite	х							
007-E1			Silicocarbonatite	х							
007-E4			Silicocarbonatite	х							
007-E8			Silicocarbonatite	х							
007-E10			Cb-bearing Neph. syenite	х							
008-A1	662009	5481013	Carbonatite	х							
008-B1			Neph. syenite	х							
008-B4			Silicocarbonatite	х							
008-C1			Cb-bearing Neph. syenite	х							
008-D1			Cb-bearing Neph. syenite	х							
008-E2			Silicocarbonatite	х							
008-E3			Cb-bearing Neph. syenite	х							
009-A2	658768	5484886	Carbonatite	х							
009-A5			Carbonatite	х							х
009-B2			Cb-bearing Neph. syenite	х							
009-D1			Silicocarbonatite	х							
009-D2			Silicocarbonatite	х							
009-E6			Silicocarbonatite	х							
009-E8			Silicocarbonatite	х							
010-D1	662126	5480819	Silicocarbonatite	х							
010-D2			Neph. syenite	х							
010-G			Ocelli facies	х							
010-G1			Ocelli facies	х	х						
010-G2			Ocelli facies		Х						
011-B1	662159	5480785	Neph. syenite	х				х	х	х	
011-B2			Neph. syenite	х							
011-B3			Cb-bearing Neph. syenite	х							
011-G1			Neph. syenite	x							
012-G	661313	5481636	Neph. syenite	x							
012-G2			Cb-bearing Neph. syenite	х							
012-G3			Neph. syenite								

012-G4			Neph. syenite		Х		
013-D1	661251	5481687	Silicocarbonatite	х	Х		
019-A	658326	5485165	Carbonatite	х			Х
019-G1			Silicocarbonatite	х			
019-G2			Neph. syenite	х			
019-E1			Cb-bearing Neph. syenite	х			
019-E2			Silicocarbonatite	х			
Sample names follow the format 12 PA YYY YY							

<sup>1</sup> Sample names follow the format *12-PA-XXX-XX*.

<sup>2</sup> Coordinates are reported as Universal Transverse Mercator (UTM) zone U18 and the National Topographic System (NTS) map sheet is 32H07.

<sup>3</sup> Whole rock geochemistry data are reported in Figs. 5 to 7 and in Table 2.

<sup>4</sup> A summary of pyrochlore compositions is reported in Table 3 and all data are available in Table S1. Data are represented in Fig. 9.

<sup>5</sup> Apatite compositions are reported in Fig. 8 and in Table 4.

<sup>6</sup> U-Pb dating analyses of apatite grains are reported in Fig. 15(d) and in Table 6.

<sup>7</sup> A summary of trace elements analyses of zircon grains is reported in Table 5 and all data are available in Table S2. Data are represented in Fig. 12.

<sup>8</sup> A summary of U-Pb dating analyses of zircon grains is reported in Table 5 and all data are available in Table S3. Data are represented in Fig. 15(a-b).

<sup>9</sup> A summary of oxygen isotopes analyses of zircon grains is reported in Table 5 and all data are available in Table S5. Data are represented in Fig. 14.

<sup>10</sup> Oxygen and carbon isotopes analyses of carbonates are reported in Fig. 13 and in Table 7.

**Table 2.** Whole-rock geochemistry of the plutonic rocks from the various facies identified in the Crevier alkaline intrusion from the central Grenville Province, ordered by rock type. Abbreviations: C = (Fe,Ca)-carbonatite; Cb-bearing NS = carbonate-bearing nepheline syenite; LOI = loss on ignition; Mg# = (Mg)/(Mg+Fe); NS = nepheline syenite; SC = Silicocarbonatite.

