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# Poisson brackets and truncations in nonlinear reduced fluid models for plasmas 

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#### Abstract

The Hamiltonian structure for an infinite class of nonlinear reduced fluid models, derived from a Hamiltonian drift-kinetic system, is explicitly provided in terms of the $N+1$ fluid moments evolving in each model of the class, with $N$ an arbitrary positive integer. This improves previous results, in which the existence of the Hamiltonian structure was shown, but the complete explicit expression for the Poisson bracket of each model of the class was not provided. We also show that, whereas the Hamiltonian functional of the fluid models can be derived from that of the drift-kinetic system, by projecting the perturbation of the distribution function onto its truncated series in terms of Hermite polynomials, this is not the case for the Poisson bracket. Indeed, the antisymmetric bilinear form obtained by means of the aforementioned projection, although, interestingly, "very similar" to the Poisson bracket of the fluid models, turns out to differ from it. The difference is found to reside in the coefficients $\mathbb{W}_{(N) l}^{m n}$ of the bilinear form, when the indices are such that $l+m+n$ is even and $l \geqslant N+1, m \geqslant N+1, n \geqslant N+1$. We show with a counterexample, related to the case $N=2$, that such bilinear form, in general, does not satisfy the Jacobi identity. We provide a physical interpretation of the set of variables $G_{0}, G_{1}, \cdots, G_{N}$, in terms of which the Poisson bracket of the fluid models exhibits a direct-sum structure, and point out an analogy between the present fluid reduction problem and the problem of the truncated quantum harmonic oscillator.


## 1. Introduction

In many circumstances, the behaviour of plasmas is influenced by the presence of a magnetic field characterized by one dominant component, nearly constant in time, which is referred to as guide field component. This can be the 5 case, for instance, in tokamaks or in coronal loops, where the toroidal component of the field is much greater than the poloidal component. In these situations, a small parameter naturally emerges, corresponding to the ratio between the characteristic amplitude of the time-dependent components of the magnetic field and the strong guide field component. Taking advantage of this small parameter, several nonlinear reduced fluid models have been derived in order to describe plasma dynamics in the presence of a strong guide field. In the context of tokamak fusion plasmas, two classical examples of such reduced models are provided by reduced magnetohydrodynamics (MHD) [1, 2] and by the Hasegawa-Wakatani model for drift-wave turbulence [3]. Reduced fluid models of the same kind also proved to be useful for investigating fundamental aspects of turbulence relevant for the solar wind, as for instance in Refs. [4, 5. Further applications of reduced fluid models include the description of nonlinear coherent structures in plasmas [6, 7, 8] and magnetic reconnection [9, 10, 11, 12].

In the plasma physics literature, such nonlinear reduced fluid models were typically derived either from two-fluid models by asymptotic expansion in terms of small parameters (see, e.g. Refs. [1, 2, 13, 14, 15), or by taking moments of gyrokinetic or drift-kinetic equations and imposing a closure relation (see, e.g. Refs. [16, 4, 17]).

From a dynamical systems perspective, a number of such models, when considered in their non-dissipative limit, were shown to possess a noncanonical Hamiltonian structure 18, 19, 15, 12, 20, 21, 22, 23, 24, 25, 26, 27, 28, (see also Ref. [29] for a review), which is the typical case for fluid models formulated from the Eulerian point of view [30]. The existence of a Hamiltonian structure is crucial for avoiding the presence of fake dissipative terms in the model, as well as for the opportunity it gives, to apply methods of Hamiltonian mechanics for the
analysis of the dynamics described by the model 30, 31. Qualitatively speaking, a Hamiltonian structure is identified by a phase space, a Poisson bracket acting on functions defined on the phase space, and a Hamiltonian, which is a prescribed function on the phase space. For the reduced fluid models under

