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Maxwell C Brown, Gwenaël Hervé, Monika Korte, Agnès Genevey. Global archaeomagnetic data: The state of the art and future challenges. *Physics of the Earth and Planetary Interiors*, 2021, 318, pp.106766. 10.1016/j.pepi.2021.106766 . insu-03280044

**HAL Id: insu-03280044**

**<https://insu.hal.science/insu-03280044>**

Submitted on 7 Jul 2021

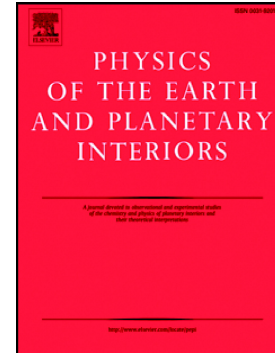
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## Journal Pre-proof

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PII: S0031-9201(21)00124-2

DOI: <https://doi.org/10.1016/j.pepi.2021.106766>

Reference: PEPI 106766

To appear in: *Physics of the Earth and Planetary Interiors*

Received date: 7 January 2021

Revised date: 17 June 2021

Accepted date: 29 June 2021

Please cite this article as: M.C. Brown, G. Hervé, M. Korte, et al., Global archaeomagnetic data: The state of the art and future challenges, *Physics of the Earth and Planetary Interiors* (2021), <https://doi.org/10.1016/j.pepi.2021.106766>

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# Global archaeomagnetic data: the state of the art and future challenges

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## Abstract

Archaeomagnetic data are fundamental for our understanding of the evolution of Earth's magnetic field on centennial to millennial timescales. From the earliest studies of the Thelliers, Aitken, Nagata and others in the 1950s and 1960s, archaeomagnetic data have been vital for extending our knowledge of the field to times prior to observational measurements. Today, many thousands of archaeomagnetic data allow us to explore the geomagnetic field in more detail than ever before. Both regional time series of archaeomagnetic data and the inclusion of archaeomagnetic data in time-varying global spherical harmonic field models have revealed a range of newly discovered field behaviour. More sophisticated approaches to developing regional curves and global models have allowed us to resolve the field in certain regions more robustly and with greater resolution than previously possible. In this review we give an overview of the widely used global

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archaeomagnetic database GEOMAGIA50, discuss the methods used to obtain archaeomagnetic data, their challenges, and explore progress over the past twenty years in developing regional secular variation curves and global spherical harmonic models of the archaeomagnetic field. We end the review by covering what we see as the “grand challenges” in archaeomagnetism, including which regions of the world should be focussed on with regards to data acquisition.

*Keywords:* archaeomagnetism, GEOMAGIA50, global models

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## 1. Introduction

Archaeomagnetism is the study of the past direction and intensity of Earth’s magnetic field recorded by any type of manmade artefact or fired material. It is dependent on archaeological discoveries and advances that lead to a better description and understanding of our history and heritage. Although it was recognized at the end of the 19th century that fired materials can record Earth’s magnetic field (Folgeraiter, 1899), it was not until the pioneering work of Émile and Odette Thellier beginning in the 1930s that the physical principles, methods and instrumentation necessary to accurately obtain the past direction and intensity of the geomagnetic field recorded by archaeological materials were developed (Thellier, 1938, 1941; Thellier and Thellier, 1959). Reviews by Thellier (1977), Le Goff et al. (2006) and Dunlop (2011) give excellent English language overviews of the Thelliers’ most important contributions to the subject.

Archaeomagnetism established itself as a research field through the 1950s and 1960s, with proponents of the subject obtaining data from fired materials from locations globally. Studies from these decades reported data from Europe (e.g., Burlatskaya, 1961; Aitken and Weaver, 1962; Belshé et al., 1963; Chelidze, 1965; Bucha, 1967; Kovacheva, 1969), Northern Africa (Athavale, 1969), India (Athavale, 1966), China (Deng and Li, 1965), Japan (Watanabe, 1958; Nagata et al., 1963; Sasajima, 1965), North America (e.g., Watanabe and Dubois, 1965; Schwarz and Christie, 1967) and South America (e.g., Nagata et al., 1965; Kitazawa and Kobayashi, 1968). Research continued through the 1970s, but it was not until the 1980s that there was a general increase in the number of studies reporting new archaeomagnetic data each year (Fig. 1); a trend that continued through to the 2010s. This has resulted in a large compilation of global data that has greatly improved our understanding of how Earth’s magnetic field has varied spatially and temporally on centennial to millennial timescales.



To date close to 700 studies reporting archaeomagnetic data have been published. The majority of studies have concentrated on specific regions, with data from Europe, the Middle East, China, Japan and North America dominating the global database (see section 2.3). A peak in productivity in the 2000s coincided with the successful European Commission funded Archaeomagnetic Applications for the Rescue of Cultural Heritage (AARCH) research and training network. Data from Europe vastly outweighs that from any other region (section 2.3). Since the early 2000s the development of temporally continuous global spherical harmonic models of the geomagnetic field (see section 4.2) and an interest in the development of the South Atlantic Anomaly on archaeomagnetic timescales has led to a number of studies focussed on obtaining data from archaeological sites in the Southern Hemisphere and equatorial regions (e.g., Tarduno et al., 2015; Hartmann et al., 2019). Significant new studies have been published for Africa (Gómez-Paccard et al., 2012b; Neukirch et al., 2012; Mitra et al., 2013; Tarduno et al., 2015; Donadini et al., 2015; Hare et al., 2018; Kapper et al., 2017, 2020; Tchibinda Madingou et al., 2020), South America (e.g., Hartmann et al., 2010, 2011, 2019; Gogutchiaichvili et al., 2011, 2015, 2019; Poletti et al., 2016; Capdepon et al., 2019; Cejudo et al., 2019; Gómez-Paccard et al., 2019) and West Oceania (Stark et al., 2010; Turner et al., 2020). These areas are ripe for expanding our global data set. However, there are limitations on the availability of archaeological materials for analyses from these areas. As archaeomagnetism is a destructive method (artefacts must be cut and often heated), there can be restrictions on the materials available for laboratory analyses.

Figure 1: Histogram of the number of archaeomagnetic studies published per year in the GEOMAGIA50.v3.4 database that contain data dated to the past 10,000 years (accurate as of December 2020). The total number of studies is 685. This excludes studies that employed archaeomagnetic dating as the sole dating method. NB (1) there are additional studies that have published data in non-tabulated form, which have not been added to the database and do not contribute to the total number of studies reported here (e.g., Aitken and Weaver, 1965; Aitken et al., 1989); (2) not all archaeomagnetic studies from Japan have been fully integrated in the current version of GEOMAGIA50 (see section 2.3).

The majority of archaeomagnetic data have been dated to within the past 3000 years, with the number of data on the whole decreasing with increasing age (section 2.4). This has led an

increasing number of studies to focus on obtaining archaeomagnetic data from materials between 6000 BCE (Before the Current Era) and 1000 BCE (e.g., Kovacheva et al., 2009a; Fanjat et al., 2013; Gallet et al., 2014, 2015; Shaar et al., 2016, 2020; Cai et al., 2020); however, almost all data are from Eurasia, limiting our global knowledge of the field at older archaeological times. Extending archaeological time series to older ages is an exciting direction of research for the coming years. Although, as with improving the global distribution of data, limitations on the materials available for analysis impact the time periods that can be studied further.

The usefulness of compiling regional and global archaeomagnetic data for understanding the evolution of the geomagnetic field was recognized early on in the development of the subject (e.g., Cook and Belshé, 1958; Watanabe, 1958; Aitken and Weaver, 1965; Kawai and Hirooka, 1967). This has continued through today, with country or regional specific archaeomagnetic data compilations (e.g., Thellier, 1981; Márton, 2003; Tema et al., 2006; Márton, 2010; Carrancho et al., 2013; Hervé et al., 2013a; De Marco et al., 2014; Kovacheva et al., 2014; Batt et al., 2017; Molina-Cardín et al., 2018; Goguitchaichvili et al., 2019; Schnepf et al., 2020b,a; Rivero-Montero et al., 2021). Such regional data sets have been used to develop secular variation (or reference) curves (see section 4.1), using evermore sophisticated mathematical approaches (recent examples include Lodge and Holme, 2009; Thébaud and Gallet, 2010; Hellio et al., 2014; Batt et al., 2017; Livermore et al., 2018; Genevey et al., 2021; Kapper et al., 2020). Compilations of global archaeointensity data have also been used to infer global dipole moment evolution (e.g., McElhinny and Senanayake, 1982; Aitken et al., 1989; Yang et al., 2000; Genevey et al., 2008; Knudsen et al., 2008; Usoskin et al., 2016).

Over the past 20 years (alongside the construction of direction and intensity curves), has been the development of temporally continuous global palaeomagnetic field models (see section 4.2). These data-based inverse models employ spherical harmonic methods initially developed to analyze and depict the present day field (e.g., Bloxham and Gubbins, 1985; Bloxham and Jackson, 1992) and the historical field (from 1590 CE onwards, based on shipboard and ground based measurements) (Jackson et al., 2000). They have been adapted to suit archaeomagnetic and palaeomagnetic data to produce maps of the geomagnetic field at Earth's surface and the core-mantle boundary (CMB). The earliest global models were developed by Hongre et al. (1998), Constable et al. (2000), Korte and Constable (2003) and Korte and Constable (2005) and combined a variety of data sources (archaeomagnetic, volcanic and sediment data). Global models

based on primarily archaeomagnetic data (but also including volcanic data) were not developed until the construction of ARCH3k.1 (Korte et al., 2009) (a three thousand year model), which was recently updated to a 10,000 year model (Constable et al., 2016). Spherical harmonic cap approaches using archaeomagnetic data have also been used to create regional models (e.g., Pavón-Carrasco et al., 2008; Pavón-Carrasco et al., 2009). Varying approaches to modelling the archaeomagnetic field have been applied since, including Licht et al. (2013), Pavón-Carrasco et al. (2014), Sanchez et al. (2016), Hellio and Gillet (2018), Arneitz et al. (2019) and Mauerberger et al. (2020).

Concurrent to regional compilations of data and the development of global models there have been continued efforts to create global databases of archaeomagnetic data. The first global archaeomagnetic databases were paper lists of results, the first likely being the historical and archaeointensity compilation of Smith (1967). With the development of digital database structures, archaeomagnetic data could be compiled and updated more easily. Early efforts included those of Burlatskaya et al. (1986), Liritzis and Lagios (1993) and Daly and Goff (1996), although the data were not available in a digital form. The first digital archaeomagnetic database that was easily accessible was the Plymouth archaeomagnetic directional database (ARCHEO97 and ARCHEO00) compiled by Don Tauxe and last released in 1999. This was one of seven International Association of Geomagnetism and Aeronomy (IAGA) databases available online to download as stand-alone programs. Two major efforts to compile all global archaeomagnetic data have been the ArcheoInt database (Genevey et al., 2008) and the GEOMAGIA50 database (Donadini et al., 2006; Korbman et al., 2008; Brown et al., 2015b). Although GEOMAGIA50 largely subsumes the data within ArcheoInt, ArcheoInt contains additional fields that place archaeomagnetic results in their archaeological context and provides greater descriptive information regarding the acquisition of the data sets. The databases can be viewed as complementary. In addition, there is the HISTMAG database of Arneitz et al. (2017), which combines historical and archaeomagnetic data. There are also numerous archaeomagnetic data in the MagIC database (described in part in Tauxe et al., 2016); however, GEOMAGIA50 is currently the primary database for archaeomagnetic data. Unlike MagIC, GEOMAGIA50 includes only average data and is not designed to include results at the specimen level or raw measurements. The site level data from GEOMAGIA50 has been used in numerous studies. In addition to being used to construct secular variation curves and global and regional field models, it has been used to

understand solar activity during the Holocene (Usoskin et al., 2016) and to calibrate cosmogenic nuclide production stacks through the use of intensity data (e.g., authigenic  $^{10}\text{Be}/^9\text{Be}$  ratios, Simon et al., 2016).

An important consideration when using archaeomagnetic data for any purpose is the reliability of the data. This includes chronological controls and archaeomagnetic components (direction and intensity), which are most commonly determined from a thermoremanent magnetization (TRM): a magnetization acquired on cooling from firing temperature to room temperature. Archaeomagnetic directions can be influenced by post-cooling displacement and magnetic refraction (section 3.1.1) and obtaining reliable archaeointensities requires that numerous factors are considered (section 3.1.2). These include thermal alteration during palaeointensity experiments (section 3.1.3), the influence of non-ideal magnetic remanence carriers (e.g., multi-domain (MD) grains, section 3.1.4), remanence anisotropy (section 3.1.5) and differences between natural and experimental cooling rates (section 3.1.6). All chronological determinations have an associated uncertainty, whether an archaeological age, determined through physical measurements (e.g., by radiocarbon dating or luminescence methods), or by a combination of approaches (section 3.2). Documenting such uncertainties is a challenge (section 5.1.1) and uncertainties should be carefully considered in any study looking to investigate field behaviour.

In this review we cover the current status of the global archaeomagnetic database (GEOMAGIA50; section 2), provide an overview of archaeomagnetic procedures, data quality, uncertainties and chronological controls (section 3) and explore advances in regional secular variation curve construction and global archaeomagnetic field modelling (section 4). The review ends with a discussion on the future challenges of the subject (section 5).

## **2. Overview of the GEOMAGIA50 archaeomagnetic database**

In the following sections we give a brief history of the GEOMAGIA50 database (section 2.1), cover the abundance of archaeomagnetic data within the most recent version of the database (GEOMAGIA50.v3.4) (section 2.2), discuss the spatial and temporal distribution of data (section 2.3 and section 2.4), and provide an overview of the archaeological materials used to obtain archaeomagnetic data (section 2.5). The methods used to obtain archaeomagnetic and age data, as well as their uncertainties, are discussed in section 3.

In this review we consider purely archaeomagnetic data. Data from volcanic materials (lava, volcanic ashes, obsidian) and speleothems (i.e. Latham et al., 1986; Trindade et al., 2018), although stored in GEOMAGIA50.v3.4, are neglected for the purpose of this study. We also restrict our analysis to materials dated between 8000 BCE and today, and we do not include materials that have been dated using archaeomagnetic dating.

### 2.1. *History of GEOMAGIA50 and its most recent compilation*

Version 1 of GEOMAGIA50 primarily focused on compiling palaeointensity data and contained data from both archaeological materials and lava flows. Directional data were added only if they accompanied intensity data. Version 1 integrated the ArchoInt database of Genevey et al. (2008) and the IAGA ARCHEO00 database (<http://www.ngdc.noaa.gov/geomag/paleo.shtml>) compiled by Don Tarling. Data from other country- or region-specific compilations were also added (see Brown et al. (2015b) for a list of compilations). Further details of version 1 of the database can be found in Donadini et al. (2006), Korhonen et al. (2008) and Brown et al. (2015b). After numerous updates since original publication, 2762 archaeomagnetic entries from 109 studies remain from version 1 in the most up-to-date version of the database.

No publication accompanied version 2 of the database; however, the data compilation is described in Donadini et al. (2009). Around 100 archaeomagnetic entries from version 1 of the database were updated in version 2. Archaeomagnetic directional results were added independently of whether they accompanied intensity data. This greatly increased the amount of data in version 2 of the database, with 3072 data from 130 studies added at this time that remain in the most recent update of the database (5834 entries from 240 studies in total).

The current version of the database is version 3, which was initially published in 2015 (Brown et al., 2015b). It marked a change from being hosted at the Scripps Institution of Oceanography, University of California-San Diego, to GFZ Potsdam (<https://geomagia.gfz-potsdam.de/>). Sediment data were also added in version 3 (Brown et al., 2015a). 1006 entries from 100 archaeomagnetic studies were added to version 3.1 of the database; 498 entries from 220 studies were added to version 3.2 (released in 2017); and 1717 entries from 109 studies in version 3.3 (released in 2019). GEOMAGIA50.v3.2 also incorporated a number of legacy studies (studies published prior to the inception of the database in 2004) that were missing

in previous versions of the database. This included 141 studies from the UK, which was part of a major revision of all UK entries (Batt et al., 2017). It also included 75 UK studies published since 2004.

The most up-to-date version of GEOMAGIA50 (v3.4) was released in December 2020. To our knowledge, it includes nearly all archaeomagnetic studies with independent age constraints published to date, with the exception of a large number of entries in the Japanese archaeomagnetic database (<http://mag.center.ous.ac.jp/en>) and some entries from HISTMAG (Arneitz et al., 2017), which have not yet been integrated into GEOMAGIA50. In total 1188 archaeomagnetic entries from 29 studies were added to GEOMAGIA50.v3.4 in 2020 and the current database contains 9981 archaeomagnetic entries from 685 studies. This is 87% of all entries within the database as a whole. This includes 528 French directional entries determined in the Thellier laboratory at Saint Maur over the past 25 years (Le Goff et al., 2020) and a re-evaluation of the French directional compilations of Thellier (1981) and Bucur (1994) (170 entries). It also contains a significant new compilation of central European archaeomagnetic data, both directional data (Schnepp et al., 2020b) and intensity data (Schnepp et al., 2020a) (188 new entries and 18 updates). Data from China have also been significantly increased with 64 entries published in Cai et al. (2020). Improvements to the Southern Hemisphere/equatorial compilation were made, with new data from Kenya (Tchibinda Madingou et al., 2020), Burkina Faso and Ivory Coast (Kapper et al., 2020), Ecuador (Herrero-Bervera et al., 2020), Colombia (Cejudo et al., 2019), Uruguay (Capdepon et al., 2019) and New Zealand (Turner et al., 2020). Changes in the distribution of data with each version of GEOMAGIA50, both globally and for Europe, are shown in Fig. 2 and Fig. 3.

