Infrared spectroscopy quantification of functional carbon groups in kerogens and coals: A calibration procedure

van T.H. Phan, Eric Quirico, Pierre Beck, Yann Le Brech, Lora Jovanovic, Corentin Le Guillou, Sylvain Bernard, Lydie Bonal, Nathalie Carrasco, Thomas Gautier, et al.

HAL Id: insu-03208201

https://insu.hal.science/insu-03208201

Submitted on 12 Oct 2021

Infrared spectroscopy quantification of functional carbon groups in kerogens and coals: A calibration procedure

Van T.H. Phan[1], Eric Quirico[1], Pierre Beck[1], Yann Le Brech[2], Lora Jovanovic[3], Corentin Le Guillou[4], Sylvain Bernard[5], Lydie Bonal[1], Nathalie Carrasco[3], Thomas Gautier[3], Jesus Raya[6]

[1]Institut de Planétologie et d'Astrophysique de Grenoble (IPAG), Université Grenoble Alpes/CNRS-INSU, UMR 5274, Grenoble F-38041,

[2]LRGP, CNRS, Université de Lorraine, 1 rue Grandville 54000 Nancy, France

[3]CNRS/INSU, LATMOS-IPSL, Guyancourt F-78280, France

[4]Unité Matériaux et Transformations, UMET, UMR-CNRS 8207, Université de Lille, France

[5]Muséum National d'Histoire Naturelle, Sorbonne Université, CNRS UMR 7590, IRD, Institut de Minéralogie, de Physique des Matériaux et de Cosmochimie, Paris, France

[6]ICS, CNRS, Université de Strasbourg, BP 296 R8 670008 Strasbourg

Corresponding author:

Van Thi Hai Phan

Email: thi-hai-van.phan@univ-grenoble-alpes.fr

ABSTRACT

The determination of the abundances of the CH_x, $C=O$ and aromatic groups in chondritic Insoluble Organic Matter (IOM) and coals by Infrared (IR) spectroscopy is a challenging issue, due to insufficient knowledge on the absorption cross-sections and their sensitivity to the molecular environment. Here, we report on a calibration approach based on a ^{13}C synthetic model material, whose composition was unambiguously determined by Direct-Pulse/Magic Angle Spinning Nuclear Magnetic Resonance (DP/MAS NMR). Ratios of the cross-sections of the CH_x, $C=O$ and aromatic groups have been determined, and the method has been applied to IOM samples extracted from four chondrites as Orgueil (CI), Murchison (CM), Tagish Lake (C2-ungrouped) and EET 92042 (CR), and to a series of coals. The estimate of the aliphatic to aromatic carbon ratio (n_{CHx}/n_{Aro}) in IOM samples from Orgueil, Murchison and Tagish Lake chondrites is in good agreement with Single-Pulse/NMR estimates earlier published, and is lower by a factor 1.3 in the case of the CR chondrite EET 92042 (but error bars overlap). In contrast, the aliphatic to carbonyl ratio ($n_{CHx}/n_{C=O}$) is overestimated for the four chondrites. These discrepancies are likely due to the control of the absorption cross-section of the $C=O$ and $C=C$ bonds by the local molecular environment. Regarding coals, the use of published NMR analyses has brought to light that the integrated cross-section ratio A_{CHx}/A_{Aro} varies with the vitrinite reflectance over an order of magnitude. Here as well, the local oxygen speciation plays a critical control on A_{Aro}, which decreases with increasing the vitrinite reflectance. We provide an analytical law that links A_{CHx}/A_{Aro} and vitrinite reflectance, which will allow the determination of n_{CHx}/n_{Aro} for any coal sample, provided its vitrinite reflectance is known.

59

1. INTRODUCTION

Infra-Red (IR) spectroscopy is a non-destructive technique well suited to characterize complex insoluble organic compounds in meteorites and coals. It provides an estimation of organic composition based on the integrated absorbance of the main functional groups (e.g., alkyl CH, CH_2 and CH_3, aromatic C=C and C-H, carbonyl/carboxylic C=O, hydroxyl -OH) and tracers of the chemical and structural evolution during diagenesis and catagenesis (Li et al., 2016; Lis et al., 2005a; Painter et al., 1981; Petersen et al., 2008; Sun, 2005; Chen et al., 2012; Xueqiu et al., 2017). In addition, micro-FTIR, and more recently, Atomic force microscopy-based infrared spectroscopy (AFM-IR), has offered new perspectives to collect spatially resolved IR spectra, thereby identifying chemical heterogeneities at the submicrometric scale (Yang et al., 2017; Jakob et al., 2019; Abarghani et al., 2020; Jubb et al., 2020). These techniques have been proved valuable for the characterization of precious extraterrestrial micrometric grains, like Interplanetary Dust Particles (IDPs), Antarctic micro-meteorites and matrix grains extracted from rare primitive chondrites (Battandier et al., 2018; Beck et al., 2010; Dartois et al., 2004; Kebukawa et al., 2018; Mathurin et al., 2019; Orthous-Daunay et al., 2013; Quirico et al., 2018; Quirico et al., 2014). Although the assignment of the main vibration bands has reached a consensus, quantifying the exact proportions of functional groups based on IR signal remains difficult (Painter et al., 1981; delaRosa et al., 1992). Indeed, only a few absorption cross-sections have been published so far (Sandford et al., 1991; Dartois et al., 2004; Russell et al., 2009; Takahama et al., 2013). In addition, due to the complex structure of kerogens and coals, a given chemical group exists within different molecular environments, and is difficult to reside in various frequency shifts, inhomogeneous broadening and variability of the absorption cross-section values (Craddock et al., 2015).

Here, we report the determination of the absorption cross-sections of the alkyl (CH_x), carbonyl (C=O) and aromatic (C=C) groups based on a H^{13}CO synthetic kerogen analogue (named H^{13}CO-tholins), whose composition is unambiguously determined by DP/MAS (Direct Polarization/Magic Angle Spinning) Nuclear Magnetic Resonance and XANES (X-ray Absorption at Near-Edge Spectroscopy) at the C K-edge. The use of a ^{13}C sample (e.g. synthesized H^{13}CO-tholins) overcomes the critical bias of undetected carbons encountered in solid-state ^{13}C NMR analysis. These absorption cross-sections have been applied to a series of infrared spectra of kerogens extracted from primitive chondrites, and to a series of coals ranging a broad range of maturity. As the relative abundances of C=O, C=C and CH_x groups in these samples are reported in several publications, the calibration protocol could be tested, and we further discuss the dependency of the cross-sections with the chemical structure.

94

2. MATERIALS AND METHODS

2.1 Sample selection and preparation

2.1.1 Production of IOM analogues from HCO-tholins

98   HCO-tholins samples are polymeric disordered carbonaceous materials that were
99   synthesized in the PAMPRE experimental reactor located at LATMOS (Guyancourt, France)
100   (Alcouffe et al., 2010; Szopa et al., 2006). This setup consists in a radio-frequency capacitively
101   coupled plasma (RF CCP) confined within a cylindrical cage, filled with an initial gaseous
102   mixture made of 7% of CH_4, 3% of CO and 90% of Ar gas, with a 3 sccm flux under the total
103   pressure of 0.43 mbar. Two types of $H^{12}CO$ and $H^{13}CO$ tholins were produced from natural ^{12}C-
104   rich and ^{13}C- rich gases (purity > 99.5%), provided by Eurisotop (Saclay, France). The power
105   delivered by the generator was maintained within 11–12 W. Grains were formed within the
106   plasma and kept under electrostatic levitation, and grew up until they deposited into a glass
107   vessel surrounding the confining cage of the plasma. Experiments were run daily (8-10 h) during
108   working hours for safety. The temperature of the electrode was 21ºC when starting the
109   discharge, and 45ºC when switching it off in the evening. At total, about 30 mg of $H^{13}CO$-tholins
110   were obtained in 63 hours, and 130 mg of $H^{12}CO$-tholins in 98 hours. The production yield, 0.50
111   to 1.3 $mg.hour^{-1}$, was much lower compared to tholins formed from N_2 : CH_4 mixtures
112   (Sciamma-O'Brien et al., 2010). The $H^{13}CO$-tholin was composed of a dominant yellow, fine-
113   grained powder and a less abundant population of dark platelets (Fig. S-1). In comparison, the
114     $H^{12}CO$-tholins did not contain any dark grains. The black platelets were constituted by direct
115   deposition on the grid, which eventually fell and mixed with the dusty tholins. Infrared
116   spectroscopy has shown that these two populations were chemically similar (Fig. S-2). Finally,
117   the differences in the color of the grains are probably due to the larger thickness of the platelets.