35 consideration, the phase space is a space of functions defined on the domain occupied by the plasma and satisfying appropriate boundary conditions. In the two-dimensional (2D) limit, where the dynamics is assumed to be invariant along the direction of the guide field, the noncanonical Poisson brackets for the reduced models are generally extensions [20 of the classical Lie-Poisson bracket of the 2D Euler equation for an incompressible fluid. The Poisson bracket for the full 3 D models consists of the sum of the 2 D bracket with a second Poisson bracket, such that some relations between the coefficients of the two brackets are satisfied, which guarantees the Jacobi identity [23. The Hamiltonian of the models, on the other hand, typically consists of a functional on the phase space, ${ }_{45}$ quadratic in the model field variables.

In Ref. 32] it was shown how an infinite class of Hamiltonian reduced fluid models can be obtained, by imposing a particular closure on the hierarchy of fluid equations obtained by taking moments of a drift-kinetic system. In Ref. [33], this result was extended to more general models accounting also for finite ing to the eigenvalues of an explicitly given, symmetric matrix denoted as $W$ (according to the notation of Ref. [32). The expression of the Poisson bracket also depends on an orthogonal matrix $U$, thanks to which, the matrix $W$ can be put in diagonal form, according to the relation $U^{T} W U=\operatorname{diag}\left(\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}\right)$, where $U^{T}$ is the transpose of $U$ and $\operatorname{diag}\left(\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}\right)$ is the diagonal matrix with elements $\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}$ on the main diagonal. The knowledge of the matrix
$U$ and of the eigenvalues of $W$ is also necessary to cast the Poisson bracket in its simplest form (putting in evidence its direct sum structure) which occurs when the bracket is expressed in terms of a set of variables $G_{0}, G_{1}, \cdots, G_{N}$, alternative to the fluid moments. Although the existence of such an orthogonal matrix $U$ is guaranteed by the spectral theorem, its explicit expression, for arbitrary $N$, was not provided in Refs. [32, 33]. Likewise, the properties of the eigenvalues of $W$ of being real is guaranteed by the spectral theorem, but an explicit formula for their expression was not given. Therefore, although the Hamiltonian struc70 ture for the class of fluid models under consideration was shown to exist, the actual expression for the Poisson bracket, for a given model, had to be found by determining, case by case, the matrix $U$ and the eigenvalues $\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}$. Of course, this deficiency limited the immediate applicability of the results of Refs. [32, 33, if not for the cases with very low $N$, where only the evolution 75 of the first few fluid moments is retained. We recall 32 that the matrix $U$ actually provides the transformation from the variables $G_{0}, G_{1}, \cdots, G_{N}$ to the fluid moments. Indeed, if we indicate the fluid moments with $g_{0}, g_{1}, \cdots, g_{N}$, such transformation is given by $g_{m}=U_{m n} G_{n}$ (see also Eq. 90) in the present paper, where the notation has been changed with respect to Ref. [32] and the matrix $U$ is indicated with $U_{N}$ ).

One of the purposes of the present paper is to remedy the above mentioned gap, by providing the explicit expression for the Poisson bracket for any model belonging to the class treated in Refs. [32, 33]. This is made possible by making use of actually rather elementary properties of Jacobi matrices and Hermite polynomials.

A second objective of this paper concerns the relation between the Hamiltonian structure of the reduced fluid models and that of the parent drift-kinetic (or gyro-kinetic) model. Indeed, the parent model describes the evolution of the generalized perturbed distribution function $g(x, y, z, v, t)$, given by a linear combination of the perturbation of the gyrocenter distribution function with the component of the magnetic vector potential along the guide field (see Eq. (5)). The function $g$ depends on the spatial coordinates $x, y, z$, as well as on time
$t$ and on the component $v$ of the velocity along the guide field. We consider generalized perturbed distribution functions that can be expanded in series as

$$
\begin{equation*}
g(x, y, z, v, t)=\sum_{n=0}^{+\infty} g_{n}(x, y, z, t) \frac{H_{n}(v)}{\sqrt{n!}} F_{e q}(v) \tag{1}
\end{equation*}
$$

