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### ► To cite this version:

Guillaume Carazzo, Stephen Tait, Audrey Michaud-Dubuy, Allan Fries, Edouard Kaminski. Transition from stable column to partial collapse during the 79 cal CE P3 Plinian eruption of Mt. Pelée volcano (Lesser Antilles). *Journal of Volcanology and Geothermal Research*, 2020, 392, pp.106764. 10.1016/j.jvolgeores.2019.106764 . insu-02920453

**HAL Id: insu-02920453**

**<https://insu.hal.science/insu-02920453>**

Submitted on 21 Jul 2022

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1 **Transition from stable column to partial collapse during the 79 cal CE P3**

2 **Plinian eruption of Mt Pelée volcano (Lesser Antilles)**

3

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## 14 **Highlights**

- 15 - We determine the eruptive sequence of the P3 Plinian eruption of Mt Pelée volcano
- 16 - New radiocarbon dating measurements provide a refined age of  $79 \pm 21$  cal CE
- 17 - We estimate key eruptive parameters of this VEI 5 eruption in the transitional regime
- 18 - Increasing eruption rate and decreasing gas content led to partial column collapse
- 19 - The transition occurred at conditions well predicted by our 1D theoretical model

20

## 21 **Keywords**

22 Mt Pelée volcano, Plinian eruption, tephra fallout, partial collapse, tephrostratigraphy,  
23 radiocarbon dating

24

## 25 **Abstract**

26 Explosive volcanic eruptions commonly form sustained Plinian columns that collapse at some  
27 stage producing dangerous pyroclastic density currents (PDC) on the ground. Numerical and  
28 laboratory models of volcanic plumes show that the conditions leading to total column  
29 collapse are strongly controlled by the amount of exsolved gas at the source and the mass  
30 eruption rate. However, column collapse is rarely total and the volcanic jet often separates in  
31 a dense collapsing part feeding PDC and a buoyant rising plume spreading volcanic gases and  
32 pyroclasts in the atmosphere. This transitional regime has been directly observed and/or  
33 inferred from the structure of the deposits for several past eruptions, but the number of cases  
34 for which the partial collapse regime is described in detail, including the 79 CE Vesuvius, and  
35 the 186 CE Taupo eruptions, remains too small to fully constrain physical models. Here, we  
36 present a detailed reconstruction of the time evolution of the P3 eruption at Mt Pelée volcano  
37 (Martinique, Lesser Antilles) that underwent partial column collapse in order to discuss the  
38 mechanisms controlling the eruption dynamics and improve the volcanological database on

39 transitional eruptions. The P3 eruptive succession consists of seven major phases that  
40 produced a total of  $1 \text{ km}^3$  dense rock equivalent (DRE) of deposits (i.e., VEI 5 event), starting  
41 with a thick pumice fall deposit ( $0.1 \text{ km}^3$  DRE) overlain by alternating pyroclastic density  
42 current (PDC) ( $0.7 \text{ km}^3$  DRE) and pumice fall deposits ( $0.2 \text{ km}^3$  DRE). We use physical  
43 models together with field data on deposit dispersal, thickness, and grain-size distribution to  
44 reconstruct the dynamical evolution of the volcanic column. Our results show that the mass  
45 eruption rate (MER) increased from  $1.2 \times 10^8$  to  $1.7 \times 10^8 \text{ kg s}^{-1}$  during the initial phase  
46 producing a 28 to 30 km-high Plinian plume. The MER later reached up to  $2.5 \times 10^8 \text{ kg s}^{-1}$   
47 and the column entered the partial collapse regime characterized by the formation of a small  
48 (i.e., 12 to 17 km-high) ash plume and contemporaneous PDCs mainly channelized in three  
49 paleo-valleys. These estimates are used together with published data on magmatic water  
50 contents in glass inclusions to decipher the mechanisms leading to partial collapse. The P3  
51 eruption column collapsed due to an increase in mass eruption rate and a decrease in gas  
52 content. A similar evolution was also inferred for the 1300 CE P1 and 280 CE P2 eruption  
53 deposits, revealing a systematic behavior in the recent Plinian eruptions of Mt Pelée volcano.  
54 The comparison of model predictions of column collapse and field data reveals a good  
55 agreement for the P1, P2, P3 and Taupo eruptions, but not for the 79 CE Vesuvius eruption  
56 where thermal disequilibrium between gas and pyroclasts most likely strongly affected the  
57 column dynamics.

58

## 59 **1. Introduction**

60 Plinian eruptions are amongst the most powerful and destructive volcanic events on Earth.  
61 These explosive eruptions generally form a sustained stable column that can rise up to 40 km  
62 in the atmosphere (e.g., the 1991 Pinatubo eruption) and spread out laterally as an umbrella  
63 cloud carrying pyroclasts at large distances from the volcano ([Koyaguchi and Tokuno, 1993](#)).

64 Abrupt or progressive changes in eruptive conditions at the base of a Plinian column can lead  
65 to a radically different behavior in which the column collapses as a turbulent fountain  
66 producing hazardous pyroclastic density currents (PDC) that may reach and destroy populated  
67 areas ([Wilson et al., 2014](#)). One important goal in physical volcanology is to identify the key  
68 processes affecting the column stability in order to improve hazard assessment.

69         Recent real-time observations of volcanic plumes and field-based studies on fallout  
70 tephra from past eruptions show that explosive eruptions commonly have a complex behavior  
71 where both a stable Plinian column and a collapsing fountain feeding PDC are  
72 contemporaneous ([Carazzo et al., 2015 and references therein](#)). This transitional regime of  
73 partial column collapse occurs for Plinian eruptions ranging in size from small to large (e.g.,  
74 the 79 CE Vesuvius ([Shea et al., 2011](#)), 186 CE Taupo ([Wilson and Walker, 1985](#)), 1150  
75 Quilotoa ([Di Muro et al., 2008](#)), 1912 Novarupta-Katmai ([Fierstein and Hildreth, 1992](#)), and  
76 1963 Agung ([Self and Rampino, 2012](#)) eruptions). Detailed mapping and analysis of eruptive  
77 products emplaced during the transitional regime provide valuable data to better understand  
78 the mechanisms controlling the partitioning of a Plinian column, and can be used to robustly  
79 test numerical and analog models of explosive volcanic eruptions ([Di Muro et al., 2004](#);  
80 [Carazzo et al., 2015](#); [Suzuki et al., 2016](#)). However, the number of natural cases for which the  
81 partial collapse of an initially stable column is described in detail (i.e., in terms of time  
82 variation in mass eruption rate, gas content, grain-size distribution and exit velocity) remains  
83 too small to fully constrain theoretical models of volcanic plumes ([Michaud-Dubuy et al.,](#)  
84 [2018](#)). With an aim of increasing the inventory of well-constrained past eruptions, we have  
85 undertaken a long-term study of the Plinian eruptions of Mt Pelée volcano (Lesser Antilles) as  
86 they systematically produced sustained columns that ultimately collapsed ([Traineau et al.,](#)  
87 [1989](#)). Because the Mt Pelée eruptions share many characteristics including similar pre-  
88 eruptive magma storage conditions ([Martel et al., 1998](#)), they can be used to decipher the

89 mechanisms controlling the behavior of a Plinian plume. In previous studies, we reconstructed  
90 the dynamics of the 1300 CE P1 (Carazzo et al., 2012), and 280 CE P2 Plinian eruptions  
91 (Carazzo et al., 2019) focusing on the causes for column collapse. We now extend this work  
92 to study the P3 eruption deposits whose structure suggests the entrance into the partial  
93 collapse regime at some stage of the eruption (Traineau et al., 1989; Wright et al., 2016),  
94 making it an excellent candidate to improve the volcanological database on transitional  
95 eruptions.

96 The P3 eruption is described in the literature (Roobol and Smith, 1976; Westercamp  
97 and Traineau, 1983; Traineau et al., 1989; Wright et al., 2016) but the exact sequence of  
98 events and the eruptive parameters are currently poorly known. Here, we use data from a new  
99 comprehensive field study together with physical models to reconstruct the detailed time  
100 evolution of this eruption and to discuss the mechanisms responsible for the transition from  
101 stable column to partial collapse. We identify the eruptive conditions leading to this  
102 transitional regime and compare the results to previous work on P3, and on younger Plinian  
103 eruptions in Martinique (P1 and P2). We then compare our estimated eruptive parameters for  
104 the column at collapse with theoretical predictions from a model of volcanic plumes.

105

## 106 **2. Overview of the P3 eruption**

107 The P3 deposits were first identified by Roobol and Smith (1976) as a thick pumice fallout  
108 unit on the western flank of Mt Pelée volcano standing above the 2406 and 2447 yr BP  
109 horizons of their stratigraphic sections. Westercamp and Traineau (1983) provided the first  
110 age, synthetic stratigraphic section and distribution map of the P3 deposits (Fig. 1a). The P3  
111 Plinian eruption is currently dated at  $2010 \pm 140$  yr BP based on the average of five  $^{14}\text{C}$  ages  
112 (Westercamp and Traineau, 1983; Traineau et al., 1989). According to these studies, the  
113 eruption began with the formation of a sustained Plinian column that covered the western

114 flank of the volcano with a meter of coarse lapilli, lithic-rich, pumice fallout deposit (phase  
115 P3<sub>1</sub>). This initial stage was immediately followed by the production of high-concentration  
116 PDC in the Grande Savane, Habitation Depaz and Falaise River paleo-valleys (Fig. 1a). The  
117 second eruptive phase identified by Westercamp and Traineau (1983) consisted of a sustained  
118 Plinian column that spread fine lapilli pumice to the south with a striking elongated axis of  
119 dispersion (phase P3<sub>2</sub>). The P3 eruption ended up with the formation of a pumice ash-rich bed  
120 (phase P3<sub>3</sub>) interbedded with high-concentration PDC deposits (Fig. 1a). Later, Traineau et al.  
121 (1989) excluded the second phase (P3<sub>2</sub>) whose distal deposits were found below tephra from  
122 an eruption older than P3<sub>1</sub>. Recent work by Michaud-Dubuy et al. (2019) confirmed that the  
123 deposits named P3<sub>2</sub> by Westercamp and Traineau (1983) are not interbedded between P3<sub>1</sub> and  
124 P3<sub>3</sub> but belong to a much older event, the 13.5 ka cal BP Bellefontaine Plinian eruption.

