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# 1 Transition from stable column to partial collapse during the 79 cal CE P3

# 2 Plinian eruption of Mt Pelée volcano (Lesser Antilles)

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#### 14 Highlights

15 - We determine the eruptive sequence of the P3 Plinian eruption of Mt Pelée volcano 16 - New radiocarbon dating measurements provide a refined age of  $79 \pm 21$  cal CE 17 - We estimate key eruptive parameters of this VEI 5 eruption in the transitional regime 18 - Increasing eruption rate and decreasing gas content led to partial column collapse 19 - The transition occurred at conditions well predicted by our 1D theoretical model 20 21 **Keywords** 22 Mt Pelée volcano, Plinian eruption, tephra fallout, partial collapse, tephrostratigraphy, 23 radiocarbon dating 24

## 25 Abstract

26 Explosive volcanic eruptions commonly form sustained Plinian columns that collapse at some 27 stage producing dangerous pyroclastic density currents (PDC) on the ground. Numerical and 28 laboratory models of volcanic plumes show that the conditions leading to total column 29 collapse are strongly controlled by the amount of exsolved gas at the source and the mass 30 eruption rate. However, column collapse is rarely total and the volcanic jet often separates in 31 a dense collapsing part feeding PDC and a buoyant rising plume spreading volcanic gases and 32 pyroclasts in the atmosphere. This transitional regime has been directly observed and/or 33 inferred from the structure of the deposits for several past eruptions, but the number of cases 34 for which the partial collapse regime is described in detail, including the 79 CE Vesuvius, and 35 the 186 CE Taupo eruptions, remains too small to fully constrain physical models. Here, we 36 present a detailed reconstruction of the time evolution of the P3 eruption at Mt Pelée volcano 37 (Martinique, Lesser Antilles) that underwent partial column collapse in order to discuss the 38 mechanisms controlling the eruption dynamics and improve the volcanological database on 39 transitional eruptions. The P3 eruptive succession consists of seven major phases that produced a total of 1 km<sup>3</sup> dense rock equivalent (DRE) of deposits (i.e., VEI 5 event), starting 40 with a thick pumice fall deposit (0.1 km<sup>3</sup> DRE) overlain by alternating pyroclastic density 41 current (PDC) (0.7 km<sup>3</sup> DRE) and pumice fall deposits (0.2 km<sup>3</sup> DRE). We use physical 42 43 models together with field data on deposit dispersal, thickness, and grain-size distribution to 44 reconstruct the dynamical evolution of the volcanic column. Our results show that the mass eruption rate (MER) increased from  $1.2 \times 10^8$  to  $1.7 \times 10^8$  kg s<sup>-1</sup> during the initial phase 45 producing a 28 to 30 km-high Plinian plume. The MER later reached up to  $2.5 \times 10^8$  kg s<sup>-1</sup> 46 47 and the column entered the partial collapse regime characterized by the formation of a small (i.e., 12 to 17 km-high) ash plume and contemporaneous PDCs mainly channelized in three 48 49 paleo-valleys. These estimates are used together with published data on magmatic water 50 contents in glass inclusions to decipher the mechanisms leading to partial collapse. The P3 51 eruption column collapsed due to an increase in mass eruption rate and a decrease in gas 52 content. A similar evolution was also inferred for the 1300 CE P1 and 280 CE P2 eruption 53 deposits, revealing a systematic behavior in the recent Plinian eruptions of Mt Pelée volcano. 54 The comparison of model predictions of column collapse and field data reveals a good 55 agreement for the P1, P2, P3 and Taupo eruptions, but not for the 79 CE Vesuvius eruption 56 where thermal disequilibrium between gas and pyroclasts most likely strongly affected the 57 column dynamics.

58

# 59 **1. Introduction**

Plinian eruptions are amongst the most powerful and destructive volcanic events on Earth.
These explosive eruptions generally form a sustained stable column that can rise up to 40 km
in the atmosphere (e.g., the 1991 Pinatubo eruption) and spread out laterally as an umbrella
cloud carrying pyroclasts at large distances from the volcano (Koyaguchi and Tokuno, 1993).

Abrupt or progressive changes in eruptive conditions at the base of a Plinian column can lead to a radically different behavior in which the column collapses as a turbulent fountain producing hazardous pyroclastic density currents (PDC) that may reach and destroy populated areas (Wilson et al., 2014). One important goal in physical volcanology is to identify the key processes affecting the column stability in order to improve hazard assessment.

69 Recent real-time observations of volcanic plumes and field-based studies on fallout 70 tephra from past eruptions show that explosive eruptions commonly have a complex behavior 71 where both a stable Plinian column and a collapsing fountain feeding PDC are 72 contemporaneous (Carazzo et al., 2015 and references therein). This transitional regime of 73 partial column collapse occurs for Plinian eruptions ranging in size from small to large (e.g., 74 the 79 CE Vesuvius (Shea et al., 2011), 186 CE Taupo (Wilson and Walker, 1985), 1150 75 Quilotoa (Di Muro et al., 2008), 1912 Novarupta-Katmai (Fierstein and Hildreth, 1992), and 1963 Agung (Self and Rampino, 2012) eruptions). Detailed mapping and analysis of eruptive 76 77 products emplaced during the transitional regime provide valuable data to better understand 78 the mechanisms controlling the partitioning of a Plinian column, and can be used to robustly 79 test numerical and analog models of explosive volcanic eruptions (Di Muro et al., 2004; 80 Carazzo et al., 2015; Suzuki et al., 2016). However, the number of natural cases for which the 81 partial collapse of an initially stable column is described in detail (i.e., in terms of time 82 variation in mass eruption rate, gas content, grain-size distribution and exit velocity) remains 83 too small to fully constrain theoretical models of volcanic plumes (Michaud-Dubuy et al., 84 2018). With an aim of increasing the inventory of well-constrained past eruptions, we have 85 undertaken a long-term study of the Plinian eruptions of Mt Pelée volcano (Lesser Antilles) as 86 they systematically produced sustained columns that ultimately collapsed (Traineau et al., 87 1989). Because the Mt Pelée eruptions share many characteristics including similar preeruptive magma storage conditions (Martel et al., 1998), they can be used to decipher the 88

mechanisms controlling the behavior of a Plinian plume. In previous studies, we reconstructed the dynamics of the 1300 CE P1 (Carazzo et al., 2012), and 280 CE P2 Plinian eruptions (Carazzo et al., 2019) focusing on the causes for column collapse. We now extend this work to study the P3 eruption deposits whose structure suggests the entrance into the partial collapse regime at some stage of the eruption (Traineau et al., 1989; Wright et al., 2016), making it an excellent candidate to improve the volcanological database on transitional eruptions.

96 The P3 eruption is described in the literature (Roobol and Smith, 1976; Westercamp 97 and Traineau, 1983; Traineau et al., 1989; Wright et al., 2016) but the exact sequence of 98 events and the eruptive parameters are currently poorly known. Here, we use data from a new 99 comprehensive field study together with physical models to reconstruct the detailed time 100 evolution of this eruption and to discuss the mechanisms responsible for the transition from 101 stable column to partial collapse. We identify the eruptive conditions leading to this 102 transitional regime and compare the results to previous work on P3, and on younger Plinian 103 eruptions in Martinique (P1 and P2). We then compare our estimated eruptive parameters for 104 the column at collapse with theoretical predictions from a model of volcanic plumes.