Facies	С	С	С	С	SC	SC	SC	SC	SC	SC	SC	SC	SC	SC
Sample <sup>1</sup>	009-A2	009-A5	008-A1	019-A	007-E1	019-E2	007-E4	007-E8	008-E2	008-B4	009-D1	009-E6	013-D1	009-D2
$SiO_2$ (wt.%)	3.20	0.34	2.34	0.16	35.31	35.69	39.95	36.76	41.98	32.52	40.19	39.30	39.57	37.49
	1.22	1.70	35.13	1.18	14.80	15.57	9.79	16.43	12.03	14.31	12.07	12.23	16.39	13.86
FeO	1.24	1.39	31.79	0.93	9.75	10.35	7.34	12.46	10.77	8.42	8.80	8.10	11.39	10.06
Fe <sub>2</sub> O <sub>3</sub>	0.16	0.17	0.16	0.15	4.07	4.18	1.72	2.72	1.94	5.05	2.84	3.07	3.86	2.79
MgO MnO	0.64	0.55	0.65	0.41	4.27	5.34 0.22	2.51	9.36	3.95 0.23	6./1 0.24	7.00 0.17	4.05	4.51 0.14	7.59 0.17
CaO	49.73	51.78	28.29	52.44	12.44	9.34	12.28	8.18	8.27	12.05	7.89	11.05	7.59	8.42
Na <sub>2</sub> O	0.58	0.09	0.12	0.07	3.22	2.72	4.83	2.19	3.46	2.10	3.49	4.01	3.38	2.63
K <sub>2</sub> O	0.24	0.03	0.43	0.01	4.55	5.25	4.56	6.72	5.58	5.59	5.60	4.08	5.37	5.96
$P_2O_5$	0.01	0.02	0.01 5.84	0.00	5.92 2.42	4.03	1.57	2.57	0.64	4.90	2.69	2.92	5.71 5.64	2.64
$Cr_2O_3$	< 0.01	< 0.01	< 0.01	< 0.01	< 0.01	0.02	< 0.01	0.05	0.02	0.02	0.04	< 0.01	< 0.01	0.03
	40.23	42.00	40.62	42.07	7.68	6.95	8.38	6.07	6.56	8.69	6.05	8.16	0.58	6.20
<u> </u>	97.82	96.96	<u>69.18</u> 4152	97.01	100.40	97.50	2736	99.89	98.88	99.43	99.45	99.30	100.00	97.77
Ba	361	403	298	360	686	716	1118	506	566	437	780	465	328	545
Nb	17	7.2	2.3	12	391	334	138	216	210	565	301	481	921	282
Ta	3.9	0.4	0.0	0.4	48	32	9.5	15	15	54	21	22	140	28
W TI	< 0.8	< 0.5	5.0 0.1	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5 0 7	0.7	< 0.5	< 0.5 0.8	< 0.5
Bi	< 0.1	< 0.1	0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1
Th	3.8	0.5	6.3	0.6	3.7	0.6	3.1	2.3	1.4	0.9	5.9	1.1	16	3.9
U V	0.8 7.0	0.3	1.3	0.1 7.0	1.0 268	0.7 289	1.7	0.4 281	0.4 119	1.0 318	0.6 234	0.1 178	12 215	1.2 221
Rb	4.0	< 1	7.0	< 1	119	143	81	139	289	270	153	142	360	201
Y	123	114	79	135	41	25	34	22	23	27	26	27	33	25
Zr	121	4.0	4.0	2.0	627	408	549	412	199	292	637	52	621	393
La Ce	243 509	248 510	542	200 598	226	138	179	118	121	151	140	141	237	140
Pr	61	62	65	75	28	16	22	14	14	18	17	17	30	17
Nd	228	234	250	280	115	63	87	56	54	70	66	64	117	64
SM Eu	43	43	44 12	54 16	63	35	16 4 8	31	10	36	34	35	21 5.8	34
Gd	35	34	29	41	17	9.0	12	8.1	7.4	9.6	8.8	9.1	15	9.1
Tb	4.7	4.6	3.6	5.4	2.0	1.2	1.6	1.0	1.0	1.2	1.1	1.1	1.8	1.1
Dy Ho	24	23	18	28	10	5.5	1.7	4.8	4.9	5.9	5.7	5.8	8.4	5.5
Er	12	10	7.5	12	3.8	2.4	3.0	2.1	2.3	2.5	2.5	2.5	3.1	2.3
Tm	1.4	1.4	0.9	1.6	0.4	0.3	0.4	0.3	0.3	0.3	0.3	0.3	0.3	0.3
Yb	8.7	8.4	5.5	9.0	2.6	1.7	2.1	1.4	1.9	1.9	1.9	1.8	1.8	1.7
ΣREE	1189	1197	1243	1414	530	317	416	274	275	347	326	1189	1197	1243
[La/Yb] <sub>N</sub>	19	20	32	22	25	25	25	25	20	25	24	19	20	32
[Gd/Yb] <sub>N</sub>	3.2	3.3	4.3	3.7	5.3	4.3	4.7	4.6	3.2	4.0	3.7	4.1	6.6	4.4
CI F	0.0	0.1	0.0	0.0	< 0.01	0.0	0.0	0.0	0.0	< 0.01	0.0	0.0	0.0	< 0.01
r Cu	< 1	12	57	4.0	10	2.0	9.0	< 1	< 1	< 1	< 1	4.0	< 1	5.0
Ni	< 1	3.0	27	< 1	18	41	6.0	255	42	37	118	12	10	157
Mo Zn	3.0	1.0	104	< 1 5 0	< 1 151	< 1 217	< 1 138	< 1 212	< 1 270	< 1 246	< 1 160	< 1 180	< 1 44	< 1 184
S	1600	2400	43100	1800	400	400	500	300	300	500	300	400	300	1400
Au	< 2	< 2	5.0	< 2	6.0	< 2	11	< 2	8.0	< 2	< 2	4.0	< 2	< 2
Ag	0.6	< 0.5	< 0.5	< 0.5	3.3	2.1	2.9	2.1	1.0	1.4	2.9	< 0.5	3.0	2.0
As Br	2.8	< 0.5	< 0.5	< 0.5 < 0.5	< 0.5 < 0.5	< 0.5	< 0.3 3.4	< 0.5	< 0.5 3.0	< 0.3 4.4	< 0.5 < 0.5	< 0.5	< 0.5	< 0.5 < 0.5
Cr	11	5.0	< 5	< 5	14	125	< 5	268	106	111	204	10	9.0	198
Sc	5.8	4.0	5.8	1.7	5.6	6.4	2.1	23	6.0	4.9	9.5	3.4	3.4	6.0
SD Se	< 0.1 < 3	0.3 < 3	< 0.1 < 3	< 0.1 < 3	< 0.1 < 3	< 0.1 < 3	< 0.1 < 3	< 0.1 < 3	< 0.1 < 3	< 0.1 < 3	< 0.1 < 3	0.4 < 3	< 0.1 < 3	< 0.1 < 3
Co	< 1	< 1	57	< 1	26	32	13	50	22	18	36	25	23	42
Ga	6.0	4.0	5.0	4.0	36	36	31	28	32	50	31	36	50	34
Ge Sn	1.0 < 1	0.9 < 1	1.0 < 1	1.1 < 1	2.4	1.8	1.4	1.9 12	1.8	1.9 5.0	1.8	1.9 4.0	2.7	1.8
Cs	0.1	< 0.1	< 0.1	< 0.1	2.3	2.8	5.0 1.4	3.7	5.0 4.4	7.8	4.3	3.3	7.5	4.7
Pb	16	14	7.0	10	4.0	< 3	4.0	< 3	< 3	5.0	< 3	7.0	6.0	7.0
Be	< 1	< 1	< 1	< 1	2.0	2.0	2.0	7.0	2.0	1.0	5.0	2.0	2.0	5.0
	3.3	0.3	0.1	0.5	14	1.6	9.0	0.0	4.0	3.4	12	1.0	9.1	/.1