where $F_{e q}$ is a Maxwellian distribution function, $H_{n}$ is the Hermite polynomial of order $n$ and the coefficients $g_{n}$ correspond to the fluid moments. The evolution of $g$ is then governed by the infinite system of evolution equations for the moments $g_{n}$, obtained by projecting the Hamiltonian drift-kinetic (or gyro- kinetic) equation for $g$ on the basis of Hermite polyomials. The $(N+1)$-moment Hamiltonian reduced fluid models, on the other hand, are obtained by imposing that the index $n$ in the series (1) goes from 0 to $N<+\infty$, which truncates the infinite system of equations of the parent model, reducing it to a closed set of $N+1$ equations. The natural question then arises, about whether the Hamiltonian and the Poisson bracket of a reduced fluid model can be derived from the Hamiltonian and the Poisson bracket of the parent model, by replacing, in the latter, the perturbation of the distribution function, given by the infinite series (11), with the series truncated at the moment of order $N$. The analysis we describe in the present paper shows that, whereas this occurs for the Hamiltonian, it is not the case for the Poisson bracket. In particular, it is the 2D component of the Poisson bracket for the fluid model, which turns out to differ from the corresponding bilinear form, obtained from the Poisson bracket of the parent model upon replacing the perturbed distribution function with its truncated series. Roughly speaking, the Poisson bracket and the bilinear form turn out to be "very similar" and the difference concerns some coefficients which, in a sense that will be made precise later, are associated with "high-order" moments, among those retained in each fluid model. Moreover, by means of a counterexample, we show that, in general, the bilinear form obtained from the truncated series, does not satisfy the Jacobi identity, and thus is not a Poisson bracket.

With the present paper, we also provide a physical interpretation of the alternative variables $G_{0}, G_{1}, \cdots, G_{N}$. Indeed, making use of properties of Hermite polynomials, we show that such variables are proportional to what we refer to
as the truncated generalized perturbed distribution function, evaluated at values of the parallel velocity coordinate equal to $\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}$. This might also boundary conditions requires some restrictions on the set of observables on the phase space.

The paper is organized as follows. In Sec. 2 we introduce the parent driftkinetic model. Section 3 describes the Hamiltonian structure of the parent drift-kinetic model and of the reduced fluid models obtained after imposing a Hamiltonian closure provided in Ref. [32]. The latter structure is expressed in terms of the variables $G_{0}, G_{1}, \cdots, G_{N}$. This Section essentially reviews already known results but formulates them in a more precise manner, with respect
to Ref. [32. In Sec. 4 we present one of our new results, consisting of the Lemma and of a Proposition, respectively, formulated in Sec. 5 .

## 2. Hamiltonian parent drift-kinetic model

We consider the following drift-kinetic model in normalized form

$$
\begin{align*}
& \frac{\partial g}{\partial t}+[\phi-v A, g]+v \frac{\partial}{\partial z}\left(g-\sqrt{\frac{2}{\beta_{e}}} F_{e q}(\phi-v A)\right)=0  \tag{2}\\
& \Delta_{\perp} \phi=\delta^{2} \sqrt{\frac{2}{\beta_{e}}} \int_{-\infty}^{+\infty} d v g  \tag{3}\\
& \Delta_{\perp} A-A=\sqrt{\frac{\beta_{e}}{2}} \int_{-\infty}^{+\infty} d v v g \tag{4}
\end{align*}
$$

where Eq. (2) corresponds to the electron drift kinetic equation, whereas Eqs. (3) and (4) correspond to the quasi-neutrality relation and to the projection of Ampère's law along the direction of a magnetic guide field, respectively.