118   Analogues of kerogens and coals were obtained from the carbonization of $H^{13}CO$-tholins
119   with the elemental composition revealed in Table S-1 (Supporting Information). To select
120   optimal conditions to get the good analogues, preliminary heating experiments were run at 200
121   °C, 300 °C, 350 °C, 400 °C, 450 °C, 500 °C and 550 °C, with a tubular furnace maintained
122   under secondary vacuum ($\sim 10^{-6}$ mbar). Small grains of samples were selected and pressed
123   between two diamond windows and put into a ceramic crucible that was placed at the middle of
124   the tube. For each experiment, the setting temperature was maintained during 1h, and the tube
125   was then let cool down freely for 3–6 hours, depending on the initial temperature setting.
126   Infrared spectra of the heated samples were collected and compared to spectra of Insoluble
127   Organic Matter (IOM) samples extracted from primitive chondrites (Fig. S-2). Finally, we select
128   a temperature of 400°C (for extensive details on the procedure, see Supporting Information).
129   The final carbonization of HCO-tholins was then operated with this set temperature, using a
130   ceramic crucible filled with 5–10 mg of the powdered samples. We used an anti-static gun
131   (Zerostat, Sigma-Aldrich) to limit sample loss due to static electricity.

132

2.1.2   Coal and IOM samples

134   A suite of coal samples with a mean maximum vitrinite reflectance (R_o) ranging 0.33–
135   2.8%   were   purchased   from   the   Penn   State   University   coal   sample   Bank
136   (http://www.energy.psu.edu) (Table S-2). The IR spectra of IOMs extracted from primitive
137   chondrites have been earlier published in Orthous-Daunay et al. (2013), including IOMs from
138   Orgueil (CI1), Murchison (CM2), EET92042 (CR2) and Tagish Lake (C2-ung), while the IR
139   spectra of coals were collected by using Bruker Vertex 70v spectrometer.

140

2.2 Analytical methods

2.2.1 FTIR measurements

The FTIR instrument used was a Bruker Hyperion 3000 infrared microscope, equipped with liquid nitrogen cooled–mercury cadmium telluride (MCT) detector. *Micro*-FTIR spectra were collected in the range of $4000 - 700$ cm^{-1} with a spectral resolution of 4 cm^{-1}. No samples required more than 200 scans (5 min) to get a satisfying signal-to-noise ratio. All measurements were done using an environmental cell with an optical path through two ZnS windows and pumped to $\sim 10^{-6}$ mbar (Beck et al., 2010). The sample (e.g., tholins, coal) was deposited onto a type IIa synthetic diamond window (3 mm diameter, 500 µm thick) and crushed with a similar window to have a thin and flat sample. The diamond window holding the crushed sample was inserted in the cell, pumped, and the window was heated at 80 °C to remove efficiently adsorbed water. Background spectra were collected on a location next to the sample just before measuring the sample.

FTIR spectra of bulk coals and HCO-tholins samples were collected on KBr pellets. Pellets were prepared using 300 mg of a KBr powder previously heated at 300 ºC under vacuum during 3 h. The sample (1 ± 0.1 mg) was ground and mixed with 300 mg KBr with a MM200 Retsch grinder equipped with an agate mortar filled with a single ball (30 Hz; 30 minutes). Finally, 13 mm diameter pellets were prepared using a press Atlas^{TM} Manuelles (Eurolabo), under 400 bars and the gentle heating at 60 ºC during 1h (cooling down during 30 minutes). Infrared spectra were collected under vacuum condition with a Bruker Vertex 70v spectrometer, in the range $4000 - 400$ cm^{-1} with 200 scans and a 4 cm^{-1} spectral resolution.

162

2.2.2 Solid-state Nuclear Magnetic Resonance

The H^{13}CO- tholins material was characterized by solid NMR using an AVANCE 750 MHz wide-bore spectrometer (Bruker^{TM}) operating at MAS rate of 60 kHz designed for 1.3 mm outside diameter (OD) zirconia rotors (closed with vespel caps). About $1 - 5$ mg of materials was put inside the rotor. Under those conditions (188.62 MHz and 60kHz spinning speed), any rotational sideband was thrown out of the spectral region of interest. This allows a proper deconvolution of spectra. Proton decoupling during acquisition was obtained by using SPINAL-64 at a 200 kHz ^{1}H RF field (Fung et al., 2000). All spectra were collected within a 250kHz spectral windows (1325.3ppm) and 16324 acquisition time domains. To get undistorted lineshapes and filter out background probes signal, the quantitative Direct Polarization (DP) was performed using a speed synchronized spin-echo (Hahn, 1950; Vieira et al., 2016; Lucherelli et al., 2019), pulse sequence P_{90}-τ-P_{180}- τ, with the following conditions: 2.75 µs and 5.5º s for $P_{90°}$ and $P_{180°}$ pulses, and echo time being set to 33.6 µs (2 rotation periods). Under such conditions, baseline problems and first order phasing are fixed, which is a prerequisite for quantitative analysis. T1 (spin-lattice relaxation time) were measured following inversion-recovery method (Kingsley, 1999) and led us to set up DP experiments with recycling delays equal to 30s. NMR spectra were processed with the Topspin software (phase and baseline, LB=25 applied). The interpretation of 1D ^{13}C NMR spectra presents a significant difficulty due to signal overlap.

5

181 Therefore, the spectral decomposition of the spectra was performed with the DMFIT Software
182 developed by Massiot et al. (2002) and previously used for biomass and biochar using Gaussian
183 profiles (Le Brech et al., 2016). The decomposition was repeated three times for the 400 ºC
184 H^{13}CO- tholins to estimate the uncertainty on the derived composition due to the processing.

185

186 ### 2.2.3   STXM-XANES analysis

187   Synchrotron-based scanning transmission X-ray microscope (STXM) analyses were
188 performed on heated H^{12}CO- and H^{13}CO-tholins to document the carbon speciation using the
189 HERMES STXM beamline at the synchrotron SOLEIL (Saint-Aubin, France) (Belkhou et al.,
190 2015; Swaraj et al., 2017). Samples were prepared either by cryo-ultramicrotomy (H^{13}CO-
191 tholins) or by depositing the powder on a Si$_3$N$_4$ membrane (H^{12}CO-tholin). Energy calibration
192 was done using the well-resolved 3p Rydberg peak of gaseous CO$_2$ at 294.96 eV (Swaraj et al.,
193 2017). XANES hypercube data (stacks) over the carbon absorption range (250 – 350 eV) were
194 collected with a spatial resolution of 35 nm at energy increments of 0.1 eV over the 283 – 291.5
195 eV region, with a dwell time of less than 1 ms per pixel to prevent irradiation damage (Wang et
196 al., 2009). Stack alignments and extraction of XANES spectra were done using the aXis2000
197 software (ver2.1n). Normalization of data and quantification of the concentrations of functional
198 groups   were   done   using   the   QUANTORXS   freeware
199 (https://github.com/CorentinLG/QuantORXS/) described in Le Guillou et al. (2018). Note that
200 although this quantification procedure provides consistent results for aromatic, ketone\phenol
201 and carboxylic\ester groups, it remains difficult to estimate the concentration of aliphatic groups
202 precisely.

203

204 ### 2.2.4   Raman setup

205   Raman measurements were collected with a HORIBA JOBIN-YVON LaRam
206 spectrometer, equipped with a 600 gr.mm^{-1} groove and operating at 532 nm (INSU national
207 Facility at ENS-Lyon, France). The laser beam was focused through a ×100 objective. The
208 power onto the sample was 380 µW, and acquisition time was 5×20s. The configuration used
209 yields a lateral resolution of ≈ 1 µm.

210

211 ### 2.2.5   Analysis of infrared spectra

212   For each spectrum, a baseline correction was done using spline functions, and the peak
213 absorbance of the C=C band at 1540 cm^{-1} was set to 1. The C=O and C=C bands in H^{13}CO-
214 tholin spectra at 1656 and 1540 cm^{-1}, respectively, were fitted with Gaussian profiles (using the
215 multi-fit peak functions of the Igor 8 software, Wavemetrics) (Fig.1).

216   The integrated cross-section is defined by:

217       $$A = \int_{v1}^{v2} \sigma dv = \int_{v1}^{v2} ln(10) \frac{a(v) \times S}{n} dv \text{ (cm/group) (Equation 1)}$$

218   where $a(v) = -\log(\frac{I}{I_0})$ is the absorbance, with I and I_0 are the intensity of the infrared beam

219   after and before the sample, respectively; n is the number of molecular oscillators and σ is the

220   absorption cross-section (cm^2), with v the wavenumber, v_1 and v_2 the edges of the spectral range

221   of the spectral component, and S is the probed area (cm^2).

222   Finally, the ratio of the absorption cross-section of two chemical groups X_1 and X_2 is:

223
$$\frac{A_{X1}}{A_{X2}} = \frac{a_{X1}}{a_{X2}} \times \frac{n_{X2}}{n_{X1}} \qquad \text{(Equation 2)}$$

224   where a_X refers to the integrated absorbance of a chemical group X as aliphatic CH_x, C=O

225   or aromatic species (Aro). a_{CHx} was calculated from the integration of the alkyl band between

226   2750 and 3000 cm^{-1}, while a_{CO} and a_{Aro} were obtained from a Gaussian fit of the bands at 1656

227   and 1540 cm^{-1}, respectively. The integrated cross-section ratios of the CH_x, C=O and Aro

228   molecular oscillators, A_{CH_X}/A_{Aro}, $A_{CH_X}/A_{C=O}$ or $A_{Aro}/A_{C=O}$ were finally calculated

229   with Equation 2, using estimates from NMR measurements on the 400 °C H[13]CO-tholin. No

230   estimation of individual A_X cross-sections was made, because NMR measurements provided

231   ratios of chemical groups abundances.