Mg# 48 41	4	44	44	48	38	57	40	59	59	47	41	57

Facies	SC	SC	SC	Ocelli	Ijolite	Cb- bearing							
6 l -	000 59	010 D1	010 C1	Tacles	007 D1	NS 009 D1	NS	NS 008 C1	NS 012 C2	NS 007 E10	NS 010 E1	NS 009 E2	NS 011 D2
SiO <sub>2</sub> (wt %)	42 38	41 51	41.51	40.00	43.81	49.64	46 45	48.91	50.53	47.12	46 77	52.84	53.93
$Al_2O_3$	14.89	13.89	16.03	14.28	17.37	21.15	17.96	17.59	19.86	16.69	17.48	19.10	21.91
$Fe_2O_{3(T)}$	10.25	15.09	4.53	7.97	3.85	4.73	3.88	1.74	3.51	7.77	6.87	4.20	2.54
FeO FeoO	6.88 2.68	2 70	3.34 0.86	6.00 1.37	2.75	3.59 0.78	2.83	0.33	2.59	5.90	5.11	3.13	1.89
MgO	3.35	4.56	1.48	2.29	0.98	0.89	1.32	0.28	0.89	1.79	2.37	0.80	0.38
MnO	0.16	0.20	0.09	0.16	0.06	0.07	0.11	0.10	0.06	0.11	0.09	0.07	0.05
CaO Na-O	9.07	6.59 3.22	12.72	12.62 5.84	6.98	4.60 9.04	8.75 8.96	10.44	5.68 8.82	6.42	7.50	5.21	4.01 9.46
K <sub>2</sub> O	3.82	6.39	5.41	3.29	4.78	4.86	2.63	3.71	3.77	4.69	3.55	4.59	4.78
TiO <sub>2</sub>	2.53	2.55	0.71	1.22	0.68	0.63	0.62	0.18	0.51	1.13	1.10	0.61	0.31
$P_2O_5$	1.67	2.22	2.23	2.79	0.16	0.59	1.08	0.23	1.89	0.99	1.44	0.99	0.40
LOI	6.26	3.39	9.62	7.54	8.76	3.48	6.72	8.49	3.36	5.37	5.23	3.45	3.06
Total	100.00	99.61	98.23	98.00	98.54	99.68	98.48	99.43	98.89	99.44	99.97	99.87	100.80
Sr (ppm) Bo	1658	1021	2785	1649	934	1248	1559	2332	954 456	1909	1588	1556	1521
Nb	429	374	74	439	121	82	277	70	413	944 97	106	70	47
Та	35	24	10	99	14	8.8	99	24	87	8.0	13	6.7	4.2
W	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	0.7	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	1.0
Bi	< 0.3	< 0.8	< 0.2	< 0.1	< 0.2	< 0.4	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1
Th	3.4	2.6	6.5	18	0.8	0.6	6.4	8.9	18	1.5	6.6	0.9	0.8
U	0.9	1.5	6.5	5.1	0.2	1.5	30	34	3.4	0.3	1.4	0.7	0.5
v Rb	182	337	102	209	30 72	43 122	00 93	10 65	54 127	86	104 77	42	23 87
Y	25	23	39	35	7.2	12	23	32	19	20	25	15	10
Zr	635	440	484	535	179	43	71	3720	520	321	546	863	349
La Ce	68 148	65 147	76 171	203	8.9 25	30 67	66 141	57	62 140	46 99	58 130	36 82	25 53
Pr	18	18	21	25	3.6	7.9	17	13	17	12	16	10	6.1
Nd	69	71	84	94	15	31	62	50	64	46	63	39	23
SM Eu	35	13	16 4 8	5.0	3.1 1.0	5.5 1.6	31	9.1 2.7	32	8.3 2.5	34	7.0	4.3
Gd	9.4	9.3	13	13	2.4	4.2	8.1	7.4	8.1	6.6	8.9	5.3	3.1
Tb	1.1	1.2	1.6	1.6	0.3	0.5	1.0	1.1	1.0	0.9	1.2	0.7	0.4
Ду Но	5.6 1.0	5.3 0.9	8.0 1.4	/.6	1.6	2.6	4.9	6.0 1.2	4.6	4.1 0.7	5.4 1.0	3.2 0.6	2.1 0.4
Er	2.3	2.1	3.4	3.1	0.7	1.2	2.1	3.4	1.7	1.9	2.4	1.4	1.0
Tm	0.3	0.2	0.4	0.4	0.1	0.2	0.3	0.5	0.2	0.2	0.3	0.2	0.1
yd Lu	0.3	0.2	2.7	2.3	0.5	0.9	0.2	5.4 0.6	0.2	0.2	0.3	0.2	0.8
ΣREE	340	339	404	462	62	153	319	273	315	229	303	188	121
[La/Yb] <sub>N</sub>	26	31	19	26	12	22	28	12	34	22	21	21	22
[Ga/YD] <sub>N</sub> Cl	4.2	5.4	3.8	4.4	3.9	3.6	4.1	0.1	<u> </u>	3.8	3.9	3.6	3.3
F	0.1	0.3	0.1	0.2	0.3	0.0	0.1	< 0.01	0.1	0.1	0.1	0.0	< 0.01
Cu	30	< 1	2.0	< 1	11	6.0	1.0	7.0	11	1.0	1.0	2.0	3.0
NI Mo	9.0 < 1	22 < 1	< 1	3.0 < 1	< 1	2.0	6.0 < 1	< 1	2.0	12 < 1	12 < 1	2.0	2.0
Zn	151	274	73	128	53	69	65	27	65	114	105	59	38
S	300	300	600	400	600	400	300	400	300	400	300	200	200
Au Ag	< 2	< 2 2 1	< 2 2 3	< 2 2.6	< 2 0 9	< 2 < 0.5	< 2 < 0.5	< 2	< 2	< 2	< 2 2.6	< 2 4 5	< 2 1 9
As	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5
Br	8.0	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	3.7	4.7	4.9	2.4	2.6	3.2	< 0.5
Cr Sc	9.0 3.2	55 4.5	< 5 1.0	6.0 4.4	< 5 1.5	< 5	< 5 2.6	7.0 1.4	< 3	8.0 2.2	18	< 5	< 5 0.4
Sb	< 0.1	0.3	0.2	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1
Se	< 3	< 3	< 3	< 3	< 3	< 3	< 3	< 3	< 3	< 3	< 3	< 3	< 3
Ga	18 36	23 45	33	11 49	3.0 21	3.0 27	6.0 45	25	5.0 43	38	32	28	2.0 24
Ge	1.5	2.2	1.5	2.6	0.7	1.2	1.7	1.3	1.9	1.4	1.5	1.4	0.9
Sn Ca	3.0	3.0	< 1	1.0	2.0	1.0	1.0	< 1	1.0	2.0	1.0	< 1	< 1
CS Pb	5.0	5.8 < 3	3.8 10	5.0 7.0	< 3	< 3	2.0 21	25	2.5 6.0	< 3	< 3	1.3 < 3	< 3
Be	2.0	2.0	1.0	2.0	2.0	2.0	4.0	2.0	3.0	2.0	1.0	1.0	2.0
Hf Ma#	11	9.3	9.4	9.0	3.6	0.6	1.8	47	8.8	6.3	7.6	8.9	4.3