105

#### 106 **2. Overview of the P3 eruption**

107 The P3 deposits were first identified by Roobol and Smith (1976) as a thick pumice fallout 108 unit on the western flank of Mt Pelée volcano standing above the 2406 and 2447 yr BP 109 horizons of their stratigraphic sections. Westercamp and Traineau (1983) provided the first 110 age, synthetic stratigraphic section and distribution map of the P3 deposits (Fig. 1a). The P3 111 Plinian eruption is currently dated at 2010  $\pm$  140 yr BP based on the average of five <sup>14</sup>C ages 112 (Westercamp and Traineau, 1983; Traineau et al., 1989). According to these studies, the 113 eruption began with the formation of a sustained Plinian column that covered the western

114 flank of the volcano with a meter of coarse lapilli, lithic-rich, pumice fallout deposit (phase 115 P3<sub>1</sub>). This initial stage was immediately followed by the production of high-concentration 116 PDC in the Grande Savane, Habitation Depaz and Falaise River paleo-valleys (Fig. 1a). The 117 second eruptive phase identified by Westercamp and Traineau (1983) consisted of a sustained 118 Plinian column that spread fine lapilli pumice to the south with a striking elongated axis of 119 dispersion (phase P3<sub>2</sub>). The P3 eruption ended up with the formation of a pumice ash-rich bed 120 (phase P3<sub>3</sub>) interbedded with high-concentration PDC deposits (Fig. 1a). Later, Traineau et al. 121 (1989) excluded the second phase  $(P3_2)$  whose distal deposits were found below tephra from 122 an eruption older than P3<sub>1</sub>. Recent work by Michaud-Dubuy et al. (2019) confirmed that the 123 deposits named P3<sub>2</sub> by Westercamp and Traineau (1983) are not interbedded between P3<sub>1</sub> and 124 P3<sub>3</sub> but belong to a much older event, the 13.5 ka cal BP Bellefontaine Plinian eruption.

125 Wright et al. (2016) studied the distal products of a Plinian eruption dated between 1800 and 2200 yr BP (their subunit d), which corresponds to the P3 eruption identified by 126 127 Westercamp and Traineau (1983). Their synthetic stratigraphic section includes a basal 128 pumice-rich low-concentration PDC (named 'surge') deposit overlain by a pumice fallout 129 deposit interbedded with high-concentration PDC (named 'flow') and low-concentration PDC 130 (named 'ash hurricane') deposits. According to their distribution map of the deposits (Fig. 1b), 131 the high-concentration PDC flowed in the Fond Canonville valley, Grande River, Bijou River, 132 and Falaise River, whereas the low-concentration PDC that elutriated from the top of the 133 current spread on hills at distances up to 20 km from the vent. Isopachs of the pumice fallout 134 deposit exhibit a SSW elongation (Fig. 1b), in good agreement with the one inferred by 135 Westercamp and Traineau (1983) (Fig. 1a). Wright et al. (2016) interpreted their complete 136 stratigraphic succession as resulting from the partial and final collapse of an initially stable 137 Plinian column.

138 The eruptive parameters of the P3 event are currently poorly constrained. Volume 139 estimates for the fallout tephra of the P3 eruption yield a relatively small value of 0.19 km<sup>3</sup> 140 dense rock equivalent (DRE) (Traineau et al., 1989). The direction of tephra dispersal during 141 the initial phase  $(P3_1)$  suggests that the column was affected by low altitude tropospheric 142 winds, and thus was certainly more than 6 km (Traineau et al., 1989), but no more than 12 km 143 in height (Wright et al., 2016). Such a large uncertainty on the maximum column height 144 makes it difficult to estimate its mass eruption rate. Overall, these sparse and loosely 145 constrained data cannot be used to understand the conditions leading to column collapse.

146

#### 147 **3. Methodology**

148 3.1. Fieldwork

149 We identified the P3 eruption deposits at 102 locations in Martinique (Fig. 2). In most cases, 150 these deposits are easily recognizable by their strong internal layering, the abundance of 151 juvenile and accidental lithic fragments in the pumice fallout units, and the presence of the P1 152 and/or P2 deposits above them (Carazzo et al., 2012; 2019). The P3 outcrops are distributed 153 all around the volcano except to the northwest where exposure is very limited due to dense 154 tropical forest and difficult conditions of access. The thickness of each layer of P3 was 155 measured at every site in order to construct isopach maps. We also excavated a standard 25 x 156 20 cm area of each fall layer and measured the major axes of the five largest lithic fragments 157 found in order to produce isopleth maps.

158

159 3.2. Radiocarbon dating

160 We dated four charcoals sampled within the P3 deposits at sites 6 (unit A-G), 97 (unit A), 178

161 (unit A), and 180 (unit A-G), and a paleosol sampled beneath unit A at site 145 (see locations

162 in Fig. 2 and unit descriptions in section 4.1) in order to refine the age of the P3 eruption.

163 Ages were determined using accelerator mass spectrometry at the LMC14 (Artemis, 164 Laboratoire de Mesure du Carbone 14, CEA Saclay, France), and calibrated using the free 165 software OxCal 4.3 (Bronk Ramsey, 2009) with the atmospheric IntCal13 calibration curve 166 commonly used for the Northern hemisphere (Reimer, 2013). The radiocarbon ages obtained 167 for our stratigraphically-constrained samples were combined with those of Westercamp and 168 Traineau (1983) and validated using the R\_combine function of Oxcal and  $\chi^2$  test prior to 169 calibration (Ward and Wilson, 1978).

170

171 3.3. Grain-size analysis

172 We carried out grain-size analyses on twenty samples from ten locations representative of the 173 seven different units (A to G) (see Table 1). The samples were dried for 24 h in an oven and 174 sieved by hand down to 60. We separated the lithic fragments from the pumices by hand in 175 the size range  $-6\phi$  to  $-4\phi$ , and used a binocular microscope to discard crystals and lithic 176 fragments in the size range  $-4\phi$  to  $-2\phi$ . Volume calculations for isomass maps for each  $\phi$ 177 interval were then used to determine the grain-size distributions of single sub-layers. The total grain-size distribution was estimated from the total mass in sieve class  $\phi$ ,  $M_{\phi}$ , by the volume 178 179 integral (Kaminski and Jaupart, 1998)

180 
$$M_{\phi} = \int_{0}^{L} h(l) C_{\phi}(l) A(l) dl,$$
 (1)

181 where h(l) is the deposit thickness,  $C_{\phi}(l)$  is the concentration of class  $\phi$  at distance l from the 182 vent, A(l)dl is the area bounded by isopachs at distances l and l+dl, and L is the distance 183 where h or  $C_{\phi}$  drop to zero. We used linear interpolations for h and  $C_{\phi}$  between the localities.

184 The cumulative frequency curves were determined assuming a power-law size 185 distribution of the rock fragments (Turcotte, 1986; Kaminski and Jaupart, 1998), where the 186 number of particles with a radius larger than r is given by:

187 
$$N(R_p \ge r_p) = \lambda r^{-D}, \tag{2}$$

188 with *D* the power-law exponent that commonly varies between 2.9 and 3.9 for Plinian 189 eruptions (Girault et al., 2014). Because fragmentation is a size-invariant process, the 190 exponent *D* can be estimated accurately using any sufficiently large range of sizes. Thus, the 191 lack of fine particles lost out at sea does not affect our estimations of *D*.

192

## 193 3.4. Eruptive parameters

194 The volume of tephra fallout produced during an explosive eruption can be inferred using 195 several methods based on the thinning trend of the deposit with distance from the source. 196 Here, we used the exponential, power-law and Weibull fits (Pyle, 1989; Bonadonna and 197 Houghton, 2005; Bonadonna and Costa, 2012) computed using the AshCalc software (Daggit 198 et al., 2014). Bearing in mind that only proximal and furthermore incompletely preserved 199 deposits are available in Martinique, the calculated volumes of fall deposits should be taken 200 as minimum estimates. The volume of dilute PDC deposits was calculated using the observed 201 relationship between the volume and area covered by dilute PDC (Dade and Huppert, 1998; 202 Calder et al., 1999), as in Carazzo et al. (2012, 2019) for the P1 and P2 eruptions. The volume 203 of dense PDC was calculated using the product of the average thickness and area covered by 204 the dense PDC. This method is bound to provide minimum estimates since only proximal and 205 furthermore incompletely preserved deposits are available in Martinique.