232   Strictly speaking, IR spectroscopy probes normal modes and not bonds. In the case of the

233   CH_x groups, the mass difference between H and C atoms is large and decouples their respective

234   motions. Thus, the corresponding bands are controlled by hydrogen vibration around a single

235   carbon atom. For the terminating carbonyl group, environmental effects from neighboring

236   chemical groups are present, leading to a moderate frequency shift and variations of the

237   absorption coefficient by up to a factor of 2. But for both CHx and C=O, the number of bonds

238   or molecular groups is equal to the number of carbon atoms and a straightforward comparison

239   with ^{13}C-NMR is possible. Things get harder in the case of aromatic species. A benzene single

240   ring (C_6H_6), in the gas or as a molecular crystal has a vibration mode at ~1500 cm^{-1} that involves

241   all carbon atoms (6). Small polycyclic hydrocarbons (1-4 rings) display a sharp narrow band in

242   the range 1700-1500 cm^{-1}, but the number of C···C bonds does not equal the number of C atoms.

243   For instance, naphthalene units have 10 C atoms and 11 C···C bonds, anthracene and

244   phenanthrene have 14 C atoms and 16 C···C bonds, and pyrene 16 C atoms and 19 C···C bonds.

245   In the case of larger polycyclic aromatics, theoretical calculations point to a broad feature

246   peaking around 1600 cm^{-1}, which is due to the E_{2g} normal mode that involves all carbon atoms

247   and phonon propagation (Mapelli et al., 1999). Immature kerogens and coals contain small (1-

248   4 rings) polycyclic aromatics inserted in a macromolecular network and display a similar broad

249   band at 1600 cm^{-1}. Altogether, the infrared absorption at 1600 cm^{-1} results from the motion of

250   all carbon atoms, but in contrast no clear connection can be established with a number of C=C

251   bonds, nor a number of rings. In this respect, the terms "aromatic ring stretch" or "C=C" used

252   in literature are somewhat misleading, and we have used the term "Aro". As such, A_{Aro} is

253   controlled by the number of carbon atoms, allowing a direct comparison with ^{13}C-NMR. In the

254   details, we expect significant variations of A_{Aro} with the size and local environment of the

255   polyaromatic specie, for instance –OH groups (see discussion below).

256   Last, the Artificial Bees algorithm was used from the ESA Package PYGMO

257   (https://esa.github.io/pygmo/) and tested two fit models: one based on 7 Gaussian and one on 7

258   pseudo-Voigt components. Each group comprised 50 simulations and the fit quality was
259   estimated with a χ^2 function. A 20000 bees' population over 50 generations was used. For each
260   of the 200 numerical simulations to get the good fit in term of χ^2 value.
261

3   RESULTS

3.1   Solid-state ^{13}C DP/MAS NMR analysis

264   Solid-state ^{13}C DP/MAS NMR spectra of the pristine and heated H^{13}CO-tholins are
265   presented in Fig. 2a and the band assignment is given in Table 1, following the fitting results in
266   Fig. 2b and 1c for the initial and 400 ℃ H^{13}CO tholins, respectively. We observe two intense
267   features at 100–160 ppm and 0 – 100 ppm in the spectra of both samples, which are assigned to
268   sp^2 and sp^{3}carbon atoms, respectively. The range 0–45 ppm usually points to C atoms in alkyl
269   chains, while the range of 45-90 ppm points to oxygenated aliphatic carbons, e.g. alcoholic,
270   carboxylic or ether groups (Pretsch et al., 2000). It is, however, challenging to separate the
271   respective contribution of these two groups of carbon atoms accurately. sp^2 carbons in the 100-
272   160 ppm range can be assigned to olefinic chains (conjugated and unconjugated alkenes) and
273   aromatic or polyaromatic species. We cannot distinguish these two types of speciation. The
274   broad and weak signal which peaks at 170 ppm points to C=O groups in esters or carboxylic
275   groups, and that at 210 ppm indicates to C=O in ketones or aldehydes. The effect of heating
276   results in a steady increase of unsaturated versus saturated carbons, and a dramatic drop of labile
277   oxygenated groups (Fig. 2a). This is in full agreement with the carbonization process that
278   promotes hydrogen loss, carbonyl loss and aromatization.

279   The relative abundances F$_X$, of the different carbon groups, was derived through a
280   decomposition procedure using a fit model based on Gaussian components (Fig. b-c; Table 1)
281   (Massiot et al., 2002, Le Brech et al., 2016). The unheated H^{13}CO- tholins appears to be mainly
282   composed of -CH$_x$ (x=1 to 3) groups (F$_{ali}$= ~ 61 %), olefinic carbons (F$_{aro}$ = ~ 32.5 %) and a
283   lower content of C=O groups (F$_{CO}$ = 5.1 %). The ^{13}C NMR spectrum of the 400 °C heated
284   H^{13}CO-tholins is dominated by aromatic carbons (F$_{aro}$ = ~ 64%) and contains less aliphatic–CH
285   groups (F$_{ali}$ = ~ 32%) and C=O groups (F$_{CO}$ = 4.1 %) (Table 1).
286

3.2   Raman Analysis

288   The Raman spectrum of the unheated H^{13}CO-tholin is dominated by a strong
289   fluorescence background, with no detectable vibrational bands. The spectra of the heated
290   H^{13}CO-tholin display the first-order carbon bands superimposed onto a high fluorescence
291   background. The spectral contrast (i.e. the intensity of the G and D bands with respect to the
292   fluorescence intensity) is low, leading to poor quality of the reduced spectra after subtraction of
293   a linear baseline (Fig. S-3). The use of a Lorentzian-Breit-Wigner-Fano fit model was
294   complicated, due to the poor spectral quality at the edge of the G and D bands. Despite these
295   troubles, these data point a G band peak position at ~ 1537 cm^{-1} and a FWHM of ~ 80 cm^{-1}. For
296   comparison with published data, the peak position of a ^{12}C-bearing sample can be estimated
297   through $\omega_{12}=\omega_{13}(\mu_{12}/\mu_{13})$, with ω the peak position and μ the reduced mass. We obtain ω_{12}=1593
298   cm^{-1}, which points to a nano-graphitic structure in terms of sp^2 structure and is entirely consistent

299 with a kerogen structure as found in chondritic IOM, coals and terrestrial kerogens (Ferrari &
300 Robertson, 2001; Quirico et al., 2018). These data support that most of the unsaturated carbon
301 atoms are present in small-sized polyaromatic units, with negligible olefinic carbon bonds.

302

3.3 XANES analysis

304 STXM-based XANES data revealed that the $H^{13}CO$-tholins heated at 400 °C is
305 homogeneous at the submicrometric level. The prominent feature at 285 – 285.5 eV of the
306 XANES spectra indicate that more than half of the carbon atoms (56%) are involved in
307 aromatic/olefinic groups (C=C groups) (Fig. 3a, Table 2). A shoulder is observed at 287.7 eV,
308 which can be assigned to the aliphatic carbons ($-CH_x$) ($\sim$ 30% of the C atoms). Additional broad
309 features at 286.2 eV and 288.5 eV are assigned to ketone/phenols (C=O and C_{Ar}-OH, $\sim$ 14 % of
310 the C atoms) and carboxylic/ester (R-COOH,R'; $\sim$ 4 % of the C atoms). These estimates are
311 consistent with NMR results. The absorption around 286.5 eV could also be assigned to the
312 nitrile functional group, but its presence is discarded by the low elemental N/C ratio (0.01)
313 (Table S-1). Even if our measurements do not report such heterogeneity in $H^{13}CO$-tholins, we
314 cannot exclude the presence of such tiny inclusions.

315 To validate their XANES quantification method, Le Guillou et al. (2018) applied both
316 XANES and DP/MAS NMR techniques to fresh lignin and lignin heated at 300 °C and 350 °C
317 given by Le Brech et al. (2016). This work suggested that XANES and NMR quantification
318 methods provide consistent results for aromatic groups, ketone\phenol groups and
319 carboxylic\ester groups. However, the aliphatic groups of lignin samples appeared
320 overestimated by about 15%. In the case of the heated $H^{13}CO$-tholins, there is a reasonable
321 agreement between XANES and NMR for the aliphatic and the carboxylic groups (Fig. 3b). In
322 contrast, the aromatic groups' abundances determined from XANES are lower than that obtained
323 from DP/MAS NMR, as 55 ± 4 % and 64 ± 1 %, respectively.