125 Wright et al. (2016) studied the distal products of a Plinian eruption dated between  
126 1800 and 2200 yr BP (their subunit d), which corresponds to the P3 eruption identified by  
127 Westercamp and Traineau (1983). Their synthetic stratigraphic section includes a basal  
128 pumice-rich low-concentration PDC (named 'surge') deposit overlain by a pumice fallout  
129 deposit interbedded with high-concentration PDC (named 'flow') and low-concentration PDC  
130 (named 'ash hurricane') deposits. According to their distribution map of the deposits (Fig. 1b),  
131 the high-concentration PDC flowed in the Fond Canonville valley, Grande River, Bijou River,  
132 and Falaise River, whereas the low-concentration PDC that elutriated from the top of the  
133 current spread on hills at distances up to 20 km from the vent. Isopachs of the pumice fallout  
134 deposit exhibit a SSW elongation (Fig. 1b), in good agreement with the one inferred by  
135 Westercamp and Traineau (1983) (Fig. 1a). Wright et al. (2016) interpreted their complete  
136 stratigraphic succession as resulting from the partial and final collapse of an initially stable  
137 Plinian column.

138 The eruptive parameters of the P3 event are currently poorly constrained. Volume  
139 estimates for the fallout tephra of the P3 eruption yield a relatively small value of 0.19 km<sup>3</sup>  
140 dense rock equivalent (DRE) (Traineau et al., 1989). The direction of tephra dispersal during  
141 the initial phase (P3<sub>1</sub>) suggests that the column was affected by low altitude tropospheric  
142 winds, and thus was certainly more than 6 km (Traineau et al., 1989), but no more than 12 km  
143 in height (Wright et al., 2016). Such a large uncertainty on the maximum column height  
144 makes it difficult to estimate its mass eruption rate. Overall, these sparse and loosely  
145 constrained data cannot be used to understand the conditions leading to column collapse.

146

### 147 **3. Methodology**

#### 148 3.1. Fieldwork

149 We identified the P3 eruption deposits at 102 locations in Martinique (Fig. 2). In most cases,  
150 these deposits are easily recognizable by their strong internal layering, the abundance of  
151 juvenile and accidental lithic fragments in the pumice fallout units, and the presence of the P1  
152 and/or P2 deposits above them (Carazzo et al., 2012; 2019). The P3 outcrops are distributed  
153 all around the volcano except to the northwest where exposure is very limited due to dense  
154 tropical forest and difficult conditions of access. The thickness of each layer of P3 was  
155 measured at every site in order to construct isopach maps. We also excavated a standard 25 x  
156 20 cm area of each fall layer and measured the major axes of the five largest lithic fragments  
157 found in order to produce isopleth maps.

158

#### 159 3.2. Radiocarbon dating

160 We dated four charcoals sampled within the P3 deposits at sites 6 (unit A-G), 97 (unit A), 178  
161 (unit A), and 180 (unit A-G), and a paleosol sampled beneath unit A at site 145 (see locations  
162 in Fig. 2 and unit descriptions in section 4.1) in order to refine the age of the P3 eruption.

163 Ages were determined using accelerator mass spectrometry at the LMC14 (Artemis,  
164 Laboratoire de Mesure du Carbone 14, CEA Saclay, France), and calibrated using the free  
165 software OxCal 4.3 (Bronk Ramsey, 2009) with the atmospheric IntCal13 calibration curve  
166 commonly used for the Northern hemisphere (Reimer, 2013). The radiocarbon ages obtained  
167 for our stratigraphically-constrained samples were combined with those of Westercamp and  
168 Traineau (1983) and validated using the R\_combine function of Oxcal and  $\chi^2$  test prior to  
169 calibration (Ward and Wilson, 1978).

170

### 171 3.3. Grain-size analysis

172 We carried out grain-size analyses on twenty samples from ten locations representative of the  
173 seven different units (A to G) (see Table 1). The samples were dried for 24 h in an oven and  
174 sieved by hand down to  $6\phi$ . We separated the lithic fragments from the pumices by hand in  
175 the size range  $-6\phi$  to  $-4\phi$ , and used a binocular microscope to discard crystals and lithic  
176 fragments in the size range  $-4\phi$  to  $-2\phi$ . Volume calculations for isomass maps for each  $\phi$   
177 interval were then used to determine the grain-size distributions of single sub-layers. The total  
178 grain-size distribution was estimated from the total mass in sieve class  $\phi$ ,  $M_\phi$ , by the volume  
179 integral (Kaminski and Jaupart, 1998)

$$180 \quad M_\phi = \int_0^L h(l) C_\phi(l) A(l) dl, \quad (1)$$

181 where  $h(l)$  is the deposit thickness,  $C_\phi(l)$  is the concentration of class  $\phi$  at distance  $l$  from the  
182 vent,  $A(l)dl$  is the area bounded by isopachs at distances  $l$  and  $l+dl$ , and  $L$  is the distance  
183 where  $h$  or  $C_\phi$  drop to zero. We used linear interpolations for  $h$  and  $C_\phi$  between the localities.

184 The cumulative frequency curves were determined assuming a power-law size  
185 distribution of the rock fragments (Turcotte, 1986; Kaminski and Jaupart, 1998), where the  
186 number of particles with a radius larger than  $r$  is given by:

187  $N(R_p \geq r_p) = \lambda r^{-D},$  (2)

188 with  $D$  the power-law exponent that commonly varies between 2.9 and 3.9 for Plinian  
189 eruptions (Girault et al., 2014). Because fragmentation is a size-invariant process, the  
190 exponent  $D$  can be estimated accurately using any sufficiently large range of sizes. Thus, the  
191 lack of fine particles lost out at sea does not affect our estimations of  $D$ .

192

### 193 3.4. Eruptive parameters

194 The volume of tephra fallout produced during an explosive eruption can be inferred using  
195 several methods based on the thinning trend of the deposit with distance from the source.  
196 Here, we used the exponential, power-law and Weibull fits (Pyle, 1989; Bonadonna and  
197 Houghton, 2005; Bonadonna and Costa, 2012) computed using the AshCalc software (Daggit  
198 et al., 2014). Bearing in mind that only proximal and furthermore incompletely preserved  
199 deposits are available in Martinique, the calculated volumes of fall deposits should be taken  
200 as minimum estimates. The volume of dilute PDC deposits was calculated using the observed  
201 relationship between the volume and area covered by dilute PDC (Dade and Huppert, 1998;  
202 Calder et al., 1999), as in Carazzo et al. (2012, 2019) for the P1 and P2 eruptions. The volume  
203 of dense PDC was calculated using the product of the average thickness and area covered by  
204 the dense PDC. This method is bound to provide minimum estimates since only proximal and  
205 furthermore incompletely preserved deposits are available in Martinique.

206 The maximum column heights associated with the fall deposits were estimated from  
207 the distribution of lithic fragments on our isopleth maps using the method of Carey and  
208 Sigurdsson (1986). Error bars were calculated using the three values of maximum height  
209 inferred from the 8, 16, and 32-mm isopleths (see Fig. 17 in Carey and Sigurdsson, 1986).  
210 The alternative method of Bonadonna and Costa (2013) based on variations of grain size with  
211 the distance from the source was also used to confirm the robustness of the estimates.

212 The mass discharge rate feeding the eruption during phases of stable column was  
213 inferred from a given column height by using the empirical relationship from [Mastin et al.](#)  
214 [\(2009\)](#) and the predictions of the model of [Girault et al. \(2016\)](#), which explicitly includes the  
215 effect of total grain-size distribution on the plume dynamics. Calculations were made for  
216 tropical atmospheric conditions and for a crystal-bearing rhyolitic magma (andesitic bulk  
217 composition) with an initial temperature of 1,150 K ([Martel and Poussineau, 2007](#)) and a  
218 dense rock density of 2,400 kg m<sup>-3</sup> ([Traineau et al., 1989](#)). The mass discharge rates for the  
219 collapsing phases were calculated using the model of [Bursik and Woods \(1996\)](#), which treats  
220 pyroclastic density currents as supercritical dilute suspensions entraining air as they  
221 propagate. [Doyle et al. \(2010\)](#) showed that this assumption is valid for tall, fine-grained  
222 column collapses, for which the flow slowly transfers its mass to the dense basal flow. The  
223 model of [Bursik and Woods \(1996\)](#) is thus relevant for collapsing phases characterized by the  
224 formation of dilute PDC deposits, and was used here.

225

## 226 **4. The P3 eruptive succession**

### 227 4.1. Stratigraphy

228 [Fig. 3a](#) shows a composite stratigraphic section of the P3 deposits, divided into seven major  
229 phases based on diagnostic stratigraphic and lithofacies associations, which we describe in  
230 more detail below.

231

#### 232 4.1.1. Unit A

233 The stratigraphic succession begins with a layer of clast-supported, coarse white pumice  
234 lapilli ([Fig. 4a-d](#)). Unit A contains juvenile and accidental lithic fragments in a total amount  
235 that typically increases from  $\approx 20$  wt.% at base to  $\approx 25$  wt.% at top ([Fig. 3c](#)). At most sites,

236 median pumice size remains relatively constant, whereas both the maximum lithic size and  
237 the sorting increase from base to top (Fig. 3b-c).

238 In the downwind direction (to the southwest), unit A has a maximum thickness of 180  
239 cm at 4.5 km from the crater (Fig. 4c). In the crosswind direction, unit A is 130 cm thick at  
240 4.6 km, and steadily thins to 50 cm within 6.1 km of the crater (Fig. 4d). The relatively  
241 widespread nature of this deposit, its uniformly decreasing thickness with distance from the  
242 source, its clast-supported framework, and its pumice and juvenile lithic fragments  
243 characteristics identify unit A as a fall deposit.