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Facies	NS	NS	NS	NS	NS	NS	NS	NS	NS	NS	NS	NS
Sample	010-D2	011-G1	019-G2	008-B1	011-B1	011-B2	007-B1	006-C3	004-C1	012-G2	012-G3	004-C2
SiO <sub>2</sub> ( <i>wt.%</i> )	49.32	54.10	54.15	53.94	53.87	58.98	54.51	59.72	56.12	59.46	56.76	56.10
$Al_2O_3$	21.59	20.14	19.86	21.59	21.10	22.36	23.32	23.01	23.44	23.13	21.59	25.27
$Fe_2O_{3(T)}$	5.53	3.02	3.61	2.89	2.41	1.08	1.80	1.06	1.22	2.24	0.73	1.36
FeO FeO	5.98	2.30	2.60	2.18	1.79	0.75	1.20	0.73	0.89	1.08	0.50	1.00
	0.95	0.49	0.75	0.49	0.44	0.25	0.41	0.23	0.24	0.39	0.18	0.20
MnO	0.95	0.45	0.90	0.42	0.55	0.23	0.47	0.22	0.13	0.02	0.03	0.03
CaO	3 54	3 48	2.61	3 24	3 20	0.02	2 45	0.02	0.02	0.05	1.66	0.30
Na <sub>2</sub> O	8 66	8 59	7.16	9.08	9.04	11 34	11.12	11.51	10.06	12.63	11.53	12.38
K <sub>2</sub> O	5.41	4.54	5.59	4.69	4.61	1.89	2.92	2.03	5.25	1.81	1.68	3.94
TiO <sub>2</sub>	1.00	0.34	0.60	0.34	0.29	0.10	0.26	0.10	0.09	0.24	0.03	0.11
$P_2O_5$	0.36	0.32	0.21	0.53	0.48	0.07	0.21	0.05	0.12	0.08	0.25	0.06
Cr <sub>2</sub> O <sub>3</sub>	< 0.01	< 0.01	< 0.01	< 0.01	< 0.01	< 0.01	< 0.01	< 0.01	< 0.01	< 0.01	< 0.01	< 0.01
LOI	2.65	2.60	4.42	2.12	2.23	1.79	2.99	1.69	1.46	0.55	3.15	0.56
Total	99.11	97.59	99.14	98.89	97.63	98.88	100.10	100.20	98.41	100.50	97.50	100.20
Sr (ppm)	1154	1282	1083	1497	1347	201	586	194	161	66	383	96
Ba	1611	991	1474	1322	1272	106	238	52	91	18	147	35
Nb	157	79	70	50	40	905	154	1850	2290	1200	945	1170
	- 0.5	21	/./	4.8	5.3 - 0.5	254	62	249	326	139	1/8	141
VV TI	< 0.5 0.5	< 0.5 0.3	1.4	< 0.5 0.3	< 0.5 0.2	1.0	0.8	< 0.3	0.0	< 0.3	1.1	< 0.5 0.5
Bi	< 0.5	< 0.5	< 0.2	< 0.5	< 0.2	< 0.5	< 0.5	0.2	0.0	0.5	< 0.2	< 0.5
Th	0.5	19	3.0	0.6	13	52	8 1	49	93	39	30	41
U	0.5	22	1.5	0.5	2.8	14	67	0.1	1.5	0.2	2.0	1.5
V	55	24	54	19	22	12	27	8.0	10	17	8.0	11
Rb	146	100	66	97	86	42	54	55	110	36	35	86
Y	10	12	7.4	8.9	10	10	8.1	4.2	2.6	3.1	7.7	1.5
Zr	69	1000	811	190	195	3360	1340	636	153	1800	2060	153
La	27	22	19	23	25	22	23	13	16	5.6	26	8.1
Ce	58	47	38	53	54	39	44	27	39	12	52	18
Pr NJ	7.1	5.7	4	6.3	6.5	3.9	5.0	2.7	3.9	1.3	5.7	1.8
Na Sm	27	21	1/	23	24	21	18	9	21	4.5	20	0.0
SIII Fu	4.0 1.4	4.1	09	4.2	4.5	2.1	5.1	1.5	2.1	0.7	5.5	1.0
Gd	3.4	3.2	2.2	3.1	33	1.6	2.5	0.4	13	0.2	23	0.5
Th	0.5	0.4	0.3	0.4	0.4	0.3	0.3	0.0	0.1	0.5	0.3	0.0
Dv	2.2	2.3	1.6	2.0	2.1	1.6	1.7	0.8	0.6	0.6	1.5	0.4
Ho	0.4	0.4	0.3	0.3	0.4	0.4	0.3	0.2	0.1	0.1	0.3	0.1
Er	0.9	1.2	0.8	0.9	1.0	1.3	0.8	0.5	0.3	0.4	0.8	0.2
Tm	0.1	0.2	0.1	0.1	0.1	0.2	0.1	0.1	0.0	0.1	0.1	0.0
Yb	0.7	1.3	0.8	0.7	0.8	1.6	0.8	0.5	0.2	0.5	0.8	0.1
Lu	0.1	0.2	0.1	0.1	0.1	0.3	0.1	0.1	0.0	0.1	0.1	0.0
<b>EREE</b>	133	110	87	119	123	87	101	55	77	27	115	36
[La/Yb] <sub>N</sub> [Cd/Vb]	27	20	1/	23	21	9.2	18	19	65	/.6	24	46
	4.1	2.0	2.4	3.3	0.1	0.8	2.4	0.1	0.5	0.8	2.3	4.0
F	0.0	< 0.01	< 0.01	0.0	< 0.01	< 0.01	< 0.01	< 0.01	< 0.01	< 0.0	< 0.01	< 0.01
r Cu	12	< 0.01	8.0	1.0	< 0.01 7 0	23	23	31	< 0.01 5 0	< 0.01 7 0	4 0	15
Ni	4.0	5.0	4.0	2.0	< 1	2.0	2.0	2.0	2.0	1.0	< 1	1.0
Мо	< 1	< 1	2.0	< 1	< 1	2.0	1.0	1.0	1.0	< 1	2.0	< 1
Zn	117	58	60	51	38	32	33	31	28	48	22	19
S	600	200	700	200	200	100	300	400	900	< 100	100	< 100
Au	< 2	< 2	< 2	< 2	< 2	< 2	4.0	< 2	< 2	< 2	< 2	< 2
Ag	< 0.5	-	4.1	0.9	0.8	-	-	3.2	0.7	-	-	0.7
As D-	< 0.5	< 0.5	0.9	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5	< 0.5
br Cr	< 0.5	< 0.5	0.2	5.4 2.5	2.5	2.9	0.4 < 5	10	10	< 0.3	10	< 0.5
Sc	13	0.8	05	04	04	23	20	<ul> <li>5</li> <li>1.6</li> </ul>	<ul> <li>3</li> <li>14</li> </ul>	<ul> <li>3</li> <li>14</li> </ul>	20	10
Sb	03	0.0	0.3	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	< 0.1	0.4
Se	< 3	< 3	< 3	< 3	< 3	< 3	< 3	< 3	< 3	< 3	< 3	< 3
Со	7.0	3.0	6.0	3.0	10	< 1	2.0	1.0	< 1	< 1	< 1	< 1
Ga	25	30	36	27	25	32	39	39	34	41	40	36
Ge	1.0	1.3	1.2	1.1	0.9	1.1	1.4	1.3	1.1	1.7	1.3	1.2
Sn	1.0	< 1	< 1	< 1	< 1	< 1	< 1	< 1	1.0	1.0	< 1	< 1
Cs	1.6	0.9	1.4	0.9	0.7	0.9	1.0	0.9	0.7	0.2	0.6	0.6
Pb	< 3	12	< 3	< 3	< 3	12	37	< 3	4.0	< 3	7.0	< 3
Ве	2.0	1.0	2.0	2.0	1.0	4.0	3.0	5.0	5.0	5.0	4.0	5.0
 	30	25	38	2.3	2.1	33	40	<u> </u>	2.7	22	24	3.2
-1-51	50	20	50	20	20	51	-10	55	- 1	2	27	0

<sup>1</sup> Sample names follow the format *12-PA-XXX-XX*.

**Table 3.** Representative magmatic pyrochlore compositions and structural formula from different syenite facies of the Crevier alkaline intrusion from the central Grenville Province obtained using EPMA. Abbreviations: apfu = atoms per formula unit; Neph. Sye. = nepheline syenite. General formula:  $A_{2-m}B_2X_{6-w}Y_{1-n}$  (Atencio *et al.*, 2010) with m, w and n corresponding to the <u>vacanciesoecupaney</u> of the different sites.