## 2.2. Overview of archaeomagnetic data

Out of the 9981 archaeomagnetic entries in GEOMAGIA50.v3.4, 5931 archaeomagnetic entries contain either declination or inclination and 4528 entries have both. The majority of entries that only have inclination are from the Russian school (e.g., Burlatskaya et al., 1986) (85% of inclination only entries). Although 5231 entries contain archaeointensity, only 651 entries contain full vector information (declination, inclination and intensity); 533 entries report intensity and inclination without declination; and 4047 entries list intensity without accompanying directions.

In addition to archaeomagnetic results, GEOMAGIA50 contains age and age uncertainty information (see section 3.2) and a variety of meta data that outline the directional, intensity and

dating methods used. It also includes the number of samples/specimens and specimen types investigated, and the types of archaeological materials the data were obtained from (section 2.5). Full details of the fields within GEOMAGIA50 are given in Brown et al. (2015b).

### 2.3. *Spatial distribution of archaeomagnetic data*

Figure 2: Geographical distribution of archaeomagnetic sites given in the six versions of GEOMAGIA50 to date (date range from 8000 BCE to 2000 CE; archaeomagnetically dated sites are excluded). Colours denote when the data were added to the database. (a) sites with directional data; (b) sites with intensity data. Some version 1 (v1) sites were updated with directional information and are shown in (a) as belonging to v1, although they were updated after the initial release of v1. If sites were removed during revisions of subsequent versions, they are not shown on the figure.

Figure 3: Geographical distribution of archaeomagnetic sites given in the six versions of GEOMAGIA50 to date for Europe and surrounding regions (date range from 8000 BCE to 2000 CE; archaeomagnetically dated sites are excluded). (a) sites with directional data; (b) sites with intensity data. See Fig 2 for legend and other details.

There is a large disparity in the global distribution of archaeomagnetic data (Fig. 2). Data from Europe dominates the database (Fig. 3 and Fig. 4a-c): 59% of all entries (including Russia), 51% without Russian data. The UK (10% of entries), France (9%), Russia (8%) and Georgia (5%) contribute to a significant portion the European entries (Fig. 3 and Fig. 4b). Many European countries individually contribute between 2% and 4% of the total number of entries. The UK comprises the largest number of all entries (961), which are primarily directional data (905). See Batt et al. (2017) for further details on the UK contribution. France is the second largest contributor with 890 entries (770 with directions, 162 with intensity). A large amount of data was added (520 entries) following the publication of Le Goff et al. (2020).

Figure 4: Pie charts of the number of archaeomagnetic entries (in brackets) in GEOMAGIA50.v3.4 by region (a-c) and by country (d-f) by data type: (a,d) directional and



intensity, (b,e) directional, and (c,f) intensity. Country plots list the top nine countries by number of entries, with all other entries grouped into a single pie segment.\*The data within GEOMAGIA50.v3.4 does not contain all known Japanese data, which total around 800.

It is worth noting that although 72% of directional entries come from Europe (omitting Russia), this region covers only 1-2% of Earth's surface (depending on the definition of Europe) (Fig. 4b). Data from regions adjacent to Europe are also dense, with the Levant (Israel, Syria, Jordan), Egypt and Iraq contributing significantly to the database. The distribution of data (directions and intensities) from Europe and these regions is shown in Fig. 3.

Outside of Europe the United States of America (7% of entries), China (6%), Japan (4%) and Mexico (3%) are the main contributors. All other nations make up 29% of entries. Although the number of Japanese entries in GEOMAGIA50 totals 370, the Japanese archaeomagnetism database of T. Hatakeyama (Okayama University, Japan) (<http://mag.center.ous.ac.jp/en>) lists 744 directional data and 59 intensity data, placing it third in the list of country entries. We aim to integrate this significant contribution with GEOMAGIA50 in the future.

Although there have been recent efforts to improve the global distribution of data, the Southern Hemisphere is currently poorly represented, with only 400 entries or 4% of all archaeomagnetic entries. 76 entries contain a direction and 340 an intensity. The disparity in data distribution is stark when it is considered that Africa and South America, which cover 9% of Earth's surface when combined (32% of the land area), provide only 7% of the entries in the database (Fig. 4a). However, the amount of Southern Hemisphere data continues to improve. In Fig. 2 we show the increase of Southern Hemisphere data with each new version of GEOMAGIA50. Notable studies that have obtained data from southern Africa are Neukirch et al. (2012), Tarduno et al. (2015) and Hare et al. (2018). Previously only one study had published data from this region (Henthorn et al., 1979) and this was not added to GEOMAGIA50 until version 3.3. A number of South American countries have garnered new data. In the first version of GEOMAGIA50, there were no entries from Brazil, Argentina, Uruguay and Chile. In the past 10 years data have been obtained from all four: Brazil, 49 intensity entries (Hartmann et al., 2010, 2011, 2019; Poletti et al., 2016), Argentina, 44 entries (e.g., Gómez-Paccard et al., 2019; Goguitchaichvili et al., 2019), Uruguay, 6 entries (Capdeponet et al., 2019) and Chile, 1 entry (Roperch et al., 2015). In addition, data have been obtained from other South American countries



south of the Equator. By far the most number of entries come from Peru, with 191 (e.g., Gunn and Murray, 1980; Yang et al., 1993). Smaller contributions come from Bolivia (13 entries) (e.g., Nagata et al., 1965; Kitazawa and Kobayashi, 1968) and Ecuador (23 entries) (Kitazawa and Kobayashi, 1968; Bowles et al., 2002; Herrero-Bervera et al., 2020).

The area between the tropics fairs better than the Southern Hemisphere, with nearly 10% of all entries coming from this latitude band. This includes the large and growing data set from Mexico (see, Hervé et al., 2019b,c; Mahgoub et al., 2019). New studies from India (Basavaiah et al., 2019; Deenadayalan et al., 2020), western Africa (Kapper et al., 2017, 2020) and eastern Africa (Osete et al., 2015) have contributed important intensity data from areas that are isolated from others globally. As we move closer to the equator the amount of available data shrinks with < 2% of database entries from between  $\pm 10^\circ$  latitude. Six studies have produced new data in this latitude band over the past 10 years, with the first archaeomagnetic data from Kenya (Tchibinda Madingou et al., 2020) and the Ivory Coast (Kapper et al., 2020), and others building on small data sets from Ecuador (Herrero-Bervera et al., 2020) and Colombia (Cejudo et al., 2019).

The spatial distribution of directional and intensity data are distinctly different (Fig. 2, Fig. 3 and Fig. 4). Countries that produce numerous directional data do not always produce large amounts of intensity data and vice versa. As stated already, the UK has the most directional entries, but few intensity entries. Conversely, China has the most abundant intensity data by country, but does not make the top ten countries for directional data. Russia, east of the Black Sea has abundant directional data, but sparser intensity data. To a lesser extent the same is true for the Ukraine, which produces the 5th most directional data (5% of all directional entries), but far fewer intensity data than other countries. India and Brazil have no directional data, but numerous intensity data. This disparity can be crucial in areas with sparse data coverage, where full vector data are particularly important for constraining field models, e.g., sites in West Africa, where few directional data have been obtained (Burkina Faso; Donadini et al., 2015), whereas intensity data are more plentiful (Mitra et al., 2013; Kapper et al., 2017, 2020). The greater abundance of intensity data can be related to the availability of material to study (see section 2.5).

#### 2.4. *Temporal distribution of archaeomagnetic data*

There is a large variability in the temporal distribution of data in GEOMAGIA50.v3.4 over the past 10,000 years (Fig. 5). Both the number of archaeomagnetic directions and intensity in

general decrease with age. This is most stark for BCE data, with 35% of all entries from this time. The number of BCE directions is substantially less (20% of total directions) than for CE (Common Era) directions. The contrast is less abrupt for archaeointensity data. Although the number of BCE intensity entries per century is in general less than for CE entries, 54% of all intensity data span 8000 BCE to 1 BCE.

There are notable spikes in the number of directional and intensity entries for certain time periods. For directional data there are peaks in the number of directional data between 100 CE and 300 CE, 700 CE and 900 CE, 1100 and 1400 BCE, and 1700 BCE and 2000 (Fig. 5a). The most populous century for directional results is the 19th (410 entries from 31 studies). Some peaks can be attributed to certain cultural periods, e.g., the high number of entries between 100 CE and 300 CE are from the peak of the Roman Empire, with the data set dominated by entries from present day England, France, Hungary, Bulgaria and Spain. Other peaks are associated with concerted research initiatives in specific countries (or by certain research groups with dedicated focuses), e.g., the 700 CE to 900 CE peak is dominated by data from France for the High Middle Ages (Le Goff et al., 2020).

There is a peak in archaeointensity age entries during the first millennium BCE, where there has been concerted efforts to characterize the Levantine intensity spike (see section 4.1.2). There are notable minor peaks in the number of intensity entries during the Neolithic, with notable studies from the Neolithic and Bronze age from China (207 entries) (see Cai et al., 2020), Iraq (179 entries) (Sakai, 1980; Nachasova and Burakov, 1995, 1998; Yutsis-Akimova et al., 2018a,b) and the rest of the Middle East (142 entries) (e.g., Kawai et al., 1972; Gallet et al., 2014; Stillinger et al., 2015; Shaar et al., 2015; Gallet et al., 2020), Bulgaria (136 entries) (e.g., Kovacheva, 1997; Kovacheva et al., 2009, 2014; Kostadinova-Avramova et al., 2020) and Spain (79 entries) (Nachasova et al., 2002, 2007; Carrancho et al., 2013).

Figure 5: Archaeomagnetic entries in GEOMAGIA50.v3.4 by age in 100 year bins. (a) directions, (b) intensity, (c) Northern Hemisphere data and (d) Southern Hemisphere data.

We note that there are very few BCE data from the Southern Hemisphere. There are only a few data per century back to 6000 BCE (Fig. 5d). In contrast the Northern Hemisphere (Fig. 5c) has 10 times or more data per century.

## 2.5. *Overview of archaeological materials*

A wide range of archaeological materials and structures can be used to obtain directional and intensity information (Fig. 6). Almost all data from archaeological material ( $\sim 99\%$ ) were recovered from baked clays that acquired a TRM roughly parallel and proportional to the ambient geomagnetic field at the time of their firing. A few other archaeological materials can carry a remanent magnetization acquired through different processes. In mural paintings, e.g., frescoes, red pigments with hematite can acquire a so-called pictorial remanent magnetization when paint is sufficiently liquid to enable hematite grains to orientate parallel to the geomagnetic field (e.g., Chiari and Lanza, 1997; Zanella et al., 2000). Through a related process, lime-plasters (e.g., Hueda-Tanabe et al., 2004) and unburnt adobe bricks (e.g., Games, 1977) can also acquire a remanent magnetization, when the plaster or the clay is mixed with water. These materials are promising, even though experimental uncertainties are generally higher than for baked clays.

Fifty types of materials and structures are listed in the current version of GEOMAGIA50; however, there are some that have been sampled more frequently than others. In Fig. 6 we list the 8 most commonly used. In some cases (13% of entries) the type of material that was used is not given in the database. There are clear differences in the materials used for directional and intensity studies. For directional analysis, in-place oriented structures are necessary. Therefore kilns, ovens and hearths, bricks, and burnt structures are frequently used. For intensity the materials do not need to be in-situ, which allows a more diverse array of materials to be pooled from. Pottery and ceramics, owing to their abundance and ease of sampling are therefore the most common for intensity analysis. Over recent years copper slags have been used owing to their magnetically appropriate characteristics for intensity experiments (Shaar et al., 2010).

Materials suitable for intensity are often easier to access, because the material has already been sampled and the collections they are from are well-studied. Sampling of these objects is also less invasive. For directional studies, it is necessary to be reactive to an archaeological excavation. In-situ structures are uncovered and maybe destroyed when working on rescue excavations. Sometimes kiln-type structures are preserved because of an obvious archaeological interest; however, sampling is invasive and possibly incompatible with heritage conservation. These issues may partially explain why the proportion of direction and intensity studies varies in different countries (section 2.3).

Figure 6: Pie charts of the number of mentions (in brackets) of the archaeomagnetic materials used to determine (a) direction and/or intensity, (b) directions, (c) intensity. The eight most used materials are shown in each subplot, remaining material types are grouped under “Other”. Note in (b), most inclination data only come from displaced bricks, making the assumption that they were fired on one of their sides. The number of directional and intensity entries do not match the number of materials given in the plots, as numerous entries were determined from multiple materials.

### 3. Experimental considerations and data quality

In this section we outline the methods that have been used to obtain archaeomagnetic data and date archaeological materials. We discuss how experimental methods and practices affect the accuracy and precision of archaeomagnetic and chronological data and address how uncertainties are represented in the database. For intensity experiments we cover alteration during heating, the influence of multi-domain grains, remanence anisotropy and the effect of cooling rate.

Dating methods applied to archaeological materials are varied and we group them into two categories: those that directly or indirectly date a material. We discuss the nuances of these methods when applied to archaeological materials, how they can be combined to create a site chronology, and their age uncertainties.

#### 3.1. Archaeomagnetic measurements

##### 3.1.1. Directions

Three approaches have commonly been used to recover directional data from archaeological materials. The first two involve stepwise removal (demagnetization) of a TRM by either heating to increasing temperatures (thermal demagnetization) or by increasing the alternating current of a field coil (alternating field (AF) demagnetization). For some entries in the database both approaches have been used in conjunction. An alternative approach is to use viscosity cleaning. Developed by Émile Thellier (see, Thellier, 1981), viscosity cleaning has proven to be as effective as a complete demagnetization in isolating directions, when a sample records a single TRM component. See Le Goff et al. (2020) for an overview of this two-step method. Unfortunately, 56% of entries in the database do not report the demagnetization method used. Of entries that do list a demagnetization method, alternating field (AF) demagnetization is

the most commonly used (33%), followed by viscosity cleaning (28%) (largely from entries from France, Le Goff et al., 2020), a mixture of AF and thermal demagnetization (23%), and solely thermal demagnetization (16%).

There are various factors that are likely to interfere with the accurate recovery of past field directions. First is the precision of the sampling and sample orientation, which is critical in archaeomagnetism where one tries to recover small directional variations. Conservation of structures and mechanical problems, such as the inward or outward sagging of the walls or a slight tilting of the kiln sole, can influence the precision and reliability of the archaeomagnetic direction. The direction recorded by a structure can be further perturbed by magnetic refraction, whereby the magnetization of a structure can distort the magnetic field recorded, in particular when the magnetization is strong (e.g., Aitken and Hawley, 1970; Hus et al., 2004). This can also result from differential cooling as, for instance, may occur in large structures (Lanos, 1987). Understanding magnetic refraction requires dense sampling across all parts of a structure. Too much localized sampling can lead to a precise but biased mean direction.

Another factor that may bias remanence directions is the anisotropy of TRM. For bricks or tiles used to mason all or parts of a kiln, this effect results in a recorded direction that may deviate from the ancient field. Taking this effect into account requires the determination of an anisotropy tensor (see details on the correction for anisotropy effects in section 3.1.5). For baked clay ovens or hearths, the degree of anisotropy is usually considered to be weak, and does not impact the remanence direction, e.g., Kovacheva et al. (2009b) and Le Goff et al. (2020). However, it should be noted that a significant shallowing of inclinations of up to  $13^\circ$  was recently documented for thin oven soles (Palencia-Cintas et al., 2017, 2021). We further note that the GEOMAGIA50 database does not yet make it possible to assess whether or not the anisotropy effect has been evaluated and taken into account in the directional studies.

On the whole the precision of directional data within GEOMAGIA50.v3.4 is variable, but is in general of statistically good quality, with 80% of entries having  $0^\circ \leq \alpha_{95} \leq 5^\circ$  (the cone of confidence at 95%; Fisher, 1953) and 90% with  $0^\circ \leq \alpha_{95} \leq 10^\circ$  (Fig. 7). Some data have particularly low  $\alpha_{95}$  (30% of entries have  $0^\circ \leq \alpha_{95} \leq 2^\circ$ ) and values of  $k$  (the precision parameter; Fisher, 1953) into the thousands. Conversely, some  $\alpha_{95}$  values are notably high and some  $k$  are very low. The precision of directional data can be difficult to quantify for some

entries as  $\alpha_{95}$  is not specified for 6% of directional entries and  $k$  is given for only 50%. We also note that the method of calculating  $\alpha_{95}$  is not always noted in publications. There are two forms of the  $\alpha_{95}$  equation; the original equation in Fisher (1953) and an approximation for a large number of samples (see, e.g., Butler, 1992). These can result in different values of  $\alpha_{95}$  if the number of samples is less than approximately 10.

Less than 2% of entries are based on the successful analysis of only one or two samples and have no associated  $\alpha_{95}$  or  $k$ . When the number of successfully measured samples is at least equal to 3,  $k$  is greater than 100 for 80% of the entries reporting  $k$  (40% of the all directional results). Any study wishing to use directional data should assess the uncertainty that they are comfortable in incorporating into the analysis.

Figure 7: Measures of uncertainty and precision (Fisher, 1953) on archaeomagnetic directional and intensity entries within GEOMAGIA50.v3.4: (a) 95% cone of confidence ( $\alpha_{95}$ ) (bin size = 1 degree); (b) precision parameter ( $k$ ) (bin size = 100  $k$ ). Only  $\alpha_{95}$  values < 20 are shown, corresponding to 5542 entries or 99% of all entries with an  $\alpha_{95}$  or 93% of all directional entries. Only  $k$  values < 4000 are plotted, totalling 2769 values (91% of all entries with  $k$ ; 47% of all directional entries). Whether  $\alpha_{95}$  is calculated using the full equation of Fisher (1953) or an approximation (see, Butler, 1992) is not noted in the database as it is commonly not stated. (c) Uncertainty on archaeointensity estimates expressed as a percentage of the archaeointensity value (bin size = 1%). Note that the uncertainties plotted here are those given by the author and result from different approaches to calculating uncertainty.

