

# Exploring the stellar properties of M dwarfs with high-resolution spectroscopy from the optical to the near-infrared (Corrigendum)

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References to the figures in the text are not correct. The corrected figure references along with the respective texts of the original paper are as follows:

– Caption of Fig. 5: References to right and left panels are inverted. The corrected figure caption is:

*Left panel:* comparison of  $T_{\text{eff}}$  versus spectral type relation as a function of metallicity from this study with [Passegger et al. \(2018; open triangle\)](#) and with [Rajpurohit et al. \(2013; open circle\)](#). *Right panel:* comparison of  $T_{\text{eff}}$  determined from this study with [Passegger et al. \(2018; open triangle\)](#) and with [Gaidos & Mann \(2014; filled square\)](#) as function of metallicity.

– Corrected text (page 5, right column): In order to explore the effect of stellar properties of M dwarfs, we have compared our results with the behavior of  $T_{\text{eff}}$ , as a function of spectral type with [Rajpurohit et al. \(2013\)](#) as shown in Fig. 5 (left panel).

– Corrected text (page 5, right column): In order to explore further, we compared  $T_{\text{eff}}$  determination with [Gaidos & Mann \(2014; hereafter G14\)](#) along with [Passegger et al. \(2018\)](#), for the common stars in our sample as shown in Fig 5 (right panel).

– Corrected text (page 6, left column): [Passegger et al. \(2018\)](#) compare the observed spectra with the synthetic spectra computed using the PHOENIX-ACES models, which are based on the PHOENIX version by [Husser et al. \(2013\)](#), and the equation of state by [Barman et al. \(2001\)](#), as shown in Fig 5 (right panel).

– Corrected text (page 6, right column): Figure 6 (left panel) shows the comparison of  $\log g$  with [Passegger et al. \(2018\)](#) indicating difference of around 0.2 to 0.6 dex with our values.

– Corrected text (page 7, left column): Figure 6 (right panel) shows the comparison of the  $[M/H]$  derived in this study with [Passegger et al. \(2018\)](#) and G14, indicating difference of around 0.2 to 0.4 dex with our values.

## References

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