Mineral	Pyrochlore										
Sample <sup>1</sup>	002-G1	002-G1	006-C2	006-C3	006-C3	006-C3	012-G4	010-G1	010-G1	010-G2	010-G2
Facios	Nanh Sua	Neph.	Neph.	Neph.	Neph.	Neph.	Neph.	Ocelli	Ocelli	Ocelli	Ocelli
racies	Nepn. Sye.	Sye.	Sye.	Sye.	Sye.	Sye.	Sye.	facies	facies	facies	facies
Analysis no. <sup>2</sup>	5-3	6-3	2-16	10-2	7-2	8-3	6-7	6-8	6-9	3-6	4-11
A-site											
UO <sub>2</sub> (wt.%)	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.41	1.94	0.00	0.16
ThO <sub>2</sub>	2.75	1.38	1.20	1.04	0.82	1.07	1.16	1.00	0.65	0.97	0.92
SnO <sub>2</sub>	0.04	0.00	0.00	0.04	0.00	0.00	0.00	0.00	0.04	0.00	0.02
$Sc_2O_3$	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
$Y_2O_3$	0.00	0.00	0.00	0.00	0.00	0.00	0.10	0.00	0.00	0.02	0.00
$La_2O_3$	0.07	0.14	0.10	0.13	0.09	0.03	0.07	0.14	0.08	0.04	0.06
Ce <sub>2</sub> O <sub>3</sub>	0.34	0.40	0.26	0.51	0.30	0.27	0.31	0.46	0.27	0.25	0.41
Pr <sub>2</sub> O <sub>3</sub>	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
MnO	0.00	0.08	0.00	0.03	0.05	0.00	0.00	0.00	0.00	0.00	0.05
CaO	13.18	13.05	13.11	12.08	14.02	13.49	14.02	14.86	13.73	15.35	12.18
SrO	1.17	1.73	1.21	0.97	1.12	0.99	0.10	1.78	0.87	0.69	1.34
BaO	0.00	0.25	0.06	0.01	0.07	0.02	0.00	0.00	0.00	0.00	0.23
PbO	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Na <sub>2</sub> O	7.51	6.39	7.80	7.87	6.32	7.83	8.06	6.76	6.64	6.83	6.50
K <sub>2</sub> O	0.02	0.17	0.03	0.05	0.21	0.01	0.06	0.01	0.01	0.00	0.08
B-site											
WO <sub>3</sub>	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Nb <sub>2</sub> O <sub>5</sub>	60.64	55.92	64.13	59.07	56.02	65.63	64.14	56.60	53.93	64.48	53.22
$Ta_2O_5$	9.04	9.49	6.63	6.88	7.15	6.07	3.76	9.57	15.64	4.77	16.06
SiO <sub>2</sub>	0.00	1.79	0.15	1.12	2.22	0.01	0.02	0.05	0.00	0.00	0.72
TiO <sub>2</sub>	2.01	1.78	0.83	4.00	2.25	1.03	4.00	1.86	2.14	2.66	1.37
ZrO <sub>2</sub>	0.00	0.09	0.10	0.00	0.00	0.00	0.03	0.00	0.06	0.00	0.53
$Al_2O_3$	0.04	0.18	0.06	0.11	0.19	0.16	0.02	0.47	0.04	0.00	0.08
$Fe_2O_3$	0.04	2.56	0.22	1.07	2.32	0.04	0.11	0.21	0.08	0.05	1.13
X/Y-Sue	1 0 2	262	1 65	4.16	4.02	4.07	4.16	1 80	2.02	176	2.07
г	4.85	5.02 0.77	4.03	4.10	4.05	4.97	4.10	4.80	5.92	4.70	5.97
Total	101.75	0.77	100.72	0.55	0.37	101.60	100.63	0.00	100.42	101.05	0.47
-O=F	2 03	1.52	1.96	1 75	1 70	2 09	1 75	2 02	1.65	2 00	1.67
Total	99.71	98.27	98.77	97.94	96.06	99.59	98.88	96.96	98.81	99.04	97.83
A-site		,	,,		,		,		,		
U <sup>4+</sup> (anfu)	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.006	0.028	0.000	0.002
Th <sup>4+</sup>	0.040	0.019	0.017	0.014	0.011	0.015	0.016	0.015	0.010	0.014	0.013
Sn <sup>4+</sup>	0.001	0.000	0.000	0.001	0.000	0.000	0.000	0.000	0.001	0.000	0.001
Sc <sup>3+</sup>	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000
Y <sup>3+</sup>	0.000	0.000	0.000	0.000	0.000	0.000	0.003	0.000	0.000	0.001	0.000
La <sup>3+</sup>	0.002	0.003	0.002	0.003	0.002	0.001	0.002	0.003	0.002	0.001	0.001
Ce <sup>3+</sup>	0.008	0.009	0.006	0.011	0.007	0.006	0.007	0.011	0.007	0.006	0.010
Pr <sup>3+</sup>	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000
Mn <sup>2+</sup>	0.000	0.004	0.000	0.002	0.003	0.000	0.000	0.000	0.000	0.000	0.003
Ca2+	0.898	0.843	0.882	0.770	0.906	0.894	0.906	1.049	0.968	1.013	0.832
Sr <sup>2+</sup>	0.043	0.060	0.044	0.033	0.039	0.036	0.003	0.068	0.033	0.025	0.050
Ba <sup>2+</sup>	0.000	0.006	0.001	0.000	0.002	0.000	0.000	0.000	0.000	0.000	0.006
ru- Net	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000
11a K <sup>+</sup>	0.920	0.013	0.950	0.907	0.759	0.939	0.942	0.004	0.048	0.010	0.803
R B-site	0.002	0.015	0.002	0.004	0.010	0.001	0.005	0.001	0.001	0.000	0.007
W <sup>6+</sup>	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000
Nb <sup>5+</sup>	1.743	1.524	1.820	1.588	1.528	1.836	1.748	1.686	1.605	1.795	1.533
Ta <sup>5+</sup>	0.156	0.156	0.113	0.111	0.117	0.102	0.062	0.171	0.280	0.080	0.278
Si <sup>4+</sup>	0.000	0.108	0.009	0.067	0.134	0.001	0.001	0.003	0.000	0.000	0.046
Ti <sup>4+</sup>	0.096	0.081	0.039	0.179	0.102	0.048	0.181	0.092	0.106	0.123	0.066
Zr <sup>4+</sup>	0.000	0.003	0.003	0.000	0.000	0.000	0.001	0.000	0.002	0.000	0.016
Al <sup>3+</sup>	0.003	0.013	0.004	0.008	0.014	0.012	0.001	0.037	0.003	0.000	0.006
Fe <sup>3+</sup>	0.002	0.116	0.010	0.048	0.105	0.002	0.005	0.010	0.004	0.002	0.054
X/Y-site											
F-	0.971	0.690	0.923	0.782	0.769	0.973	0.793	1.000	0.816	0.927	0.800
OH-	0.029	0.310	0.077	0.218	0.231	0.027	0.207	0.000	0.184	0.073	0.200
Total	4.918	4.704	4.905	4.745	4.725	4.892	4.884	5.017	4.898	4.873	4.726
A	1.92	1.70	1.90	1.74	1.72	1.89	1.88	2.02	1.90	1.87	1.73
о ОЦ	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00
F	0.03	0.51	0.08	0.22	0.23	0.03	0.21	0.00	0.18	0.07	0.20
1	0.97	0.09	0.92	0.70	0.77	0.97	0.19	1.00	0.62	0.95	0.60