In Eqs. (22)-(4), the dynamical variable $g$ is defined by

$$
\begin{equation*}
g(x, y, z, v, t)=f(x, y, z, v, t)-\sqrt{\frac{2}{\beta_{e}}} v F_{e q}(v) A(x, y, z, t) \tag{5}
\end{equation*}
$$

where $f$ is the actual perturbation of the electron gyrocenter distribution function, averaged with respect to the magnetic moment. We will refer to the
function $g$ as to the generalized perturbed distribution function. The field $A$ is related to the normalized magnetic field $\mathbf{B}$ by

$$
\begin{equation*}
\mathbf{B}(x, y, z, t)=\nabla A(x, y, z, t) \times \hat{z}+\hat{z} \tag{6}
\end{equation*}
$$

where $\hat{z}$ is the unit vector along the $z$ direction of a Cartesian coordinate system $x, y, z$. We note that the model assumes the presence of a strong uniform magnetic guide field along the $z$ direction (corresponding to the second term on the right-hand side of Eq. (6)). The function $F_{e q}$ is the Maxwellian equilibrium distribution function, whose explicit expression reads

$$
\begin{equation*}
F_{e q}(v)=\frac{1}{\sqrt{2 \pi}} \mathrm{e}^{-\frac{v^{2}}{2}} \tag{7}
\end{equation*}
$$

The field $\phi=\phi(x, y, z, t)$, on the other hand, corresponds to the electrostatic potential. The independent variables in Eqs. (2)-(4) are given by the spatial Cartesian coordinates $x, y$, and $z$, by the coordinate $v$, representing the velocity coordinate along the direction of the guide field, and by the time $t$. The spatial coordinates belong to the domain

$$
\begin{equation*}
\mathcal{D}=\left\{(x, y, z) \in \mathbb{R}^{3} \mid-L_{x} \leqslant x \leqslant L_{x},-L_{y} \leqslant y \leqslant L_{y},-L_{z} \leqslant z \leqslant L_{z}\right\} \tag{8}
\end{equation*}
$$

with $L_{x}, L_{y}$ and $L_{z}$ positive real numbers. On the other hand, for the parallel velocity and time coordinates one has $-\infty<v<+\infty$ and $t \geqslant 0$, respectively. Periodic boundary conditions over the domain $\mathcal{D}$ are imposed on the fields $g, A$ and $\phi$, whereas we will assume that $g \rightarrow 0$ sufficiently fast, as $v \rightarrow \pm \infty$, in such a way that all integrals, with respect to $v$ and involving $g$, converge.

Two parameters are present in the system and are defined as

$$
\begin{equation*}
\beta_{e}=8 \pi \frac{n_{0} T_{0 e}}{B_{0}^{2}}, \quad \delta^{2}=\frac{m_{e}}{m_{i}}, \tag{9}
\end{equation*}
$$

where $n_{0}$ and $T_{0 e}$ are the uniform equilibrium particle density and electron temperature, respectively, with the temperature expressed in energy units. We indicated with $B_{0}$ the (dimensional) amplitude of the magnetic guide field, whereas $m_{e}$ and $m_{i}$ are the electron and ion mass, respectively. The perpendicular Laplacian operator $\Delta_{\perp}$ and the canonical Poisson bracket [, ], on the other hand, are
defined by

$$
\begin{gather*}
\Delta_{\perp} f=\frac{\partial^{2} f}{\partial x^{2}}+\frac{\partial^{2} f}{\partial y^{2}}  \tag{10}\\
{[f, g]=\frac{\partial f}{\partial x} \frac{\partial g}{\partial y}-\frac{\partial g}{\partial x} \frac{\partial f}{\partial y}} \tag{11}
\end{gather*}
$$

respectively, for two functions $f$ and $g$.
As above anticipated, Eqs. (2)- (4) are expressed in terms of normalized quantities. The normalization of the adopted variables is given explicitly by