324

3.4 Infrared spectroscopy

326 The infrared spectra of the initial and 400 °C $H^{12}CO$ and $H^{13}CO$-tholins are presented in
327 Fig. 4. The heterogeneity of the 400 °C $H^{13}CO$-tholins was checked in both the KBr pellet and
328 Micro-FTIR spectra (Supporting Information, Fig. S-4). The broadband at 3600 – 3100 cm^{-1}
329 corresponds to the stretching mode of – OH groups in alcohols, carboxylic groups or phenol
330 (C_{aro}-OH). The peak at 3100 – 3000 cm^{-1} is assigned to aromatic C-H. The region of 2800–3000
331 cm^{-1} displays four main peaks, including the antisymmetric and symmetric stretching modes of
332 methyl (CH_3) and methylene (CH_2) groups and possibly of the CH group, as a component
333 centered at $\sim$ 2900 cm^{-1}. However, this spectral range is highly congested and more components
334 are present (see below). The spectral region 1800 – 1500 cm^{-1} displays bands at 1656 and 1540
335 cm^{-1}, and 1708 cm^{-1} and 1608 cm^{-1} for the $H^{13}CO$ and $H^{12}CO$-tholins, respectively. These
336 features (e.g., 1656 and 1540 cm^{-1}) are assigned to the stretching modes of the carbonyl (C=O)
337 and aromatic (C=C), respectively (Painter et al., 1981). In the region 1500 – 900 cm^{-1}, bending,
338 deformation and out-of-plane modes of CH_2, CH_3 and C-O bonds are present at 1450, 1375 and
339 1300 cm^{-1}. We will follow here Painter et al. (1981) which state that the broad feature peaking

340 at ~ 1200 cm^{-1} does not provide any significant compositional information, in particular on the
341 abundance of ether groups (-C-O-C-).

342

3.4.1 *Spectral decomposition of the CH$_x$ stretching region*

344   The region between 3000 – 2750 cm^{-1} results from the contribution of symmetric and
345 antisymmetric stretching modes of CH$_2$, CH$_3$, possibly of -CH, and contains information on the
346 abundance of alkyl groups and the ratio of CH$_2$ and CH$_3$ groups. However, the spectral
347 congestion makes it difficult to identify the exact number of components, spectral profiles and
348 physical interpretation. Pure alkanes at a low temperature usually display several components
349 ascribed to the d$^-$ antisymmetric stretch of chain CH$_2$, the symmetric stretch of terminal CH$_2$
350 (d$^+\omega$) and two components arising from a Fermi resonance between the symmetric stretch of
351 chain CH$_2$ and the overtone of the H-C-H bending mode, reported as d$^+$ and d$^+_{FR}$ (the first being
352 dominating by the fundamental mode) (Fig. 5) (MacPhail et al., 1984); other components are
353 the r^-_a and r^-_b in-plane and out-of-plane asymmetric stretch of –CH$_3$, and the symmetric stretch
354 in Fermi resonance with the first overtone of the CH$_3$ bending mode, named r$^+$ and r$^+_{FR}$. Pure
355 alkanes measured at low-temperature display peaks and shoulders in agreement with this
356 assignment. However, the different components cannot be guessed for a similar compound in
357 the liquid state at room temperature (Fig. 5). Even more, critical issues arise in the case of natural
358 complex mixtures or solids, such as broader components, spectral shapes changes and peak
359 position shifts. In the case of coals and chondritic IOM, the presence of oxygen as ether or
360 carbonyl groups likely modifies the spectral distribution, making uncertain the spectral
361 decomposition of the bands (Orthous-Daunay et al., 2013; Quirico et al., 2018).

362   To explore more quantitatively this congestion issue, we ran optimization simulations of
363 the fit of the alkyl band of an IOM sample extracted from the Orgueil (CI1) chondrite (Orthous-
364 Daunay et al., 2013). We used the Artificial Bees algorithm from the ESA Package PYGMO
365 (https://esa.github.io/pygmo/) and tested two fit models: one based on 7 Gaussian and one on
366 seven pseudo-Voigt components. Two components aimed to fit the small features at ~ 3000 and
367 2790 cm^{-1}, due to the aromatic C-H stretch and a combination mode from fundamental modes
368 in the fingerprint region (Fig. 6). The fit model of the aliphatic bands then consisted of five
369 components, as usually reported in the literature (Dartois et al., 2007, 2004; Orthous-Daunay et
370 al., 2013). We ran two groups of simulations: one keeping constant the Full Width at Half
371 Maximum FWHM of 35 cm^{-1} for all the components, and another one where FWHM is let free
372 within the range 20-40 cm^{-1}. Each group comprised 50 simulations, and the fit quality was
373 estimated with a χ^2 function. A 20000 bees' population over 50 generations were used. For each
374 of the 200 numerical simulations, the resulting fit was excellent in terms of χ^2 value (Fig. 6a).
375 The results are summarized in Fig. 6b-c. Two parameters are of particular interest: (1) the ratio
376 of the integrated absorbance of the antisymmetric stretching components of CH$_2$ and CH$_3$
377 (a$_{CH2}$/a$_{CH3}$) and (2) the sum of these values. The first estimates the ratio of the number of CH$_2$
378 over CH$_3$ groups, while the second is a measure of the alkyl content (a$_{CH2}$+a$_{CH3}$). Note, however,
379 that for this latter the abundance of the CH group is not measured regarding the assignment of
380 the central component peaking at ~2900 cm^{-1} to the overtone of a bending mode in Fermi
381 resonance with a fundamental C-H stretch (Dartois et al., 2007; 2004). In the case of FWHM

382 forced to 35 cm^{-1}, we obtain 1.36 ± 0.07 and 1.52 ± 0.28 for a_{CH2}/a_{CH3} and 57.0 ± 1.8 and 59.0
383 ± 2.7 for $a_{CH2}+a_{CH3}$, for the Gaussian and pseudo-Voigt models, respectively (Fig. 6b). If FWHM
384 is let free within 20–40 cm^{-1}, we obtain 1.37 ± 0.79 and 1.66 ± 0.90 for a_{CH2}/a_{CH3} and 50.4 ± 6.3
385 and 52.8 ± 7 for $a_{CH2}+a_{CH3}$, for the Gaussian and pseudo-Voigt models, respectively. The
386 dispersion is much larger, and the absolute values of the sum are also significantly different
387 from those obtained by forcing FWHM. To make things clear on the impact in terms of chemical
388 information, we can estimate the n_{CH2}/n_{CH3} of Orgueil IOM, using the A_{CH2} and A_{CH3} values
389 from Dartois et al. (2007). We find that n_{CH2}/n_{CH3} ranges between 0.5 and 4.5, while NMR
390 measurements provide a value ranging 1.9-2.6 (Gardinier et al. 2000; assuming that ether
391 bridges are CH$_3$-O-CH$_2$:CH$_2$-O-CH$_2$=1:1 or CH$_2$-O-CH$_2$, respectively. This result is basically
392 of little value. In practice, things may not be as worse, provided that the fit model is similar to
393 that used to derive the absorption cross-sections. Finally, due to the sensitivity of the fit models
394 on derived parameters, and to propose an easy procedure allowing cross-comparison of data
395 collected by independent groups, we prefer to use the integrated absorbance of the whole alkyl
396 band, in place of $a_{CH2}+a_{CH3}$. The calibration presented below will lose accuracy when the
397 n_{CH2}/n_{CH3} ratio departs from that of the alkane's standards, but as discussed below, the impact
398 is very limited.

399

400 ### 3.4.2 *Determining the $_{CHx}/A_{C=C}$ and $A_{CHx}/A_{C=O}$ ratios*

401 The ratio of the absorption cross-section of two chemical groups X_1 and X_2 is calculated
402 from the Equation 1. The term n_{X2}/n_{X1} is provided by the NMR analysis, and the integrated
403 absorbances of the alkyl, carbonyl and aromatic groups are measured in the spectra of the 400
404 °C H^{13}CO-tholins. The uncertainty values were estimated by considering fit uncertainties, the
405 baseline correction, sample heterogeneity (Fig. S-4) and the uncertainty on NMR values of the
406 relative abundance of the three chemical groups. We finally obtain the ratio values of integrated
407 cross-section: A_{CHx}/A_{Aro} = 1.578 ± 0.183 and $A_{CHx}/A_{C=O}$ = 0.116 ± 0.006 (Table 3).

408

409 ## 4 DISCUSSION

410 ### 4.1 Achievements and limitations

411 The composition of the H^{13}CO-tholin determined by NMR is very reliable for the
412 aliphatic and aromatic contents, due to intense DP/MAS features at 38 and 130 ppm that are
413 unambiguously assigned to alkyl and conjugated carbons. Unlike CP/MAS data, these spectra
414 do not suffer from selective signal enhancement and are suitable for quantification. The
415 assignment is also robust for oxygen-bearing groups, but the features at 169 and 200 ppm are
416 weak and suffer from a higher uncertainty in their relative abundances. The new calibration of
417 H^{13}CO-tholin allows determining the absorption cross-section ratios of the CH$_x$, C=O and Aro
418 groups.