<u>A-site</u> <del>V</del> vacancv <sup>3</sup>	0.08	0.30	0.10	0.26	0.28	0.11	0.12	-0.02	0.10	0.13	0.27	
% vacancy	4.1	14.8	4.8	12.8	13.8	5.4	5.8	-0.9	5.1	6.3	13.7	
Semilar and fillen the formet 12 DA VVV VV												

<sup>1</sup> Sample names follow the format *12-PA-XXX-XX*.
 <sup>2</sup> All analyses are reported in Table S1.
 <sup>3</sup> OH has been calculated to fill the Y site.

Mineral	Apatite
n	195
Ca-site	
CaO ( <i>wt.%</i> )	54.90
	(0.13)
Na <sub>2</sub> O	0.36
	(0.02)
MgO	0.04
	(0.06)
FeO	0.10
	(0.04)
SrO	0.92
	(0.02)
MnO	0.03
1.0	(0.00)
$La_2O_3$	0.15
C. O	(0.01)
$Ce_2O_3$	0.54
NJO	(0.02)
$\mathbf{M}\mathbf{u}_2\mathbf{U}_3$	(0.01)
P_site	(0.01)
$\mathbf{P}_{\mathbf{O}}$	43.03
1 205	(0.15)
SiO	0.02
5102	(0.02)
<b>SO</b> <sub>3</sub>	0.01
	(0.00)
F-site	· · · · ·
F	3.16
	(0.04)
Cl	0.01
	(0.00)
Total	103.41
	(0.15)
Atoms per form	ula unit
calculated on th	e basis of 12.5 O
Ca-sue	1 955
Ca	(0.014)
Na	0.057
1 100	(0.003)
Mg	0.005
8	(0.007)
Fe	0.007
	(0.003)
Sr	0.044
	(0.001)
Mn	0.002
	(0.000)
La	0.007
~	(0.001)
Ce	0.010
	(0.001)
Nd	0.006
	(0.000)

**Table 4:** Apatite compositions from carbonatite facies (sample 12-PA-007-A6) of the Crevier alkaline intrusion from the central Grenville Province obtained using EPMA. They are reported with confidence interval of 95% in brackets (n = number of analyses).

P-site	
Р	3.005
	(0.006)
Si	0.002
	(0.002)
S	0.001
	(0.000)
F-site	
F	0.825
	(0.010)
Cl	0.001
	(0.000)
Ca-site	4.993
P-site	3.008
F-site	0.827

Zone <sup>a</sup>	LREE- Ca-Fe content	Ca (ppm)	Fe (ppm)	ΣLREE (ppm)	ΣHREE (ppm)	Sr (ppm)	Ti (ppm)	Temp. (°C) <sup>b</sup>	Th (ppm)	U (ppm)	Th/U	Ce/ Ce*c	Eu/ Eu*c	Sm <sub>N</sub> <sup>c</sup> /Yb <sub>N</sub> <sup>c</sup>	δ <sup>18</sup> Ο (‰)	±2σ	<sup>207</sup> Pb/ <sup>206</sup> Pb age (Ma)	±2σ	Pb-Pb conc. (%)
(Sub-)euhed	ral grains																		
Grain I Core															7.74	0.55			
Grain 2							•		10.50										
Core	High	394	1381	94	980	94	20	815	1950	297	6.57	3.09	0.92	0.09			968	22	99
Grain 5	High	1771	38/6	121	1316	210	230	1167	2180	610	3 52	3 36	0.02	0.00			951	10	96
Core	High	1604	5262	112	1202	128	346	1241	2180	727	2.87	3 53	0.92	0.09			912	23	86
Core	High	3377	8758	112	1202	312	623	1373	2000	1006	2.07	2.71	1.00	0.10			713	15	58
Core		5511	0,00		1221	512	020	1010	22	1000	2.10	2.71	1.00	0.11			,15	10	20
Grain 4																			
Core																	917	11	67
Core	High	3376	2890	111	1147	410	200	1134	2137	526	4.06	4.16	0.92	0.09			807	14	84
Core	High	4260	2091	144	1345	648	130	1061	2348	369	6.36	4.27	0.93	0.10					
Rim	High	6890	8623	147	1342	879	554	1344	2871	817	3.52	3.31	0.96	0.11					
Grain 5																			
Core	High	1580	1702	78	670	133	196	1131	921	853	1.08	2.40	0.95	0.12			850	26	101
Rim	High	187	126	2.5	44	5.1	9.4	740	16	31	0.52	2.14	0.88	0.01					
Rim	High	3815	3487	166	1290	360	503	1322	2621	971	2.70	2.26	0.94	0.12					
Rim	High	1862	9/3	55	414	357	251	1177	421	765	0.55	1.38	0.94	0.06					
Grain 6															4.01	0.50			
Core															4.81	0.56			
Grain 9															2.00	0.46	065	27	100
Rim	High	278	130	14	216	18	24	835	246	243	1.01	1 32	1.04	0.02	2.90	0.40	905	57	100
Rim	High	1204	1758	102	1062	40	24	1153	240	634	3.81	3 33	0.97	0.02	3.05	0.47			
Grain 10	mgn	1201	1750	102	1002	10	221	1100	2115	051	5.01	5.55	0.77	0.10	5.05	0.17			
Core	High	2690	3940	187	1268	151	78	983	4274	1240	3.45	2.57	0.92	0.15			914	12	68
Rim	High	5681	6798	178	1147	388	105	1027	2973	1102	2.70	2.10	0.91	0.14					
Grain 11	0																		
Core	High	580	990	20	198	69	32	866	469	178	2.63	2.42	0.93	0.06	5.70	0.57			
Core																			
Core	High	1603	9440	159	1139	205	178	1114	2519	843	2.99	2.48	0.88	0.11	5.66	0.55			
Core	High	312	1288	12	105	42	25	838	113	75	1.52	2.23	0.93	0.06					
Grain 12																			
Core	High	4244	2808	230	928	296	651	1384	1872	1031	1.82	1.28	0.94	0.13	2.1.4	0.42			
Rim	High	223	/4	23	227	19	31	866	263	87	3.03	2.14	0.89	0.05	3.14	0.43			
Grain 21																			
Core																			
Rim	High	2572	1485	182	1306	213	251	1177	2942	595	4 94	1.80	0.99	0.12					
Grain 22	mgn	2012	1105	102	1500	215	201	11//	2712	575	1.91	1.00	0.77	0.12					
Core																	590	16	70
Grain 23																			
Core	High	2404	2471	204	1304	209	268	1189	3220	850	3.79	2.19	0.97	0.13	5.03	0.54			
Core	e																		
Grain 24																			
Core															7.62	0.57			
Rim	High	8101	2742	176	1417	1062	914	1473	2361	1084	2.18	2.48	0.97	0.10	6.54	0.56			
								http	://www.	petroloa	v.oupiou	urnals.or	a/						

**Table 5:** Summary of the results obtained on zircon grains from the carbonate-bearing nepheline syenite (sample 12-PA-011-B11) of the Crevier alkaline intrusion from the central Grenville Province. The zircon grains are grouped according to the classification of Figs. 10 and 11. Abbreviation: Temp. = temperature.