244

#### 245 4.1.2. Unit B

246 Unit B consists of light brown fine ash material in which pumice, lithic fragments and crystals  
247 are dispersed in a matrix of dense, angular, glass fragments (Fig. 4b, c, e). Most clasts range  
248 in size from fine to coarse ash but a few large pumices are present (Fig. 3b). The total amount  
249 of lithic fragments in unit B commonly reaches  $\approx 28$  wt.%. Where unit B is present, the  
250 contact with the underlying layer is sharp, which suggests that there was no time break  
251 between deposition of the two units.

252 Unit B varies irregularly in thickness with distance from the source, and ranges from  
253 18 to 35 cm at 3.8 km from the eruptive center (Fig. 4), but is absent beyond 7 km from the  
254 vent. Where thick, unit B contains a few lenses of coarse lapilli, angular to sub-rounded  
255 pumices. At some sites in valleys, pumice lapilli and blocks are rounded and account for a  
256 large proportion in the matrix of white fine ash. The relatively limited dispersal of this unit, its  
257 irregular thickness, matrix-supported framework and grain types (juvenile and accidental  
258 lithic fragments), as well as the nature of the lower contact, indicate that unit B is a low-  
259 concentration pyroclastic density current deposit (Branney and Kokelaar, 2002).

260

261 4.1.3. Unit C

262 Unit C is a relatively thin blanket of clast-supported, white pumice lapilli (Fig. 4b, c, e)  
263 containing juvenile and accidental lithic fragments in a total amount that reaches  $\approx 25$  wt.% in  
264 most studied outcrops (Fig. 3c). At most field sites, unit C is easily distinguishable from unit  
265 A, the former being distinctly richer in lithic fragments and coarse lapilli pumice. The  
266 maximum lithic size in unit C is, however, always lower than the maximum lithic size at the  
267 top of unit A (Fig. 3c).

268 Unit C was dispersed in approximately the same direction as unit A, and is the thickest  
269  $\approx 3.7$  km southwest of the volcano in the downwind direction, where it reaches 22 cm. This  
270 layer steadily thins with distance from the vent, and is still 10 cm thick at 6.2 km in the  
271 crosswind direction. At most sites, the upper part of unit C was eroded by the pyroclastic  
272 density current that overrode it and deposited unit D (see below). The relatively widespread  
273 nature of unit C, its uniformly decreasing thickness with distance from the source, its clast-  
274 supported framework, and its pumice and juvenile lithic fragments characteristics identify it  
275 as a fall deposit.

276

277 4.1.4. Unit D

278 Unit D is a relatively thin layer of fine ash and small rounded pumice lapilli dispersed in a  
279 matrix of dense, angular, glass fragments (Fig. 4a, c, f). At distal sites, this layer is thin,  
280 stratified and unconsolidated. The total amount of lithic fragments in unit D commonly  
281 reaches  $\approx 20$  wt.% (Fig. 3c). Where unit D is present, the contact with the underlying layer is  
282 sharp and exhibits some erosion of unit C.

283 Unit D varies irregularly in thickness with distance from the vent, and ranges from 5 to  
284 17 cm at 3.8 km from the volcano (Fig. 4). This layer is, however, absent beyond 7 km from  
285 the vent. These irregular thicknesses, the relatively limited dispersal of this deposit, the clast

286 and matrix types, and the nature of the lower contact indicate that unit D is a low-  
287 concentration pyroclastic density current deposit.

288

#### 289 4.1.5. Unit E

290 Unit E is a layer of clast-supported, white pumice lapilli (Fig. 4b, c, e) containing juvenile and  
291 accidental lithic fragments in a total amount that typically reaches  $\approx 18$  wt.% (Fig. 3c). At  
292 most sites, both the median grain-size and the maximum lithic size are lower than those  
293 measured in unit C (Fig. 3c).

294 Material was dispersed in approximately the same direction as units A and C to the  
295 southwest of the volcano. The maximum thickness is 13 cm at 3.7 km, and steadily thins to 9  
296 cm within 4.8 km from the vent in the downwind direction. The full thickness of unit E is  
297 however poorly preserved due to significant erosion by the overlying pyroclastic density  
298 current that overrode it and deposited unit F (see below). The relatively widespread nature of  
299 this deposit, its uniformly decreasing thickness with distance from the source, its clast-  
300 supported framework, and its pumice and juvenile lithic fragments characteristics identify unit  
301 E as a fall deposit.

302

#### 303 4.1.6. Unit F

304 Unit F is a matrix-supported layer of fine to coarse ash containing rounded pumice lapilli and  
305 lithic fragments (Fig. 4b, f), the latter of which typically reaches  $\approx 15$  wt.% (Fig. 3c). Where  
306 unit F is present, the contact with the underlying layer is sharp and exhibits some erosion of  
307 unit E.

308 Unit F varies irregularly in thickness with distance from the vent depending on the  
309 topography, ranges from 6 to 75 cm at 3.8 km from the volcano, and is absent 7 km from the  
310 vent (Fig. 4). The relatively limited dispersal of this unit, its irregular thickness, matrix-

311 supported framework and grain types (juvenile and accidental lithic fragments), as well as the  
312 nature of the lower contact, indicate that unit F is a low-concentration pyroclastic density  
313 current deposit.

314

#### 315 4.1.7. Unit G

316 Unit G is the uppermost layer in the P3 sequence and consists of a blanket of fine to coarse  
317 white pumice lapilli (Fig. 4a, b, f). This unit contains juvenile and accidental lithic fragments  
318 in a total amount that reaches  $\approx 20$  wt.% at most studied outcrops (Fig. 3c) with a maximum  
319 lithic size similar to that of unit C (Fig. 3c).

320 In the downwind direction (to the southwest), unit G has a maximum thickness of 80  
321 cm at 1.4 km from the crater, and steadily thins to 20 cm within 4.8 km of the vent. In the  
322 crosswind direction, unit G is 30 cm thick at 5 km, and 15 cm at 6.4 km from the vent. The  
323 relatively widespread nature of this deposit, its uniformly decreasing thickness with distance  
324 from the source, its clast-supported framework, and its pumice and juvenile lithic fragment  
325 characteristics identify unit G as a fall deposit.

326

#### 327 4.2. Age of the P3 eruption

328 Radiocarbon dates and calibrated ages are reported in [Supplementary material Table S1](#). Two  
329 samples provide the exact same radiocarbon age of  $1,870 \pm 30$  yr BP (sites 145 and 178 on  
330 the western flank). A third sample taken near location 178 gives a radiocarbon age of  $1,915 \pm$   
331  $30$  yr BP (site 97). The oldest age is found near the town of Le Carbet where the charcoal  
332 sampled within the pumice fallout deposit provides an age of  $2,030 \pm 30$  yr BP. On the  
333 eastern flank of the volcano, the charcoal sampled at location 6 has an age of  $1,795 \pm 30$  yr  
334 BP. These five new radiocarbon ages are in good agreement with a compilation of twelve  $^{14}\text{C}$   
335 ages determined in previous studies ([Walker, 1973](#); [Roobol and Smith, 1976](#); [Traineau, 1982](#);

336 [Westercamp and Traineau, 1983](#)) ranging from 1,800 to 2,150 yr BP ([Supplementary material](#)  
337 [Table S1](#)). Combining these seventeen values using the R\_combine function of OxCal ([Bronk](#)  
338 [Ramsey, 2009](#)) provides an age of  $1,926 \pm 11$  yr BP for the P3 eruption. We note that the  $\chi^2$   
339 test fails here most likely due to the large number of individual measurements. However, a  
340 reasonable correction on the error bar of each dating by a factor of 2 is enough to successfully  
341 pass the  $\chi^2$  test, reinforcing the confidence that all these otherwise stratigraphically-  
342 constrained samples belong to the same eruption. Our new refined radiocarbon age for the P3  
343 eruption stands in the lower error bars of the  $2,010 \pm 140$  yr BP proposed by [Traineau et al.](#)  
344 ([1989](#)). Correcting our new estimate with the IntCal13 calibration curve ([Reimer, 2013](#))  
345 available in OxCal gives a mean calendar age of  $79 \pm 21$  cal CE (i.e., 31-37 cal CE at 1.7%  
346 probability and 51-125 cal CE at 93.7% probability).

347

#### 348 4.3. Grain-size distribution of selected samples

349 Twenty samples from the P3 units were analyzed ([Table 1](#)) in order to discuss the evolution of  
350 the grain-size distribution during the entire P3 eruption at a single outcrop ([Fig. 3](#)), and to  
351 determine the total grain-size distribution for the lapilli fraction of unit A deposits (see section  
352 4.6). The objective is twofold: to confirm the fall or flow nature of the P3 units, and to  
353 estimate the amount of gas trapped in pumice fragments at the beginning of the eruption (see  
354 section 6.3). A detailed analysis of the pumice textures is beyond the scope of this paper and  
355 can be found in [Martel and Poussineau \(2007\)](#).

356 The fourteen samples from unit A display the typical fallout characteristics with  
357 median diameter ranging from  $0.7\phi$  to  $-3.3\phi$ , and sorting ranging from 1.4 to 2.7. These  
358 values are relatively close to those obtained by [Bardintzeff et al. \(1989\)](#) who measured  
359 median diameter and sorting ranging from  $-0.3\phi$  to  $-3\phi$ , and 1.7 to 2.8, respectively. Most of  
360 the samples collected on land belong to proximal deposits since the medial and distal deposits

361 were lost at sea. The grain-size distribution of individual samples is systematically bimodal  
362 (Fig. 3b) and the amount of ash particles (< 2 mm) always increases from base ( $\approx 9$  wt.%) to  
363 top ( $\approx 36$  wt.%). The top of unit A is thus always more poorly sorted than the base, but the  
364 median grain-size remains approximately constant throughout the unit (Fig. 3c).