The maximum column heights associated with the fall deposits were estimated from the distribution of lithic fragments on our isopleth maps using the method of Carey and Sigurdsson (1986). Error bars were calculated using the three values of maximum height inferred from the 8, 16, and 32-mm isopleths (see Fig. 17 in Carey and Sigurdsson, 1986). The alternative method of Bonadonna and Costa (2013) based on variations of grain size with the distance from the source was also used to confirm the robustness of the estimates.

212 The mass discharge rate feeding the eruption during phases of stable column was inferred from a given column height by using the empirical relationship from Mastin et al. 213 214 (2009) and the predictions of the model of Girault et al. (2016), which explicitly includes the 215 effect of total grain-size distribution on the plume dynamics. Calculations were made for 216 tropical atmospheric conditions and for a crystal-bearing rhyolitic magma (andesitic bulk 217 composition) with an initial temperature of 1,150 K (Martel and Poussineau, 2007) and a dense rock density of 2,400 kg m<sup>-3</sup> (Traineau et al., 1989). The mass discharge rates for the 218 219 collapsing phases were calculated using the model of Bursik and Woods (1996), which treats 220 pyroclastic density currents as supercritical dilute suspensions entraining air as they 221 propagate. Doyle et al. (2010) showed that this assumption is valid for tall, fine-grained 222 column collapses, for which the flow slowly transfers its mass to the dense basal flow. The 223 model of Bursik and Woods (1996) is thus relevant for collapsing phases characterized by the 224 formation of dilute PDC deposits, and was used here.

225

#### 226 **4. The P3 eruptive succession**

4.1. Stratigraphy

Fig. 3a shows a composite stratigraphic section of the P3 deposits, divided into seven major phases based on diagnostic stratigraphic and lithofacies associations, which we describe in more detail below.

231

232 4.1.1. Unit A

The stratigraphic succession begins with a layer of clast-supported, coarse white pumice lapilli (Fig. 4a-d). Unit A contains juvenile and accidental lithic fragments in a total amount that typically increases from  $\approx 20$  wt.% at base to  $\approx 25$  wt.% at top (Fig. 3c). At most sites, median pumice size remains relatively constant, whereas both the maximum lithic size andthe sorting increase from base to top (Fig. 3b-c).

In the downwind direction (to the southwest), unit A has a maximum thickness of 180 cm at 4.5 km from the crater (Fig. 4c). In the crosswind direction, unit A is 130 cm thick at 4.6 km, and steadily thins to 50 cm within 6.1 km of the crater (Fig. 4d). The relatively widespread nature of this deposit, its uniformly decreasing thickness with distance from the source, its clast-supported framework, and its pumice and juvenile lithic fragments characteristics identify unit A as a fall deposit.

244

245 4.1.2. Unit B

Unit B consists of light brown fine ash material in which pumice, lithic fragments and crystals are dispersed in a matrix of dense, angular, glass fragments (Fig. 4b, c, e). Most clasts range in size from fine to coarse ash but a few large pumices are present (Fig. 3b). The total amount of lithic fragments in unit B commonly reaches  $\approx 28$  wt.%. Where unit B is present, the contact with the underlying layer is sharp, which suggests that there was no time break between deposition of the two units.

252 Unit B varies irregularly in thickness with distance from the source, and ranges from 253 18 to 35 cm at 3.8 km from the eruptive center (Fig. 4), but is absent beyond 7 km from the 254 vent. Where thick, unit B contains a few lenses of coarse lapilli, angular to sub-rounded 255 pumices. At some sites in valleys, pumice lapilli and blocks are rounded and account for a 256 large proportion in the matrix of white fine ash. The relatively limited dispersal of this unit, its 257 irregular thickness, matrix-supported framework and grain types (juvenile and accidental 258 lithic fragments), as well as the nature of the lower contact, indicate that unit B is a low-259 concentration pyroclastic density current deposit (Branney and Kokelaar, 2002).

261 4.1.3. Unit C

Unit C is a relatively thin blanket of clast-supported, white pumice lapilli (Fig. 4b, c, e) containing juvenile and accidental lithic fragments in a total amount that reaches  $\approx 25$  wt.% in most studied outcrops (Fig. 3c). At most field sites, unit C is easily distinguishable from unit A, the former being distinctly richer in lithic fragments and coarse lapilli pumice. The maximum lithic size in unit C is, however, always lower than the maximum lithic size at the top of unit A (Fig. 3c).

268 Unit C was dispersed in approximately the same direction as unit A, and is the thickest 269  $\approx$ 3.7 km southwest of the volcano in the downwind direction, where it reaches 22 cm. This 270 layer steadily thins with distance from the vent, and is still 10 cm thick at 6.2 km in the 271 crosswind direction. At most sites, the upper part of unit C was eroded by the pyroclastic 272 density current that overrode it and deposited unit D (see below). The relatively widespread 273 nature of unit C, its uniformly decreasing thickness with distance from the source, its clast-274 supported framework, and its pumice and juvenile lithic fragments characteristics identify it 275 as a fall deposit.

276

277 4.1.4. Unit D

Unit D is a relatively thin layer of fine ash and small rounded pumice lapilli dispersed in a matrix of dense, angular, glass fragments (Fig. 4a, c, f). At distal sites, this layer is thin, stratified and unconsolidated. The total amount of lithic fragments in unit D commonly reaches  $\approx 20$  wt.% (Fig. 3c). Where unit D is present, the contact with the underlying layer is sharp and exhibits some erosion of unit C.

Unit D varies irregularly in thickness with distance from the vent, and ranges from 5 to 17 cm at 3.8 km from the volcano (Fig. 4). This layer is, however, absent beyond 7 km from the vent. These irregular thicknesses, the relatively limited dispersal of this deposit, the clast and matrix types, and the nature of the lower contact indicate that unit D is a low-concentration pyroclastic density current deposit.

288

289 4.1.5. Unit E

Unit E is a layer of clast-supported, white pumice lapilli (Fig. 4b, c, e) containing juvenile and accidental lithic fragments in a total amount that typically reaches  $\approx 18$  wt.% (Fig. 3c). At most sites, both the median grain-size and the maximum lithic size are lower than those measured in unit C (Fig. 3c).

294 Material was dispersed in approximately the same direction as units A and C to the 295 southwest of the volcano. The maximum thickness is 13 cm at 3.7 km, and steadily thins to 9 296 cm within 4.8 km from the vent in the downwind direction. The full thickness of unit E is 297 however poorly preserved due to significant erosion by the overlying pyroclastic density 298 current that overrode it and deposited unit F (see below). The relatively widespread nature of 299 this deposit, its uniformly decreasing thickness with distance from the source, its clast-300 supported framework, and its pumice and juvenile lithic fragments characteristics identify unit 301 E as a fall deposit.

302

303 4.1.6. Unit F

304 Unit F is a matrix-supported layer of fine to coarse ash containing rounded pumice lapilli and 305 lithic fragments (Fig. 4b, f), the latter of which typically reaches  $\approx 15$  wt.% (Fig. 3c). Where 306 unit F is present, the contact with the underlying layer is sharp and exhibits some erosion of 307 unit E.

308 Unit F varies irregularly in thickness with distance from the vent depending on the 309 topography, ranges from 6 to 75 cm at 3.8 km from the volcano, and is absent 7 km from the 310 vent (Fig. 4). The relatively limited dispersal of this unit, its irregular thickness, matrixsupported framework and grain types (juvenile and accidental lithic fragments), as well as the
nature of the lower contact, indicate that unit F is a low-concentration pyroclastic density
current deposit.