419 Nevertheless, the calibration derived from the H^{13}CO-tholins suffers from limitations
420 and biases that are intrinsic to infrared spectroscopy. First, as we estimate the alkyl abundance
421 through the integration of all the alkyl bands (CH$_x$ stretching in the range of 3000 – 2800 cm^{-1}),

422   the validity of the $H^{13}CO$-tholins as a calibration standard relies on a similar distribution of the
423   alkyl groups in the kerogen and coal samples that are analyzed. Hence, the absorption cross-
424   section of CH_2 and CH_3 are not identical, e.g. for the anti-symmetric stretching modes $A_{CH3} =$
425    $(11.8 \pm 0.76) \times 10^{-18}$ and $A_{CH2} = (7.7 \pm 1.5) \times 10^{-18}$ cm/group (Dartois et al., 2007; Sandford et
426   al., 1991). Considering a variation range of 1.5 – 3.0 for n_{CH2}/n_{CH3} that is observed in chondritic
427   IOM (Orthous-Daunay et al., 2013), we can calculate the variation of the sum of the integrated
428   absorbance of the anti-symmetric stretching modes of CH_2 and CH_3, $a=a_{CH2}+a_{CH3}$, as:

429
$$\frac{a_{1.5}}{a_3} = \frac{1 + \frac{1}{1.5}\frac{A_{CH3}}{A_{CH2}}}{1 + \frac{1}{3}\frac{A_{CH3}}{A_{CH2}}} = 1.18$$

430   This means that n_{CH2}/n_{CH3} estimates on IOM from thermally processed chondrites
431   obtained with our calibration are underestimated by 20% at maximum. Second, a critical issue
432   with the 1600 cm^{-1} band is the sensitivity of its absorption cross-section to the oxygen molecular
433   environment, in particular the phenolic (-OH) and possibly the ether (C-O-C) groups (Painter et
434   al., 1981). This means that the kerogen to be analyzed should be as similar as possible to the
435   calibration standard in terms of oxygen speciation around aromatic carbons. In this respect, the
436   400°C $H^{13}CO$-tholins do not fit mature samples that have large polyaromatic units, and which
437   are known to have a lower aromatic cross-section. This issue is probably more critical for coals,
438   due to the large O/C variations upon maturity.

439

440   **4.2   IOM composition of primitive chondrites**

441   A way to test our calibration is to apply our method to samples whose composition has
442   been previously determined by NMR. We report here the composition ($n_{CHx}/n_{C=O}$ and $n_{CHx}/n_{C=C}$)
443   of the IOMs from the Orgueil, Murchison, EET 92042 and Tagish Lake chondrites, derived
444   from the spectra published by Orthous-Daunay et al. (2013) (Fig. S-5). These compositions are
445   compared with the NMR estimates of Cody & Alexander, (2005) and Gardinier et al. (2000)
446   (Fig. 7). It is first necessary to discuss those NMR data, which display some inconsistencies
447   between them.

448   Two types of NMR measurements are presented in these publications: Single-pulse (SP)
449   MAS and Cross Polarized (CP) MAS. Single-pulse (SP) MAS NMR is usually presented as
450   suitable for quantification (Baccile et al., 2014), but in the case of natural samples ($^{13}C/^{12}C \sim$
451   1%), the signal is low, long acquisition time is required, and eventually the spectra still contain
452   significant noise. In the case of broad features, band distortion can happen due to dephasing
453   processes before spectrum acquisition, leading to significant errors in the quantification of band
454   integral. In addition, the baseline may depart from an ideal flat line, thereby perturbing the
455   profile of broad features. Altogether, these artifacts may lead to significant errors. These issues
456   can be fixed, at least partly, by using spin echo DP/MAS experimental configuration, as done
457   here on the standard samples. By applying two subsequent pulses (P_{90}-τ-P_{180}- τ, with τ being a
458   multiple of the MAS frequency), a rephasing of the signal is obtained, and broad features are
459   not anymore distorted (Hahn, 1950; Vieira et al., 2016; Lucherelli et al., 2019). Cross
460   polarization Magic Angle Spinning (CP/MAS) analysis is based on the magnetization transfer

from hydrogen to carbon atoms. It provides spectra with an improved signal-to-noise ratio, but the quantification is not straightforward (Baccile et al., 2014; dela Rosa et al., 1992). Thus, the intensity of the signal arising from a carbon atom depends on its hydrogen environment, and the signal depends on the contact time. It requires an estimate of the fraction of undetected carbon atoms through calculations constrained by the elemental composition of the sample. The CP/MAS estimation is then much less direct than DP/MAS, and may suffer from error propagation on speciation quantitation and elemental composition, as well as a non-comprehensive identification of all involved carbon speciation.

Cody and Alexander, (2005) and Gardinier et al. (2000) both report CP/MAS NMR data that provide quantitative estimates of the composition of IOM from Orgueil and Murchison. In the case of the Murchison chondrite, the ratio of aliphatic to aromatic carbon atoms F_{ali}/F_{aro} estimated by Gardinier et al. (2000) is slightly higher than that of Cody and Alexander (2005) but it lies within the error bars (Fig. 8). In the case of Orgueil, the F_{ali}/F_{aro} ratio from is lower than that of Cody and Alexander (2005), but the error bars also overlap (Fig. 8). In contrast, there is some significant disagreement regarding carbon atoms in oxygenated functions, in particular for Orgueil (Fig. 8). These differences may be due to different experimental procedures. Gardinier et al. (2000) investigated the spectral dependence upon contact time, applied a correction accounting for undetected carbon atoms based on the bulk H/C ratio, determined from the combustion experiment. The abundance of undetected aromatic carbon was found significant, ranging from 20 – 29%. We note that the H/C value (0.7 in their study is higher than that measured by Alexander et al. (2007, 2013) (0.588 ± 0.016 and 0.618, respectively), which may have an impact on the derived abundances. Cody and Alexander (2005) did not run systematic measurements as a function of contact time, but this latter was set to 4.5 ms to fit with SP/MAS spectra. This quantitative interpretation of CP/MAS spectra without investigations at variable contact time is not usual, and estimates might suffer from biases (S. Derenne, personal communication; see also). In fact, in the details, the estimates derived by Cody and Alexander (2005) from SP/MAS and CP/MAS show significant differences. Due to these complex issues with CP/MAS, we have restricted to the SP/MAS data from Cody and Aleander (2005) for comparison with our FTIR data.

We observe that $n_{CHx}/n_{C=C}$ estimated from infrared data following equation 2 fits well NMR data in the case of Orgueil, Murchison and Tagish Lake. $n_{CHx}/n_{C=C}$ in EET 92042 is underestimated by a factor of ~ 1.3, though the error bars overlap (Fig. 8). The first explanation of this moderate discrepancy is the chemical control of the A_{Aro} absorption cross-section by oxygenated groups (Painter et al., 1981). The O/C ratio in EET 92042 (O/C=16.6 ± 0.2) is lower than those of Murchison (18.3 ± 0.3) and Orgueil (18.1 ± 0.2). These values are quite similar. According to NMR data, the main oxygen speciation in chondritic IOM is carbonyl (ketone, ester) and ether bridges. The estimations from SP/MAS NMR spectra are challenging due to the faintness of the features and the poor signal-to-noise ratio, and they do not point to significant variations between EET 92042, and Orgueil and Murchison. Altogether, there is no evidence of a chemical control of the A_{Aro} absorption cross-section. A second tentative explanation is that the SP/MAS data are not quantitative enough, due to the biases abovementioned. The only way to firmly solve this issue would be to collect DP/MAS NMR spectra (MAS synchronized spin echoes) of IOMs from those three chondrites.

504 Regarding carbonylated species, we observe that $n_{CHx}/n_{C=O}$ is systematically higher than
505 NMR estimates, though the hierarchy Orgueil > Murchison > Tagish Lake is similar to NMR
506 estimates (Fig. 8). The number of bonds is equal to the number of carbon atoms in the case of
507 CH_x groups and carbonyl; therefore, the difference may be due here to different absorption
508 cross-sections of the C=O group between these IOMs and the HCO-tholin standard. As for the
509 C=C bonds, the local molecular environment plays a role in the control of the absorption cross-
510 section of the C=O stretching mode. Such variations have been observed for small molecules in
511 the solid-state, for instance, amorphous ices of formic acid and acetaldehyde, whose carbonyl
512 absorption cross-sections differ by a factor of 2 (Bouilloud et al., 2015; Hudson & Ferrante,
513 2020). Nevertheless, as mentioned in the previous section, the re-analysis of these IOM samples
514 by DP/MAS NMR (MAS synchronized spin echoes) is required to get a firm conclusion.

515 *To sum up, the $H^{13}CO$-tholins standard provides a suitable calibration for estimating*
516 *$n_{CHx}/n_{C=C}$ in IOMs derived from FTIR measurement compared with SP/MAS NMR. In contrast,*
517 *$n_{CHx}/n_{C=O}$ is overestimated in all classes of chondrites compared with the SP/MAS NMR from*
518 Cody and Aleander (2005).