### 3.1.2. Archaeointensity determinations

The linearity at low fields (< 150  $\mu$ T) between geomagnetic field strength and the intensity of a TRM acquired on cooling in this field is the physical basis for intensity estimates. A detailed description of the protocols is beyond the scope of this article as there are numerous approaches and derivatives that can be used (e.g., Dunlop, 2011; Tauxe and Yamazaki, 2015; Tauxe et al., 2018), but we give an overview of those used for archaeological entries in GEOMAGIA50 and review the different experimental strategies used to detect and/or possibly

mitigate various effects that influence the intensity measurements.

GEOMAGIA50.v3.4 lists 25 palaeointensity methods and variants; however, these can be primarily classed into five main types, as listed in Fig. 8a. Thellier-type approaches that use double heating steps to impart a laboratory induced TRM as proposed by Thellier and Thellier (1959) make up 87% of all intensity entries in the database. Of the Thellier-type approaches, the original Thellier and Thellier (1959) method has been used more than any other method, followed by the Coe-Thellier approach (Coe, 1967). The IZZI protocol (Yu et al., 2004) has increasingly been used over recent years as the revised order of the in-field and zero-field steps during the experiment aids in the identification of non-ideal (MD) grains that can bias intensity estimates (section 3.1.4). It currently makes up 11% of Thellier-type entries, but we anticipate it will be used increasingly over coming years. Other Thellier-type variants, such as that of Aitken et al. (1988), MT4 of Leonhardt et al. (2004), and the two specimen approach of Domen (1977), make up only a minor contribution to the database.

Figure 8: Pie charts of the number of entries (in brackets) within GEOMAGIA50.v3.4 associated with different palaeointensity methods. (a) All palaeointensity methods (note that the total number of entries exceeds 5231 as multiple palaeointensity entries were derived from measurements using one or more methods). Thellier method, by type are given in (b): Original Thellier-Thellier method (Thellier and Thellier, 1959), Coe-Thellier (Coe, 1967), Aitken (Aitken et al., 1988), IZZI (Yu et al., 2004), MT4 method (Leonhardt et al., 2004), the two specimen approach of Domen (1977) and other non-specific Thellier-based methods. Note that for three entries two Thellier-type methods were used for the mean intensity given in the entry, therefore the individual mentions of Thellier-type methods totals 4610. Shaw methods include the original procedure (Shaw, 1974) and modified versions by Kono (1978), Rolph and Shaw (1985), Shaw et al. (1995), Tsunakawa and Shaw (1994), Yamamoto et al. (2003). Triaxe method is that of Le Goff and Gallet (2004). Microwave methods are based on versions of the Thellier-type approaches listed above (see, e.g., Hill and Shaw, 1999, 2007). The multispecimen entries include both Dekkers and Bönhel (2006) and Fabian and Leonhardt (2010) approaches. Other methods are the approach of Walton (1977) and the calibrated pseudo-Thellier method (de Groot et al., 2013).

The remaining 13% of palaeointensity estimates were determined by variants of the Shaw



(1974) method (5%), the Triaxe approach (Le Goff and Gallet, 2004) (4%) (derived from a technique proposed by Wilson (1961)), microwave variants of Thellier-type protocols (Shaw et al., 1999; Hill and Shaw, 1999, 2007; Stark et al., 2010) (2%) and the two variants of the multispecimen parallel differential partial TRM (pTRM) method (Dekkers and Böhnel, 2006; Fabian and Leonhardt, 2010) (1%). The calibrated pseudo-Thellier method (de Groot et al., 2013) and the approach of Walton (1977) contribute less than 1% of entries.

### 3.1.3. *Checking and/or correcting for thermal alteration*

As noted above, most archaeointensity data have been obtained using protocols derived from the original Thellier and Thellier (1959) method. Its principle is based on the stepwise thermal demagnetization of the natural remanent magnetization (NRM, assumed to be a TRM) and its progressive replacement by a new TRM acquired in a laboratory field whose direction and intensity are controlled. The ratio between the remaining NRM and the partial TRM acquired after each heating/cooling step, with data usually displayed on an Arai-Nagata diagram (Nagata et al., 1963), allows an estimation of the past geomagnetic field intensity. The comparison of NRM lost to TRM gained requires the magnetic mineralogy of the specimen to remain unchanged during the thermal treatment. In order to assess alteration, Thellier suggested as early as 1946 a partial-TRM check (a pTRM check) (Thellier, 1946). During the stepwise heating-cooling cycle, additional pTRM acquisition steps are added. After a number of heating steps, a lower temperature step is repeated and the pTRMs compared. This is done multiple times throughout the experiment, e.g., after every three heating steps, the first step of the three will be repeated. This alteration test is now common and always required for modern intensity studies using the Thellier method and derivatives (i.e. Coe, 1967; Aitken et al., 1988; Yu et al., 2004). It is important to underline that different approaches have been used to calculate the degree of alteration at each pTRM check (commonly expressed as a percentage) and the associated cut-off values to accept or reject a check or an intensity determination. 44% of Thellier-type intensity entries are accompanied by a pTRM check; however, this number hides the variability in the statistical cutoffs used (see Genevey et al., 2008; Paterson et al., 2014).

Monitoring magnetic susceptibility during heating has been used to check for the stability of the magnetic mineralogy; however, it must be noted that slight changes in susceptibility may not relate to changes in remanence carrying minerals or the formation of new remanence carriers



(rather changes in the susceptibility of magnetic minerals that do not have the capacity to hold or acquire a remanence). This approach was for example used for the part of the Bulgarian data set acquired in the 70s and 80s (Kovacheva et al., 2014). Susceptibility monitoring was only used for 1.5% of intensity entries in the database.

Instead of rejecting samples for which alteration is judged too strong, another possibility is to correct for this effect. This was proposed by Burakov and Nachasova (1985), with a protocol that additionally takes into account anisotropy of TRM. Several sets of data were acquired using this protocol (26 studies spanning 1986 to the present day). This protocol which has not been used in other laboratories is viewed with caution.

For  $\sim 30\%$  of database entries listing the use of a Thellier-type protocol, no alteration test was performed to check or correct for alteration: the linearity of the data points in the Arai-Nagata diagram over a large proportion of the unblocking temperatures was judged sufficient to testify of the absence of this effect. This concerns mainly studies published before the 1990s.

In the Triaxe method (Le Goff and Gallet, 2004) measurements are made continuously in temperature, through successive series of heating and cooling, in zero field or laboratory field. The stability of the magnetic mineralogy is assessed by checking the stability of the ratio between the demagnetized NRM fraction and the acquired TRM fraction at each increasing temperature step. This approach corresponds in a similar way to testing the linearity in an Arai-Nagata diagram, but the steps are spaced only  $5^\circ\text{C}$  apart: the data are therefore numerous (e.g., 60 data for a  $300^\circ\text{C}$  temperature interval) and the linearity is thus finely checked and also assessed through specific linearity tests (Le Goff and Gallet, 2004).

To mitigate the risk of magnetic alteration, alternative methods have been developed. From the oldest to the most recent: the Shaw technique and derivatives (Shaw, 1974; Tsunakawa and Shaw, 1994; Yamamoto et al., 2003), the microwave technique (e.g., Walton et al., 1996; Hill and Shaw, 1999) and the multispecimen protocol and adaptations (Dekkers and Böhnel, 2006; Fabian and Leonhardt, 2010).

Most data obtained with the Shaw technique were acquired between 1975 and 1995 (e.g., Liritzis and Thomas, 1980; Shaw et al., 1995), but the method has seen a revival in recent years (Kitahara et al., 2018, 2020) in the form of the modified Tsunakawa-Shaw approach (Yamamoto et al., 2003). The Shaw method involves only one heating in which the sample is heated above its Curie temperature allowing the acquisition of a full TRM. Prior to heating the NRM is stepwise

demagnetized using increasing alternating field (AF) steps. After heating the sample is again demagnetized using the AF steps as for the NRM. The linear relationship of the demagnetized NRM to TRM is then used to calculate an estimate of palaeointensity. Alteration is assessed through a comparison of coercivity spectra. Changes in an AF demagnetized anhysteretic magnetization (ARM) given before and after heating are compared. Later modifications to the method incorporated corrections to take into account alteration to the pre- and post-heating ARM spectra (Kono, 1978; Rolph and Shaw, 1985).

The microwave method follows the protocols of Thellier and Thellier and modified variants, e.g., the perpendicular single heating method (Hill and Shaw, 2007), but thermal demagnetization is replaced by microwave demagnetization. The rationale is that microwave power should limit the rise in temperature of the sample matrix and reduce the possibility of alteration. However, some conversion to thermal energy to heat the matrix is likely and pTRM-checks test are now integrated in the microwave technique. Recent studies have also included checks for evaluating the cooling rate effect (e.g., Poletti et al., 2013; Ertepinar et al., 2020).

The multispecimen parallel differential partial pTRM method (Dekkers and Böhnell, 2006) started life as essentially a very simple method. Multiple specimens from a site were heated at the same temperature (below the temperature of alteration, but high enough for an appreciable decrease in NRM), but with a different field for each specimen aligned with the specimens NRM. However, shortcomings in the method were evident and the method was expanded upon by Fabian and Leonhardt (2010). It was elaborated upon to correct for differences in the fraction of the pTRM imparted in each specimen, a specimen's domain state, and included a step to monitor alteration.

#### *3.1.4. Checking or correcting for the presence of multi-domain grains*

Another possible factor for the failure of intensity determinations is linked to the presence of MD grains for which the laws of reciprocity and additivity of the partial TRMs are not obeyed (Néel, 1949). Although the influence of MD grains on volcanic palaeointensity estimates has been investigated in detail, it has received less attention in archaeomagnetic studies. This is primarily a result of the different grain size distributions found in archaeomagnetic materials compared with volcanic rocks: archaeomagnetic materials are commonly dominated by pseudo-single domain grains, which are not effected by pTRM tails, whereas volcanic rocks frequently contain a MD

fraction where pTRM tails are significant (where a pTRM-tail results from a non-reciprocity between the blocking and unblocking temperatures). The influence of MD grains can be recognized on Arai-Nagata diagrams as a concave-up curve, whose misinterpretation can lead to underestimates or overestimates of intensity depending on which portion of the curve was used to calculate palaeointensity (e.g., Levi, 1977; Dunlop, 2011). The linearity of the data in the Arai diagram was often considered as a sufficient criterion to, if not exclude, at least consider that the proportion of MD grains is too small to critically affect the intensity determination. The presence of MD grains is now more directly investigated with either rock magnetic measurements, such as hysteresis curves, backfield curves and first order reversal curves (see, e.g., Day et al., 1977; Dunlop, 2002; Roberts et al., 2019), or through additional tests implemented during Thellier-type methods and microwave protocols, such as pTRM-tail checks (aiming at testing the independence of pTRM; Riisager and Riisager, 2001) and additivity checks (Krása et al., 2003). Only 5% of intensity entries in the database list an MD check.

The IZZI protocol (Yu et al., 2004), a variant on the Thellier method, was designed to accentuate the influence of MD tails, evident by pronounced zig-zagging in the Arai-Nagata plot. However, this method is sensitive to the direction of the laboratory field relative to the orientation of the NRM leading to over- or under-estimation of the pTRM-tail and with the field aligned with the direction of the NRM, MD tails can be suppressed.

In comparison to other protocols, the MSP-DSC method of Fabian and Leonhardt (2010) has the advantage to (partially) correct intensities for domain state effect. The Triaxe protocol (Le Goff and Gallet, 2004) mitigates the spurious effect of large grains because the laboratory TRM is almost a full one, mimicking the acquisition of the original TRM. The Shaw derivative of Yamamoto et al. (2003) aims to remove all MD contributions by incorporating a low-temperature demagnetization step after each remanence acquisition and prior to AF demagnetization.

### 3.1.5. TRM anisotropy

An important parameter that may affect intensity determinations when analyzing baked clay artefacts is anisotropy of TRM (already touched upon in section 3.1.1). This anisotropy arises from the stretching of clay during the process of shaping an object, resulting in a preferential alignment of magnetic grains in the clay matrix (e.g., Rogers et al., 1979; Aitken et al., 1981). This effect may be particularly intense for pottery fragments and thin tiles and to a lesser extent to thick

bricks, with biases up to several dozens of micro Tesla (e.g., Genevey et al., 2008; Hervé et al., 2017; Gómez-Paccard et al., 2019). Conversely, it has been observed that this effect is generally less critical when analysing fragments made of clay, which are coarsely assembled, as they are usually taken from in situ structures (e.g., Kovacheva et al., 2009b).

Figure 9: Pie charts of the intensity entries in the GEOMAGIA5.v3.4 database noting (a) remanence anisotropy corrections and (b) cooling rate corrections. Number of uses of an approach are given in brackets. In (a) TRM = thermoremanent magnetization; ARM = anhysteretic remanent magnetization; IRM = isothermal remanent magnetization; NRM  $D_{lab}$  parallel = laboratory field applied parallel to specimen natural remanent magnetization (NRM) direction during the palaeointensity method; other corrections are generally approaches that were insufficiently defined in a publication). See more details in section 3.1.5 and Section 3.1.6,

For 38% of archaeointensity entries in the database remanence anisotropy was not investigated (Fig. 9a). In some cases data were obtained from less anisotropic materials and no measure of anisotropy was pursued. In a small number of entries where anisotropy was estimated, a correction was not necessary. This is most likely as the anisotropy was not considered to be significant. Different approaches have been proposed to evaluate remanence anisotropy. Determination of a TRM anisotropy tensor for each analysed sample allows to evaluate the importance of this effect and to accurately correct the raw intensity determinations (Veitch et al., 1984; Selkin and Tauxe, 2000). This approach was used for 30% of entries considering anisotropy in the database. The drawback of this approach is the time-consuming multiple heating steps (usually six), which increases the risk of mineralogical alteration. Aligning the laboratory field direction with the original NRM (25% of entries considering anisotropy) is an adequate alternative, as long as the degree of anisotropy is not too strong to bias significantly the direction (e.g., Aitken et al., 1981). Ideally, the laboratory field direction should be aligned with the ancient ambient field. This is achieved with the Triaxe protocol and MSP protocols where the direction of the laboratory field is adjusted so a TRM is imparted parallel to the primary TRM (see, Le Goff and Gallet, 2004). To minimize the effect of TRM anisotropy, Morales et al. (2009) proposed to average the intensity values obtained for 6 specimens from the same fragment: here the specimens are oriented in such a way that the TRM is acquired in 6 orthogonal directions relative to a fixed

arbitrary orientation. However, Poletti et al. (2016) and Hervé et al. (2019b) demonstrated that this approach results in larger standard deviations and possibly significant inaccuracies as high as 10-15  $\mu\text{T}$ .

As an alternative to the full determination of the TRM anisotropy tensor, it has been suggested to use other tensors to evaluate and correct for anisotropy; namely tensors of magnetic susceptibility (AMS; 14% of anisotropy assessed entries), anhysteretic remanent magnetization (ARM; 8%) or isothermal remanent magnetization (IRM; <1%). These substitutes are often quicker and easier to implement and avoid the six additional heatings during the thermal protocol. However, the respective ellipsoids significantly differ in their shape and anisotropy degree from TRM ellipsoids (e.g., Chauvin et al., 2000). AMS can underestimate TRM anisotropy by several dozens of percent (Gómez-Paccard et al., 2019). In 22% of entries other types of anisotropy corrections have been applied, but either a method was not listed in the database or the method was not described in the publication.

### 3.1.6. Cooling rate effect

Another possible biasing factor for intensity determinations is the cooling rate dependence of TRM intensity (Fox and Aitken, 1980). Ideally, to avoid such systematic bias, the cooling duration used for the acquisition of the laboratory TRM should be chosen to be identical to the original one when the primary TRM was recorded by the archaeological object. This is rarely possible as the original cooling time is usually long, ranging typically from half a day to a few days (with the notable exception of the slags, Shaar et al., 2010), while the laboratory cooling time is faster, generally from 0.5 up to 2 hours, depending of the type of oven and the size of the specimens.

For Thellier-Thellier data, the cooling rate effect can be evaluated through a comparison of the TRM acquired with a rapid cooling time (the one used routinely during the experiment) and a slow cooling time chosen to be close to the original one (e.g., Chauvin et al., 2000; Leonhardt et al., 2006; Poletti et al., 2013). This is performed for 28% of the intensity entries in the database and comprises 80% of entries that used a cooling rate correction (Fig. 9b). Precisely evaluating the duration of the past cooling is the main difficulty of the correction protocol. Experimental archaeology has provided constraints on this issue (e.g., Morales et al., 2011; Calvo-Rathert et al., 2019; Genevey et al., 2016; Schnepf et al., 2016; Hervé et al., 2019a; Jones et al., 2020).

Archaeological information concerning, e.g., the estimated size of kilns, their morphology, and the type of firing (open or closed), may also help to assess the original cooling conditions. Another approach is to measure the cooling rate effect on TRM acquisition with increasingly slow cooling duration (therefore exploring different conditions of cooling) and to infer from this the error that would be made by under or over estimating the original cooling rate (Genevey et al., 2003; Hartmann et al., 2010).

A different possibility is to apply a fixed correction for all samples from the same archaeomagnetic site, usually 5% or 10%. This “educated guess” concerns 7% of intensity data or 20% of entries which applied a cooling rate correction (Fig. 9b). This approach is based on the assumption that all fragments of the same archaeological object show the same TRM intensity dependence on cooling rate. Experimental studies have, however, pointed out that this effect is variable from one sample to another and (as predicted by theory) that the TRM intensity increases following a logarithmic law as a function of the ratio between an increasingly slow cooling time and a fixed rapid one (Genevey et al., 2008; Hervé et al., 2019a). To avoid applying an educated guess correction to all fragments, it has been suggested to estimate at least for part of the collection the cooling rate effect and to apply an average correction to the other fragments (Kostadinova-Avramova and Jordanova, 2019).