Grain 25																			
Core																	965	39	93
Grain 26																			
Core																	756	113	80
Grain 27																			
Rim	High	9249	2839	199	1129	1204	547	1342	2626	763	3.44	1.84	1.00	0.13					
Rim																			
Rim Contra 20																			
Grain 28																	060	28	00
Rim	High	2240	401	205	887	247	45	911	2379	352	676	1.52	0.98	0.16			900	20	77
Grain 29	mgn	2240	401	205	007	247	45	711	2517	552	0.70	1.52	0.70	0.10					
Core																	855	30	66
Grain 30																			
Core	High	6705	9156	272	1475	320	1536	1628	2514	2607	0.96	1.83	0.98	0.17	6.53	0.55	946	18	66
Core	High	5319	6407	234	1022	311	1136	1534	2037	1887	1.08	1.44	0.96	0.20					
Rim	High	2065	1993	131	1098	190	296	1209	2146	624	3.44	2.26	0.93	0.10					
Rim	High	3294	2384	100	912	428	395	1268	1456	572	2.55	2.33	0.91	0.10	6.97	0.56			
Grain 31																			
Rim	High	2597	2274	162	515	366	130	1062	2065	659	3.13	1.39	1.00	0.16	5.48	0.54			
Rim																			
Rim																			
Grain 32																			
Core															5.19	0.55	962	32	100
Grain 33	· · ·	4101	2720	120	702		150	1201	2077	(00	2 00	5.01	0.00	0.10	1.00	0.50		10	
Core	High	4181	3720	138	793	555	459	1301	2077	698	2.98	5.81	0.29	0.10	4.28	0.58	802	13	44
Grain 34	IIIh	17571	10424	20.4	1511	2460	1000	1522	2707	1250	2.05	2.25	1.01	0.14	7.01	0.54			
Rim	High	1/5/1	10434	204	1511	2460	1088	1522	3/0/	1256	2.95	2.55	1.01	0.14	7.81	0.54			
Core	High	7524	3721	193	1351	1086	244	1171	2116	385	5 49	1.96	0.96	0.11	4 64	0.56			
Core	mgn	7524	5721	175	1551	1000	244	11/1	2110	505	5.47	1.90	0.70	0.11	4.04	0.50			
Grain 35																			
Core																	963	38	93
Rim	High	2595	5005	143	651	181	390	1266	2620	693	3.78	1.82	0.99	0.19					
Rim	c																		
Grain 36																			
Rim																			
Core	High	1931	1466	118	973	170	138	1071	2377	938	2.53	14.04	0.29	0.10	5.81	0.59	941	17	94
Grain 37																			
Core	High	4625	3765	301	1691	609	613	1369	5524	1415	3.90	1.79	0.97	0.15			952	19	101
Grain 38																			
Core															7.34	0.53			
Core															0.15	0.54			
Intermedia	e grains																		
Grain 15	3																		
Core	High	381	116	50	881	46	8.9	736	701	484	1.45	4.05	0.94	0.06			961	23	100
Core	High	631	274	40	672	49	42	901	690	337	2.05	4.90	0.92	0.06					
Core	High	4223	2141	115	1082	284	331	1231	2187	636	3.44	3.09	0.94	0.09					
Rim	High	236	127	22	365	23	9.0	736	257	221	1.16	3.59	0.95	0.05					
Rim	Low	124	39	5.1	43	11	6.9	712	32	34	0.93	2.18	1.28	0.05					
Rim	Low	167	78	5.4	125	7.7	13	770	53	56	0.94	3.56	1.13	0.02					
Grain 19																			
Core																	962	65	100
Core	High	2473	1086	203	392	301	234	1164	696	289	2.41	1.28	1.08	0.22					
Rim	Low	66	20	1.3	35	1.1	2.3	626	8.2	17	0.47	3.09	0.93	0.02					
Rim	Low	106	23	3.9	36	5.8	7.5	719	17	24	0.71	1.07	1.19	0.04			970	91	98