519

4.3 Composition of ligno-cellulosic coals

521 We report now the composition of a series of coals of varying ranks from lignite to semi-
522 anthracite (Table S-2, Fig. S-5). The majority of coals plot within the Type III kerogen band on
523 the "Van Krevelen" diagram (Fig. 9a). Due to spectral congestion, the fit in the region 1800-
524 1500 cm^{-1} region is more complicated than for IOMs. We used four spectral components due to
525 ketones /quinone, aromatic branched with phenolic -OH (or -O-), and aromatic with a carboxylic
526 group (Chen et al., 2012; Painter et al., 1981; Petersen et al., 2008) (Fig. S-6). The abundance
527 of aromatic groups is then estimated through the two last components, and the C=O by the first
528 two components. Our results of the evolution trends are consistent with previous estimations
529 based on integrated abundance derived from IR spectra (Chen et al., 2012). No NMR data are
530 available for our coals, but NMR data are available for Chinese series of coals (Wei & Tang,
531 2018) and for the so-called Argonne Premium coals, a subset of the Penn State Coals Data bank
532 (dela Rosa et al., 1992). Figure 9b-c reports the FTIR and NMR CP/MAS estimates of Wei and
533 Tang (2018) plotted against the vitrinite reflectance. We observe two opposite trends: while the
534 NMR estimates point to a progressive decrease of $n_{CHx}/n_{C=C}$, the FTIR estimates show an
535 increase up to the vitrinite reflectance of 1.73% and then a decline for the semi-anthracite
536 samples. In contrast, n_{CO}/n_{CHx} is consistent with NMR. This situation is opposite to what he had
537 with IOMs (Fig. 8).

538 The NMR estimates of Wei and Tang (2018) are, at least qualitatively, in agreement with
539 the progressive aromatization of the coals during diagenesis and catagenesis (Chen et al., 2012).
540 This supports the view that the ratio of the cross-section A_{CHx}/A_{Aro} varies with the vitrinite
541 reflectance, as previously reported (Painter et al., 1981; Dela Rosa et al., 1992). It also shows
542 that any semi-quantitative analysis based on IR measurements leads to inaccurate estimates of
543 the chemical evolution. NMR data collected by Wei & Tang (2018) were obtained in CP/MAS
544 with a 3 ms contact time. The publication does not report measurements over variable time

14

contact, nor the acquisition of single pulse spectra. In contrast, Dela Rosa et al. (1992) report varying contact time CP/MAS measurements on Argonne Premium coals and one single pulse (Bloch decay) analysis of Pittsurgh n°8 that fits the corresponding CP/MAS estimate. Muntean et al. (1990) also report consistent CP/MAS and single pulse estimates in the case of Pocahontas N°3, provided that the delay time remains < 75 µs. However, the agreement between both types of measurements may be worse in the case of O-rich low-rank coals (Dela Rosa et al., 1992). We have recalculated the aromaticity of the coals from Wei and Tang (2018) following the definition of Dela Rosa et al. (1992), and found that the difference between estimates of the two sets of data does not exceed 5%.

Finally, using those data and keeping in mind that uncertainties might be worse for the more immature coals, we have calculated the cross-section ratio $A_{CHx}/A_{C=C}$ upon the vitrinite reflectance (Fig. 9d). This curve is best fitted with a 6th order polynomial (Equation 3):

$$\frac{A_{CHx}}{A_{C=C}} = \sum_{i=0}^{6} K_i(x - x_0)^i$$

(Equation 3)

with $K_0=0.64599$; $K_1=0.23482$; $K_2=6.3733$; $K_3=-16.939$; $K_4=23.446$; $K_5=-14.846$; $K_6=3.7045$ and $X_0=0.3$. This formula applies in the range $R_o = 0.3 – 2 \%$, provided that a similar fit model is used to decompose the coal infrared spectra. We observe that $A_{CHx}/A_{C=C}$ increases and $A_{C=C}$ decreases, upon increasing the vitrinite reflectance R_o, consistently with control of the absorption cross-section by oxygenated groups. The increase is quite high, around one order of magnitude, which means that semi-quantitative sets of infrared data must be imperatively corrected. However, further investigations combining DP/MAS NMR and FTIR measurements on the same series of coals should provide more accurate data.

5   CONCLUSION

We have tested a protocol for quantifying the abundance of the CHx, C=O and C=C groups in chondritic IOMs and coals, based on a ^{13}C model material characterized by DP/MAS NMR and FTIR. The ratios of the cross-sections of these groups have been determined, and the method has been applied to IOM samples extracted from CI, CM and CR chondrites, and to a series of coals. The estimate of $n_{CHx}/n_{C=C}$ in IOM from CI and CM chondrites is fair, but $n_{CHx}/n_{C=O}$ is overestimated. The estimate of $n_{CHx}/n_{C=C}$ in the CR chondrite is lower than the estimation by SP/MAS published earlier. These discrepancies are due to the control of the absorption of the C=O and C=C bonds by the local oxygenated chemical structure. In this regard, the oxygen speciation in CR chondrites may be significantly different than that of CI and CM chondrites. However, new NMR analyses of IOM from CR chondrites are necessary to test the reproducibility of existing data. Regarding coals, the use of published NMR analyses shows that the ratio $A_{CHx}/A_{C=C}$ varies with the vitrinite reflectance over an order of magnitude. Here as well, the local oxygen speciation plays a critical control on A_{Aro}, which decreases with increasing the vitrinite reflectance. We propose an analytical law for A_{CHx}/A_{Aro}, which should be helpful to determine the aromaticity of any coal sample, provided its vitrinite reflectance is known.

584

595 **Supplementary data**

596 Supplementary data related to this article can be found at the Supporting Information part.

597

598 **Reference**

599 Abarghani, A., Gentzis, T., Shokouhimehr, M., Liu, B., & Ostadhassan, M. (2020). Chemical
600   heterogeneity of organic matter at nanoscale by AFM-based IR spectroscopy. *Fuel*,
601   *261*(116454). https://doi.org/10.1016/j.fuel.2019.116454

602 Alcouffe, G., Cavarroc, M., Cernogora, G., Ouni, F., Jolly, A., Boufendi, L., & Szopa, C.
603   (2010). Capacitively coupled plasma used to simulate Titan's atmospheric chemistry.
604   *Plasma Sources Science and Technology*, *19*(1). https://doi.org/10.1088/0963-
605   0252/19/1/015008

606 Baccile, N., Falco, C., & Titirici, M. M. (2014). Characterization of biomass and its derived
607   char using 13C-solid state nuclear magnetic resonance. *Green Chemistry*, *16*(12), 4839–
608   4869. https://doi.org/10.1039/c3gc42570c

609 Battandier, M., Bonal, L., Quirico, E., Beck, P., Engrand, C., Duprat, J., & Dartois, E. (2018).
610   Characterization of the organic matter and hydration state of Antarctic micrometeorites:
611   A reservoir distinct from carbonaceous chondrites. *Icarus*, *306*, 74–93.
612   https://doi.org/10.1016/j.icarus.2018.02.002

613 Beck, P., Quirico, E., Montes-hernandez, G., L., B., J., B., Orthous-Daunay, F.-R., … O., B.
614   (2010). Hydrous mineralogy of CM and CI chondrites from infrared spectroscopy and
615   their relationship with low Albedo asteroid. In *41st Lunar and Planetary Science*.

616 Belkhou, R., Stanescu, S., Swaraj, S., Besson, A., Ledoux, M., Hajlaoui, M., & Dalle, D.
617   (2015). HERMES: A soft X-ray beamline dedicated to X-ray microscopy. *Journal of
618   Synchrotron Radiation*, *22*, 968–979. https://doi.org/10.1107/S1600577515007778

619 Bouilloud, M., Fray, N., Bénilan, Y., Cottin, H., Gazeau, M. C., & Jolly, A. (2015).
620   Bibliographic review and new measurements of the infrared band strengths of pure
621   molecules at 25 K: H2O, CO2, CO, CH4, NH3, CH3OH, HCOOH and H2CO. *Monthly
622   Notices of the Royal Astronomical Society*, *451*(2), 2145–2160.
623   https://doi.org/10.1093/mnras/stv1021

624 Chen, Y., Mastalerz, M., & Schimmelmann, A. (2012). Characterization of chemical
625   functional groups in macerals across different coal ranks via micro-FTIR spectroscopy.
626   *International Journal of Coal Geology*, *104*, 22–33.
627   https://doi.org/10.1016/j.coal.2012.09.001

628 Cody, G. D., & Alexander, C. M. O. D. (2005). NMR studies of chemical structural variation
629   of insoluble organic matter from different carbonaceous chondrite groups. *Geochimica et*
630   *Cosmochimica Acta*, *69*(4), 1085–1097. https://doi.org/10.1016/j.gca.2004.08.031

631 Craddock, P. R., Doan, T. V. Le, Bake, K., Polyakov, M., Charsky, A. M., & Pomerantz, A.
632   E. (2015). Evolution of Kerogen and Bitumen during Thermal Maturation via Semi-
633   Open Pyrolysis Investigated by Infrared Spectroscopy. *Energy and Fuels*, *29*, 2197–
634   2210. https://doi.org/10.1021/ef5027532

635 Dartois, E., Geballe, T. R., Pino, T., Cao, A.-T., Jones, A., Deboffle, D., … D'Hendecourt, L.
636   (2007). IRAS 08572+3915: Constraining the aromatic versus aliphatic content of
637   interstellar HACs. *Astronomy & Astrophysics*, *463*, 635–640.
638   https://doi.org/10.1051/0004-6361

639 Dartois, E., Munoz Caro, G. M., Deboffle, D., & D'Hendecourt, L. (2004). Diffuse
640   interstellar medium organic polymers Photoproduction of the 3.4, 6.85 and 7.25 μm
641   features. *Astronomy & Astrophysics*, *423*, 33–36.