365 The six samples from units B to G collected at outcrop 163 show that units C, E, and  
366 G have typical fallout characteristics similar to unit A. The median diameter for these units  
367 ranges between  $-0.1\phi$  and  $-3\phi$ , and they have moderate sorting values from 2.5 to 2.7 (Fig.  
368 3c). The amount of ash-sized particles is similar to or larger than in unit A, ranging from  $\approx 28$   
369 wt.% to  $\approx 55$  wt.%. Units B, D, and F have flow features with median diameter ranging from  
370  $1.7\phi$  and  $0.4\phi$ , and moderate sorting ranging from 2.2 to 2.6 (Fig. 3c). The amount of ash-  
371 sized particles is larger than in the fall deposits, ranging from  $\approx 62$  wt.% to  $\approx 85$  wt.%.

372

#### 373 4.4. Isopach maps

374 Fig. 5 shows a stratigraphic correlation of P3 outcrops along three different dispersal axes  
375 (i.e., north to south, northwest to southeast, and northeast to southwest, see Fig. 2 for  
376 localization). The complete sequence can be found up to 5 km from the vent (sites 13, 163,  
377 101, 78 in Fig. 5). Thickness measurements at each location are reported on isopach maps for  
378 the cumulative pumice fallout phases A+C+E+G (Fig. 6a), the pumice fallout phase A (Fig.  
379 6b), the cumulative pyroclastic density current phases B+D+F (Fig. 6c), and the cumulative  
380 pumice fallout phases C+E+G (Fig. 6d). These maps are later used to calculate the total grain-  
381 size distribution of unit A (see section 4.6), and the volume of deposits (see section 5.1).

382 The cumulative isopach map of the pumice fallout phases A+C+E+G shows  
383 ellipsoidal contour patterns indicating fallout dispersion towards the southwest (Fig. 6a). We  
384 find a similar direction of dispersion for the pumice fallout phase A, and the cumulative  
385 pumice fallout phases C+E+G (Fig. 6b, d). This direction of dispersion is in relatively good

386 agreement with the results of [Traineau et al. \(1989\)](#) and [Wright et al. \(2016\)](#) ([Fig. 1](#)), and is  
387 characteristic of the wind profiles in the Lesser Antilles during the wet season ([Michaud-  
388 Dubuy et al., 2019](#)). In the upwind direction to the northeast and in the crosswind direction to  
389 the southeast, the entire P3 succession is reduced to a single layer of clast-supported, white,  
390 angular, fine pumice lapilli containing abundant juvenile and accidental lithic fragments,  
391 making it difficult to distinguish the pumice fallout units A, C, E, and G. At locations near the  
392 vent, unit A is however clearly visible under the cumulative pumice fallout of units C+E+G  
393 ([Fig. 4d](#)), which allows to build an individual isopach map for the pumice fallout phase A  
394 ([Fig. 6b](#)). We note that the northwestern arms of our isopach maps are not well constrained  
395 due to a lack of outcrops (in our study as in previous ones) in a region extremely difficult to  
396 access due to the presence of dense tropical forest on steep hills.

397 Units B, D, and F are widespread on the southwest flank of Mt Pelée volcano and the  
398 low-concentration PDC deposits almost reached the sea there ([Fig. 6c](#)). The high-  
399 concentration PDC deposits are channelized in paleo-valleys and now form large plateaux in  
400 the Grande Savane, Habitation Depaz and Ajoupa Bouillon areas (see locations in [Fig. 1a](#)), an  
401 observation consistent with the results of [Westercamp and Traineau \(1983\)](#) and [Traineau et al.  
402 \(1989\)](#). However, contrary to [Wright et al. \(2016\)](#), we found no evidence of high-  
403 concentration PDC deposits in the Bijou River, Grande River, and Fond Canonville areas (see  
404 locations in [Fig. 1b](#)).

405

#### 406 4.5. Isopleth maps of air fall deposits

407 [Fig. 7](#) shows the isopleth maps built from the measurements of the major axes of the five  
408 largest fragments found at the base and top of unit A, and at the base of unit C. The isopleth  
409 map of the base of unit A is well constrained thanks to the good preservation of the deposit  
410 ([Fig. 7a](#)). However, isopleth contours for the top of unit A and the base of unit C are

411 constrained by only a few points from localities where the two units are separated by the  
412 presence of unit B and/or by clear variation in grain-size characteristics (Fig. 7b, c). Isopleth  
413 maps for units E and G could not be constructed due to the poor preservation of these  
414 deposits. However, in both cases, we were able to draw a single isopleth curve (i.e., the 15  
415 mm-isopleth for unit E and 25 mm-isopleth for unit G) that is not reported in Fig. 7. The main  
416 direction of dispersion to the southwest (Fig. 7) is consistent with the one inferred from our  
417 isopach maps (Fig. 6). Isopleth contours of the top of unit A are clearly more extended in the  
418 crosswind direction than those of the base of units A and C.

419

#### 420 4.6. Total grain-size distribution (*Unit A*)

421 We calculated the total grain-size distribution (TGSD) of unit A using the method of  
422 Kaminski and Jaupart (1998) (see section 3.3 for calculation details). The interbedded nature  
423 and important erosion of the subsequent units made the reconstruction of the TGSD for units  
424 B to G impossible. The TGSD of unit A is bimodal with a primary fine mode at  $2\phi$  and a  
425 secondary coarse mode at  $-2\phi$  (Supplementary material Figure S1). The median diameter and  
426 sorting for the TGSD of unit A are  $-0.2\phi$  and 2.3, respectively. The amounts of ash ( $< 2 \mu\text{m}$ )  
427 and fine ash particles ( $< 63 \mu\text{m}$ ) reach 59 wt.% and less than 1 wt.%, respectively. The latter  
428 value should be taken with caution because most of the fine ash particles settled at sea and  
429 cannot be taken into account in our grain-size analyses. From these results, we infer that the  
430 power law coefficient  $D$  that fully characterizes the TGSD at the beginning of the P3 eruption  
431 (i.e., unit A) is  $D = 3.3 \pm 0.1$  (Supplementary material Figure S1). The time evolution of the  
432 value of  $D$  after the phase A remains unknown, but this lack of information will not affect our  
433 calculations of the stable plume / collapsing fountain transition that occurred at the end of  
434 phase A (see section 6.3).

435

## 436 5. Eruptive dynamics

### 437 5.1. Erupted volumes

438 The estimates of volume of tephra fall from phase A using the three integration  
439 techniques introduced in section 3 give  $0.282 \text{ km}^3$  with the exponential method,  $0.316 \text{ km}^3$   
440 with the power law, and  $0.306 \text{ km}^3$  with the Weibull function (Fig. 8a). Integration of the  
441 exponential, power-law, and Weibull fits for the cumulative pumice fallout C+E+G yields  
442 volumes of  $0.334$ ,  $0.317$ , and  $0.355 \text{ km}^3$ , respectively (Fig. 8b). Because the latter estimates  
443 are based on three proximal isopach contours only they remain relatively poorly constrained  
444 (Fig. 6d). We thus calculated the volume of the cumulative pumice fallout A+C+E+G in order  
445 to achieve a better estimate of the total volume of tephra fall. Integrating the exponential,  
446 power-law, and Weibull fits yields  $0.735$ ,  $0.692$ , and  $0.712 \text{ km}^3$ , respectively (Fig. 8c). We  
447 thus retain a minimum volume of  $0.30 \pm 0.01 \text{ km}^3$  for unit A, and  $0.41 \pm 0.02 \text{ km}^3$  for the  
448 cumulative pumice fall units C+E+G, and  $0.71 \pm 0.02 \text{ km}^3$  for the total volume of tephra fall.  
449 Based on fallout deposit and magma densities of  $1,000$  and  $2,400 \text{ kg m}^{-3}$ , respectively  
450 (Traineau et al., 1989), we infer the cumulative DRE volume of fall units to be  $0.3 \text{ km}^3$ .

451 The volume of PDC deposits is calculated using the area covered by the dense and  
452 dilute PDC deposits. The deposits corresponding to the dilute PDC of phases B, D, and F  
453 cover  $42$ ,  $20$ , and  $36 \text{ km}^2$ , respectively (Fig. 6c), which yield bulk volumes of  $0.025 \text{ km}^3$  for  
454 unit B,  $0.010 \text{ km}^3$  for unit D, and  $0.016 \text{ km}^3$  for unit G (Calder et al., 1999). The cumulative  
455 dense PDC deposits of the phases B, D, and F cover  $7 \text{ km}^2$  in Grande Savane,  $9.5 \text{ km}^2$  in  
456 Habitation Depaz, and  $18 \text{ km}^2$  in Ajoupa Bouillon (see locations in Fig. 1a), with an average  
457 thickness of  $15$ ,  $30$ , and  $2 \text{ m}$ , respectively (Traineau, 1982). This yields a minimum volume of  
458  $0.67 \pm 0.02 \text{ km}^3$  DRE for the dense PDC deposits, and gives a total volume of (dense + dilute)  
459 PDC deposits of  $0.72 \pm 0.02 \text{ km}^3$  DRE.

460 The final estimate of the total volume of the P3 eruption (units A+B+C+D+E+F+G) is  
461 thus  $1.02 \pm 0.06 \text{ km}^3$  DRE, and the total mass of tephra emitted is estimated to be  $2.4 \pm 0.1 \times$   
462  $10^{12}$  kg. The P3 eruption thus stands as a VEI 5 event (Newhall and Self, 1984) with a  
463 magnitude of 5.4 (Pyle, 2000).

464

### 465 5.2. Column heights and exit velocities (*airfall units*)

466 The estimates of maximum height using our isopleth maps (Fig. 7) with the model of Carey  
467 and Sigurdsson (1986) give  $28.3 \pm 3.3$  km for the base of unit A,  $29.7 \pm 1.9$  km for the top of  
468 unit A, and  $16.9 \pm 0.4$  km for the base of unit C. Following the approach of Bonadonna and  
469 Costa (2013), a Weibull fit gives a maximum height of  $28.1 \pm 1$  km for the base of unit A,  
470  $30.0 \pm 0.5$  km for the top of unit A, and  $17.2 \pm 0.5$  km for the base of unit C (Supplementary  
471 material Figure S2). The lack of information about the distribution of lithic sizes in unit E and  
472 G prevents us from providing a robust estimate for the maximum height. Based on a single  
473 isopleth curve in each case, we calculate maximum column heights of  $12 \pm 4$  km for unit E  
474 and  $13 \pm 4$  km for unit G using the model of Carey and Sigurdsson (1986).