314

315 4.1.7. Unit G

Unit G is the uppermost layer in the P3 sequence and consists of a blanket of fine to coarse white pumice lapilli (Fig. 4a, b, f). This unit contains juvenile and accidental lithic fragments in a total amount that reaches  $\approx 20$  wt.% at most studied outcrops (Fig. 3c) with a maximum lithic size similar to that of unit C (Fig. 3c).

In the downwind direction (to the southwest), unit G has a maximum thickness of 80 cm at 1.4 km from the crater, and steadily thins to 20 cm within 4.8 km of the vent. In the crosswind direction, unit G is 30 cm thick at 5 km, and 15 cm at 6.4 km from the vent. The relatively widespread nature of this deposit, its uniformly decreasing thickness with distance from the source, its clast-supported framework, and its pumice and juvenile lithic fragment characteristics identify unit G as a fall deposit.

326

327 4.2. Age of the P3 eruption

328 Radiocarbon dates and calibrated ages are reported in Supplementary material Table S1. Two 329 samples provide the exact same radiocarbon age of  $1.870 \pm 30$  yr BP (sites 145 and 178 on 330 the western flank). A third sample taken near location 178 gives a radiocarbon age of  $1,915 \pm$ 331 30 yr BP (site 97). The oldest age is found near the town of Le Carbet where the charcoal 332 sampled within the pumice fallout deposit provides an age of  $2,030 \pm 30$  yr BP. On the 333 eastern flank of the volcano, the charcoal sampled at location 6 has an age of  $1,795 \pm 30$  yr 334 BP. These five new radiocarbon ages are in good agreement with a compilation of twelve  ${}^{14}C$ 335 ages determined in previous studies (Walker, 1973; Roobol and Smith, 1976; Traineau, 1982;

336 Westercamp and Traineau, 1983) ranging from 1,800 to 2,150 yr BP (Supplementary material 337 Table S1). Combining these seventeen values using the R combine function of OxCal (Bronk 338 Ramsey, 2009) provides an age of 1,926 ± 11 yr BP for the P3 eruption. We note that the  $\chi^2$ 339 test fails here most likely due to the large number of individual measurements. However, a 340 reasonable correction on the error bar of each dating by a factor of 2 is enough to successfully pass the  $\chi^2$  test, reinforcing the confidence that all these otherwise stratigraphically-341 342 constrained samples belong to the same eruption. Our new refined radiocarbon age for the P3 343 eruption stands in the lower error bars of the 2,010  $\pm$  140 yr BP proposed by Traineau et al. 344 (1989). Correcting our new estimate with the IntCal13 calibration curve (Reimer, 2013) 345 available in OxCal gives a mean calendar age of  $79 \pm 21$  cal CE (i.e., 31-37 cal CE at 1.7%) 346 probability and 51-125 cal CE at 93.7% probability).

347

### 348 4.3. Grain-size distribution of selected samples

Twenty samples from the P3 units were analyzed (Table 1) in order to discuss the evolution of the grain-size distribution during the entire P3 eruption at a single outcrop (Fig. 3), and to determine the total grain-size distribution for the lapilli fraction of unit A deposits (see section 4.6). The objective is twofold: to confirm the fall or flow nature of the P3 units, and to estimate the amount of gas trapped in pumice fragments at the beginning of the eruption (see section 6.3). A detailed analysis of the pumice textures is beyond the scope of this paper and can be found in Martel and Poussineau (2007).

The fourteen samples from unit A display the typical fallout characteristics with median diameter ranging from  $0.7\phi$  to  $-3.3\phi$ , and sorting ranging from 1.4 to 2.7. These values are relatively close to those obtained by Bardintzeff et al. (1989) who measured median diameter and sorting ranging from  $-0.3\phi$  to  $-3\phi$ , and 1.7 to 2.8, respectively. Most of the samples collected on land belong to proximal deposits since the medial and distal deposits were lost at sea. The grain-size distribution of individual samples is systematically bimodal (Fig. 3b) and the amount of ash particles (< 2 mm) always increases from base ( $\approx 9$  wt.%) to top ( $\approx 36$  wt.%). The top of unit A is thus always more poorly sorted than the base, but the median grain-size remains approximately constant throughout the unit (Fig. 3c).

The six samples from units B to G collected at outcrop 163 show that units C, E, and G have typical fallout characteristics similar to unit A. The median diameter for these units ranges between -0.1 $\phi$  and -3 $\phi$ , and they have moderate sorting values from 2.5 to 2.7 (Fig. 3c). The amount of ash-sized particles is similar to or larger than in unit A, ranging from  $\approx 28$ wt.% to  $\approx 55$  wt.%. Units B, D, and F have flow features with median diameter ranging from 1.7 $\phi$  and 0.4 $\phi$ , and moderate sorting ranging from 2.2 to 2.6 (Fig. 3c). The amount of ashsized particles is larger than in the fall deposits, ranging from  $\approx 62$  wt.% to  $\approx 85$  wt.%.

372

#### 373 4.4. Isopach maps

374 Fig. 5 shows a stratigraphic correlation of P3 outcrops along three different dispersal axes 375 (i.e., north to south, northwest to southeast, and northeast to southwest, see Fig. 2 for 376 localization). The complete sequence can be found up to 5 km from the vent (sites 13, 163, 377 101, 78 in Fig. 5). Thickness measurements at each location are reported on isopach maps for 378 the cumulative pumice fallout phases A+C+E+G (Fig. 6a), the pumice fallout phase A (Fig. 379 6b), the cumulative pyroclastic density current phases B+D+F (Fig. 6c), and the cumulative 380 pumice fallout phases C+E+G (Fig. 6d). These maps are later used to calculate the total grain-381 size distribution of unit A (see section 4.6), and the volume of deposits (see section 5.1).

The cumulative isopach map of the pumice fallout phases A+C+E+G shows ellipsoidal contour patterns indicating fallout dispersion towards the southwest (Fig. 6a). We find a similar direction of dispersion for the pumice fallout phase A, and the cumulative pumice fallout phases C+E+G (Fig. 6b, d). This direction of dispersion is in relatively good

386 agreement with the results of Traineau et al. (1989) and Wright et al. (2016) (Fig. 1), and is 387 characteristic of the wind profiles in the Lesser Antilles during the wet season (Michaud-388 Dubuy et al., 2019). In the upwind direction to the northeast and in the crosswind direction to 389 the southeast, the entire P3 succession is reduced to a single layer of clast-supported, white, 390 angular, fine pumice lapilli containing abundant juvenile and accidental lithic fragments, 391 making it difficult to distinguish the pumice fallout units A, C, E, and G. At locations near the 392 vent, unit A is however clearly visible under the cumulative pumice fallout of units C+E+G 393 (Fig. 4d), which allows to build an individual isopach map for the pumice fallout phase A 394 (Fig. 6b). We note that the northwestern arms of our isopach maps are not well constrained 395 due to a lack of outcrops (in our study as in previous ones) in a region extremely difficult to 396 access due to the presence of dense tropical forest on steep hills.

397 Units B, D, and F are widespread on the southwest flank of Mt Pelée volcano and the 398 low-concentration PDC deposits almost reached the sea there (Fig. 6c). The high-399 concentration PDC deposits are channelized in paleo-valleys and now form large plateaux in 400 the Grande Savane, Habitation Depaz and Ajoupa Bouillon areas (see locations in Fig. 1a), an 401 observation consistent with the results of Westercamp and Traineau (1983) and Traineau et al. 402 (1989). However, contrary to Wright et al. (2016), we found no evidence of high-403 concentration PDC deposits in the Bijou River, Grande River, and Fond Canonville areas (see 404 locations in Fig. 1b).