642 Dela Rosa, L., Pruski, M., Lang, D., Gerstein, B., & Solomon, P. (1992). Characterization of
643   the Argonne Premium Coals by using 1H and 13C NMR and FT-IR spectroscopies.
644   *Energy and Fuels*, *6*(4), 460–468. https://doi.org/10.1021/ef00034a016

645 Ferrari, A. C., & Robertson, J. (2001). Resonant Raman spectroscopy of disordered,
646   amorphous, and diamondlike carbon. *Physical Review B - Condensed Matter and*
647   *Materials Physics*, *64*(7), 1–13. https://doi.org/10.1103/PhysRevB.64.075414

648 Fung, B. M., Khitrin, A. K., & Ermolaev, K. (2000). An Improved Broadband Decoupling
649   Sequence for Liquid Crystals and Solids. *Journal of Magnetic Resonance*, *142*(1), 97–
650   101. https://doi.org/10.1006/jmre.1999.1896

651 Gardinier, a, Derenne, S., Robert, F., Behar, F., Largeau, C., & Maquet, J. (2000). Solid state
652   CP/MAS 13C NMR of the insoluble organic matter of the Orgueil and Murchison
653   meteorites: quantitative study. *Earth and Planetary Science Letters*, *184*, 9–21.
654   https://doi.org/http://dx.doi.org/10.1016/S0012-821X(00)00300-9

655 Guillou, C. Le, Bernard, S., Pena, F. De, & Brech, Y. Le. (2018). XANES-Based
656   Quantification of Carbon Functional Group Concentrations. *Analytical Chemistry*, *90*,
657   8379–8386. research-article. https://doi.org/10.1021/acs.analchem.8b00689

658 Hahn, E. L. (1950). Spin Echoes. *Phys Rev Lett*.

659 Hudson, R. L., & Ferrante, R. F. (2020). Quantifying acetaldehyde in astronomical ices and
660   laboratory analogues: IR spectra, intensities, 13C shifts, and radiation chemistry.
661   *Monthly Notices of the Royal Astronomical Society*, *492*, 283–293.
662   https://doi.org/10.1093/mnras/stz3323

663 Jakob, D. S., Wang, L., Wang, H., & Xu, X. G. (2019). Spectro-Mechanical Characterizations
664   of Kerogen Heterogeneity and Mechanical Properties of Source Rocks at 6 nm Spatial
665   Resolution. *Analytical Chemistry*, *91*(14), 8883–8890.
666   https://doi.org/10.1021/acs.analchem.9b00264

667 Jubb, A. M., Birdwell, J. E., Hackley, P. C., Hatcherian, J. J., & Qu, J. (2020). Nanoscale
668   Molecular Composition of Solid Bitumen from the Eagle Ford Group across a Natural
669   Thermal Maturity Gradient. *Energy and Fuels*, *34*(7), 8167–8177.
670   https://doi.org/10.1021/acs.energyfuels.0c00963

671 Kebukawa, Y., Alexander, C. M. O. D., & Cody, G. D. (2011). Compositional diversity in
672   insoluble organic matter in type 1, 2 and 3 chondrites as detected by infrared
673   spectroscopy. *Geochimica et Cosmochimica Acta*, *75*(12), 3530–3541.

18

674   https://doi.org/10.1016/j.gca.2011.03.037

675   Kebukawa, Y., Kobayashi, H., Urayama, N., Baden, N., & Kondo, M. (2018). Nanoscale
676   infrared imaging analysis of carbonaceous chondrites to understand organic-mineral
677   interactions during aqueous alteration. *PNAS*, 1–6.
678   https://doi.org/10.1073/pnas.1816265116

679   Kingsley, P. B. (1999). Methods of measuring spin-lattice (T1) relaxation times: An
680   annotated bibliography. *Concepts in Magnetic Resonance*, *11*(4), 243–276.
681   https://doi.org/10.1002/(SICI)1099-0534(1999)11:4<243::AID-CMR5>3.0.CO;2-C

682   Le Brech, Y., Raya, J., Delmotte, L., Brosse, N., Gadiou, R., & Dufour, A. (2016).
683   Characterization of biomass char formation investigated by advanced solid state NMR.
684   *Carbon*, *108*, 165–177. https://doi.org/10.1016/j.jrp.2009.07.007

685   Le Guillou, C., Bernard, S., De La Pena, F., & Le Brech, Y. (2018). XANES-Based
686   Quantification of Carbon Functional Group Concentrations. *Analytical Chemistry*,
687   *90*(14), 8379–8386. https://doi.org/10.1021/acs.analchem.8b00689

688   Li, W., Zhu, Y. M., Wang, G., & Jiang, B. (2016). Characterization of coalification jumps
689   during high rank coal chemical structure evolution. *Fuel*, *185*, 298–304.
690   https://doi.org/10.1016/j.fuel.2016.07.121

691   Lis, G. P., Mastalerz, M., Schimmelmann, A., Lewan, M. D., & Stankiewicz, B. A. (2005).
692   FTIR absorption indices for thermal maturity in comparison with vitrinite reflectance R0
693   in type-II kerogens from Devonian black shales. *Organic Geochemistry*, *36*(11), 1533–
694   1552. https://doi.org/10.1016/j.orggeochem.2005.07.001

695   Lucherelli, M. A., Raya, J., Edelthalhammer, K. F., Hauke, F., Hirsch, A., Abellán, G., &
696   Bianco, A. (2019). A Straightforward Approach to Multifunctional Graphene. *Chemistry
697   - A European Journal*, *25*(57), 13218–13223. https://doi.org/10.1002/chem.201903165

698   MacPhail, R. A., Strauss, H. L., Snyder, R. G., & Eiliger, C. A. (1984). C-H stretching modes
699   and the structure of n-alkyl chains. 2. Long, all-trans chains. *Journal of Physical
700   Chemistry*, *88*(3), 334–341. https://doi.org/10.1021/j150647a002

701   Mapelli, C., Castiglioni, C., Zerbi, G., & Mullen, K. (1999). Common force field for graphite
702   and polycyclic aromatic hydrocarbons. *Physical Review B - Condensed Matter and
703   Materials Physics*, *60*(18), 12710–12725. https://doi.org/10.1103/PhysRevB.60.12710

704   Massiot, D., Fayon, F., Capron, M., King, I., Le Calvé, S., Alonso, B., … Hoatson, G. (2002).
705   Modelling one- and two-dimensional solid-state NMR spectra. *Magnetic Resonance in
706   Chemistry*, *40*(1), 70–76. https://doi.org/10.1002/mrc.984

707   Mathurin, J., Dartois, E., Pino, T., Engrand, C., Duprat, J., Deniset-besseau, A., … Dazzi, A.
708   (2019). Nanometre-scale infrared chemical imaging of organic matter in ultra-
709   carbonaceous Antarctic micrometeorites (UCAMMs). *Astronomy & Astrophysics*, *160*,
710   1–9.

711   Orthous-Daunay, F. R., Quirico, E., Beck, P., Brissaud, O., Dartois, E., Pino, T., & Schmitt,
712   B. (2013). Mid-infrared study of the molecular structure variability of insoluble organic
713   matter from primitive chondrites. *Icarus*, *223*, 534–543.
714   https://doi.org/10.1016/j.icarus.2013.01.003

715   Painter, P. P., Snyder, R. W., Starsinic, M., Coleman, M. M., Kuehn, D. W., & Davis, A.
716   (1981). Concerning the application of FT-IR to the study of coal: A critical assessment
717   of band assignements and the application of spectral analysis programs. *Applied
718   Spectroscopy*, *35*(5), 475–485. https://doi.org/10.1366/0003702814732256

719   Petersen, H. I., Rosenberg, P., & Nytoft, H. P. (2008). Oxygen groups in coals and alginite-

rich kerogen revisited. *Coal Geology*, *74*, 93–113.
https://doi.org/10.1016/j.coal.2007.11.007

Pretsch, E., Bühlmann, P., & Affolter, C. (2000). *Structure Determination of Organic Compounds*. *Structure Determination of Organic Compounds*. Springer.
https://doi.org/10.1007/978-3-662-04201-4

Quirico, E, Bonal, L., Beck, P., Alexander, C. M. O. D., Yabuta, H., Nakamura, T., …
Schmitt-kopplin, P. (2018). Prevalence and nature of heating processes in CM and C2-
ungrouped chondrites as revealed by insoluble organic matter. *Geochimica et Cosmochimica Acta*, *241*, 17–37. https://doi.org/10.1016/j.gca.2018.08.029

Quirico, Eric, Orthous-Daunay, F. R., Beck, P., Bonal, L., Brunetto, R., Dartois, E., …
Duprat, J. (2014). Origin of insoluble organic matter in type 1 and 2 chondrites: New
clues, new questions. *Geochimica et Cosmochimica Acta*, *136*, 80–99.
https://doi.org/10.1016/j.gca.2014.03.025

Russell, L. M., Takahama, S., Liu, S., Hawkins, L. N., Covert, D. S., Quinn, P. K., & Bates,
T. S. (2009). Oxygenated fraction and mass of organic aerosol from direct emission and
atmospheric processing measured on the R/V Ronald Brown during
TEXAQS/GoMACCS 2006. *Journal of Geophysical Research Atmospheres*, *114*(7), 1–
15. https://doi.org/10.1029/2008JD011275

Sandford, S. A., Allamandola, L. J., Tielens, A. G. G. M., Sellgren, K., Tapia, M., &
Pendleton, Y. (1991). The interstellar C-H stretching band near 3.4 microns: constraints
on the composition of organic material in the diffuse interstellar medium. *The Astrophysical Journal*, *371*, 607–620.