$$
\begin{align*}
& x=\frac{\tilde{x}}{d_{e}}, \quad y=\frac{\tilde{y}}{d_{e}}, \quad z=\frac{\tilde{z}}{L}, \quad v=\frac{\tilde{v}}{v_{t e}}, \quad t=\frac{v_{t e}}{L} \tilde{t}, \\
& g=\frac{L}{d_{e}} \frac{v_{t e}}{n_{0}} \tilde{g}, \quad F_{e q}=\frac{v_{t e}}{n_{0}} \tilde{F}_{e q}, \quad \phi=\frac{L}{d_{e}} \frac{c \tilde{\phi}}{v_{t e} B_{0} d_{e}}, \quad A=\frac{L}{d_{e}} \frac{\tilde{A}}{B_{0} d_{e}}, \tag{12}
\end{align*}
$$

where $c$ is the speed of light, $d_{e}=c \sqrt{m_{e} /\left(4 \pi e^{2} n_{0}\right)}$ is the electron skin depth, $L$ is a characteristic scale length of variation along the guide field direction, be relevant for tokamak devices and the solar corona. The same model was also adopted in Ref. 38 for the description of drift-Alfvén vortices in plasmas. In the same reference [38], the Hamiltonian structure of the model was also given.

We point out that, although the results presented in the present paper apply to the model (22) (4), they can be extended to more sophisticated Hamiltonian ${ }^{85}$ parent models such as those treated in Refs. [32, 33, which account for further physical ingredients such as multiple species, parallel magnetic perturbations, finite Larmor radius effects and equilibrium temperature anisotropies. Our choice
for the model (22-(4) for the present article is mainly due to its relative simplicity, although, at the same time, the Hamiltonian structure of this system possesses all the fundamental features of the Hamiltonian structures of the more general parent models of Refs. [32, 33].

## 3. Hamiltonian structure of the parent drift-kinetic model and of the family of fluid models in terms of the variables $G_{0}, G_{1}, \cdots, G_{N}$

In this Section we review, although formulated in a more precise setting, the Hamiltonian structure of the parent model (2)-(4), as well as that of the family of fluid models that can be derived from it by means of a specific closure preserving the Hamiltonian structure. However, the Hamiltonian structure of the fluid models that we review in this Section, as anticipated in Sec. 1 is the one presented in Ref. [32, which is not expressed in terms of the fluid moments but in terms of alternative variables $G_{0}, G_{1}, \cdots, G_{N}$, which are linear combinations of the fluid moments, but with coefficients whose expressions are not known in general, so far.

Before proceeding with the review of such Hamiltonian structures, we find it appropriate to introduce some preliminary definitions. The purpose of the following Sec. 3.1 is to formulate a more precise setting, with respect to previous references such as Refs. [32, 33, for the drift-kinetic and the reduced fluid models. Also, we will recall a few notions, such as that of functional derivative, which will be repeatedly used throughout the paper. The readers already familiar with these subjects can of course skip these parts and go directly to Sec. 3.2

### 3.1. Preliminaries

We first introduce the space $\mathcal{F}$ of smooth, periodic and square integrable functions on $\mathcal{D}$. This space will include the fluid Hermite moments of order
greater than zero, and is given by
$\mathcal{F}=\left\{h: \mathbb{R}^{3} \rightarrow \mathbb{R} \mid h \in L^{2}(\mathcal{D}) \cap C^{\infty}(\mathcal{D})\right.$,
$\left.h(x, y, z)=h\left(x+2 L_{x}, y, z\right)=h\left(x, y+2 L_{y}, z\right)=h\left(x, y, z+2 L_{z}\right) \quad \forall(x, y, z) \in \mathbb{R}^{3}\right\}$.