Another important question is at what temperature to estimate the effect of cooling rate. In particular, Hervé et al. (2019a) showed that too high of a temperature could greatly overestimate this effect and therefore underestimate the intensity value. This appears to depend on the magnetic mineralogy of the material analysed (see also Kostadinova-Avramova and Jordanova, 2019).

The cooling rate effect is a challenging parameter to estimate and many studies have not explored this question (over 70% of entries in the database). However, some of these data were obtained with a relatively slow cooling time as part of routine intensity experiments (for example for Bulgarian dataset; Kovacheva et al., 2014): the cooling rate effect is therefore expected to affect them less strongly.

Optimally, we would like to be able to dispense with the question of the cooling rate effect. It has been observed experimentally that the Triaxe protocol accounts for cooling rate (Le Goff and Gallet, 2004; Genevey et al., 2009; Hartmann et al., 2010; Hervé et al., 2017; Salnaia et al., 2017). The multispecimen parallel differential pTRM method also seems to be insensitive to cooling rate (e.g., Schnepf et al., 2016; Calvo-Rathert et al., 2019), possibly because in this technique all

pTRMs are acquired at medium temperatures. However, this question still needs to be further explored (Schnepp et al., 2020a).

### 3.1.7. Intensity uncertainties

On the whole archaeomagnetic data within GEOMAGIA50.v3.4 have reasonably well constrained uncertainties (Fig. 7c). The majority of estimates have an uncertainty of less than 10% of the intensity estimate (60% of intensity entries that report uncertainties), i.e. a few  $\mu\text{T}$  on most measurements. Some intensity measurements, however, have high uncertainties, ranging up to 40  $\mu\text{T}$ . They require careful evaluation prior to their inclusion in reference curves or for field modelling.

A caveat to all intensity uncertainties in the database is that they have been calculated in a variety of ways. Uncertainties may be reported as standard deviations (to 1 or 2  $\sigma$ ), standard errors or they could be weighted. The type of intensity uncertainty is not noted in the database. Care must therefore be taken when using intensity uncertainties when constructing field models and reference curves and using this field as a selection criteria.

## 3.2. Dating methods

Dating and its accuracy and precision are key elements for any archaeomagnetic study. For archaeological artefacts, the dating methods used and listed in GEOMAGIA50 are based on archaeological or historical constraints, or chronometric methods involving mainly radioisotopic and physicochemical measurements (Fig. 10). See Aitken (2014) for an overview of scientific dating methods. We briefly describe the most salient aspects of these methods and their caveats in section 3.2.1.

The archaeological approach remains the most common and concerns almost 60% of the database. Behind the term “archaeological dating” is often hidden the use of a relative chronology, which itself is constrained by elements of absolute dating. The different types of dating methods are clearly complementary and the quality of the two approaches cannot be simply ranked, i.e. scientific dating does not always outrank archaeological observations, it depends on the specific context and an understanding of an archaeological site. The importance of sampling in close collaboration with an archaeologist is paramount for selecting materials whose TRM acquisition can be dated with the maximum precision and confidence. Two categories of methods to date



TRM acquisition are distinguished here, either direct, i.e. directly concerning the analyzed material itself, or indirect, i.e. the material is dated by association with another dating element.

Figure 10: Pie chart of the 8 most commonly used methods to date archaeological materials in GEOMAGIA50.v3.4. Note, the numbers in brackets do not sum to the total number of entries in the database, as numerous entries have been dated using multiple methods. “Other” methods include whether accelerator mass spectrometry was used to obtain radiocarbon ages (133 entries, with frequent overlap with the calibrated and uncalibrated radiocarbon age entries), and if optically stimulated luminescence (OSL; 9 entries) or rehydroxylation (5 entries) were used.

### 3.2.1. Direct dating of TRM

One of the main sources to directly date TRM are document archives. A well-known example is the eruption of Vesuvius first described by Pliny the Younger, which destroyed the city of Herculaneum and Pompeii in 79 CE (Evans, 1991; Evans and Hoye, 2005). But more commonly, these archives are used, for example, to precisely date the edification of religious or civil buildings (e.g., Schnepf et al., 2003; Crote et al., 2015; Salnaia et al., 2017; Genevey et al., 2019) or short periods of activity of ceramics workshops (e.g., Genevey et al., 2009). Other objects, such as some amphoras, can be precisely dated directly through the identification of stamps (Ben-Yosef et al., 2017).

Among the chronometric methods used in archaeology, thermoluminescence (TL) and optically stimulated luminescence (OSL) are directly associated to the TRM acquisition. A firing above 400 °C is time-zero of the method as at above this temperature the electron traps in quartz or feldspar grains in baked clays are emptied (Aitken, 1985). From this moment, traps progressively fill again under irradiation from the surrounding environment (mainly related to  $^{40}\text{K}$ ,  $^{238}\text{U}$ ,  $^{235}\text{U}$  and  $^{232}\text{Th}$  radioactive isotopes). In spite of the advantage of dating the same instance as the TRM acquisition, luminescence methods constitute only  $\sim 3\%$  of entries in the database. However, this method has been used in recent studies (e.g., Gómez-Paccard et al., 2012a; Schnepf et al., 2003; Kondopoulou et al., 2015; Cai et al., 2015; Aidona et al., 2021). Accurate luminescence dating requires a careful reconstitution of the radioactive environment of the baked clay since the last firing. The resulting long measurement time limits the use of the techniques (Roberts et al., 2015). Another caveat of luminescence methods are age uncertainties of  $\pm 5\text{-}10\%$  ( $1\sigma$ ), corresponding



to  $\pm 100$ -200 years for a 0 CE baked clay for example. However, this can be reduced if multiple TL measurements are made. It is worth noting that OSL does not always provide a direct dating of the TRM acquisition because time-zero of this technique can also be the last exposure to sunlight, offering the possibility to date the deposit of sedimentary layers around the studied baked clays.

Another method to directly date baked clay artefacts was proposed by Wilson et al. (2009). It is based on the process of rehydroxylation (RHX) of fired-clay ceramics after production. Similar to luminescence methods, the principle is to start from a zero point by heating a sample up to  $\sim 500^\circ\text{C}$  (dehydroxylation) and then monitor precisely the sample's weight gain in known environmental conditions over several weeks (through rehydroxylation). This allows the kinetics of the rehydroxylation process to be determined. Although promising for archaeologists, and in turn for archaeomagnetists, the relationship between mass gain and time has proved more complex than initially thought, with kinetics that appear to depend on the nature and/or firing conditions of the ceramic (in addition to the environmental conditions), and the applicability of the RHX method appears clearly compromised (Bowen et al., 1971; J. Goff and Gallet, 2014, 2015). So far it has only been applied to two archaeomagnetic studies, both on Spanish ceramics (Nachasova and Burakov, 2012; Burakov and Nachasova, 2013).

### 3.2.2. *Indirect dating of TRM*

As mentioned in the introduction to this section, the archaeological approach remains the most used method of indirect dating. Archaeological dating is however a very generic term that integrates many different elements. The first is stratigraphy, which is essential for building chronologies for ancient multi-layered sites in the Middle East (Shaar et al., 2011; Gallet et al., 2020) and Eastern Europe (e.g., Kostadinova-Avramova et al., 2014). Elements such as coins, fragments of ceramics or metallic artefacts (e.g., swords and fibulae) are also key for dating, if the evolution of their typology is well known. Together, these elements make it possible to define a post quem and ante quem terminus terminus (lower and upper age limit respectively) for an archaeological level or artefact. It is also important to understand whether there has been any mixing of the layers in the stratigraphy, which can limit chronological control. The central question is to precisely understand how the object analysed for archaeomagnetism is reliably related to these chronological constraints. This question is far from trivial, e.g., for settlements occupied over a long period. For intensity determination, one way to overcome this issue is to work directly

on dated pottery fragments, i.e. those whose shape or decoration is recognized and can be linked to a known local/regional typo-chronology.

The relative chronology given by the stratigraphy is fixed to the calendar scale by historical events or chronometric methods. Their precision and reliability are mainly related to the state of the art of archaeological research in the region for a certain period. For example, in Western Europe, precise typo-chronologies are firm for the Roman period (0-500 CE), but are “floating” for the Neolithic period (6000-2000 BCE). These typo-chronologies, and more generally archaeological dating, are also likely to evolve according to the progress of knowledge. This is not a weakness insofar as the archaeomagnetic results remain accurate. However, it is important that dates associated with archaeomagnetic measurements reflect revisions to archaeological ages. For some regions there have been recent revisions to GEOMAGIA50 to accommodate new age information, e.g., Bulgaria (Kovacheva et al., 2014), United Kingdom (Batt et al., 2017), Greece (De Marco et al., 2014), USA (Bowles et al., 2002; Jones et al., 2020) and France (Le Goff et al., 2020). It must also be recognized that there are likely ages within GEOMAGIA50 that do not reflect advances in archaeological age determinations for specific times, regions or sites. Work can continue on sites for years to decades and the archaeomagnetic aspect of the excavation/project may not be the primary objective; new ages may come to light after the final publication of the archaeomagnetic work.

Another common method used to indirectly date archaeological materials is radiocarbon dating. Approximately 15% of entries have used radiocarbon dating as the sole chronological control or in conjunction with other dating methods. Charcoals from carbonaceous or ashy layers that are related to the last use of a kiln/fireplace or located in different horizons of the stratigraphy have frequently been used for dating (e.g., Shaar et al., 2015), but other materials such as seeds and bones have also been used. In comparison to the typo-chronological approach, its advantage is to give a precise date bound by experimentally derived uncertainties. However, the significance of this date relative to the TRM acquisition is not guaranteed. For example, the date can be affected by an old carbon/wood effect. Radiocarbon dates the formation of the organic cell and dating charcoals from reused woods or central tree rings can result in earlier dates up to a few centuries. A preliminary anthracological study is useful to identify such samples and select, if possible, materials with a short lifetime as burnt twigs, grasses or seeds.

A limitation of the method is that the abundance of radiocarbon within a sample can not be

simply related to a specimen's age, through comparison to a decay product, as for example in  $^{40}\text{K}/^{39}\text{Ar}$  dating; nitrogen produced by the decay of  $^{14}\text{C}$  is not captured by the majority of materials (Reimer et al., 2020). Radiocarbon dating is based on measuring the amount of  $^{14}\text{C}$  still present in the sample, but the initial concentration of atmospheric radiocarbon has varied through time and this variation must be accounted for in the calculation of a final radiocarbon age (also known as a calendar age). This process is called calibration and there has been a sustained effort by the radiocarbon community over the past 40 years to develop curves of atmospheric radiocarbon variations that can be used to transfer  $^{14}\text{C}$  ages based on the measurement of radiocarbon present in a specimen, expressed in years Before Present (0 BP = 1950 CE) to an age on a calendar timescale in calibrated BCE/CE. The last versions of calibration curves being (for the Northern Hemisphere; Reimer et al., 2020), SHCal20 (for the Southern Hemisphere; Hogg et al., 2020) and Marine20 (for the oceans; Heaton et al., 2020). As atmospheric radiocarbon variations vary rapidly and non-linearly, this leads to highly variable and complex calibration curves. This in turn results in calibrated radiocarbon ages that have a non-Gaussian error and in some cases result in very broad uncertainties with multiple age ranges. Plateau effects at certain periods result in irreducible date intervals of several centuries, such as 8200-7600 BCE, 4300-4000 BCE, 3400-2900 BCE, 800-400 BCE or the past four centuries.

Finally, we underline that the best way to minimize the risk that the true date of the TRM acquisition is not included in the given interval of age is to combine several chronometric and/or archaeological dates. This is often done by an archaeologist who has an overarching understanding of the site and its positioning in the regional fabric. More recently, mathematical techniques such as Bayesian chronological modelling (e.g., Bronk Ramsey, 2009; Lanos and Philippe, 2018) have brought additional insights into developing archaeological chronologies, especially for sites with complex stratigraphies (e.g., Shaar et al., 2011).

### 3.2.3. Age uncertainties in GEOMAGIA50

Age uncertainties (expressed here as age ranges to accommodate the multimodal age probability distributions of calibrated radiocarbon ages) vary widely within the database, ranging from 0 years for some historically and archaeologically dated entries (1.4% of data) to 2900 years for an archaeologically dated oven from Germany (Schnepp et al., 2020b) (Fig. 11). Approximately 6% of data (627 entries) have an age range  $\leq 10$  years;  $\sim 30\%$  have an age range

$\leq 50$  years; and  $\sim 50\%$  of entries have age ranges of 100 years or less. Nearly all age ranges are less than 500 years ( $\sim 90\%$ ). There are spikes in the age ranges, with ranges of 100, 200, 300, 400 and 500 being more populous than others (Fig. 11b). The majority of these ranges are from archaeological dated materials and are assignments to specific centuries or across multiple centuries. In general, there is no correlation between age and age range. It is important to note that age ranges can be reported at differing precisions (e.g., 1 or 2 standard deviations) and they do not have the same form. For example, some age distributions will follow a normal distribution (e.g., uncalibrated radiocarbon ages and luminescence techniques), some a multimodal distribution (calibrated radiocarbon ages) and others a uniform distribution (e.g., archaeological ages assigned to a specific archaeological period). For a specific age within an age range, this means there will be differing probabilities of this age depending on the dating method used.

Figure 11: Archaeomagnetic age ranges within GEOMAGNETA50.v3.4, binned by (a) 100 year age ranges and (b) 10 year age ranges. (a) the full span of age ranges; (b) truncated to age ranges  $\leq 500$  years.

## 4. Archaeomagnetic field reconstructions

As described in section 2.3 and section 2.4 archaeomagnetic data are inhomogeneous in space and time. Furthermore, little can be garnered about the large scale geomagnetic field from individual data. Regional or global compilations of data are therefore necessary to gain a greater understanding of the temporal and spatial evolution of the field. This section will give an overview of the two main approaches to reconstructing the geomagnetic field on centennial to millennial time scales: regional secular variation curves and global spherical harmonic models.

### 4.1. *Regional secular variation curves*

The potential to combine individual archaeomagnetic data from different locations into composite archaeomagnetic curves for dating purposes was recognized in the 1950s (e.g., Cook and Belshé, 1958; Watanabe, 1958) and a variety of reference curves have been obtained for several parts of the world since then (see Korte et al. (2019) for a detailed review). Because the geomagnetic field cannot be considered purely dipolar, field variations at one location (or in one region) are not representative of the evolution of the field as a whole. Smaller-scale non-dipole

contributions lead to deviations from a dipolar geometry, resulting in variations in direction and intensity that can vary from one region to another. Combining all global archaeomagnetic data into composite curves will not fully capture the evolution of the field and may obscure regional field structures. Therefore it has been common to develop regional archaeomagnetic curves. It is generally assumed that data within a radius of several 100 to a few 1000 km reflect similar field variations and can be combined to form a reference curve for a region (e.g., Tarling, 1989; Tema and Lanos, 2020). A review by Korte et al. (2019) included an investigation of the spatial correlation length of geomagnetic variations and the possible influence of the distance to a curve for dating accuracy. However, strict guidelines cannot be given owing to the complex spatial and temporal evolution of the geomagnetic field over short time scales.

Archaeomagnetic reference curves of field directions, intensity, or all three field components have been developed over a number of decades for several European countries (e.g., Kovacheva et al., 2009a; Tema and Lanos, 2020; Schnepp et al., 2020b,a), Japan (e.g., Watanabe, 1958; Nagata et al., 1963; Kitazawa, 1970; Sakai and Hirooka, 1986), China (e.g., Wei et al., 1982, 1986; Batt et al., 1998; Yang et al., 1993; Shaw et al., 1995), and the United States of America (e.g., Watanabe and Dubois, 1965; Sternberg, 1989a; Hagstrum and Blinman, 2010; Jones et al., 2020) (see Constable and Korte (2015) for a more detailed list with comprehensive references). Several curves have been frequently updated with new data as they become available, e.g., France (e.g., Thellier, 1981; Bucur, 1994; Chauvin et al., 2000; Genevey and Gallet, 2002; Genevey et al., 2009, 2016; Gallet et al., 2002; Hervé et al., 2013a,b; Le Goff et al., 2020). With efforts to improve data coverage for other regions, there are now curves for China (Cai et al., 2017), the Near East (Gallet et al., 2015; Sullinger et al., 2015; Shaar et al., 2020; Livermore et al., 2021), Mexico (Soler Arechalde et al., 2019; Mahgoub et al., 2019) and South America (Goguitchaichvili et al., 2019).

#### *4.1.1. Approaches to curve construction*

The first step to building a reference curve is to relocate the distributed data to a central location (also known as a reference location) to eliminate differences that result in directions or intensity at different locations purely from a dipole field geometry. For directions, this is commonly done by using the conversion-via-pole (CVP) method, whereby a directional pair with one set of geographic coordinates is transformed to a virtual geomagnetic pole (VGP) and the

subsequent VGP is then transformed to a new directional pair using the geographic coordinates selected for the reference curve (Shuey et al., 1970; Noel and Batt, 1990). Alternative approaches have also been used, which assume an axial dipole (a dipole aligned with Earth's rotation axis where the geographic and magnetic pole are coincident) (Aitken and Hawley, 1966; Thellier, 1981), though this method is no longer common. In addition, average curves have been calculated using VGPs and not relocated directions (e.g., in the USA; Sternberg, 1989b; Lengyel and Eighmy, 2002). It is important to note that the non-dipolar nature of the archaeomagnetic field means that all relocation methods have an associated uncertainty (Shuey et al., 1970; Casas and Incoronato, 2007).

For archaeointensities, curves are typically constructed using intensity transformed to either a virtual dipole moment (VDM) or a virtual axial dipole moment (VADM) (e.g., Daly and Goff, 1996; Yang et al., 2000). A VDM is analogous to a VCP, as it uses inclination assuming a tilted dipole. VADM assumes a geocentric axial dipole configuration and allows intensity data lacking inclination to be compared. As with relocated directions, there is an intrinsic uncertainty when calculating a V(A)DM assuming a dipole field configuration when the field can have noticeable non-dipolar components. This can result in a dispersion between sites that is a reflection of non-dipolar field behaviour and not related to issues with how estimates of archaeointensity were obtained.