- 405
- 406 4.5. Isopleth maps of air fall deposits

Fig. 7 shows the isopleth maps built from the measurements of the major axes of the five largest fragments found at the base and top of unit A, and at the base of unit C. The isopleth map of the base of unit A is well constrained thanks to the good preservation of the deposit (Fig. 7a). However, isopleth contours for the top of unit A and the base of unit C are 411 constrained by only a few points from localities where the two units are separated by the 412 presence of unit B and/or by clear variation in grain-size characteristics (Fig. 7b, c). Isopleth 413 maps for units E and G could not be constructed due to the poor preservation of these 414 deposits. However, in both cases, we were able to draw a single isopleth curve (i.e., the 15 415 mm-isopleth for unit E and 25 mm-isopleth for unit G) that is not reported in Fig. 7. The main 416 direction of dispersion to the southwest (Fig. 7) is consistent with the one inferred from our 417 isopach maps (Fig. 6). Isopleth contours of the top of unit A are clearly more extended in the 418 crosswind direction than those of the base of units A and C.

419

420 4.6. Total grain-size distribution (*Unit A*)

421 We calculated the total grain-size distribution (TGSD) of unit A using the method of 422 Kaminski and Jaupart (1998) (see section 3.3 for calculation details). The interbedded nature 423 and important erosion of the subsequent units made the reconstruction of the TGSD for units 424 B to G impossible. The TGSD of unit A is bimodal with a primary fine mode at  $2\phi$  and a 425 secondary coarse mode at -20 (Supplementary material Figure S1). The median diameter and 426 sorting for the TGSD of unit A are  $-0.2\phi$  and 2.3, respectively. The amounts of ash (< 2 ) 427 and fine ash particles (< 63 µm) reach 59 wt.% and less than 1 wt.%, respectively. The latter 428 value should be taken with caution because most of the fine ash particles settled at sea and 429 cannot be taken into account in our grain-size analyses. From these results, we infer that the 430 power law coefficient D that fully characterizes the TGSD at the beginning of the P3 eruption 431 (i.e., unit A) is  $D = 3.3 \pm 0.1$  (Supplementary material Figure S1). The time evolution of the 432 value of D after the phase A remains unknown, but this lack of information will not affect our 433 calculations of the stable plume / collapsing fountain transition that occurred at the end of 434 phase A (see section 6.3).

#### 436 **5. Eruptive dynamics**

#### 437 5.1. Erupted volumes

438 The estimates of volume of tephra fall from phase A using the three integration techniques introduced in section 3 give 0.282 km<sup>3</sup> with the exponential method, 0.316 km<sup>3</sup> 439 with the power law, and 0.306 km<sup>3</sup> with the Weibull function (Fig. 8a). Integration of the 440 441 exponential, power-law, and Weibull fits for the cumulative pumice fallout C+E+G yields 442 volumes of 0.334, 0.317, and 0.355 km<sup>3</sup>, respectively (Fig. 8b). Because the latter estimates 443 are based on three proximal isopach contours only they remain relatively poorly constrained 444 (Fig. 6d). We thus calculated the volume of the cumulative pumice fallout A+C+E+G in order 445 to achieve a better estimate of the total volume of tephra fall. Integrating the exponential, power-law, and Weibull fits yields 0.735, 0.692, and 0.712 km<sup>3</sup>, respectively (Fig. 8c). We 446 thus retain a minimum volume of  $0.30 \pm 0.01$  km<sup>3</sup> for unit A, and  $0.41 \pm 0.02$  km<sup>3</sup> for the 447 cumulative pumice fall units C+E+G, and  $0.71 \pm 0.02 \text{ km}^3$  for the total volume of tephra fall. 448 449 Based on fallout deposit and magma densities of 1,000 and 2,400 kg m<sup>-3</sup>, respectively (Traineau et al., 1989), we infer the cumulative DRE volume of fall units to be 0.3 km<sup>3</sup>. 450

451 The volume of PDC deposits is calculated using the area covered by the dense and 452 dilute PDC deposits. The deposits corresponding to the dilute PDC of phases B, D, and F cover 42, 20, and 36 km<sup>2</sup>, respectively (Fig. 6c), which yield bulk volumes of 0.025 km<sup>3</sup> for 453 454 unit B, 0.010 km<sup>3</sup> for unit D, and 0.016 km<sup>3</sup> for unit G (Calder et al., 1999). The cumulative 455 dense PDC deposits of the phases B, D, and F cover 7 km<sup>2</sup> in Grande Savane, 9.5 km<sup>2</sup> in Habitation Depaz, and 18 km<sup>2</sup> in Ajoupa Bouillon (see locations in Fig. 1a), with an average 456 457 thickness of 15, 30, and 2 m, respectively (Traineau, 1982). This yields a minimum volume of  $0.67 \pm 0.02$  km<sup>3</sup> DRE for the dense PDC deposits, and gives a total volume of (dense + dilute) 458 PDC deposits of  $0.72 \pm 0.02 \text{ km}^3 \text{ DRE}$ . 459

The final estimate of the total volume of the P3 eruption (units A+B+C+D+E+F+G) is thus  $1.02 \pm 0.06$  km<sup>3</sup> DRE, and the total mass of tephra emitted is estimated to be  $2.4 \pm 0.1$  x  $10^{12}$  kg. The P3 eruption thus stands as a VEI 5 event (Newhall and Self, 1984) with a magnitude of 5.4 (Pyle, 2000).

- 464
- 465 5.2. Column heights and exit velocities (*airfall units*)

466 The estimates of maximum height using our isopleth maps (Fig. 7) with the model of Carey 467 and Sigurdsson (1986) give  $28.3 \pm 3.3$  km for the base of unit A,  $29.7 \pm 1.9$  km for the top of 468 unit A, and  $16.9 \pm 0.4$  km for the base of unit C. Following the approach of Bonadonna and Costa (2013), a Weibull fit gives a maximum height of  $28.1 \pm 1$  km for the base of unit A, 469 470  $30.0 \pm 0.5$  km for the top of unit A, and  $17.2 \pm 0.5$  km for the base of unit C (Supplementary 471 material Figure S2). The lack of information about the distribution of lithic sizes in unit E and 472 G prevents us from providing a robust estimate for the maximum height. Based on a single 473 isopleth curve in each case, we calculate maximum column heights of  $12 \pm 4$  km for unit E 474 and  $13 \pm 4$  km for unit G using the model of Carey and Sigurdsson (1986).

We also use data on the decrease of lithic sizes with distance from the vent in Supplementary material Figure S2 to estimate the minimum exit velocity of the volcanic plume at the vent. Extrapolating the Weibull fits down to  $A^{1/2} = 0.01$  km to find the maximum lithic size at the vent, we infer minimum exit velocities of  $210 \pm 10$  m s<sup>-1</sup> for the base of unit A,  $220 \pm 10$  m s<sup>-1</sup> for the top of unit A, and  $220 \pm 10$  m s<sup>-1</sup> for the base of unit C.

480

481 5.3. Mass eruption rates and duration

We use our new estimates of maximum column height to calculate the mass eruption rate (MER) feeding the plume during the P3 eruption. The model of Girault et al. (2016) used with their complex wind profile, which is close to the average wind profiles in the Lesser Antilles,

yields maximum MERs of  $10^8$  kg s<sup>-1</sup> for the beginning of phase A,  $1.5 \times 10^8$  kg s<sup>-1</sup> for the end 485 of phase A, 3 x  $10^7$  kg s<sup>-1</sup> for phase C, 1.4 x  $10^7$  kg s<sup>-1</sup> for phase E, and 1.7 x  $10^7$  kg s<sup>-1</sup> for 486 phase G. The empirical formula of Mastin et al. (2009) built on observations of 28 Plinian 487 eruptions yields maximum MERs of  $1.4 \times 10^8 \text{ kg s}^{-1}$  for the beginning of phase A,  $1.9 \times 10^8$ 488 kg s<sup>-1</sup> for the end of phase A,  $1.8 \times 10^7$  kg s<sup>-1</sup> for phase C,  $4.2 \times 10^6$  kg s<sup>-1</sup> for phase E, and 5.9 489 490  $x \ 10^6$  kg s<sup>-1</sup> for phase G. Based on these two series of estimates, we retain peak MERs of 1.2  $\pm 0.2 \times 10^8$  kg s<sup>-1</sup> for the beginning of phase A,  $1.7 \pm 0.2 \times 10^8$  kg s<sup>-1</sup> for the end of phase A, 491  $2.4 \pm 0.6 \text{ x } 10^7 \text{ kg s}^{-1}$  for phase C,  $9.1 \pm 4.9 \text{ x } 10^6 \text{ kg s}^{-1}$  for phase E, and  $1.2 \pm 0.5 \text{ x } 10^7 \text{ kg s}^{-1}$ 492 493 for phase G. Combined with masses of fallout deposits (section 5.1), these MER provide 494 minimum durations of ~45 min for phase A, ~1h45 for phase C, ~3h for phase E, and ~2h45 495 for phase G.