475 We also use data on the decrease of lithic sizes with distance from the vent in  
476 Supplementary material Figure S2 to estimate the minimum exit velocity of the volcanic  
477 plume at the vent. Extrapolating the Weibull fits down to  $A^{1/2} = 0.01$  km to find the maximum  
478 lithic size at the vent, we infer minimum exit velocities of  $210 \pm 10 \text{ m s}^{-1}$  for the base of unit  
479 A,  $220 \pm 10 \text{ m s}^{-1}$  for the top of unit A, and  $220 \pm 10 \text{ m s}^{-1}$  for the base of unit C.

480

### 481 5.3. Mass eruption rates and duration

482 We use our new estimates of maximum column height to calculate the mass eruption rate  
483 (MER) feeding the plume during the P3 eruption. The model of Girault et al. (2016) used with  
484 their complex wind profile, which is close to the average wind profiles in the Lesser Antilles,

485 yields maximum MERs of  $10^8 \text{ kg s}^{-1}$  for the beginning of phase A,  $1.5 \times 10^8 \text{ kg s}^{-1}$  for the end  
486 of phase A,  $3 \times 10^7 \text{ kg s}^{-1}$  for phase C,  $1.4 \times 10^7 \text{ kg s}^{-1}$  for phase E, and  $1.7 \times 10^7 \text{ kg s}^{-1}$  for  
487 phase G. The empirical formula of [Mastin et al. \(2009\)](#) built on observations of 28 Plinian  
488 eruptions yields maximum MERs of  $1.4 \times 10^8 \text{ kg s}^{-1}$  for the beginning of phase A,  $1.9 \times 10^8$   
489  $\text{kg s}^{-1}$  for the end of phase A,  $1.8 \times 10^7 \text{ kg s}^{-1}$  for phase C,  $4.2 \times 10^6 \text{ kg s}^{-1}$  for phase E, and  $5.9$   
490  $\times 10^6 \text{ kg s}^{-1}$  for phase G. Based on these two series of estimates, we retain peak MERs of  $1.2$   
491  $\pm 0.2 \times 10^8 \text{ kg s}^{-1}$  for the beginning of phase A,  $1.7 \pm 0.2 \times 10^8 \text{ kg s}^{-1}$  for the end of phase A,  
492  $2.4 \pm 0.6 \times 10^7 \text{ kg s}^{-1}$  for phase C,  $9.1 \pm 4.9 \times 10^6 \text{ kg s}^{-1}$  for phase E, and  $1.2 \pm 0.5 \times 10^7 \text{ kg s}^{-1}$   
493 for phase G. Combined with masses of fallout deposits (section 5.1), these MER provide  
494 minimum durations of  $\sim 45$  min for phase A,  $\sim 1\text{h}45$  for phase C,  $\sim 3\text{h}$  for phase E, and  $\sim 2\text{h}45$   
495 for phase G.

496 The MER during the collapsing phases B, D, and F is more difficult to assess. The  
497 model of [Bursik and Woods \(1996\)](#) yields a mass eruption rate of  $\sim 2.3 \pm 0.8 \times 10^8 \text{ kg s}^{-1}$  for  
498 the three phases B, D, and F. We note that this value does not take into account topographic  
499 slopes but [Doyle et al. \(2010\)](#) showed that the runout distance of such a flow mostly depends  
500 on the initial collapse height, which is imposed by the MER rather than by ground slopes.  
501 Combined with masses of PDC deposits (section 5.1), this MER provides minimum durations  
502 of  $\sim 50$  min for phase B,  $\sim 25$  min for phase D, and  $\sim 50$  min for phase F.

503

## 504 **6. Discussion**

505 Our revisit of the P3 deposits shows that the eruption started with the formation of a stable  
506 column that spread pyroclasts to the southwest of the volcano (phase A). This initial phase  
507 was immediately followed by a partial collapse of the column producing alternating tephra  
508 fallout (phases C, E, G) and pyroclastic density currents (phases B, D, E). We now compare  
509 our results to previous work on the P3 eruption, and on the younger P1 and P2 eruptions of

510 Mt Pelée volcano. We then focus our discussion on the causes of partial column collapse  
511 during the P3 eruption by using a physical model of volcanic plumes.

512

### 513 6.1 Comparison with previous studies of P3 eruption

514 Our new stratigraphic data regarding the P3 succession largely differs from those of previous  
515 studies (Westercamp and Traineau, 1983; Traineau et al., 1989; Wright et al., 2016). The P3<sub>1</sub>  
516 phase identified by Westercamp and Traineau (1983) corresponds to our eruptive unit A, and  
517 their P3<sub>3</sub> phase most likely corresponds to our entire sequence from unit B to unit G. The  
518 distal products identified by Wright et al. (2016) were attributed to the P3 eruption based on a  
519 single radiocarbon age determined from a charcoal sampled at the top of their stratigraphic  
520 section (see their figure 6). We revisited their outcrop locations 2 and 3 along a main road,  
521 and found that their stratigraphic section lies beneath the 13.5 ka cal BP Bellefontaine Plinian  
522 eruption (Michaud-Dubuy et al., 2019). To further reinforce our observation, we sampled a  
523 paleosol at the base of their stratigraphic section and obtained a radiocarbon age of  $17,750 \pm$   
524  $100$  yr BP, which indicates that the distal deposits studied by Wright et al. (2016) are much  
525 older than the P3 eruption.

526 The total volume of the P3 eruption (units A+B+C+D+E+F+G) is calculated to be  
527  $1.02 \pm 0.06$  km<sup>3</sup> DRE, which is one order of magnitude larger than the 0.19 km<sup>3</sup> DRE  
528 previously estimated (Traineau et al., 1989). We attribute this important discrepancy to the  
529 large volume of PDC deposits that were not taken into account in Traineau et al. (1989), and  
530 to the improvement of the reconstruction techniques used here to calculate the total volume of  
531 a pumice fall deposit (Bonadonna and Costa, 2013). The maximum column height is  
532 estimated to be ~ 30 km at the end of phase A, which is much higher than the previous  
533 estimates (> 6 km by Traineau et al., 1989 and < 12 km by Wright et al., 2016). Again we  
534 attribute this important difference to the better quality of our isopleth maps and to the

535 improvement of the reconstruction techniques used to estimate the maximum height from the  
536 distribution of lithic fragments in the field.

537

## 538 6.2 Comparison with younger Plinian eruptions in Martinique (P1 and P2)

539 The three latest Plinian eruptions at Mt Pelée volcano (P1, P2 and P3) have similar  
540 depositional successions that can be interpreted to result from the partial collapse of an  
541 initially stable Plinian column (Traineau et al., 1989; Carazzo et al., 2012, 2019; Wright et al.,  
542 2016). The total DRE volume estimated for P3 ( $\sim 1 \text{ km}^3$ ) is larger than those of the P1 ( $\sim 0.2$   
543  $\text{km}^3$ ) and P2 ( $\sim 0.8 \text{ km}^3$ ) eruptions. The maximum column height during the phase of sustained  
544 Plinian column is also larger for P3 (28-30 km) than for the P1 (19-22 km) and P2 (22-26 km)  
545 eruptions. Consequently, the mass eruption rate is estimated to be larger for P3 ( $1.4 \times 10^8 \text{ kg}$   
546  $\text{s}^{-1}$ ) than for P1 ( $3.6 \times 10^7 \text{ kg s}^{-1}$ ) and P2 ( $1.1 \times 10^8 \text{ kg s}^{-1}$ ). The minimum eruption durations  
547 are 11h for P3, 7h for P2, and 5h for P1. The total grain-size distribution is however relatively  
548 similar for the three events, with power law exponent  $D$  of 3.3, 3.4-3.5, and 3.2-3.3 for P3,  
549 P2, and P1, respectively. Our estimates of exsolved gas contents for P3 (2-2.9 wt.% - see  
550 section 6.3 for calculations details) are larger than those of P1 (1.6-2.1 wt.%) and P2 (1.7-2.1  
551 wt.%). The minimum exit velocities inferred from the distribution of lithic fragments for P3  
552 ( $210\text{-}220 \text{ m s}^{-1}$ ) are also larger than those of P1 ( $150\text{-}165 \text{ m s}^{-1}$ ) and P2 ( $180\text{-}200 \text{ m s}^{-1}$ ), a  
553 result consistent with our estimates of free gas contents. Table 2 summarizes the eruptive  
554 parameters retrieved for the last three Plinian eruptions at Mt Pelée volcano. P3 stands as a  
555 large Plinian eruption more powerful and voluminous (VEI 5, M 5.4) than the P1 (VEI 4, M  
556 4.6) and P2 (VEI 4, M 5.2) eruptions. However, because these three eruptions have much in  
557 common, in particular similar pre-eruptive conditions (Martel et al., 1998; Martel and  
558 Poussineau, 2007), they can be used to test the performance of theoretical models in  
559 predicting the conditions for column collapse.

560

### 561 6.3 The stable plume / pyroclastic fountain transition during a sustained Plinian eruption

562 The P3 eruption underwent partial column collapse after an initial phase of stable Plinian  
563 plume. [Wilson et al. \(1980\)](#) showed that the mass eruption rate and the amount of free gas in  
564 the volcanic mixture at the vent both strongly control the transition between the stable Plinian  
565 plume and the collapsing fountain regimes. Our estimates of the MER show a progressive  
566 increase over time from  $\sim 1.2 \times 10^8 \text{ kg s}^{-1}$  (base of unit A) to  $\sim 2.5 \times 10^8 \text{ kg s}^{-1}$  (unit B+C) ([Fig.](#)  
567 [9](#)), an evolution consistent with the increase in maximum lithic sizes throughout unit A ([Fig.](#)  
568 [3c](#)). Other powerful eruptions underwent a similar evolution leading to column collapse, such  
569 as the 79 CE Vesuvius ([Shea et al., 2011](#)), and the 186 CE Taupo ([Walker, 1980](#)) eruptions.