Different data fitting and smoothing methods have been employed to derive regional secular variation curves. Some early curves relied on hand drawn fits through the data (see, Thellier, 1981; Clark et al., 1989). However, through time increasingly sophisticated mathematical approaches have been used to construct curves, applying methods that not only derive single curves through time (or through inclination and declination), but calculate uncertainties. Simple interpolation of individual field components (or means across time interval bins) with or without the estimation of curve uncertainties has been common (e.g., Sternberg, 1989a; Yang et al., 2000), but over the past thirty years a variety of mathematical approaches have been taken. Methods such as bivariate extensions of Fisher Statistics (Le Goff et al., 1992) continue to be used to produce curves for Europe, e.g., France (Hervé et al., 2013a; Le Goff et al., 2020). Recently bootstrap or Bayesian methods have been used to obtain curves, uncertainty estimates, and/or probability distributions (Thébault and Gallet, 2010; Hellio et al., 2014; Livermore et al., 2018). The Bayesian method of Lanos (2004) and Lanos et al. (2005) is notable as it produces consistent curves for all

three field components and provides curve uncertainties that consider uncertainties on the archaeomagnetic data and ages, and the data distribution. This approach has been used in the creation of a number of archaeomagnetic curves across Europe, e.g., Austria and Germany (Schnepp and Lanos, 2005), and Bulgaria (Kovacheva et al., 2014). The method of Livermore et al. (2018), which is published alongside open source code, also produces curve uncertainties for intensity and the posterior sample age distributions as a direct output.

#### 4.1.2. *Examples of regional field variations*

There are two features noticed in several regions that have received attention in the last decades: archaeomagnetic jerks and intensity spikes. Rapid changes in directional variations seen in Bauer plots (declination against inclination) associated with an increase in intensity in French and Middle East data have been named “archaeomagnetic jerks” (Gallet et al., 2003, 2005, 2009). Gallet et al. (2005) suggested that if archaeomagnetic jerks are global features, they could be associated with episodes of a tilted and enhanced dipole. Later, Gallet et al. (2009) noted that they may correspond to maximum geomagnetic field hemispheric asymmetry, leading to most-eccentric dipole events, related to the dynamics of flux patches at mid- to high-latitudes. If archaeomagnetic jerks are regional, they may result from a recurring non-dipole field structure that influences Western Europe. Using the global field model CALS7k.2 (Korte and Constable, 2005), Dumberry and Bloxham (2006) inferred that archaeomagnetic jerks are associated with a change in the dominant azimuthal flow direction at the top of the outer core below Europe. It is important to note that there is a clear difference in timescales between archaeomagnetic and geomagnetic jerks (e.g., Manda and Olsen, 2009). Archaeomagnetic jerks do not appear unusually rapid compared to what we know from the present field. An archaeomagnetic jerk may last 100–200 years, whereas a geomagnetic jerk lasts  $\sim 1$  year.

In contrast, the intensity variations during geomagnetic intensity spikes during the Iron Age derived from archaeological materials in the Levant (e.g., Ben-Yosef et al., 2009; Shaar et al., 2011), are much faster than field changes observed for recent and historical times. The intensity of the field was also far greater than seen today, exceeding twice today’s field strength (Shaar et al., 2016, 2018) (Fig 12). Livermore et al. (2021) suggest that six intensity spikes are required by the Levant data sets. Increases in intensity were also associated with a directional anomaly (most notably in inclination) and the combined directional-intensity anomaly is referred to as the



Levantine Iron Age Anomaly (Shaar et al., 2018). As with archaeomagnetic jerks, it is currently unclear whether the Iron Age anomaly is regional or global in extent. Data from areas surrounding the Levant (e.g., Georgia, Turkmenistan, Uzbekistan, Cyprus, Greece, Bulgaria and Egypt) indicate there is some evidence of an increase in intensity at the same (or similar) times to the Levantine spikes; however, although intensity reaches twice the present day field's at these locations, the increase spans a broader time (a few hundred years) (Fig. 12). This may be a true representation of the field behaviour during the Iron Age or it may reflect dating inaccuracies.

There is some evidence globally of an intensity increase coincident with the Iron Age anomaly and of spikes at other times (see, Korte and Constable, 2018). Their origin is under discussion (Livermore et al., 2014; Davies and Constable, 2017; Korte and Constable, 2018; Troyano et al., 2020) and difficult to explain given our current knowledge of the geodynamo. To further understand the driving mechanisms that generate high intensity spikes more high-quality archaeomagnetic data from several regions are necessary to fully characterize their regional or global behaviour.

Figure 12: Virtual axial dipole moment (VADM) for the Levant and surrounding regions calculated from data in GEOMAGIA50. 3.4. Note that the data shown in (a) do not reflect the most recent interpretations (compilations) of Shaar et al. (2016), Shaar et al. (2020) and Livermore et al. (2021) as GEOMAGIA50 includes all data and a data selection protocol is not applied. The reader is referred to the above compilations for rationales related to data selection.

#### 4.2. *Global archaeomagnetic field models*

Given its source in Earth's outer core, the geomagnetic field is a global phenomenon and any studies that aim to decipher its driving processes must consider the global evolution of the field. In addition, variations in global field strength, expressed as a dipole moment are also of interest, e.g., in the context of estimating geomagnetic shielding against solar wind, galactic cosmic ray production, atmospheric ionization and solar activity (e.g., Usoskin et al., 2006, 2008, 2010, 2016). A range of global archaeomagnetic field models have been derived over the past decades, from which maps of the field can be generated for Earth's surface (e.g., Fig. 13) and core-mantle boundary (CMB) (e.g., Fig. 14). In the following sections we give an overview of the history of global archaeomagnetic field models, how modelling approaches have evolved over the



past twenty years, and the current state of the art. We discuss how data selection and data uncertainties influence global models and how these have been treated in the most recent models. We describe some of the major findings that global modelling has facilitated, but also note caveats to the modelling approaches.

#### 4.2.1. Dipole moment reconstructions

According to the International Geomagnetic Reference Field (IGRF 13th generation; Alken et al., 2020) the present core field is dominated to about 93% by a dipole centered in the middle of the Earth and tilted with respect to Earth's rotation axis, and to about 91% percent by an axial dipole, i.e. a dipole aligned with the rotation axis. For a purely axially aligned dipole field, the global dipole moment can be determined from a single intensity value and the latitude of the observation. For the moment of a tilted dipole, the inclination at the observation site is additionally required (see, e.g., Merrill et al., 1996). However, when non-dipole field contributions are present, any dipole moment values determined in this way are biased depending on the strength of the local non-dipole field. It is often assumed that non-dipole field contributions average out when enough individual VDMs or VADM are averaged in space and/or time, so that such an average V(A)DM is considered a valid approximation of the actual dipole moment (e.g., Merrill et al., 1996). However, the validity of this assumption is unclear and at least for short intervals or a strongly inhomogeneous global data distribution, the resulting averaged V(A)DM is likely biased.

Several V(A)DM reconstructions from archaeomagnetic data (which also in general include volcanic data) span the past 10 to 50 kyr (McElhinny and Senanayake, 1982; Yang et al., 2000; Genevey et al., 2008; Knudsen et al., 2008; Valet et al., 2008; Usoskin et al., 2016). Genevey et al. (2008) showed that based on the ArcheoInt database, VADM or mixed VADM/VDM curves from Eurasia differ notably from curves for the rest of the world and they constructed global curves for the past 3 kyr using equally weighted regional curves to avoid biasing from a heterogeneous data distribution. Knudsen et al. (2008) used GEOMAGIA50 (version 1) for a VADM reconstruction over the past 50 kyr, in time windows increasing from 500 years for the past 4 kyr to 4 kyr prior to 24 ka, and noted that field strength through the Holocene is higher than during the preceding 40 kyr.

#### 4.2.2. History of archaeomagnetic field models

Although early attempts to construct global archaeomagnetic field models date back to the early 1970s (e.g., Márton, 1970; Braginskiy and Burlatskaya, 1979), they have only received considerable attention over the past 20 years, when the data basis had become large enough to allow for more spatial detail and temporally continuous reconstructions. The recent history of purely archaeomagnetic field models is closely linked to models including palaeomagnetic sediment records in addition to archaeomagnetic and also volcanic data. A surge of models spanning back to 2 to 12 ka followed the publication of Hongre et al. (1998). This includes a series of 100-year snapshot models for the past 3000 years by Constable et al. (2000), and its first continuous equivalent (Korte and Constable, 2003). Several recent reviews include overviews of all these models (Constable and Korte, 2015; Korte and Constable, 2018; Korte et al., 2019), and we focus on archaeomagnetic models (including volcanic data, but no sediment records) in the following discussion.

The first such models were ARCH3k.1 and ARCH3k\_cst.1 (Korte et al., 2009) for the time interval 1000 BCE to 1990 CE. ARCH3k.1 was initially based on all available archaeomagnetic and volcanic data that the authors were aware of (9605 values), with iterative outlier rejection. ARCH3k\_cst.1 was based on a smaller dataset (6211 values) with prior data selection, excluding data with directional uncertainty  $\alpha_{95} > 10^\circ$  intensity uncertainty  $\sigma_{VADM} > 2 \times 10^{22} \text{ Am}^2$ , and age uncertainty  $\sigma_{Age} > 100 \text{ yr}$ . Licht et al. (2013) presented model A\_FM based on 9660 data, spanning 1000 BCE to 2000 CE. Similar to the two ARCH3k models, this was part of a study comparing archaeomagnetic data only models to models including sediment records. A model derived with the main purpose of archaeomagnetic dating is SHA.DIF.14k (Pavón-Carrasco et al., 2014), spanning nearly the past 14 kyr based on 12779 data and following from a series of regional European models by the same group (e.g., Pavón-Carrasco et al., 2009). ARCH10k.1 was derived mainly as a starting model for a reconstruction including sediment records for 8000 BCE to 1990 CE (Constable et al., 2016). A model with somewhat improved Southern Hemisphere data coverage due to recent efforts to improve the global data coverage and a data weighting scheme according to archaeomagnetic quality criteria, named SHAWQ2k, was presented by Campuzano et al. (2019). New modelling methods were explored for models AmR, spanning 1200 BCE to 2000 CE in 40-year snapshots (Sanchez et al., 2016), COV-ARCH, a continuous model for the past 3 kyr (Hellio and Gillet, 2018), BIGMUDI4k.1, an iterative approach simultaneously inverting palaeomagnetic, archaeomagnetic and historical records for the past 4,000 years (Arneitz et al.,

2019), and a proof-of-concept model for the past 1000 years (Mauerberger et al., 2020).

#### 4.2.3. *Range of modelling approaches*

Most global geomagnetic field models, whether covering recent, historical, archaeo- or palaeomagnetic times, are based on series of spherical harmonic (SH) functions that are fit to the data by mathematical inversion techniques. The geomagnetic field is conveniently described by a series of coefficients that scale with field contributions that can be described by a (tilted) dipole, quadrupole, octupole and increasingly shorter wavelength parts. Moreover, when assuming that Earth's mantle is electrically insulating, the SH representation can be downward-continued to provide an image of the field morphology at the top of Earth's outer core, the CMB. For continuous models over certain time intervals the coefficients are smoothly varying time-dependent functions, mostly based on cubic B-splines when constructing historical to millennial scale models (see, e.g., Korte and Constable, 2003; Korte et al., 2009).

As a result of uncertainties in data and age (see section 3) a model cannot and should not fit all data exactly, and some form of smoothing constraint is implemented in the modelling. The simplest form is a truncation of the SH expansion at low degrees to limit the spatial variability of the model and a temporal parameterisation allowing only slow temporal changes. However, most modellers prefer a more flexible form of regularization, where the model parameterisation allows for more variability than expected to be resolved by the data, and the fit to the data is traded off against additional smoothness constraints in space and time.

Methodological differences among most archaeo- and palaeomagnetic SH models mainly lie in the choice and strength of smoothing constraints, and the treatment of outlying data. Helliou and Gillet (2018) in a new approach used statistical information about geomagnetic field evolution from satellite and observatory observations in temporal cross-covariance functions as a constraint in a Bayesian modelling frame. The method results in an ensemble of models with statistically coherent errors on the parameters. Arneitz et al. (2019) also used a Bayesian approach when directly combining archaeomagnetic data with historical observations.

Two recent studies investigate new methods for snapshots in time as a step towards improved continuous models. Sanchez et al. (2016) use statistics from a numerical dynamo simulation as mean and covariance background constraints, which avoids subjective choices of regularization parameters and provides an improved understanding what global spatial resolution

can be retrieved from the data. Mauerberger et al. (2020) implemented a Bayesian non-parametric approach, assuming the geomagnetic potential to be a Gaussian process rather than using SH basis functions. The method provides realistic regional model uncertainties depending on data distribution.

New modelling approaches provide additional relevant information on model resolution and model uncertainties. Models in general agree for regions or parameters that are well constrained by data, as can be seen in the maps of Fig. 13 and Fig. 14, where the different models appear more similar (panels in a) and have smaller uncertainties (panels in b) in the Northern, than the Southern Hemisphere. This can also be seen in time series from Europe and South Africa, where there is a better agreement of the models for Paris (dense data) than South Africa (sparse data) (Fig. 15). Differences in the data basis, outlier treatment, and how uncertainties are weighted have a stronger influence on the models than the method used (Sanchez et al., 2016; Korte and Constable, 2018). When creating a field model (and assessing site dependent output), the underlying data basis should be considered, especially how well a model is constrained for a certain region and time or for a certain purpose. Moreover, all available models are smoothed representations of the actual field variability in both space and time. The amplitudes of rapid field changes in field models are not fully resolved and are likely underestimated.

#### *4.2.4. Influence of data selection and distribution on global models*

Data selection, weighting and distribution have a significant influence on the output of global models. Data selection follows two philosophies. The first is to use all available data without any prior selection, hoping that the signal to noise ratio will increase with the number of available data. The second philosophy is to make a prior data selection by imposing a set of quality criteria. This makes sense when the quality of the data is well understood and the information is available in global databases, e.g., studies from France (Le Goff et al., 2020) or the Levant (Shaar et al., 2016, 2020). However, this is currently not the case in many other regions of the world, where results are sparse and/or many of the results have been obtained decades ago, before some of the modern laboratory methods and tests providing modern quality criteria existed. It is worth noting that Korte et al. (2009) performed a comparison of models with and without prior data selection based on data and dating uncertainties and found no notable improvement when data selection was imposed.

Well distributed global data are the most relevant ingredient for an overall good global model. Recent models providing improved uncertainty estimates (Sanchez et al., 2016; Helliö and Gillet, 2018; Mauerberger et al., 2020) quantify what has been qualitatively stated before (Korte et al., 2009): with the presently available data distribution (more precisely the scarcity of Southern Hemisphere data) archaeomagnetic field models provide limited information about the Southern Hemisphere geomagnetic field.

In Fig. 13 and Fig. 14 we show intensity at Earth's surface and the radial field at the CMB for one snapshot in time. Models based on archaeomagnetic (and volcanic) data (ARCH10k.1, SHAWQ2k, AmR and COV-ARCH) and archaeomagnetic, volcanic and sediment data (CAL10k.2; Constable et al., 2016) and (COV-LAKE; Helliö and Gillet, 2018) can produce models with some broad similarities in intensity at Earth's surface and radial field for the Southern Hemisphere, e.g., lower intensity patches extending across the Indian Ocean and southern Atlantic ocean. However, the precise locations and morphologies of intensity and radial field patches are different. This can be seen in the model COV-LAKE, which incorporates sediment data and includes more Southern Hemisphere data than its counterpart archaeomagnetic model (COV-ARCH); the use of sediment data results in different global intensity and radial field morphologies. Using sediment data in addition to archaeomagnetic data, but applying the same modelling approach, (e.g., COV-ARCH and COV-LAKE), results in reduced uncertainties on the model output for the Southern Hemisphere (Fig. 13b).

Figure 13: (a) Maps of intensity at Earth's surface at 900 CE from six global field models: ARCH10k.1 (Constable et al., 2016), SHAWQ2K (Campuzano et al., 2019), CAL10k.2 (Constable et al., 2016), AmR (Sanchez et al., 2016), and COV-ARCH and COV-LAKE (Helliö and Gillet, 2018). (b) Maps of intensity uncertainty for AmR, COV-ARCH and COV-LAKE.

Figure 14: Maps of (a) the radial field ( $B_r$ ) and (b) its uncertainty for the core-mantle boundary at 900 CE for different field models. See Fig. 13 for model references.

Figure 15: Time series of magnetic declination ( $D$ , top panels), inclination ( $I$ , middle panels) and intensity ( $F$ , bottom panels) for the past 3000 years at two locations as predicted by six different global magnetic field models. All models agree closely most of the time for Paris, where data

coverage is good (a), whereas notable differences exist for South Africa, where there are limited data (b). The included models are the archaeomagnetic models ARCH10k.1 (brown), SHAWQ2k (yellow), AmR (red) and COV-ARCH (black), and the two models additionally including sediment records CALS10k.2 (green) and COV-LAKE (blue). We note that owing to the regularization applied in global modelling, all curves are reduced in temporal resolution in comparison to regional curves, e.g., intensity curves for Paris (Livermore et al., 2018).

The lack of Southern Hemisphere data was explicitly considered by earlier versions of the SHA.DIF.14k model, which were European models based on regional rather than global basis functions (Pavón-Carrasco et al., 2010). Few archaeomagnetic (and volcanic) data in the Southern Hemisphere highlight the importance of using sediment records from the Southern Hemisphere to constrain the field in this region. Several other models (not discussed in detail here) mitigate the problem by including high resolution (mainly lacustrine) sediment records (see Constable and Korte (2015) and Korte et al. (2019) for reviews of these).