The MER during the collapsing phases B, D, and F is more difficult to assess. The model of Bursik and Woods (1996) yields a mass eruption rate of  $\sim 2.3 \pm 0.8 \times 10^8 \text{ kg s}^{-1}$  for the three phases B, D, and F. We note that this value does not take into account topographic slopes but Doyle et al. (2010) showed that the runout distance of such a flow mostly depends on the initial collapse height, which is imposed by the MER rather than by ground slopes. Combined with masses of PDC deposits (section 5.1), this MER provides minimum durations of ~50 min for phase B, ~25 min for phase D, and ~50 min for phase F.

503

### 504 **6. Discussion**

505 Our revisit of the P3 deposits shows that the eruption started with the formation of a stable 506 column that spread pyroclasts to the southwest of the volcano (phase A). This initial phase 507 was immediately followed by a partial collapse of the column producing alternating tephra 508 fallout (phases C, E, G) and pyroclastic density currents (phases B, D, E). We now compare 509 our results to previous work on the P3 eruption, and on the younger P1 and P2 eruptions of 510 Mt Pelée volcano. We then focus our discussion on the causes of partial column collapse511 during the P3 eruption by using a physical model of volcanic plumes.

512

513 6.1 Comparison with previous studies of P3 eruption

514 Our new stratigraphic data regarding the P3 succession largely differs from those of previous 515 studies (Westercamp and Traineau, 1983; Traineau et al., 1989; Wright et al., 2016). The P31 516 phase identified by Westercamp and Traineau (1983) corresponds to our eruptive unit A, and 517 their P3<sub>3</sub> phase most likely corresponds to our entire sequence from unit B to unit G. The 518 distal products identified by Wright et al. (2016) were attributed to the P3 eruption based on a 519 single radiocarbon age determined from a charcoal sampled at the top of their stratigraphic 520 section (see their figure 6). We revisited their outcrop locations 2 and 3 along a main road, 521 and found that their stratigraphic section lies beneath the 13.5 ka cal BP Bellefontaine Plinian 522 eruption (Michaud-Dubuy et al., 2019). To further reinforce our observation, we sampled a 523 paleosol at the base of their stratigraphic section and obtained a radiocarbon age of  $17,750 \pm$ 524 100 yr BP, which indicates that the distal deposits studied by Wright et al. (2016) are much 525 older than the P3 eruption.

526 The total volume of the P3 eruption (units A+B+C+D+E+F+G) is calculated to be  $1.02 \pm 0.06 \text{ km}^3$  DRE, which is one order of magnitude larger than the 0.19 km<sup>3</sup> DRE 527 528 previously estimated (Traineau et al., 1989). We attribute this important discrepancy to the 529 large volume of PDC deposits that were not taken into account in Traineau et al. (1989), and 530 to the improvement of the reconstruction techniques used here to calculate the total volume of 531 a pumice fall deposit (Bonadonna and Costa, 2013). The maximum column height is 532 estimated to be  $\sim 30$  km at the end of phase A, which is much higher than the previous 533 estimates (> 6 km by Traineau et al., 1989 and < 12 km by Wright et al., 2016). Again we 534 attribute this important difference to the better quality of our isopleth maps and to the improvement of the reconstruction techniques used to estimate the maximum height from thedistribution of lithic fragments in the field.

537

538 6.2 Comparison with younger Plinian eruptions in Martinique (P1 and P2)

539 The three latest Plinian eruptions at Mt Pelée volcano (P1, P2 and P3) have similar 540 depositional successions that can be interpreted to result from the partial collapse of an 541 initially stable Plinian column (Traineau et al., 1989; Carazzo et al., 2012, 2019; Wright et al., 542 2016). The total DRE volume estimated for P3 ( $\sim 1 \text{ km}^3$ ) is larger than those of the P1 ( $\sim 0.2$ km<sup>3</sup>) and P2 (~0.8 km<sup>3</sup>) eruptions. The maximum column height during the phase of sustained 543 544 Plinian column is also larger for P3 (28-30 km) than for the P1 (19-22 km) and P2 (22-26 km) eruptions. Consequently, the mass eruption rate is estimated to be larger for P3 (1.4 x  $10^8$  kg 545  $s^{-1}$ ) than for P1 (3.6 x 10<sup>7</sup> kg  $s^{-1}$ ) and P2 (1.1 x 10<sup>8</sup> kg  $s^{-1}$ ). The minimum eruption durations 546 547 are 11h for P3, 7h for P2, and 5h for P1. The total grain-size distribution is however relatively 548 similar for the three events, with power law exponent D of 3.3, 3.4-3.5, and 3.2-3.3 for P3, 549 P2, and P1, respectively. Our estimates of exsolved gas contents for P3 (2-2.9 wt.% - see 550 section 6.3 for calculations details) are larger than those of P1 (1.6-2.1 wt.%) and P2 (1.7-2.1 551 wt.%). The minimum exit velocities inferred from the distribution of lithic fragments for P3  $(210-220 \text{ m s}^{-1})$  are also larger than those of P1 (150-165 m s<sup>-1</sup>) and P2 (180-200 m s<sup>-1</sup>), a 552 553 result consistent with our estimates of free gas contents. Table 2 summarizes the eruptive 554 parameters retrieved for the last three Plinian eruptions at Mt Pelée volcano. P3 stands as a 555 large Plinian eruption more powerful and voluminous (VEI 5, M 5.4) than the P1 (VEI 4, M 556 4.6) and P2 (VEI 4, M 5.2) eruptions. However, because these three eruptions have much in 557 common, in particular similar pre-eruptive conditions (Martel et al., 1998; Martel and 558 Poussineau, 2007), they can be used to test the performance of theoretical models in 559 predicting the conditions for column collapse.

560

561 6.3 The stable plume / pyroclastic fountain transition during a sustained Plinian eruption

562 The P3 eruption underwent partial column collapse after an initial phase of stable Plinian 563 plume. Wilson et al. (1980) showed that the mass eruption rate and the amount of free gas in 564 the volcanic mixture at the vent both strongly control the transition between the stable Plinian 565 plume and the collapsing fountain regimes. Our estimates of the MER show a progressive increase over time from ~1.2 x  $10^8$  kg s<sup>-1</sup> (base of unit A) to ~2.5 x  $10^8$  kg s<sup>-1</sup> (unit B+C) (Fig. 566 567 9), an evolution consistent with the increase in maximum lithic sizes throughout unit A (Fig. 568 3c). Other powerful eruptions underwent a similar evolution leading to column collapse, such 569 as the 79 CE Vesuvius (Shea et al., 2011), and the 186 CE Taupo (Walker, 1980) eruptions.