Due to the periodicity assumption, an element $h \in \mathcal{F}$ can be represented as Fourier series in the following way:

$$
\begin{equation*}
h(\mathbf{x})=\sum_{\mathbf{k} \in \mathscr{K}} h_{\mathbf{k}} \mathrm{e}^{i \mathbf{k} \cdot \mathbf{x}} \tag{14}
\end{equation*}
$$

where $\mathbf{x}=(x, y, z) \in \mathbb{R}^{3}$ and $\mathbf{k}=\left(k_{x}, k_{y}, k_{z}\right)$ is an element of the lattice $\mathscr{K}$ defined by

$$
\begin{equation*}
\mathscr{K}=\left\{\left(\frac{\pi m}{L_{x}}, \frac{\pi n}{L_{y}}, \frac{\pi p}{L_{z}}\right),(m, n, p) \in \mathbb{Z}^{3}\right\} . \tag{15}
\end{equation*}
$$

In Eq. 14, the Fourier coefficients $h_{\mathbf{k}}$ are given by the Fourier transform

$$
\begin{equation*}
h_{\mathbf{k}}=\frac{1}{8 L_{x} L_{y} L_{z}} \int_{\mathcal{D}} d^{3} x h(\mathbf{x}) \mathrm{e}^{-i \mathbf{k} \cdot \mathbf{x}} \tag{16}
\end{equation*}
$$

It is now appropriate to introduce the subspace $\mathcal{F}_{0}$ of $\mathcal{F}$, consisting of the elements of $\mathcal{F}$ possessing zero spatial average with respect to $x$ and $y$, over each plane $z=$ constant. This will be the space of the fluid Hermite moment of order 0 and is given by

$$
\begin{equation*}
\mathcal{F}_{0}=\left\{h \in \mathcal{F} \mid h_{\mathbf{k}}=0 \quad \text { for } \mathbf{k}=\left(0,0, \pi p / L_{z}\right) \quad \text { with } p \in \mathbb{Z}\right\} . \tag{17}
\end{equation*}
$$

The specificity of the moment of order 0 comes from the quasi-neutrality relation (3) which, as will be seen, implies that the moment of order 0 have zero spatial average in the plane perpendicular to the guide field.

Given a positive integer $N \in \mathbb{Z}_{>0}$, we also introduce the space

$$
\begin{equation*}
\mathcal{F}_{N}=\mathcal{F}_{0} \times \underbrace{\mathcal{F} \times \cdots \times \mathcal{F}}_{N \text { times }}, \tag{18}
\end{equation*}
$$

which will correspond to the phase space of a fluid model in which the highestorder moment retained in the evolution equations is the one of order $N$ (note that such fluid model actually evolves in time $N+1$ moments, given that the lowest order moment is the one of order 0 ).

Next, we introduce the space $\mathcal{G}$ of the generalized perturbed distribution function:

$$
\begin{align*}
& \mathcal{G}=\left\{g: \mathbb{R}^{4} \rightarrow \mathbb{R} \left\lvert\, g(\mathbf{x}, v)=\sum_{n=0}^{+\infty} g_{n}(\mathbf{x}) \frac{H_{n}(v)}{\sqrt{n!}} F_{e q}(v)\right.,\right. \\
& \left.g_{0} \in \mathcal{F}_{0}, \quad g_{i} \in \mathcal{F} \quad \text { for } i \in \mathbb{Z}_{>0}, \quad H_{n}(v)=(-1)^{n} \mathrm{e}^{\frac{v^{2}}{2}} \frac{d^{n}}{d v^{n}} \mathrm{e}^{-\frac{v^{2}}{2}} \quad \text { for } n \in \mathbb{Z}_{\geqslant 0}\right\} . \tag{19}
\end{align*}
$$

which will be the phase space of the parent drift-kinetic system. Note that, once the dynamical equations are introduced, and thus the time parameter $t$ is added, the dependence on time of $g$ is contained in the dependence on time of the coefficients $g_{n}$. According to a quite standard practice in $\delta f$ driftkinetic and gyrokinetic theory (see, for instance, Refs. [37, [39, 16]), we are considering generalized perturbed distribution functions that can be expressed as a series in Hermite polynomials $H_{n}$, multiplied by the equilibrium distribution function $F_{e q}$, which guarantees a sufficiently rapid decay of $g$ as $v \rightarrow \pm \infty$. The orthogonality relation