#### 4.2.5. *Major findings*

The main applications of purely archaeomagnetic models are for dating purposes and for the calibration of relative intensities obtained from sediments. The advantage of models over regional reference curves for archaeomagnetic dating lies in that models can generate directional and intensity curves for any location without the need for re-location of data. Pavón-Carrasco et al. (2011) presented a convenient Matlab tool to obtain age probability density functions from any combination of declination, inclination and intensity data. Estimated age ranges tend to be smaller if more than one field component is available. A range of published field models and reference curves are implemented in the published version of the tool, and additional ones can be incorporated by the user.

As absolute field strength cannot be retrieved from sediments, archaeological materials and volcanic rocks are the only sources available for obtaining palaeointensity. Based on our current compilation of archaeomagnetic data, both dipole moment reconstructions and global models show that the dipole moment was high around 2 to 3 ka and greater than today's field (e.g., Constable and Korte, 2015). The Holocene maximum seems high compared to the preceding 40 kyr, as noted by Knudsen et al. (2008) and the long-term palaeomagnetic average (Tauxe, 2006;

Yamamoto and Tsunakawa, 2005). Geomagnetic intensity spikes, on the other hand, might be linked to strong dipole moment variations (Korte and Constable, 2018; Hervé et al., 2021), but their origin is not fully understood.

Studies of global field characteristics, such as symmetry (e.g., Constable et al., 2016) or the field morphology at the CMB (Dumberry and Finlay, 2007; Nilsson et al., 2020), with relevance for the theoretical understanding of the geodynamo, are preferably based on models including sediment records, which provide improved data coverage, in particular for the Southern Hemisphere. Asymmetries seen in the modern field have been found to persist over at least 10 kyr: the field is weaker, but more variable on average in the Southern Hemisphere compared with the Northern Hemisphere, and secular variation tends to be stronger in the Atlantic and Indian Oceans compared with the Pacific (Constable et al., 2016). Although the magnetic flux morphology at the CMB changes notably with time, there are preferred or recurrent long-term patterns evident in time-averaged models, in particular nearly symmetrical patches of intense flux at high latitudes in both hemispheres (for more details see, e.g., Amari et al. (2011) and the review by Constable and Korte (2015)). Terra-Nova et al. (2017) more recently found recurring positions, but no preferred direction of motion and some correlation of flux evolution with lower mantle heterogeneities, supporting hypotheses of mantle control on the geodynamo (Bloxham and Gubbins, 1987; Bloxham, 2002). Although both westward and eastward azimuthal flow motions seem to occur in the core over archaeomagnetic times (Dumberry and Finlay, 2007; Wardinski and Korte, 2008), recent studies show a clear dominance of westward drift, with rates between  $0.07^\circ/\text{yr}$  (Nilsson et al., 2014, 2020) and  $0.25^\circ/\text{yr}$  (Hellio and Gillet, 2018; Nilsson et al., 2020).

The South Atlantic Anomaly (SAA) is an area stretching from southern Africa over the Atlantic to South America where the geomagnetic field intensity is notably lower than at comparable latitudes. It is known to have deepened and moved westward from about 1700 onwards from historical data (Mandea et al., 2007; Hartmann et al., 2009). It is linked to the growth of patches of reversed flux at the CMB in the Southern Hemisphere (Gubbins and Bloxham, 1987; Terra-Nova et al., 2016) and has been discussed as a trigger for geomagnetic field reversals (Gubbins and Bloxham, 1987; Tarduno et al., 2015). A unique connection to reversals is unclear as recent modelling of the field for times prior to the Holocene (Brown et al., 2018; Panovska et al., 2019) suggests that features similar to the SAA maybe recurrent and do not necessarily lead to reversals. It is of great interest to know the longevity of the SAA, and whether it



is a recurrent feature of the field, as it may be linked to structures at the CMB that influence core flow and hence geomagnetic field generation (Tarduno et al., 2015; Tarduno, 2018). SHAWQ2k and other models indicate that reverse flux appeared in the Southern Hemisphere as early as 900 CE east of Africa and evolve into the SAA (Campuzano et al., 2019) (Fig. 14).

## 5. Future challenges

Challenges for the future include addressing current inadequacies in the GEOMAGIA50 database, how to improve the temporal and spatial distribution of archaeomagnetic data, and advances in geomagnetic field modelling.

### 5.1. GEOMAGIA50

In previous sections we outlined some of the issues in using data from GEOMAGIA50 for modelling purposes; especially for data selection. One challenge is to homogenize the definition of intensity uncertainties (section 3.1.7) or at least indicate how it was calculated in a new database field. Furthermore, a large number of directional entries are missing  $k$  values and some are missing  $\alpha_{95}$ . All entries lacking these data need to be reassessed and the data added if missing; however, it is likely, especially for  $k$ , that the values were not given in the original publications.

A major deficiency in GEOMAGIA50 is the treatment of chronological metadata and we outline these issues and possible solutions in section 5.1.1. The definition of numbers of samples and specimens also require greater clarification, because they can differ for displaced and in-situ archaeological materials, and for lava flows (section 5.1.2).

Keeping the database up-to-date and useful for the scientific community remains a challenge. Given limited resources, it is not feasible to release a new update of the database when each new archaeomagnetic study is published. Instead over coming years, we intend to release an update at the end of each year containing all the new studies published that year. We note, for example, that the recent studies of Shaar et al. (2020) and Troyano et al. (2020) have yet to be included in version 3.4. of the database, but will be available in the 2021 release.

#### 5.1.1. Archival of chronological data

Section 3.2.3 highlighted how complex the estimation of age uncertainties can be. All

methods have caveats and the reliability of the age information is intimately linked to knowledge of the archaeological context. It is a difficult task to develop a hierarchy of dating methods and deduce/calculate dates, while preserving the complexity of the dating process in the database. This may only be partly achievable in future revisions to the database. It requires more metadata which de facto increases the complexity of the database and its searchability. Ideally, additional fields could be added to GEOMAGIA50 giving more specific age information that could be used to sieve data for model or curve construction. However, owing to the range and complexity of dating methods, this is impractical. An alternative would be to add an accompanying text field to each entry describing the chronological controls, e.g. the cultural name or period and a short description of the dating results (as is given in ArcheoInt database (Genevey et al., 2008)). This approach would not allow automatic data selection and would require manual assessments of the quality of data prior to modelling. Such functionality is not currently available and would require the assessment of all articles in GEOMAGIA50.

Although attempts at incorporating greater radiocarbon information were made in version 2 of the database, there are numerous complications to successful implementation. Radiocarbon dates are currently entered in accordance with information provided in published articles, but not consistently and sometimes ambiguously. Ages may be calibrated or not and calibrated dates can be reported with symmetrical or asymmetrical bounds (reflecting the calibrated age's non-gaussian distribution) and at 68 or 95% confidence. It is not always clear at what level uncertainty is reported.

Calibration can result in significant shifts in age. It is therefore important that calibration is clearly documented. As the database stands, there are still uncalibrated ages that are used for the date of an entry. An initial goal will be to reexamine all the uncalibrated ages and, if there is enough information available, calibrate those ages. As noted in section 3.2, uncalibrated ages are just one of a number of methods that may have been used at a site, e.g., uncalibrated ages may have been combined with archaeological ages or relative stratigraphic ages. In these cases, the dates can not be simply recalibrated, as the radiocarbon ages may form only one aspect of the structural framework for the stratigraphy. This requires that every paper with radiocarbon ages is reexamined in detail to see what scope there is for age re-evaluation.

Documenting calibration of radiocarbon dates is also problematic. GEOMAGIA50 contains studies from the 1960s onwards and numerous improvements have been made in

calibration methods since this time and new calibration curves are published every few years (e.g. Reimer et al., 2013, 2020) (section 3.2.2). Therefore, there are likely some differences in field variations that are not geomagnetic in origin, but rather they relate to changes in calibration curves. Although there is minimal revision for the past 10,000 years between the generations of calibration curves since 1998 (boundaries of the 95% intervals of date are generally modified by 5-10 years at the most), calibrated ages can be shifted up to several centuries for older periods (Reimer et al., 2020). Even if changes are small, radiocarbon ages should be recalibrated to keep them mutually consistent and to remove differences in field variations that stem from non-geomagnetic origins. At a minimum, the experimental radiocarbon ages should be reported accompanied by the calibration method.

The uncertainties on calibrated ages are not treated ideally in the database. Currently only  $\pm$  uncertainties are given. Although this allows asymmetric uncertainties, i.e. the maximum and minimum ages at two standard deviations resulting from calibration, there is no way to record the full multi-modal probability distribution of a calibrated age. This is a significant limitation of the database, as it is important to take into account the irregular shape of the probability density function of calibrated radiocarbon dates in geomagnetic modelling and regional secular variation curves (e.g., Hellio et al., 2014; Lanos, 2004; Hervé and Lanos, 2018; Tema et al., 2017; Yutsis-Akimova et al., 2018a). Systematically storing uncalibrated ages (when available) and updating all radiocarbon ages and uncertainties after calibration is a major undertaking, but is an aim for the future.

### 5.1.2. Clarification of site-sample-specimen hierarchy

Reporting of the number of samples and specimens used for directional and intensity analysis requires evaluation. The database aims to follow the standard palaeomagnetic hierarchy, whereby each entry in the database is considered to be a site, a group of data related to a geological unit or archaeological context that has a unique age. A sample is treated as part of the site that was removed for further analysis. A specimen is a subdivision of a sample and it is this that palaeomagnetic or archaeomagnetic measurements are made on. In palaeomagnetism, this hierarchy works well. A site would be, e.g., a lava flow; a sample, a palaeomagnetic core drilled out of the lava flow; and a specimen, the subdivision of the sample (core) that was measured and the directional or intensity data were obtained from. The number of samples and specimens maybe

similar, e.g., if one specimen was taken from each sample or the number of specimens could be more if multiple specimens were measured from a sample. In this case, the specimen numbers are averaged to give a sample mean and it is always the sample mean and number of samples that are used to calculate the site mean direction.

In archaeointensity studies this may work differently. For example, a piece of pottery may be related to an instance in time, but might not belong to a context with other pieces of pottery of the same age (a context being a site, analogous, e.g., to a lava flow). In this case the piece of pottery could be treated as both a site or a sample. The piece of pottery can be further divided for measurement and these divisions could be considered to be either samples or specimens. If the piece of pottery were treated as if it were to belong to a context with many other pieces of pottery, then it is just one sample of possibly many, therefore the number of samples would be one and there would be multiple specimens. If the piece of pottery is treated as a site, then the number of samples would be multiple and equal to the number of specimens.

In archaeomagnetic directional studies that use “in situ” structures, i.e. ones that have remained in their original position since firing, there is obviously no ambiguity; a site can be equated with the structure it derives from. In this case, and for directional determination, one should emphasize the importance of working on several independently oriented samples for deriving a mean direction.

Both approaches have their advantages, depending on how the data are to be used or from a consistency point of view (rigid use of the site, sample, specimen hierarchy). For example, obtaining multiple measurements of intensity from a piece of pottery can provide an accurate mean intensity determination, valuable for secular variation or curve construction. This would require that the piece of pottery is treated as a site and that the number of specimens is treated as the number of samples. In the database, the number of archaeointensity measurements used for the mean value is therefore treated in the same way as the number of samples from, e.g., a lava flow: they would have equal value. However, this leads to a mismatch in how different archaeological entries may be treated in the database (e.g., an in-situ structure versus a piece of pottery) and between archaeological and palaeomagnetic hierarchies, e.g., a lava flow would no longer be equal to a context, if the piece of pottery becomes the site. Treating a single artefact as belonging to a single context, regardless of whether there are other artefacts maintains the logic of the hierarchy, but may result in data not being included in further analyses if the number of samples is listed as

one, e.g., if the data are filtered by the number of samples.

The fact that there is no common system of hierarchy is a current weakness of the database. In the future, we propose (as in ArcheoInt) that  $N$  corresponds to the number of thermal units, i.e. the number of units that can be considered to have been magnetized at the same time. It is fixed to 1 for a lava flow, an archaeological in-situ structure and a single fragment of pottery, or is higher than 1 for a group of baked clay fragments. A second number,  $n$ , would be the number of individual values from which the average and its uncertainty are calculated. Because the averages are not calculated homogeneously between data entries, either at the sample level or the specimen level, this solution has the disadvantage to mix samples and specimens, but it would clearly make it easier for data selection.

## 5.2. *Improvements in data distribution*

As noted in Section 4.2.4 the most important factor in improving global models and our overall understanding of the global archaeomagnetic field is more data from regions with sparse data distributions. With  $\sim 50\%$  of all data coming from Europe (Section 2.3), there is significant room for improving data coverage. Africa is a clear target given its large area and rich archaeological history. Studies on Burkina Faso, Ivory Coast, Mali, Ethiopia, Kenya, Zimbabwe and South Africa over the past 10 years have made significant steps in improving data coverage. However, an increased emphasis should be placed on developing archaeomagnetic research projects in Africa. Similarly, South America is well positioned to provide useful data. Data from both southern Africa and South America will be key to unraveling the long term evolution of the South Atlantic Anomaly. The Indian subcontinent has a rich archaeological history and yet no archaeomagnetic directional data have been produced. Given its unique position, new directional data could aid in understanding of westward drift of the field across the Indian Ocean. Australia, New Zealand and Pacific islands also have the potential to provide further archaeomagnetic data. This would be especially valuable as it could improve our understanding of field variations in the Southern Hemisphere and the Pacific, which are greatly lacking in data. It would be interesting to investigate/confirm the persistence of lower field variability in the western Pacific (Constable et al., 2016). Finally, the large amount of Japanese directional data not in the database should be added. There are hundreds of entries in this data set and it will be of great interest to see how this influences our understanding of field evolution in eastern Asia.

Acquiring data in regions with few data presents practical challenges. Access to in-situ structures is crucial for full vector studies. However, sampling must take place shortly after excavation and this is not always practical. In most countries archaeomagnetists are limited by several constraints (e.g., travel, time, funding, and export licences). This explains why the spatial distribution of directions is fairly close to palaeomagnetic laboratories, as is the case with Europe. A solution would be to develop local laboratories and/or networks of researchers trained in archaeomagnetic sampling, as well as to collaborate with and train local archaeologists. However, such efforts take time to implement, so we may only see a gradual increase in the amount of data from poorly represented areas.

Another aspect of improving the data distribution is to extend the database further back in time through the Iron Age, Bronze Age and the Neolithic. There is a tendency to focus on more dramatic field changes, e.g., spikes in intensity during the transition from the Iron to Bronze age in the Levant and surrounding areas; however, all times (or “quiet times”) are equally valuable to study, as all field variations relate to the underlying geodynamo process. Furthermore, more detailed descriptions of field evolution through time will allow for the development of more accurate archaeomagnetic dating curves.

A consideration of the types of archaeological materials used to extend our knowledge of field variations to older times is also required. Baked clays are less frequent back through time, as are in situ structures, which are more likely to suffer eventual post-displacements. The baking degree of older clay-based materials is also usually lower, resulting in a less stable mineralogy prone to alteration and therefore less favourable for obtaining archaeointensity results. Increasing the success rate of archaeointensity experiments on such materials is a major challenge that could be overcome in the next decade by new approaches, such as scanning magnetometry and computed tomography (de Groot et al., 2018). Furthermore, a better understanding of the magnetic mineralogy of archaeological materials will aid in this research (e.g., Lopez-Sanchez et al., 2020). To recover intensity variations in the early Holocene or even in the Pleistocene prior to apparition of ceramic production, an alternative to baked clays is the study of heated rock artefacts, such as burnt cherts (Kapper et al., 2014; Zeigen et al., 2019).

Beyond the acquisition of data for older periods and/or for regions that are still poorly documented, an important challenge in archaeomagnetism remains to better understand the issue of data dispersion. Dispersion is characteristic for many data sets and hinders our ability to finely

trace geomagnetic field variations through time. Besides age uncertainty, the sources of dispersion are more numerous for intensity data (undetected alteration, MD effects, uncorrected TRM anisotropy and cooling rate) than for directions. One major issue is the cooling rate correction because its absence can potentially result in a systematic overestimation. One could think to apply a correction factor to uncorrected data. On average, the correction factor seems to be 5-10% (Genevey et al., 2008). However, defining a suitable rate of correction is difficult, because it depends on the specific rock magnetic properties of a specimen and the equipment and protocols used in each laboratory.

Furthermore, what is the precision with which the direction and/or intensity of the geomagnetic field can be retrieved? Does dispersion reflect our current limitations in the acquisition of data? To better constrain these questions, one can ask whether there would be an interest in revisiting regions with a good data coverage and an active archaeological research, in order to acquire new precisely dated data, for example from the past few centuries where chronological constraints can be extremely tight. By limiting the influence of age uncertainties, this could help solve the above issues, which are crucial in using archaeomagnetism for archaeological purposes and for refining our knowledge of the evolution of the geomagnetic field.

### 5.3. *Global geomagnetic field modelling*

From a field modelling point of view, the above-mentioned improvements to the underlying data basis are paramount for improving the temporal resolution and full global spatial reliability of models. Both these aspects are relevant when using field models to infer geodynamo processes in the core, or when using their predictions for regional reference curves. As noted in section 4.2.4, the lack of archaeomagnetic data in the Southern Hemisphere and equatorial areas can be partly compensated for by using sediment data. Similarly, a recently renewed interest in speleothems may produce new high resolution time series for the Holocene in the coming years from locations where it is not possible to obtain archaeological or sediment data. Lascu and Feinberg (2011) give a detailed overview of the potential for speleothems to recover detailed field variations, with the study of Trindade et al. (2018) providing a detailed Holocene record from Brazil and other studies resolving other geomagnetically interesting times in great detail (e.g., Lascu et al., 2016; Chou et al., 2018).