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Anhedral g Grain 7	rains																		
Core	Low	170	55	1.5	39	0.9	2.6	632	12	29	0.40	2.12	0.97	0.01	5.08	0.53	923	40	102
Core	High	187	90	2.0	27	2.8	5.5	693	12	47	0.25	3 31	0.95	0.02	2.00	0.00	946	72	100
Core	Low	174	56	1.0	22	1.5	2.6	634	7 3	28	0.26	3 39	1.01	0.02			210	/ _	100
Rim	High	193	62	0.6	25	0.5	1.6	600	2.8	20	0.12	3 51	0.54	0.02					
Rim	High	178	388	3.8	23	13	8.0	725	10	25	0.49	1 31	0.98	0.04	5 75	0.57			
Grain 8	mgn	170	500	5.0	21	15	0.0	125	10	21	0.47	1.51	0.70	0.04	5.75	0.57			
Rim															1 29	0.43			
Grain 13															1.29	0.45			
Core	Low	67	21	0.7	13	0.2	2.0	615	15	15	0.31	19.61	0.68	0.02	2 72	0.44	031	53	07
Core	Low	69	21	0.5	17	0.2	2.0	626	3.3	15	0.22	0.30	0.00	0.02	2.72	0.41	942	32	00
Grain 14	LOW	09	22	0.5	17	0.1	2.4	020	5.5	15	0.22	9.59	0.01	0.01	2.07	0.41	942	52	"
Core	Low	76	21	37	160	1.8	2.1	617	24	41	0.58	3.86	0.70	0.01			967	30	100
Core	Low	70 68	21	1.2	40	1.0	2.1	607	6.0	41	0.38	2.60	0.79	0.01			907	30	100
Core	LOW	08	21	1.5	49	0.8	1.0	007	0.9	10	0.42	5.45	0.81	0.01			061	27	00
Core	Low	65	20	0.0	22	0.6	17	604	4.5	14	0.22	2 60	0.91	0.01			901	27	99
Core	Low	65	20	0.9	52	0.0	1./	604	4.5	14	0.55	3.08	0.81	0.01			985	22	97
Core	Low	111	19	3.2	52	10	1.0	600	10	21	0.50	1.41	0.85	0.02			985	33	97
Core																			
Grain 16	T	122	25	0.5	12	10	<b>C</b> 1	(0 <b>7</b>	20	26	1 1 2	1.52	1.05	0.07			1027	(2)	0.2
Core	Low	132	25	8.5	42	10	5.1	687	30	26	1.13	1.53	1.05	0.06			1037	63	92
Rim	Low	74	23	4.4	122	3.5	2.4	626	58	54	1.08	3.95	0.83	0.02					
Rim	Ŧ	(2)	20		1.50	0.0	2.5	(21	25	10	0.72	6.05	0.05	0.00					
Rim	Low	63	20	4.1	159	0.8	2.5	631	35	48	0.73	6.27	0.85	0.02				4.0	
Core	Low	81	20	4.3	45	5.8	2.6	634	26	26	1.00	1.94	0.93	0.04			1018	40	86
Rim	Low	109	20	5.4	58	7.0	2.3	624	15	17	0.84	1.57	0.98	0.03					
Grain 17	_																		
Core	Low	65	20	0.9	29	0.8	2.1	618	6.5	15	0.42	6.32	0.92	0.02					
Core	Low	213	23	1.6	32	1.3	3.3	652	6.0	20	0.30	6.30	1.04	0.01					
Core	Low	59	18	1.6	30	1.4	2.1	616	5.0	16	0.30	2.80	0.89	0.02	2.47	0.44	910	62	97
Core	Low	97	39	2.6	32	7.8	11	752	10	32	0.31	2.29	0.92	0.04					
Core															3.96	0.41	1026	26	93
Rim	Low	193	42	4.2	40	17	14	781	18	22	0.83	1.70	1.00	0.04	2.40	0.46			
Rim	Low	62	19	0.9	31	0.7	2.1	619	3.7	15	0.24	5.66	0.93	0.01			1026	26	93
Grain 18																			
Core	Low	64	20	5.1	174	4.2	3.0	644	76	110	0.69	4.92	0.79	0.01	2.22	0.45	951	37	100
Core																			
Core																			
Core															2.31	0.42	962	47	99
Core	Low	62	20	5.4	267	0.5	1.9	611	40	73	0.55	12.90	0.82	0.02			956	23	100
Core	High	118	63	16	456	8.3	11	755	183	203	0.90	4.25	0.92	0.03	3.10	0.45	918	10	96
Core																			
Core																	959	44	99
Core																	970	68	94
Rim	Low	123	22	1.7	33	25	3.4	653	7.5	17	0.43	2.13	0.89	0.02	1.45	0.46			
Rim	Low	78	22	1.1	66	0.2	1.3	584	14	33	0.43	25.83	0.80	0.01	2.22	0.45	951	37	100
Grain 20																			
Core	Low	65	20	1.4	38	1.8	2.1	616	5.0	16	0.32	2.86	0.82	0.01	3.00	0.42			
Core																			
Core	Low	63	20	1.6	55	1.4	2.8	640	7.1	21	0.34	2.92	0.86	0.01					
Core	Low	72	19	3.1	39	4.0	6.5	707	15	24	0.64	1.83	0.97	0.03					
Rim															1.99	0.44			

<sup>a</sup> Zonation ordered from cores towards rims in each grain.

b Calculation between Ti and Ca+Fe suggest that the Ti content of (sub-)euhedral zircon grains was not only controlled by the zircon lattice (and associated thermal conditions), but very likely also by interstitial domains, causing an overestimation and therefore unrealistic zircon crystallization temperatures. See text for details.

<sup>c</sup> Chondrite normalization after McDonough and Sun (1995).

Analysis no.	Concentrat	ions	Isotopic ratio	S		
	Pb (ppm)	U (ppm)	<sup>238</sup> U/ <sup>206</sup> Pb	2σ	<sup>207</sup> Pb/ <sup>206</sup> Pb	2σ
Ap-1	3.8	4.2	3.46	0.14	0.50	0.01
Ap-2	3.7	4.4	3.70	0.16	0.46	0.01
Ap-3	18	1.1	0.36	0.02	0.92	0.02
Ap-4	7.8	0.9	0.53	0.03	0.87	0.03
Ap-5	3.3	4.6	4.21	0.18	0.45	0.01
Ap-6	8.9	1.1	0.59	0.04	0.88	0.04
Ap-7	3.5	4.1	3.65	0.16	0.47	0.01
Ap-8	3.4	3.5	3.38	0.15	0.49	0.01
Ap-9	3.7	3.6	3.32	0.14	0.52	0.01
Ap-10	3.5	3.5	3.21	0.14	0.50	0.01
Ap-11	3.5	3.6	3.37	0.16	0.48	0.01
Ap-12	4.0	4.1	3.25	0.15	0.51	0.01
Ap-13	3.5	4.1	3.57	0.15	0.45	0.01
Ap-14	4.3	4.4	3.32	0.15	0.48	0.02
Ap-15	4.7	4.3	3.12	0.16	0.51	0.02
Ap-16	4.2	4.1	3.36	0.11	0.50	0.01
Ap-17	4.1	4.5	3.22	0.21	0.51	0.01
Ap-18	4.2	4.7	3.45	0.11	0.50	0.01
Ap-19	3.1	2.9	3.02	0.15	0.49	0.03
Ap-20	2.9	2.7	3.05	0.17	0.49	0.02
Ap-21	13	0.6	0.34	0.01	0.90	0.01
Ap-22	11	0.8	0.45	0.02	0.87	0.02
Ap-23	3.3	3.8	3.35	0.13	0.45	0.01
Ap-24	3.6	4.1	3.33	0.17	0.50	0.01
Ap-25	8.5	0.7	0.49	0.02	0.89	0.02
Ap-26	3.6	5.0	3.92	0.18	0.46	0.02

**Table 6.** U-Pb dating of apatite from carbonatite facies (sample 12-PA-007-A6) of the Crevier alkaline intrusion from the central Grenville Province.

**Table 7:**  $\delta^{18}O_{V-SMOW}$  and  $\delta^{13}C_{V-PDB}$  composition of calcite from carbonatite facies (samples 12-PA-019-A, 12-PA-007-A6 and 12-PA-009-A5) of the Crevier alkaline intrusion from the central Grenville Province.

Sample	Facies	δ <sup>18</sup> O (‰ V-SMOW)	2σ	δ <sup>13</sup> C (‰ PDB)	2σ
12-PA-019-A	Carbonatite	+7.05	0.20	-6.42	0.10
12-PA-007-A6	Carbonatite	+6.73	0.20	-6.78	0.10
12-PA-009-A5	Carbonatite	+7.35	0.20	-6.14	0.10