$$
\begin{equation*}
\int_{-\infty}^{+\infty} d v H_{m}(v) H_{n}(v) F_{e q}(v)=n!\delta_{m n} \tag{20}
\end{equation*}
$$

permits to express the coefficients $g_{n}$ of the expansion in Eq. 19), in terms of $g$, as

$$
\begin{equation*}
g_{n}=\frac{1}{\sqrt{n!}} \int_{-\infty}^{+\infty} d v H_{n} g, \quad n \geqslant 0 \tag{21}
\end{equation*}
$$

For given $g \in \mathcal{G}$ and $n \geqslant 0$, an element $g_{n}$ defined by Eq. 21) will be referred to as fluid moment, or simply as moment of $g$ of order $n$. We recall that the first Hermite polynomials correspond to $H_{0}=1, H_{1}=v, H_{2}=v^{2}-1, H_{3}=$ $v^{3}-3 v, \ldots$ and that the first four moments are proportional to fluctuations of density, parallel canonical momentum, parallel temperature and parallel heat flux, respectively, of the electron gyrocenters. In particular, following Eq. (21), one has

$$
\begin{align*}
& g_{0}=\int_{-\infty}^{+\infty} d v g, \quad g_{1}=\int_{-\infty}^{+\infty} d v v g  \tag{22}\\
& g_{2}=\frac{1}{\sqrt{2}} \int_{-\infty}^{+\infty} d v\left(v^{2}-1\right) g, \quad g_{3}=\frac{1}{\sqrt{6}} \int_{-\infty}^{+\infty} d v\left(v^{3}-3 v\right) g, \cdots \tag{23}
\end{align*}
$$

Evidently, any element $g \in \mathcal{G}$ also admits a representation in Fourier series, with respect to $\mathbf{x}$, according to:

$$
\begin{equation*}
g(\mathbf{x}, v)=\sum_{\mathbf{k} \in \mathscr{K}} \tilde{g}_{\mathbf{k}}(v) \mathrm{e}^{i \mathbf{k} \cdot \mathbf{x}} \tag{24}
\end{equation*}
$$

with $\mathbf{x} \in \mathbb{R}^{3}, v \in \mathbb{R}$ and where the Fourier coefficients $\tilde{g}_{\mathbf{k}}$, following the expressions (16) (although the tilde symbol in this case was also added, in order to avoid possible confusion with the fluid moments $g_{n}$ ) and 19 , read

$$
\begin{equation*}
\tilde{g}_{\mathbf{k}}=\sum_{n=0}^{+\infty} \frac{g_{n_{\mathbf{k}}}}{\sqrt{n!}} H_{n} F_{e q}, \quad \mathbf{k} \in \mathscr{K} \tag{25}
\end{equation*}
$$

Equations

$$
\begin{align*}
& \Delta_{\perp} \phi=\delta^{2} \sqrt{\frac{2}{\beta_{e}}} \int_{-\infty}^{+\infty} d v g  \tag{26}\\
& \Delta_{\perp} A-A=\sqrt{\frac{\beta_{e}}{2}} \int_{-\infty}^{+\infty} d v v g \tag{27}
\end{align*}
$$

can be solved in Fourier space with respect to $\phi$ and $A$, for a given drift-kinetic generalized perturbed distribution function $g \in \mathcal{G}$. The solutions for the electromagnetic potentials in terms of drift-kinetic generalized perturbed distribution function (for which we use the subscript $d k$ ) are two elements $\phi, A \in \mathcal{F}$ given by

$$
\begin{equation*}
\phi=\phi_{d k}[g], \quad A=A_{d k}[g], \tag{28}
\end{equation*}
$$