570 The time evolution of total gas content feeding the column at the vent is more difficult  
571 to assess. Petro-geochemical measurements provide quantitative estimates of pre-eruptive  
572 magma storage conditions, including temperature (875-900°C), pressure ( $\sim 200 \pm 50 \text{ MPa}$ ),  
573 and water content ( $\sim 5.8 \pm 0.5 \text{ wt.}\%$ ) ([Martel et al., 1998](#); [Martel and Poussineau, 2007](#)). Here,  
574 we use the magmatic water content measured in glass inclusions available in the literature as a  
575 total volatile content in the melt ( $n_0$ ). Correcting this value from the presence of crystals and  
576 lithic fragments ([Table 3](#)), we find that the initial gas content ( $x_0$ ) in the magma decreased  
577 from 4.1 to 2.9 wt.% during the eruption ([Table 3](#)). As in [Kaminski and Jaupart \(1998\)](#) and  
578 [Carazzo et al. \(2012; 2019\)](#), we calculate the mass fraction of exsolved gas in the mixture at  
579 fragmentation level ( $x_f$ ) for a threshold vesicularity of 70 %, which assumes closed-system  
580 conditions consistent with degassing models built on U-series measurements in the P3  
581 eruption products ([Villemant et al., 1996](#)). We find that  $x_f$  decreases from 2.9 to 2.1 wt.%  
582 during the eruption ([Fig. 9](#)). To strengthen these estimates, we calculate the theoretical exit  
583 velocities at the vent assuming that the volcanic mixture decompresses freely into the  
584 atmosphere ([Woods and Bower, 1995](#)). For this, we infer the effective amount of free gas in

585 the volcanic mixture at the base of the column that is modulated by gas entrapment by  
586 pumices as a function of their size distribution (Kaminski and Jaupart, 1998). We use the  
587 power law exponent  $D = 3.3$  (for unit A) estimated in section 4.6, and a typical open porosity  
588 of 60-70 % (Michaud-Dubuy et al., 2018) to calculate an amount of free gas of 2.4-2.7 wt.%  
589 for phase A. These estimates provide exit velocities after decompression in the range 215-230  
590  $\text{m s}^{-1}$  (Table 3) that are consistent with those estimated from the distribution of the lithic  
591 fragments collected at the base and top of unit A (see section 5.2). This relatively good  
592 agreement at least for phase A reinforces the confidence that our estimates of  $x_f$  provide  
593 reasonable values of free gas content during the P3 eruption.

594 [Fig. 10](#) compares the estimated values of MER and exsolved gas content in the column  
595 at collapse with the theoretical 1D predictions for tropical conditions (Michaud-Dubuy et al.,  
596 2018). The conditions at the beginning of the P3 eruption formed a stable plume as predicted  
597 by the model. During phase A, the combined effect of a slightly increasing MER and a  
598 decreasing gas content led the eruption to conditions close to the plume/fountain transition.  
599 The eruption reached a peak in MER during the concomitant phases B and C, which most  
600 likely induced the partial column collapse. During the partial column collapse stage of the  
601 eruption (units B to G), the MER did not change significantly but the gas content slowly  
602 decreased until the end of the eruption. The fairly good agreement between our field data and  
603 the theoretical model can be taken as an indirect confirmation of the proposed evolution of the  
604 total gas content feeding the column, even if a direct confirmation is not possible.  
605 Furthermore, our results suggest that the stable column underwent partial collapse soon after  
606 the beginning of the eruption (i.e., within less than a hour), an evolution that was also inferred  
607 from the analysis of the P1 and P2 Plinian eruption deposits (Carazzo et al., 2012; 2019). In  
608 these three eruptions, the transition occurred at conditions well predicted by our theoretical  
609 model of volcanic plumes ([Fig. 10](#)).

610 Comparing these results with other historical eruptions requires gathering well-  
611 constrained values of exsolved gas contents, mass eruption rates and power-law exponent  $D$   
612 of the TGSD (Michaud-Dubuy et al., 2018). There are however only a few examples of such  
613 well-documented events, and we choose here to compare our results to the 186 CE Taupo  
614 (VEI 7) and 79 CE Vesuvius (VEI 5) eruptions. Fig. 10 shows that the model successfully  
615 predicts the transition for the Taupo eruption but fails to explain the transition for the  
616 Vesuvius eruption. Furthermore, one can note that the collapsing phase of the 79 CE Vesuvius  
617 eruption is characterized by the same amount of gas and MER at the source as the stable  
618 phase (end of phase A) of the P3 eruption. The only difference between the two cases is the  
619 total grain-size distribution. Indeed, as discussed in Michaud-Dubuy et al. (2018), the 79 CE  
620 Vesuvius eruption is characterized by a specific total grain-size distribution ( $D = 3.0$ ) where  
621 pyroclasts are in average not large enough to be lost by sedimentation but not small enough to  
622 be in thermal equilibrium with the gas. The hypothesis of thermal equilibrium between the  
623 gas and pyroclasts that is used in our 1D model of volcanic plumes, valid for the fine grain-  
624 size distribution of the beginning of the P3 eruption ( $D = 3.3$ ), may therefore not hold  
625 anymore for this eruption and may explain the peculiar collapse conditions.

626

## 627 **7. Conclusion**

628 We have presented a new comprehensive field study of the P3 eruption, and a detailed  
629 reconstruction of the mechanisms controlling the eruptive dynamics. New radiocarbon  
630 measurements, averaged with those previously available, provide a refined age of  $79 \pm 21$  cal  
631 CE for this event. The eruption started with the formation of a 28 to 30 km-high stable  
632 column (phase A), which partially collapsed due to the combined effect of increasing MER  
633 and decreasing gas content (phase B). In the partial collapse regime, the eruptive mixture  
634 formed a 12 to 17 km-high marginally stable column (phases C, E, and G) punctuated with

635 the production of PDC fed by the partially collapsing column (phases B, D, and F). Our  
636 calculations show that P3 was a large Plinian eruption (VEI 5,  $M = 5.4$ ) evolving close to the  
637 plume/fountain transition. The total volume of tephra reached  $1 \text{ km}^3$  DRE, and the mass  
638 eruption rate was of the order of  $10^8 \text{ kg s}^{-1}$  during a period of at least 11h.

639 The P3 eruption is more powerful than the last two Plinian eruptions in Martinique but  
640 the three eruptions underwent similar evolution towards a marginally stable volcanic column.  
641 This recurrent phenomenon provides a great opportunity to look carefully at the parameters  
642 affecting column stability, to constrain theoretical models of volcanic plumes, and thus  
643 improve volcanic disaster forecasting. Petrological and geochemical measurements on the P3  
644 eruptive products may help to further investigate the complexity of magma degassing,  
645 crystallization and fragmentation in the volcanic conduit ([Gurioli et al., 2005](#); [Martel and](#)  
646 [Poussineau, 2007](#); [Shea et al., 2012; 2014](#)), and thus refine the estimated values of exsolved  
647 gas content at the vent. The global consistency between our theoretical predictions and field  
648 observations of column collapse for a number of Plinian eruptions confirms that current  
649 theoretical 1D models of volcanic plumes predict correctly the conditions of collapse when  
650 the total grain-size distribution is dominated by fine fragments (i.e., obeys a power law  
651 distribution with an exponent strictly larger than 3). Our results can thus be used to assess the  
652 dominant hazard during a Plinian scale event, which may pass from pumice fall to PDC as  
653 eruptive conditions vary.

654 Mt Pelée volcano is currently quiescent but the reconstruction of its eruptive history  
655 over the past 5,000 years shows that a Plinian eruption is a major potential scenario in the  
656 future. Such an event with the magnitude of the P3 eruption would disrupt the lives of more  
657 than 376,000 people living in Martinique and possibly impact other volcanic islands  
658 depending on the wind speed and direction.

659

660 **Acknowledgements**

661 We warmly thank the staff of the Mt Pelée volcanological observatory (OVSM) for field and  
662 administrative assistance, J.-P. Dumoulin, L. Beck, E. Delque-Kolic and C. Moreau (LMC14,  
663 CNRS UMS2572) who performed the <sup>14</sup>C dating, G. Delaviel-Anger and T. Crouzal for their  
664 help in the field, and A. Jegouzo for enthusiastic discussions about the P3 eruption. This work  
665 was partially funded by the Institut National des Sciences de l'Univers – Centre National de la  
666 Recherche Scientifique programme CT3-ALEA, INSU-CNRS Artemis 2016 and 2017 for <sup>14</sup>C  
667 dating, and CASAVA (ANR contract ANR-09-ANR-RISK-002) and RAVEX (ANR contract  
668 ANR-16-CE03-0002). This study contributes to the IdEx project "Université de Paris" (ANR-  
669 18-IDEX-0001).

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850

851 **Table captions**

852

| <b>Sample</b> | <b>Site</b> | <b>Unit</b> | <b>Subunit</b> | <b>Altitude<br/>(m)</b> | <b>Distance<br/>from the<br/>vent (km)</b> | <b>Thickness<br/>(cm)</b> |
|---------------|-------------|-------------|----------------|-------------------------|--|---------------------------|
| 1             | 163         | A           | (0-40 cm)      | 269                     | 3.7  | 160                       |
| 2             | 163         | A           | (40-60 cm)     | 269                     | 3.7  | 160                       |
| 3             | 163         | A           | (60-100 cm)    | 269                     | 3.7  | 160                       |
| 4             | 163         | A           | (100-140 cm)   | 269                     | 3.7  | 160                       |
| 5             | 163         | A           | (140-160 cm)   | 269                     | 3.7  | 160                       |
| 6             | 163         | B           | Bulk           | 269                     | 3.7  | 18                        |
| 7             | 163         | C           | Bulk           | 269                     | 3.7  | 20                        |
| 8             | 163         | D           | Bulk           | 269                     | 3.7  | 15                        |
| 9             | 163         | E           | Bulk           | 269                     | 3.7  | 13                        |
| 10            | 163         | F           | Bulk           | 269                     | 3.7  | 6                         |
| 11            | 163         | G           | Bulk           | 269                     | 3.7  | 37                        |
| 12            | 88          | A-G         | Bulk           | 466                     | 4.3  | 45                        |
| 13            | 66b         | A-G         | Bulk           | 371                     | 4.9  | 65                        |
| 14            | 142         | A-G         | Bulk           | 234                     | 5.8  | 60                        |
| 15            | 194         | A-G         | Bulk           | 280                     | 5.8  | 55                        |
| 16            | 114         | A-G         | Bulk           | 195                     | 6.3  | 105                       |
| 17            | 137         | A-G         | Bulk           | 19                      | 6.5  | 100                       |
| 18            | 191         | A-G         | Bulk           | 35                      | 6.5  | 46                        |
| 19            | 144         | A-G         | Bulk           | 47                      | 6.6  | 80                        |
| 20            | 145         | A-G         | Bulk           | 540                     | 11.9                                       | 25                        |

853 **Table 1:** Sampling of the P3 deposits for grain-size analysis.