Continuing the efforts to improve the treatment of data and dating errors and translating



them into realistic model errors through methodological developments is also of interest for both these cases. Improved global archaeomagnetic field models may contribute to answering open questions about, e.g., the maximum possible rate of geomagnetic field change and the influences of lowermost mantle structure on the geodynamo (which is reflected in magnetic field morphology). They will also likely contribute to improved predictions of future geomagnetic field evolution by assimilation of data-based models into numerical simulations (e.g., Fournier et al., 2010; Tangborn and Kuang, 2018).

## Acknowledgments

We thank the archaeomagnetic community for providing data when requested, especially Elisabeth Schnepf, Patrick Arneitz and Mary Kovacheva. We acknowledge Fabio Donadini, Kimo Korhonen and Lauri Pesonen for the development and design of GEOMAGIA50. GH was supported by the Labex LsScArBx (Bordeaux Archaeological Sciences Labex, ANR-10-LABX-52). Development of GEOMAGIA50 was supported by DFG SPP-1488. This is Institute for Rock Magnetism contribution 2003.

## Conflict of interest

I declare no conflicts of interests on behalf of all authors.

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### Highlights

- Review of all global archaeomagnetic data;
- Overview of different types of data and their number;
- Discussion on archaeomagnetic and dating methods;
- Overview of geomagnetic field models for the Holocene;
- Future challenges for archaeomagnetism.



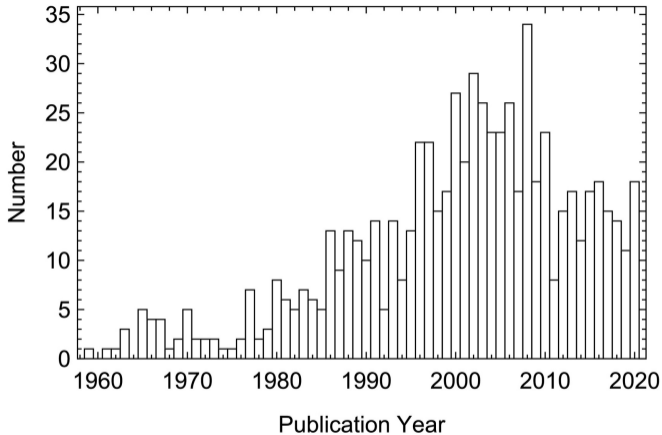


Figure 1

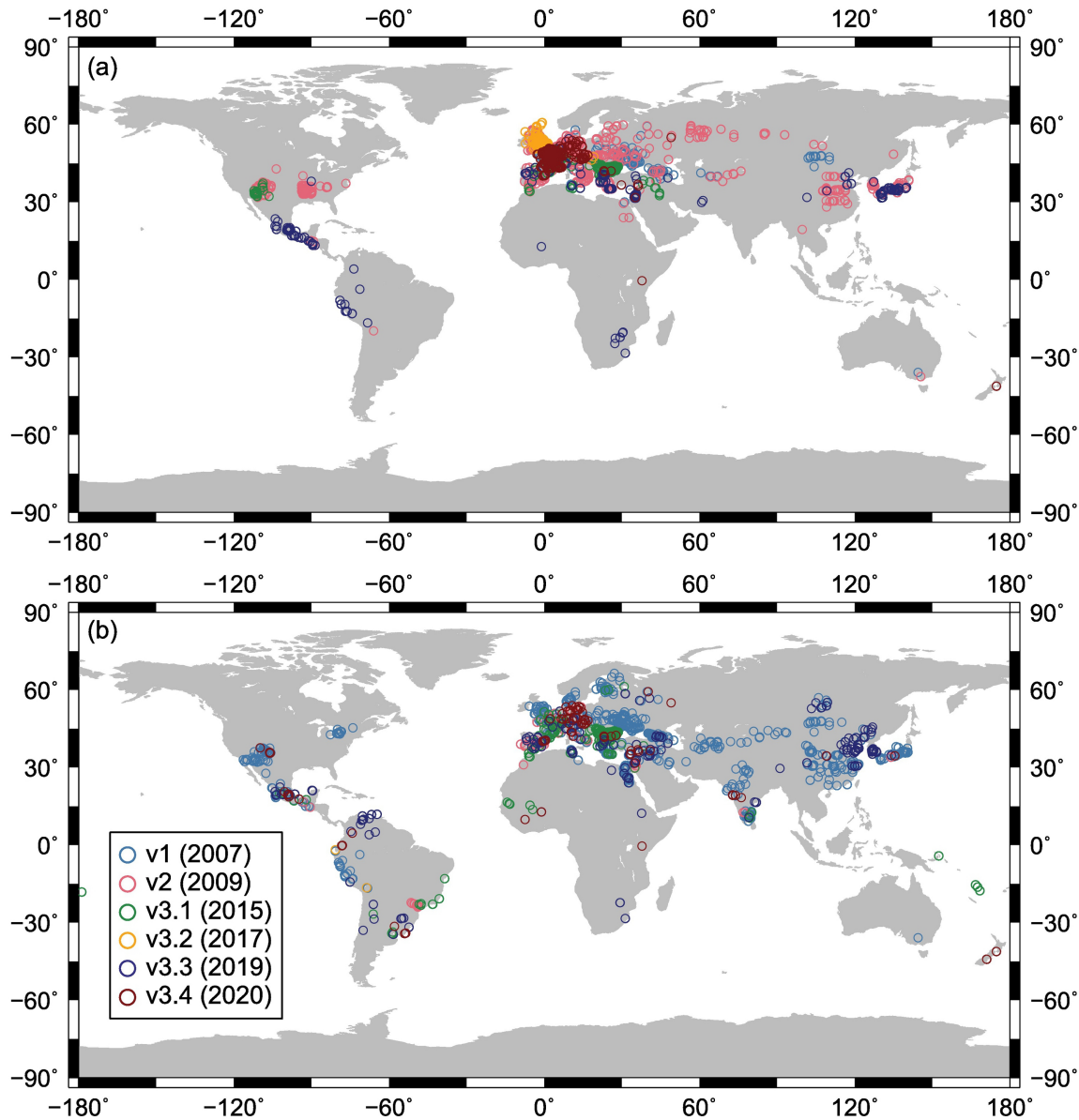


Figure 2

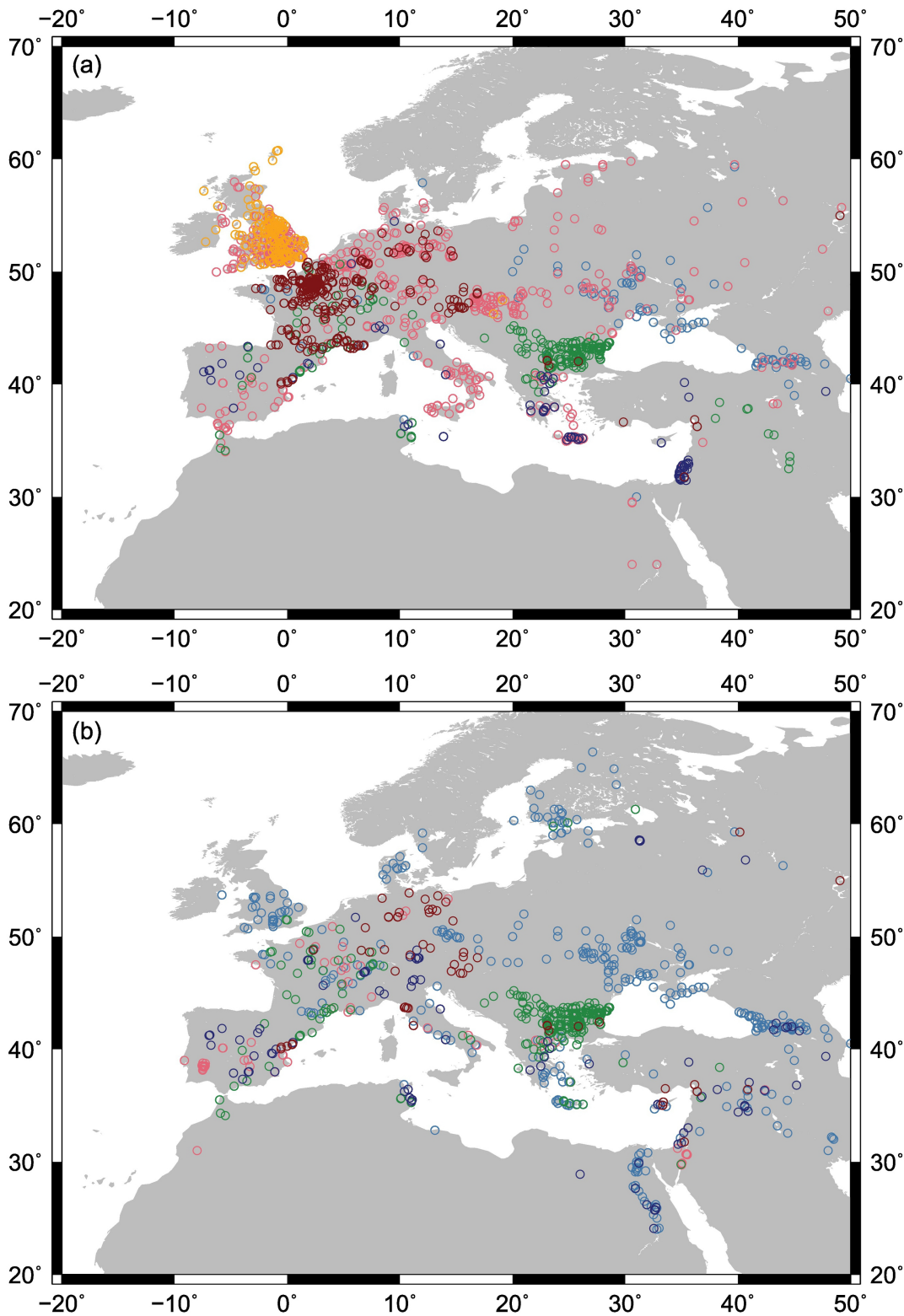
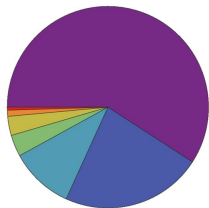
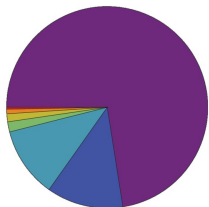


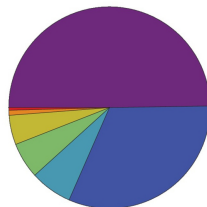
Figure 3

**(a)**

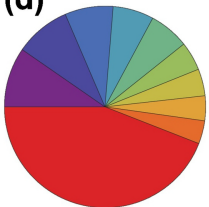
Europe (5905)  
 Asia (2262)  
 North America (1019)  
 South America (344)  
 Africa (305)  
 Central America (98)  
 Oceania/Pacific (48)

**(b)**

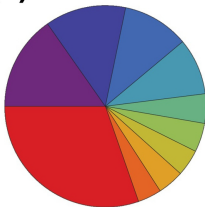
Europe (4301)  
 Asia (729)  
 North America (668)  
 Africa (96)  
 Central America (76)  
 South America (43)  
 Oceania/Pacific (18)

**(c)**

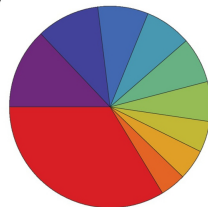
Europe (2603)  
 Asia (1664)  
 North America (352)  
 South America (301)  
 Africa (247)  
 Oceania/Pacific (42)  
 Central America (22)

**(d)**

UK (961)  
 France (890)  
 Russia (774)  
 USA (656)  
 China (642)  
 Georgia (461)  
 Greece (431)  
 Spain (398)  
 Japan (370\*)  
 Other Countries (4398)

**(e)**

UK (905)  
 France (770)  
 Russia (634)  
 USA (538)  
 Ukraine (280)  
 Bulgaria (275)  
 Hungary (253)  
 Japan (250\*)  
 Germany (226)  
 Other Countries (1800)

**(f)**

China (547)  
 Georgia (428)  
 Greece (343)  
 Bulgaria (316)  
 Spain (310)  
 Russia (270)  
 Mexico (210)  
 Iraq (207)  
 Ukraine (174)  
 Other Countries (2426)

Figure 4

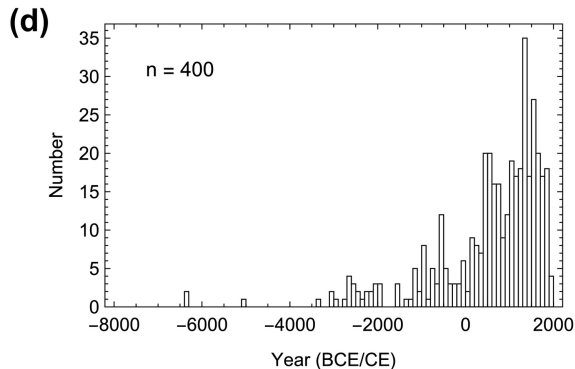
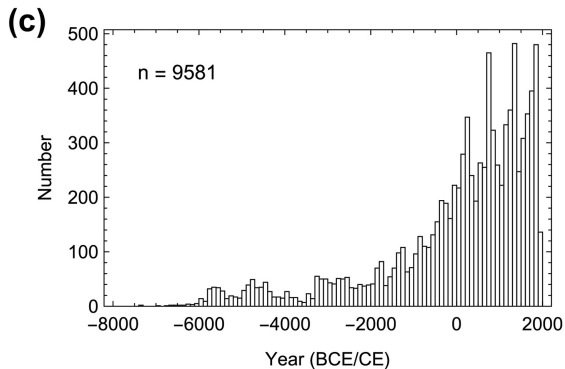
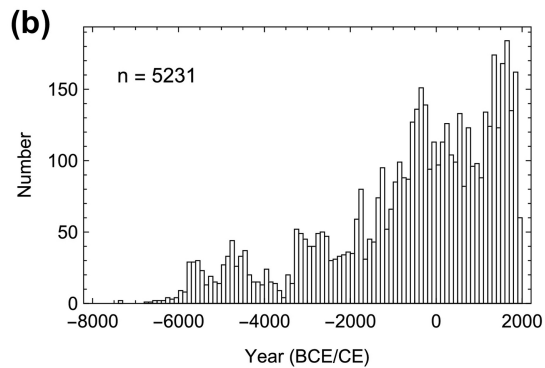
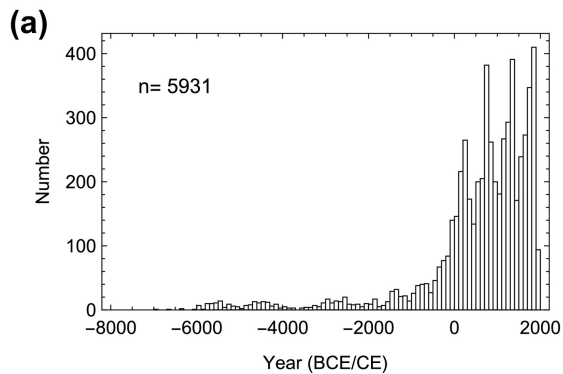


Figure 5

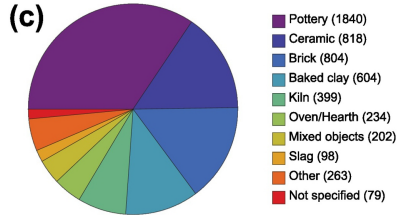
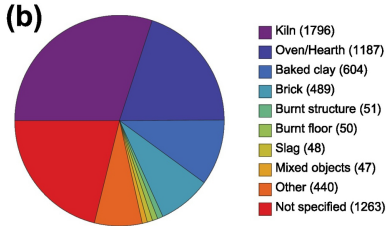
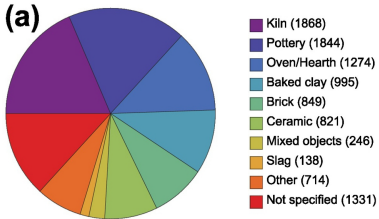


Figure 6

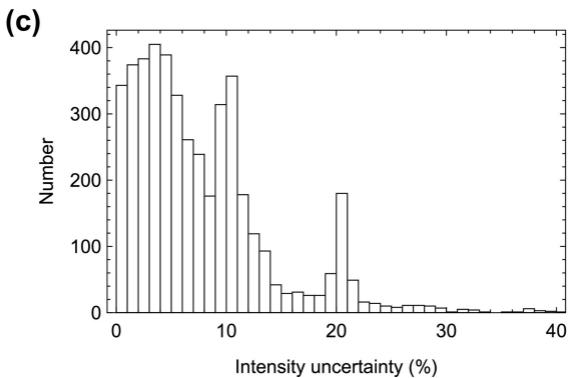
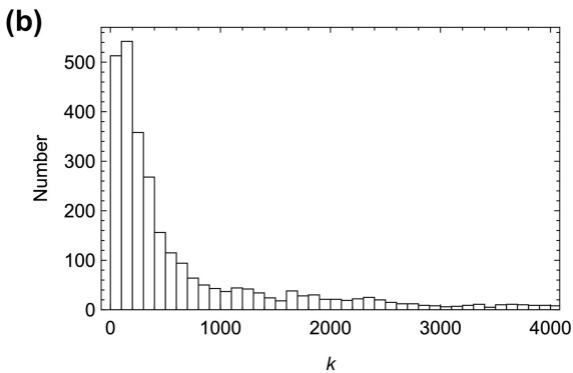
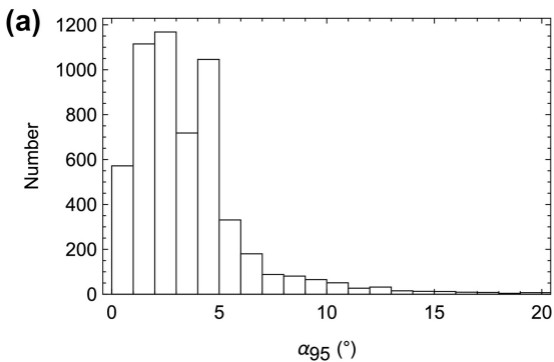