Sciamma-O'Brien, E., Carrasco, N., Szopa, C., Buch, A., & Cernogora, G. (2010). Titan's
atmosphere: An optimal gas mixture for aerosol production? *Icarus*, *209*(2), 704–714.
https://doi.org/10.1016/j.icarus.2010.04.009

Sun, X. (2005). The investigation of chemical structure of coal macerals via transmitted-light
FT-IR microspectroscopy. *Spectrochimica Acta - Part A: Molecular and Biomolecular Spectroscopy*, *62*, 557–564. https://doi.org/10.1016/j.saa.2005.01.020

Swaraj, S., Belkhou, R., Stanescu, S., Rioult, M., Besson, A., & Hitchcock, A. P. (2017).
Performance of the HERMES beamline at the carbon K-edge. *Journal of Physics: Conference Series*, *849*(1). https://doi.org/10.1088/1742-6596/849/1/012046

Szopa, C., Cernogora, G., Boufendi, L., Correia, J. J., & Coll, P. (2006). PAMPRE : A dusty
plasma experiment for Titan's tholins production and study. *Planetary and Space Science*, *54*, 394–404. https://doi.org/10.1016/j.pss.2005.12.012

Takahama, S., Johnson, A., Russell, L. M., Takahama, S., & Johnson, A. (2013).
Quantification of Carboxylic and Carbonyl Functional Groups in Organic Aerosol
Infrared Absorbance Spectra. *Aerosol Science and Technology*, *47*(3), 310–325.
https://doi.org/10.1080/02786826.2012.752065

Vieira, M. A., Gonçalves, G. R., Cipriano, D. F., Schettino, M. A., Silva Filho, E. A., Cunha,
A. G., … Freitas, J. C. C. (2016). Synthesis of graphite oxide from milled graphite
studied by solid-state 13C nuclear magnetic resonance. *Carbon*, *98*, 496–503.
https://doi.org/10.1016/j.carbon.2015.11.037

Wang, J., Morin, C., Li, L., Hitchcock, A. P., Scholl, A., & Doran, A. (2009). Radiation
damage in soft X-ray microscopy. *Journal of Electron Spectroscopy and Related Phenomena*, *170*(1–3), 25–36. https://doi.org/10.1016/j.elspec.2008.01.002

Wei, Q., & Tang, Y. (2018). 13C-NMR study on structure evolution characteristics of high-

766   organic-sulfur coals from typical chinese areas. *Minerals*, *8*(2).
767   https://doi.org/10.3390/min8020049

768   Xueqiu, H., Xianfeng, L., Baisheng, N., & Dazhao, S. (2017). FTIR and Raman spectroscopy
769   characterization of functional groups in various rank coals. *Fuel*, *206*, 555–563.
770   https://doi.org/10.1016/j.fuel.2017.05.101

771   Yabuta, H., Naraoka, H., Sakanishi, K., & Kawashima, H. (2005). Solid-state 13C NMR
772   characterization of insoluble organic matter from Antarctic CM2 chondrites: Evaluation
773   of the meteoritic alteration level. *Meteoritics and Planetary Science*, *40*(5), 779–787.
774   https://doi.org/10.1111/j.1945-5100.2005.tb00979.x

775   Yang, J., Hatcherian, J., Hackley, P. C., & Pomerantz, A. E. (2017). Nanoscale geochemical
776   and geomechanical characterization of organic matter in shale. *Nature Communications*,
777   *8*(2179). https://doi.org/10.1038/s41467-017-02254-0

778
779

780 Table 1. Chemical shifts, corrected relative abundances of the different functional carbon
781 groups in the initial and heated H^{13}CO-tholins derived from Figure 1b-c.

782 Table 2. The functional carbon groups concentrations given by the XANES-based
783 quantification (Le Guillou et al., 2018).

784 Table 3. Integrated cross-section ratio calculation of the 400 ℃ H^{13}CO-tholins based on the
785 quantitative NMR and the infrared spectral decomposition.

786

787   Figure 1. (a) The unheated and heated at 400 ℃ H^{13}CO-Tholins recorded with a sample
788   spinning rate of 60 kHz and interscan delay of 30 seconds in the comparison with the IOM
789   Murchison and Orgueil (Cody et al., 2005, 2002, Gardinier et al., 2000). Deconvolution of the
790   solid state ^{13}C DP/MAS NMR spectra of (b) the initial H^{13}CO-tholins (c) and the 400 ℃ H^{13}CO-
791   tholins

792   Figure 2. Raman spectrum of the 400°C H^{13}CO- tholin. The G and D bands at ~1537 cm^{-1} and
793   ~ 1310 cm^{-1}, respectively reveal a polyaromatic structure typical of coals and kerogens. The
794   low signal-to-noise ratio and spectral artifacts at the edge are due to the high fluorescence
795   background only partly corrected by the baseline correction.

796   Figure 3. (a) Normalized XANES spectra of the 400 ℃ H^{13}CO and H^{12}CO-tholins dominated
797   by aliphatic (-CH$_x$), C=C and C=O functional groups and (b) XANES and NMR methods
798   correlation of the 400 ℃ H^{13}CO-tholins with the re-produced data from Le Guillou et al. (2018)
799   and Le Brech et al. (2016).

800   Figure 4. Corrected infrared spectra of H^{12}CO and H^{13}CO-tholins samples. Highlighted
801   attributions (CH$_3$ and CH$_2$ anti-symmetric modes and C=O and C=C stretching modes) are
802   those used for quantitative study.

803   Figure 5. Detailed components of CH$_x$ mode in different phases (liquid, crystalline, amorphous)
804   in an example of alkane C$_{20}$H$_{42}$ (modified from MacPhail et al., 1984).

805   Figure 6. (a) Fit of the alkyl band of an IOM extracted from the Orgueil chondrite, obtained
806   through the artificial bees' algorithm. This fit was achieved with the 7 Gaussian model and all
807   components FWHM forced to 35 cm^{-1}. (b)-(c) Illusion of uncertainty of spectral decomposition
808   methods in different scenarios: Gaussian fix a FWHM of 35 cm^{-1} (Blue); Voigt fix a FWHM of
809   20 – 40 cm^{-1} (Yellow); Gaussian free a FWHM range of 20 - 40 cm^{-1} (Green), Voigt free a
810   FWHM range of 20 - 40 cm^{-1} (Red).

811   Figure 7. Normalized spectral intensity of the organic functional groups derived from the
812   calibration of C=O and C=C and CH$_x$ groups using the DP/MAS NMR and FTIR calibration
813   for the 400 ℃ H^{13}CO- tholins on (a) Murchison (CM2), (b) Orgueil (CI1), (c) EET 92042 (CR2)
814   and (d) Tagish Lake (C2-ung) IOM in the comparison with the SP/MAS ^{13}C NMR spectra
815   (Cody et al., 2005) and CP/MAS ^{13}C NMR (Gardinier et al., 2000).

816   Figure 8. Comparisons plots between (a) n$_{CHx}$/n$_{C=C}$ and n$_{CHx}$/n$_{C=O}$ ratios (b) n$_{CHx}$/n$_{C=C}$ and
817   n$_{C=O}$/n$_{C=C}$ of IOM functional C groups (Murchison (CM2), Orgueil (CI1), EET92042 (CR2)
818   and Tagish Lake (C2-ung)) derived from FTIR-NMR quantification and the previous NMR's
819   measurement (Cody & Alexander, 2005; Gardinier et al., 2000; Yabuta et al., 2005).

820   Figure 9. Comparisons plots between (a) H/C and O/C atomic ratios of studied coals (b)
821   Normalized intensity of C=O, n$_{C=O + COOH(R)}$ vs. vitrinite reflectance, (c) n$_{CHx}$/n$_{C=C}$ vs. vitrinite
822   reflectance and (d) Calculated A$_{CHx}$/A$_{C=C}$ vs. vitrinite reflectance derived from FTIR-NMR
823   quantification and the previous NMR's measurement (Wei & Tang, 2018).