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| <b>Parameters</b>     | <b>P3 (79 CE)</b>                             | <b>P2 (280 CE)</b>                       | <b>P1 (1300 CE)</b>                      |
|-----------------------|---|--|--|
| Total DRE volume      | <b>1.02 km<sup>3</sup></b>                    | 0.77 km <sup>3</sup>                     | 0.16 km <sup>3</sup>                     |
| Erupted mass          | <b>2.4 x 10<sup>12</sup> kg</b>               | 1.7 x 10 <sup>12</sup> kg                | 4 x 10 <sup>11</sup> kg                  |
| VEI                   | <b>5</b>                                      | 4  | 4  |
| Magnitude             | <b>5.4</b>                                    | 5.2                                      | 4.6                                      |
| Maximum column height | <b>30 km</b>                                  | 26 km                                    | 22 km                                    |
| Maximum PDC runout    | <b>7 – 10.3 km</b>                            | 7 – 11.5 km                              | 4.5 – 8 km                               |
| Maximum MER (fall)    | <b>1.4 x 10<sup>8</sup> kg s<sup>-1</sup></b> | 1.1 x 10 <sup>8</sup> kg s <sup>-1</sup> | 3.6 x 10 <sup>7</sup> kg s <sup>-1</sup> |
| Maximum MER (PDC)     | <b>2.5 x 10<sup>8</sup> kg s<sup>-1</sup></b> | 3.1 x 10 <sup>8</sup> kg s <sup>-1</sup> | 9 x 10 <sup>7</sup> kg s <sup>-1</sup>   |
| Maximum intensity     | <b>11.4</b>                                   | 11.5                                     | 11                                       |
| Duration              | <b>&gt; 11 h</b>                              | > 7 h                                    | > 5 h                                    |
| TGSD                  | <b>D = 3.3</b>                                | D = 3.4 - 3.5                            | D = 3.2 – 3.3                            |
| Minimum exit velocity | <b>210 - 220 m s<sup>-1</sup></b>             | 180 – 200 m s <sup>-1</sup>              | 150 – 165 m s <sup>-1</sup>              |
| Total gas content     | <b>2 - 2.9 wt.%</b>                           | 1.8 – 2.2 wt.%                           | 1.6 – 2.1 wt.%                           |

865 **Table 2:** Summary of the estimated eruptive parameters for the P3 eruption, and comparison  
866 with P1 and P2 ([Carazzo et al., 2012; 2019](#)).

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| <b>Phase</b> | $n_0$<br>(wt.%) | $\alpha_c$ | $x_0$<br>(wt.%) | $x_f$<br>(wt.%) | $U_{free}$<br>(m s <sup>-1</sup> ) | $U_{min}$<br>(m s <sup>-1</sup> ) |
|--------------|-----------------|------------|-----------------|-----------------|------------------------------------|-----------------------------------|
| Early A      | 5.8             | 29         | 4.1             | 2.9             | 230                                | 210                               |
| Late A       | 5.8             | 38         | 3.6             | 2.5             | 215                                | 220                               |
| B-C          | 5.8             | 42         | 3.4             | 2.3             | -                                  | 220                               |
| D-E          | 5.8             | 45         | 3.2             | 2.2             | -                                  | -                                 |
| F-G          | 5.8             | 49         | 2.9             | 2.1             | -                                  | -                                 |

875 **Table 3:** Gas contents and deduced exit velocities.  $n_0$  is the total volatile content of the melt  
876 ([Martel et al., 1998](#)),  $\alpha_c$  is the percentage of crystals in the melt plus lithics in the flow,  $x_0$  is  
877 the mass fraction of gas in the magma assuming complete degassing,  $x_f$  is the mass fraction of  
878 exsolved gas in the gas+pyroclasts mixture at fragmentation for a threshold vesicularity of  
879 70%,  $U_{free}$  is the supersonic velocity after decompression calculated using  $x_0$ ,  $D = 3.3$  and an  
880 open porosity of 60-70% ([Michaud-Dubuy et al., 2018](#)), and  $U_{min}$  is the eruptive velocity  
881 deduced from the distribution of the lithic fragments (section 5.2).

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893 **Figure captions**

894

895 **Fig. 1:** Current maps for the P3 eruption. **a** Isopach map (in centimeters) from [Westercamp](#)  
896 [and Traineau \(1983\)](#). (1) Dashed and (2) dotted lines correspond to isopachs for the pumice  
897 fallout deposit P3<sub>1</sub> and P3<sub>2</sub>, respectively. (3) The light grey and (4) dark areas give the  
898 distribution of the low-concentration PDC (surge), and high-concentration PDC (flow)  
899 deposits P3<sub>3</sub>, respectively. **b** Isopach map (in centimeters) from [Wright et al. \(2016\)](#). (1)  
900 Dashed lines correspond to isopachs for a pumice fallout deposit. (2) The light and (3) dark  
901 grey areas give the distribution of a low-concentration PDC (surge) deposit, and a low-  
902 concentration PDC (ash hurricane) deposit, respectively. (4) The dark areas correspond to  
903 high-concentration PDC (flow) deposits.

904

905 **Fig. 2:** Overview of our field area in Martinique (inset). White circles and numbers refer to  
906 localities where P3 deposits are present. Black dots show outcrops where P3 deposits are  
907 absent (due to erosion) and/or too deeply buried under recent eruption deposits. The dotted  
908 lines give the stratigraphic correlation reported in [Fig. 5](#).

909

910 **Fig. 3:** **a** Composite stratigraphic section showing typical unit thickness and grain-size  
911 characteristics of the P3 eruptive sequence. **b** Grain-size distribution selected samples  
912 representing the different eruptive units, including five stratigraphic heights in unit A, at  
913 location 163. **c** Variations of the median grain-size (blue circles), sorting (green circles),  
914 juvenile lithic content (grey squares), accidental lithic content (dark squares), total lithic  
915 content (white squares), and lithic size (red squares) along the stratigraphic height at location  
916 163.

917

918 **Fig. 4:** Representative photographs of outcrops of the P3 deposits in Martinique at sites **a** 57,  
919 **b** 78, **c** 13, **d** 65, and **e**, **f** 163. See [Fig. 2](#) for location and distance from the source. All scale  
920 bars are 20 cm long.

921

922 **Fig. 5:** Stratigraphic logs of representative sections of deposits of the P3 eruption. See [Fig. 2](#)  
923 for outcrop locations.

924

925 **Fig. 6:** Isopach maps (in centimeters) for **a** the cumulative pumice fallout (units A+C+E+G),  
926 **b** the pumice fallout unit A, **c** the cumulative PDC deposits (units B+D+F) and **d** the  
927 cumulative pumice fallout (units C+E+G). Open circles indicate measured sample locations.  
928 Bold numbers in panel (c) indicate the thicknesses of high-concentration PDC deposits.

929

930 **Fig. 7:** Isopleth maps (in millimeters) for the lithic fragments sampled **a** at the base of unit A,  
931 **b** at the top of unit A, and **c** the base of unit C. Open circles indicate measured sample  
932 locations. Direction of dispersal axes is consistent with those inferred from the isopach maps.