Figure 7



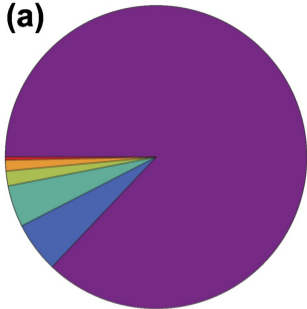
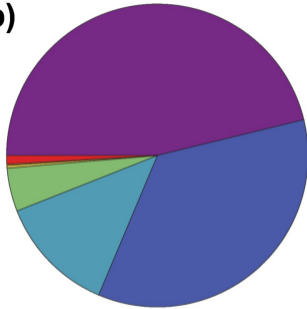
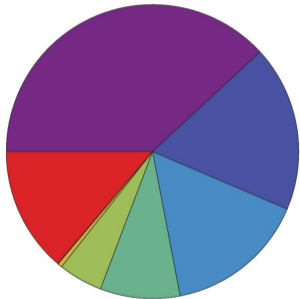
**(a)****(b)**

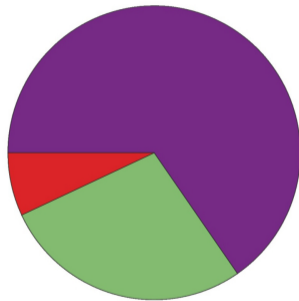
Figure 8

**(a)**



- No correction (2046)
- TRM tensor (979)
- NRM  $B_{lab}$  parallel (836)
- Susceptibility tensor (473)
- ARM tensor (262)
- IRM tensor (26)
- Correction not applied (4)
- Other correction (738)

**(b)**



- No correction (3418)
- Experiment (1448)
- Guess (364)

Figure 9

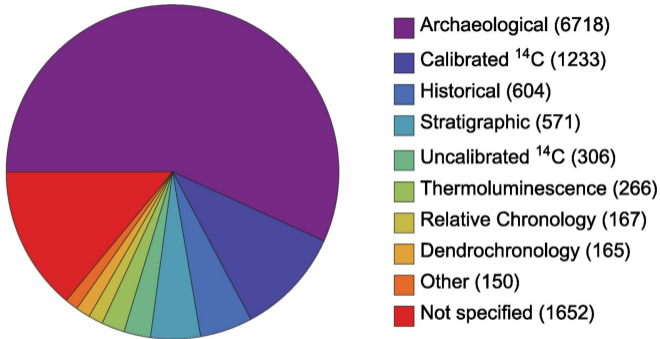


Figure 10

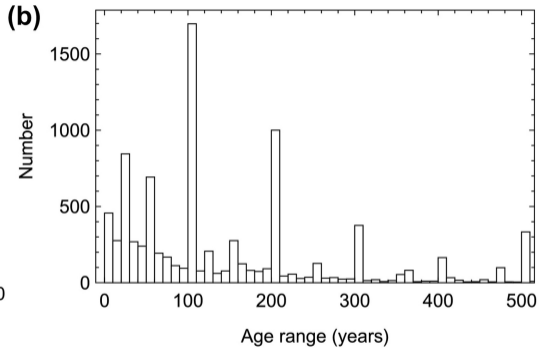
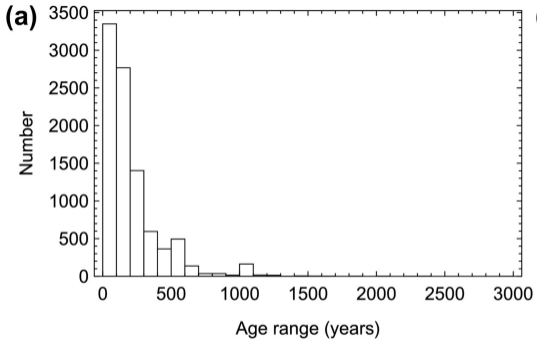


Figure 11

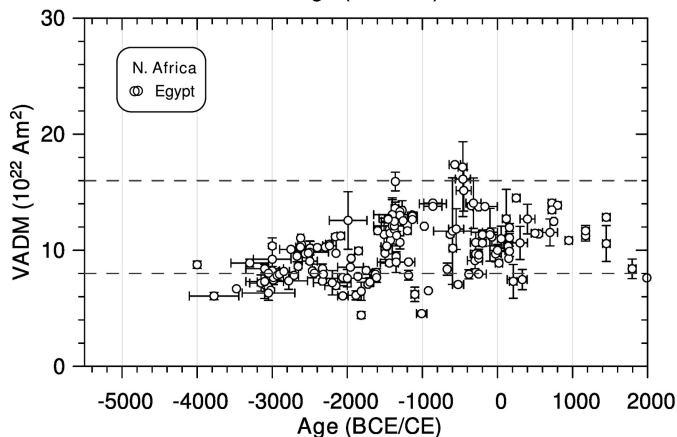
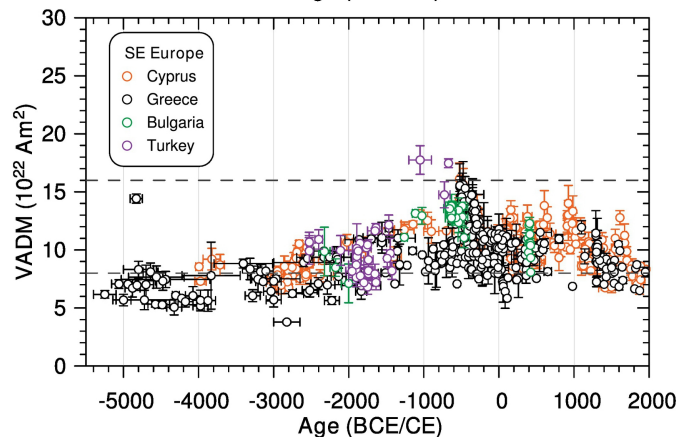
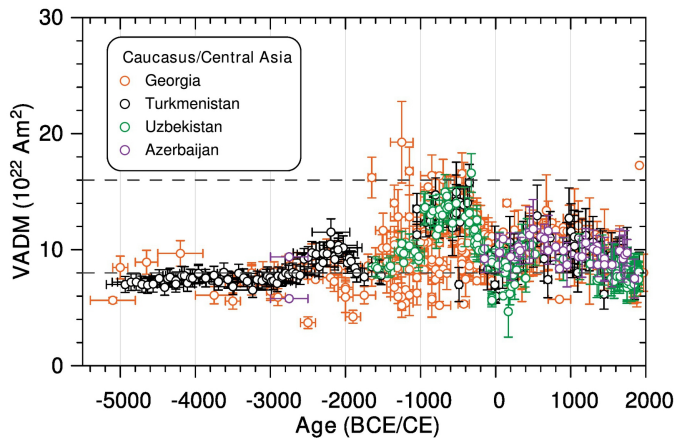
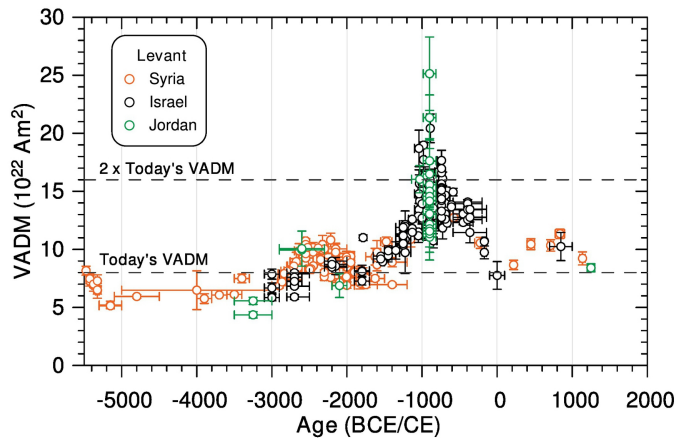


Figure 12

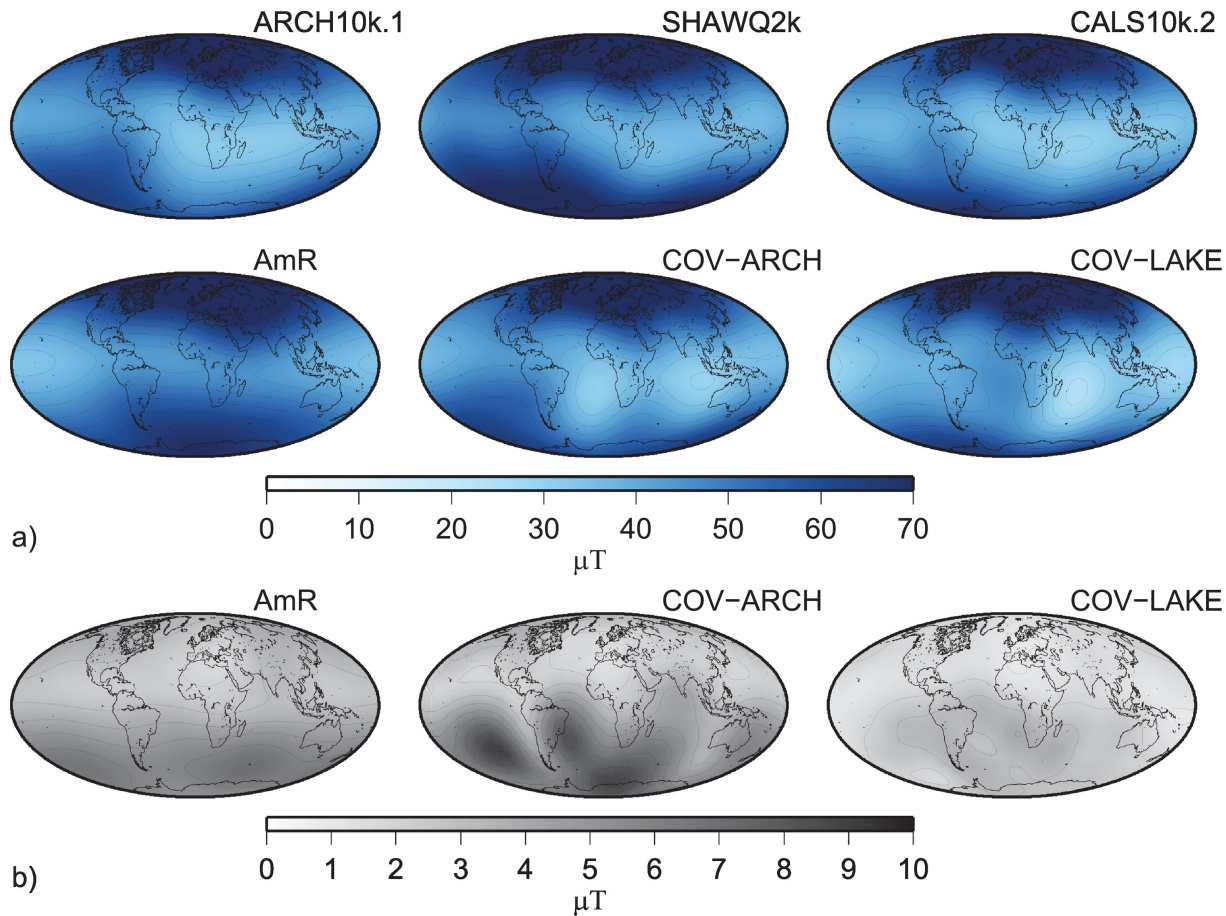


Figure 13

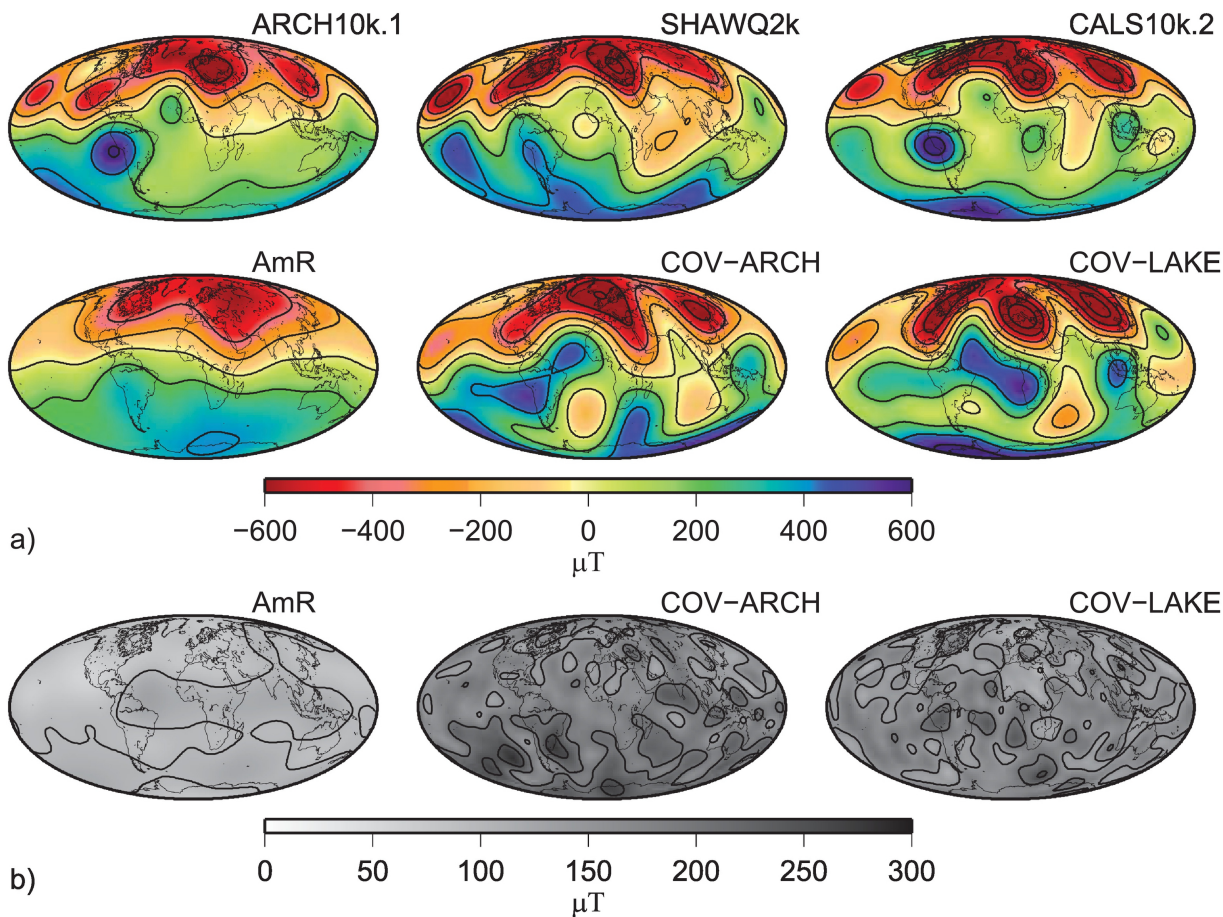


Figure 14



Paris (48.86 N, 2.35 E)

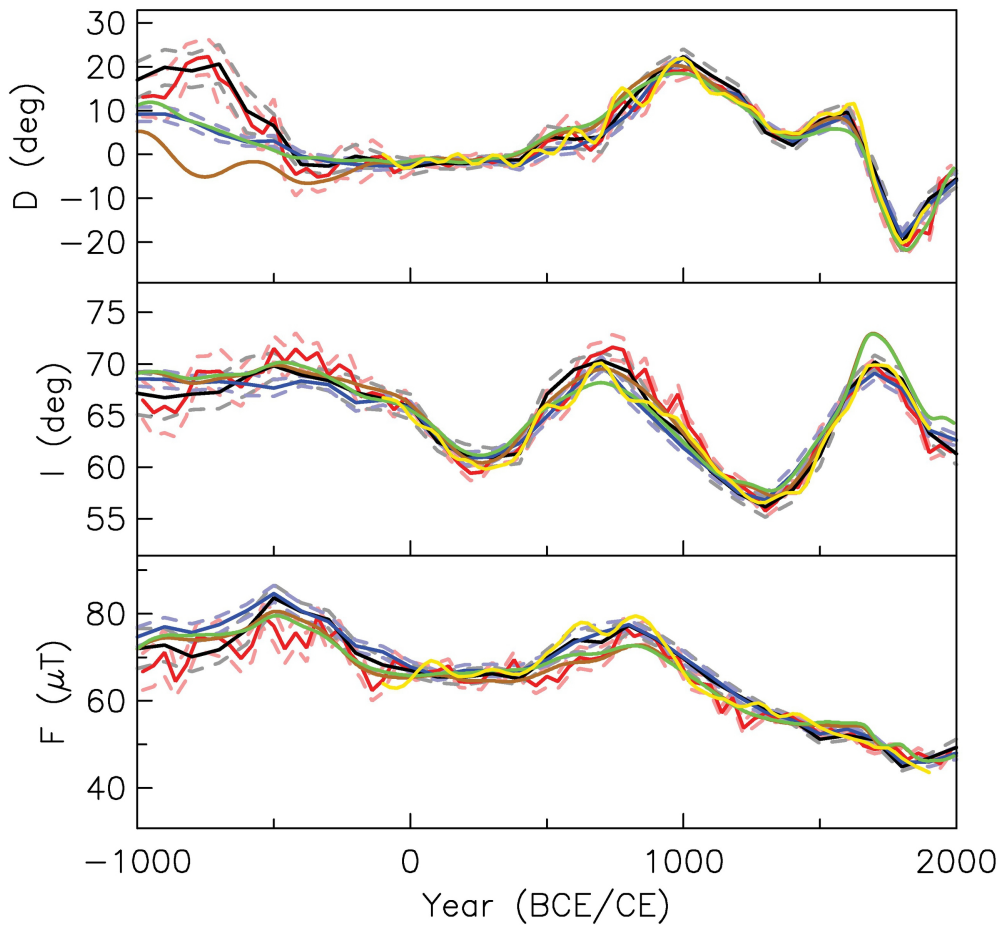


Figure 15