570 The time evolution of total gas content feeding the column at the vent is more difficult 571 to assess. Petro-geochemical measurements provide quantitative estimates of pre-eruptive 572 magma storage conditions, including temperature (875-900°C), pressure (~200 ± 50 MPa), 573 and water content (~ $5.8 \pm 0.5$  wt.%) (Martel et al., 1998; Martel and Poussineau, 2007). Here, 574 we use the magmatic water content measured in glass inclusions available in the literature as a 575 total volatile content in the melt  $(n_0)$ . Correcting this value from the presence of crystals and 576 lithic fragments (Table 3), we find that the initial gas content  $(x_0)$  in the magma decreased 577 from 4.1 to 2.9 wt.% during the eruption (Table 3). As in Kaminski and Jaupart (1998) and 578 Carazzo et al. (2012; 2019), we calculate the mass fraction of exsolved gas in the mixture at 579 fragmentation level  $(x_f)$  for a threshold vesicularity of 70 %, which assumes closed-system 580 conditions consistent with degassing models built on U-series measurements in the P3 581 eruption products (Villemant et al., 1996). We find that  $x_f$  decreases from 2.9 to 2.1 wt.% 582 during the eruption (Fig. 9). To strengthen these estimates, we calculate the theoretical exit 583 velocities at the vent assuming that the volcanic mixture decompresses freely into the 584 atmosphere (Woods and Bower, 1995). For this, we infer the effective amount of free gas in

585 the volcanic mixture at the base of the column that is modulated by gas entrapment by 586 pumices as a function of their size distribution (Kaminski and Jaupart, 1998). We use the 587 power law exponent D = 3.3 (for unit A) estimated in section 4.6, and a typical open porosity 588 of 60-70 % (Michaud-Dubuy et al., 2018) to calculate an amount of free gas of 2.4-2.7 wt.% 589 for phase A. These estimates provide exit velocities after decompression in the range 215-230 590 m s<sup>-1</sup> (Table 3) that are consistent with those estimated from the distribution of the lithic 591 fragments collected at the base and top of unit A (see section 5.2). This relatively good 592 agreement at least for phase A reinforces the confidence that our estimates of  $x_f$  provide 593 reasonable values of free gas content during the P3 eruption.

594 Fig. 10 compares the estimated values of MER and exsolved gas content in the column 595 at collapse with the theoretical 1D predictions for tropical conditions (Michaud-Dubuy et al., 596 2018). The conditions at the beginning of the P3 eruption formed a stable plume as predicted 597 by the model. During phase A, the combined effect of a slightly increasing MER and a 598 decreasing gas content led the eruption to conditions close to the plume/fountain transition. 599 The eruption reached a peak in MER during the concomitant phases B and C, which most 600 likely induced the partial column collapse. During the partial column collapse stage of the 601 eruption (units B to G), the MER did not change significantly but the gas content slowly 602 decreased until the end of the eruption. The fairly good agreement between our field data and 603 the theoretical model can be taken as an indirect confirmation of the proposed evolution of the 604 total gas content feeding the column, even if a direct confirmation is not possible. 605 Furthermore, our results suggest that the stable column underwent partial collapse soon after 606 the beginning of the eruption (i.e., within less than a hour), an evolution that was also inferred 607 from the analysis of the P1 and P2 Plinian eruption deposits (Carazzo et al., 2012; 2019). In 608 these three eruptions, the transition occurred at conditions well predicted by our theoretical 609 model of volcanic plumes (Fig. 10).

610 Comparing these results with other historical eruptions requires gathering well-611 constrained values of exsolved gas contents, mass eruption rates and power-law exponent D612 of the TGSD (Michaud-Dubuy et al., 2018). There are however only a few examples of such 613 well-documented events, and we choose here to compare our results to the 186 CE Taupo 614 (VEI 7) and 79 CE Vesuvius (VEI 5) eruptions. Fig. 10 shows that the model successfully 615 predicts the transition for the Taupo eruption but fails to explain the transition for the 616 Vesuvius eruption. Furthermore, one can note that the collapsing phase of the 79 CE Vesuvius 617 eruption is characterized by the same amount of gas and MER at the source as the stable 618 phase (end of phase A) of the P3 eruption. The only difference between the two cases is the 619 total grain-size distribution. Indeed, as discussed in Michaud-Dubuy et al. (2018), the 79 CE 620 Vesuvius eruption is characterized by a specific total grain-size distribution (D = 3.0) where 621 pyroclasts are in average not large enough to be lost by sedimentation but not small enough to 622 be in thermal equilibrium with the gas. The hypothesis of thermal equilibrium between the 623 gas and pyroclasts that is used in our 1D model of volcanic plumes, valid for the fine grain-624 size distribution of the beginning of the P3 eruption (D = 3.3), may therefore not hold 625 anymore for this eruption and may explain the peculiar collapse conditions.

626

## 627 7. Conclusion

We have presented a new comprehensive field study of the P3 eruption, and a detailed reconstruction of the mechanisms controlling the eruptive dynamics. New radiocarbon measurements, averaged with those previously available, provide a refined age of  $79 \pm 21$  cal CE for this event. The eruption started with the formation of a 28 to 30 km-high stable column (phase A), which partially collapsed due to the combined effect of increasing MER and decreasing gas content (phase B). In the partial collapse regime, the eruptive mixture formed a 12 to 17 km-high marginally stable column (phases C, E, and G) punctuated with the production of PDC fed by the partially collapsing column (phases B, D, and F). Our calculations show that P3 was a large Plinian eruption (VEI 5, M = 5.4) evolving close to the plume/fountain transition. The total volume of tephra reached 1 km<sup>3</sup> DRE, and the mass eruption rate was of the order of 10<sup>8</sup> kg s<sup>-1</sup> during a period of at least 11h.

639 The P3 eruption is more powerful than the last two Plinian eruptions in Martinique but 640 the three eruptions underwent similar evolution towards a marginally stable volcanic column. 641 This recurrent phenomenon provides a great opportunity to look carefully at the parameters 642 affecting column stability, to constrain theoretical models of volcanic plumes, and thus 643 improve volcanic disaster forecasting. Petrological and geochemical measurements on the P3 644 eruptive products may help to further investigate the complexity of magma degassing, 645 crystallization and fragmentation in the volcanic conduit (Gurioli et al., 2005; Martel and 646 Poussineau, 2007; Shea et al., 2012; 2014), and thus refine the estimated values of exsolved 647 gas content at the vent. The global consistency between our theoretical predictions and field 648 observations of column collapse for a number of Plinian eruptions confirms that current 649 theoretical 1D models of volcanic plumes predict correctly the conditions of collapse when 650 the total grain-size distribution is dominated by fine fragments (i.e., obeys a power law 651 distribution with an exponent strictly larger than 3). Our results can thus be used to assess the 652 dominant hazard during a Plinian scale event, which may pass from pumice fall to PDC as 653 eruptive conditions vary.

Mt Pelée volcano is currently quiescent but the reconstruction of its eruptive history over the past 5,000 years shows that a Plinian eruption is a major potential scenario in the future. Such an event with the magnitude of the P3 eruption would disrupt the lives of more than 376,000 people living in Martinique and possibly impact other volcanic islands depending on the wind speed and direction.