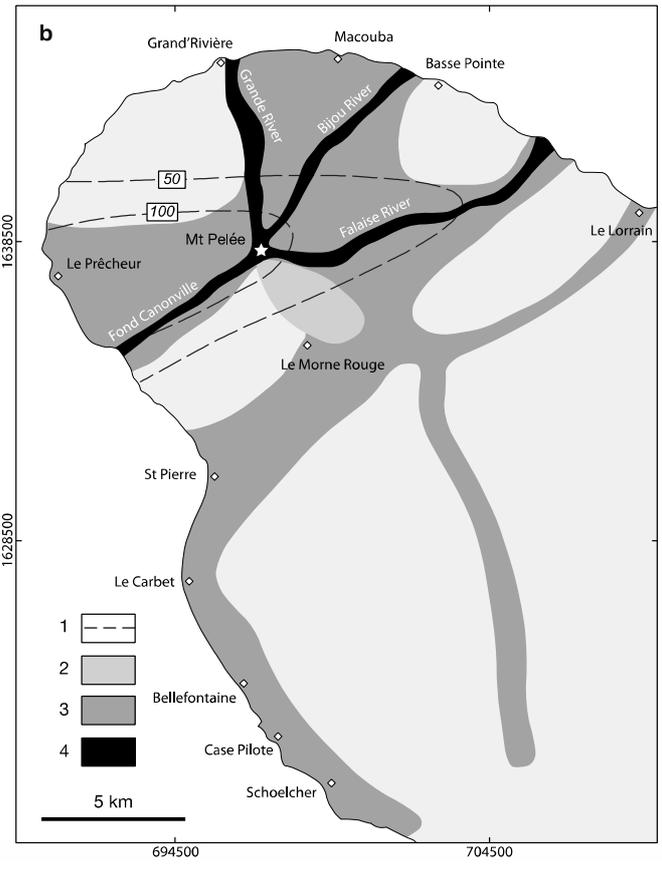
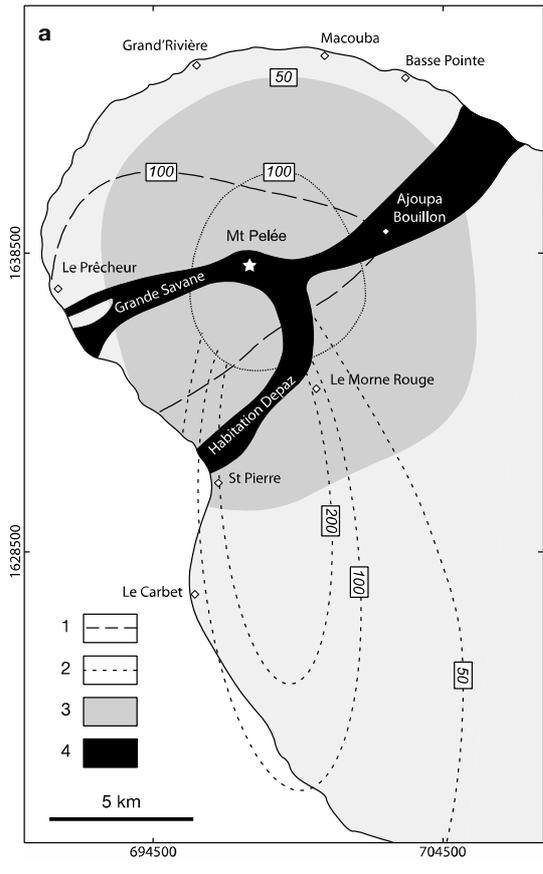
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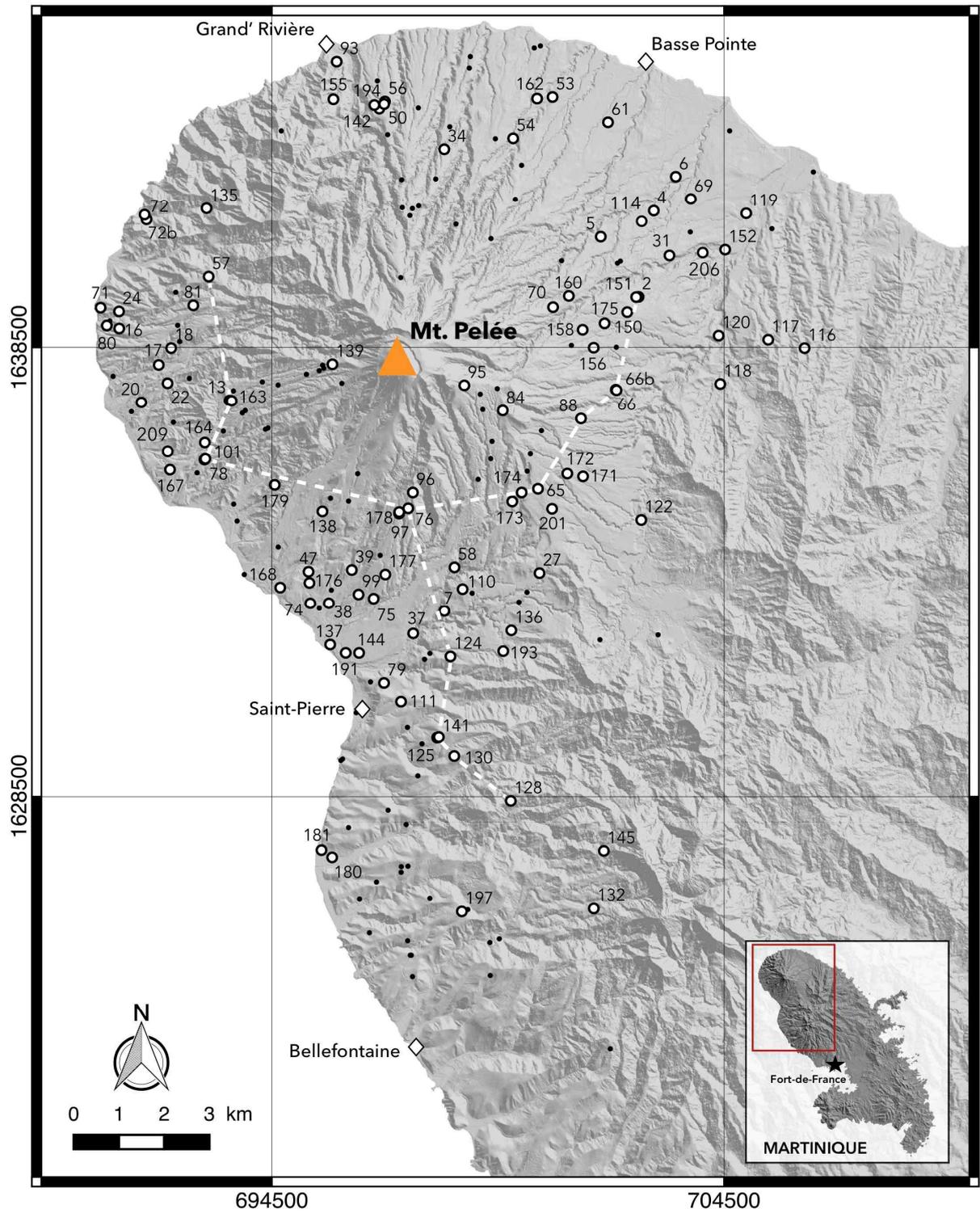
934 **Fig. 8:** Deposit thinning profiles generated from the isopach maps for **a** unit A, **b** the  
935 cumulative pumice fallout C+E+G, and **c** the cumulative pumice fallout A+C+E+G  
936 represented by semi-log plots of square root of isopach areas (in kilometers) versus thickness  
937 (in meters). Thinning trends are approximated by exponential (purple dashed line), power-law  
938 (blue dotted line), and Weibull (red solid line) fits.

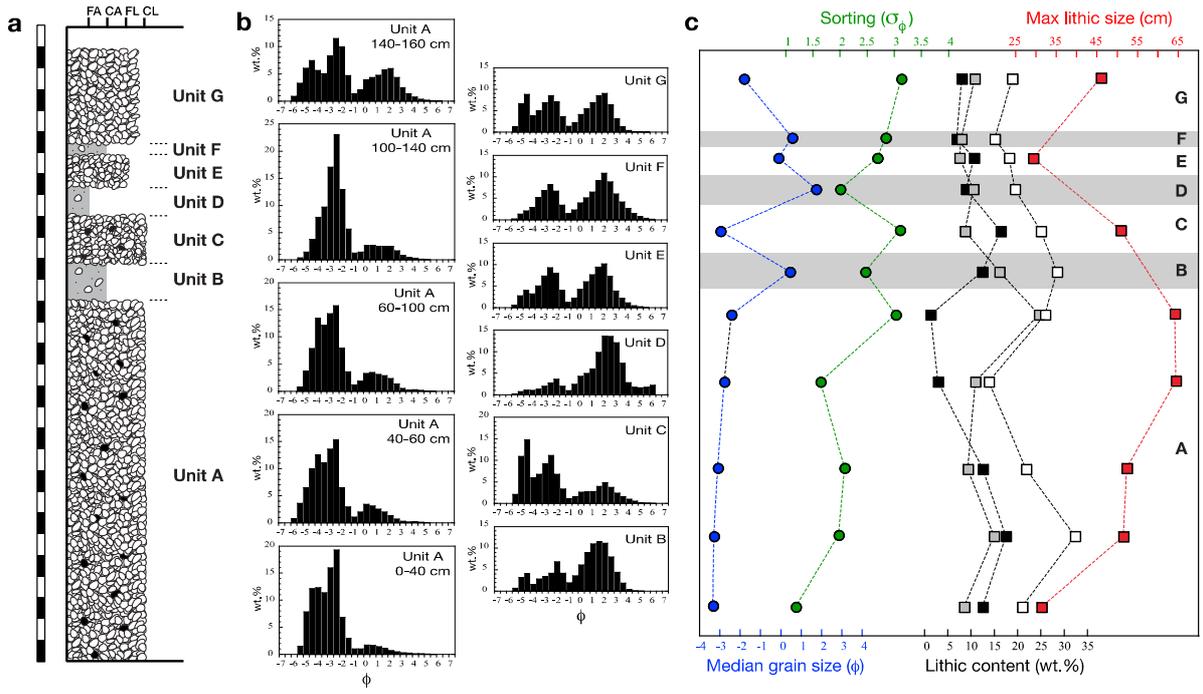
939

940 **Fig. 9:** Summary of the time evolution of the P3 eruptive parameters. **a** Cumulative erupted  
941 mass (in kg). **b** Maximum column height (in km). **c** Mass eruption rate (in  $\text{kg s}^{-1}$ ). **d** Total  
942 exsolved gas content,  $x_f$  (in wt.%).

943 **Fig. 10:** Transition diagram depicting the P3 eruptive events and depositional units  
944 (diamonds), and comparison with the P1 (squares), P2 (triangles), Taupo (circles) and  
945 Vesuvius (inverted triangles) eruptions. Numbers associated with diamonds correspond to  
946 different phases of the P3 eruption: 1: beginning of phase A, 2: end of phase A, 3:  
947 simultaneous phases B and C, 4: simultaneous phases D and E, and 5: simultaneous phases F  
948 and G. Dashed line corresponds to the maximum mass eruption rate feeding the column  
949 before collapse as a function of the total gas content at the column base (using [Michaud-](#)  
950 [Dubuy et al. \(2018\)](#) model for tropical conditions and  $D = 3.3$ ). Open, gray, and dark symbols  
951 correspond to geological data inferred for the stable column phase, partial column collapse  
952 phase, and total column collapse phase, respectively. Data for the P1, P2, Taupo and Vesuvius  
953 eruptions are compiled in [Carazzo et al. \(2012, 2019\)](#) and [Michaud-Dubuy et al. \(2018\)](#),  
954 respectively.









North West

South

South

North East