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# **Table captions**

	Sample	Site	Unit	Subunit	Altitude (m)	Distance from the vent (km)	Thickness (cm)
	1	163	А	(0-40 cm)	269	3.7	160
	2	163	А	(40-60 cm)	269	3.7	160
	3	163	А	(60-100 cm)	269	3.7	160
	4	163	А	(100-140 cm)	269	3.7	160
	5	163	А	(140-160 cm)	269	3.7	160
	6	163	В	Bulk	269	3.7	18
	7	163	С	Bulk	269	3.7	20
	8	163	D	Bulk	269	3.7	15
	9	163	Е	Bulk	269	3.7	13
	10	163	F	Bulk	269	3.7	6
	11	163	G	Bulk	269	3.7	37
	12	88	A-G	Bulk	466	4.3	45
	13	66b	A-G	Bulk	371	4.9	65
	14	142	A-G	Bulk	234	5.8	60
	15	194	A-G	Bulk	280	5.8	55
	16	114	A-G	Bulk	195	6.3	105
	17	137	A-G	Bulk	19	6.5	100
	18	191	A-G	Bulk	35	6.5	46
	19	144	A-G	Bulk	47	6.6	80
	20	145	A-G	Bulk	540	11.9	25
853	Table 1: Sa	ampling of	the P3 depos	sits for grain-size an	alysis.		
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	Parameters	P3 (79 CE)	P2 (280 CE)	P1 (1300 CE)		
	Total DRE volume	1.02 km <sup>3</sup>	0.77 km <sup>3</sup>	0.16 km <sup>3</sup>		
	Erupted mass	2.4 x 10 <sup>12</sup> kg	$1.7 \ge 10^{12} \text{ kg}$	4 x 10 <sup>11</sup> kg		
	VEI	5	4	4		
	Magnitude	5.4	5.2	4.6		
	Maximum column height	30 km	26 km	22 km		
	Maximum PDC runout	7 – 10.3 km	7 – 11.5 km	4.5 – 8 km		
	Maximum MER (fall)	1.4 x 10 <sup>8</sup> kg s <sup>-1</sup>	1.1 x 10 <sup>8</sup> kg s <sup>-1</sup>	$3.6 \text{ x } 10^7 \text{ kg s}^{-1}$		
	Maximum MER (PDC)	2.5 x 10 <sup>8</sup> kg s <sup>-1</sup>	3.1 x 10 <sup>8</sup> kg s <sup>-1</sup>	9 x 10 <sup>7</sup> kg s <sup>-1</sup>		
	Maximum intensity	11.4	11.5	11		
	Duration	> 11 h	> 7 h	> 5 h		
	TGSD	D = 3.3	D = 3.4 - 3.5	D = 3.2 - 3.3		
	Minimum exit velocity	210 - 220 m s <sup>-1</sup>	180 – 200 m s <sup>-1</sup>	150 – 165 m s <sup>-1</sup>		
	Total gas content	2 - 2.9 wt.%	1.8-2.2 wt.%	1.6 - 2.1 wt.%		
865	Table 2: Summary of the estimated eruptive parameters for the P3 eruption, and comparison					
866	with P1 and P2 (Carazzo et al., 2012; 2019).					
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	Phase	$n_0$	Юс	x <sub>0</sub>	$x_f$	$U_{\it free}$	$U_{min}$	
		(wt.%)		(wt.%)	(wt.%)	$(m s^{-1})$	(m s <sup>-1</sup> )	
-	Early A	5.8	29	4.1	2.9	230	210	
	Late A	5.8	38	3.6	2.5	215	220	
	B-C	5.8	42	3.4	2.3	-	220	
	D-E	5.8	45	3.2	2.2	-	-	
	F-G	5.8	49	2.9	2.1	-	-	

**Table 3**: Gas contents and deduced exit velocities.  $n_0$  is the total volatile content of the melt (Martel et al., 1998),  $\alpha_c$  is the percentage of crystals in the melt plus lithics in the flow,  $x_0$  is the mass fraction of gas in the magma assuming complete degassing,  $x_f$  is the mass fraction of exsolved gas in the gas+pyroclasts mixture at fragmentation for a threshold vesicularity of 70%,  $U_{free}$  is the supersonic velocity after decompression calculated using  $x_0$ , D = 3.3 and an open porosity of 60-70% (Michaud-Dubuy et al., 2018), and  $U_{min}$  is the eruptive velocity deduced from the distribution of the lithic fragments (section 5.2).

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895 Fig. 1: Current maps for the P3 eruption. a Isopach map (in centimeters) from Westercamp 896 and Traineau (1983). (1) Dashed and (2) dotted lines correspond to isopachs for the pumice 897 fallout deposit P31 and P32, respectively. (3) The light grey and (4) dark areas give the 898 distribution of the low-concentration PDC (surge), and high-concentration PDC (flow) 899 deposits P3<sub>3</sub>, respectively. **b** Isopach map (in centimeters) from Wright et al. (2016). (1) 900 Dashed lines correspond to isopachs for a pumice fallout deposit. (2) The light and (3) dark 901 grey areas give the distribution of a low-concentration PDC (surge) deposit, and a low-902 concentration PDC (ash hurricane) deposit, respectively. (4) The dark areas correspond to 903 high-concentration PDC (flow) deposits.

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905 Fig. 2: Overview of our field area in Martinique (inset). White circles and numbers refer to 906 localities where P3 deposits are present. Black dots show outcrops where P3 deposits are 907 absent (due to erosion) and/or too deeply buried under recent eruption deposits. The dotted 908 lines give the stratigraphic correlation reported in Fig. 5.

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910 Fig. 3: a Composite stratigraphic section showing typical unit thickness and grain-size 911 characteristics of the P3 eruptive sequence. b Grain-size distribution selected samples 912 representing the different eruptive units, including five stratigraphic heights in unit A, at 913 location 163. c Variations of the median grain-size (blue circles), sorting (green circles), 914 juvenile lithic content (grey squares), accidental lithic content (dark squares), total lithic 915 content (white squares), and lithic size (red squares) along the stratigraphic height at location 916 163.

Fig. 4: Representative photographs of outcrops of the P3 deposits in Martinique at sites a 57,
b 78, c 13, d 65, and e, f 163. See Fig. 2 for location and distance from the source. All scale
bars are 20 cm long.

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Fig. 5: Stratigraphic logs of representative sections of deposits of the P3 eruption. See Fig. 2for outcrop locations.

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Fig. 6: Isopach maps (in centimeters) for a the cumulative pumice fallout (units A+C+E+G),
b the pumice fallout unit A, c the cumulative PDC deposits (units B+D+F) and d the
cumulative pumice fallout (units C+E+G). Open circles indicate measured sample locations.
Bold numbers in panel (c) indicate the thicknesses of high-concentration PDC deposits.

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Fig. 7: Isopleth maps (in millimeters) for the lithic fragments sampled a at the base of unit A,
b at the top of unit A, and c the base of unit C. Open circles indicate measured sample
locations. Direction of dispersal axes is consistent with those inferred from the isopach maps.

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Fig. 8: Deposit thinning profiles generated from the isopach maps for a unit A, b the cumulative pumice fallout C+E+G, and c the cumulative pumice fallout A+C+E+G represented by semi-log plots of square root of isopach areas (in kilometers) versus thickness (in meters). Thinning trends are approximated by exponential (purple dashed line), power-law (blue dotted line), and Weibull (red solid line) fits.

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940 Fig. 9: Summary of the time evolution of the P3 eruptive parameters. a Cumulative erupted 941 mass (in kg). b Maximum column height (in km). c Mass eruption rate (in kg s<sup>-1</sup>). d Total 942 exsolved gas content,  $x_f$  (in wt.%). 943 Fig. 10: Transition diagram depicting the P3 eruptive events and depositional units 944 (diamonds), and comparison with the P1 (squares), P2 (triangles), Taupo (circles) and 945 Vesuvius (inverted triangles) eruptions. Numbers associated with diamonds correspond to 946 different phases of the P3 eruption: 1: beginning of phase A, 2: end of phase A, 3: 947 simultaneous phases B and C, 4: simultaneous phases D and E, and 5: simultaneous phases F 948 and G. Dashed line corresponds to the maximum mass eruption rate feeding the column 949 before collapse as a function of the total gas content at the column base (using Michaud-950 Dubuy et al. (2018) model for tropical conditions and D = 3.3). Open, gray, and dark symbols 951 correspond to geological data inferred for the stable column phase, partial column collapse 952 phase, and total column collapse phase, respectively. Data for the P1, P2, Taupo and Vesuvius 953 eruptions are compiled in Carazzo et al. (2012, 2019) and Michaud-Dubuy et al. (2018), 954 respectively.



