with $\phi_{d k}[]: \mathcal{G} \rightarrow \mathcal{F}$ and $A_{d k}[]: \mathcal{G} \rightarrow \mathcal{F}$ linear operators acting on $g \in \mathcal{G}$ in the following way:

$$
\begin{equation*}
\phi_{d k}[g](\mathbf{x})=\sum_{\mathbf{k} \in \mathscr{K}} \phi_{d k}[g]_{\mathbf{k}} \mathrm{e}^{i \mathbf{k} \cdot \mathbf{x}}, \quad A_{d k}[g](\mathbf{x})=\sum_{\mathbf{k} \in \mathscr{K}} A_{d k}[g]_{\mathbf{k}} \mathrm{e}^{i \mathbf{k} \cdot \mathbf{x}} \tag{29}
\end{equation*}
$$

where

$$
\begin{align*}
& \phi_{d k}[g]_{\mathbf{k}}=-\delta^{2} \sqrt{\frac{2}{\beta_{e}}} \int_{-\infty}^{+\infty} d v \frac{\tilde{g}_{\mathbf{k}}}{k_{\perp}^{2}}, \quad \text { for } \quad \mathbf{k} \in \mathscr{K} \backslash\left\{\left(0,0, \pi p / L_{z}\right), p \in \mathbb{Z}\right\}  \tag{30}\\
& \phi_{d k}[g]_{\mathbf{k}}=\phi_{0_{p}}, \quad \text { for } \quad \mathbf{k} \in\left\{\left(0,0, \pi p / L_{z}\right), p \in \mathbb{Z}\right\}  \tag{31}\\
& A_{d k}[g]_{\mathbf{k}}=-\sqrt{\frac{\beta_{e}}{2}} \int_{-\infty}^{+\infty} d v v \frac{\tilde{g}_{\mathbf{k}}}{1+k_{\perp}^{2}}, \quad \text { for } \quad \mathbf{k} \in \mathscr{K} . \tag{32}
\end{align*}
$$

In Eq. (31), $\phi_{0_{p}} \in \mathbb{C}$, for $p \in \mathbb{Z}$, is a family of complex constants, such that The choice of the arbitrary constants $\phi_{0_{n}}$, for $n \in \mathbb{Z}_{\geqslant 0}$, fixes the $z$-dependence of the value of the spatial average of $\phi$, with respect to $x$ and $y$, on planes $z=$ constant

In Eqs. (30)-(32) we also introduced the perpendicular wave number $k_{\perp}$ defined as $k_{\perp}=\sqrt{k_{x}^{2}+k_{y}^{2}}$.

The solutions (30)-(32) permit then to express, at any time $t$, the electromagnetic potentials $\phi$ and $A$ in terms of $g$ in Eqs. (3)-(4).

Using Eq. 25) and the orthogonality relation 20, one can also express the solutions for $\phi$ and $A$ as the images of linear operators (with the subscript $f l$ to indicate that $\phi$ and $A$, in this case, are expressed in terms of fluid variables) $\phi_{f l}[]: \mathcal{F}_{0} \rightarrow \mathcal{F}$ and $A_{f l}[]: \mathcal{F} \rightarrow \mathcal{F}$, acting on the fluid moments $g_{0}$ and $g_{1}$, respectively. More precisely, if we define, for $u \in \mathcal{F}_{0}$ and $w \in \mathcal{F}$ :

$$
\begin{equation*}
\phi_{f l}[u](\mathbf{x})=\sum_{\mathbf{k} \in \mathscr{K}} \phi_{f l}[u]_{\mathbf{k}} \mathrm{e}^{i \mathbf{k} \cdot \mathbf{x}} \tag{33}
\end{equation*}
$$

where

$$
\begin{align*}
& \phi_{f l}[u]_{\mathbf{k}}=-\delta^{2} \sqrt{\frac{2}{\beta_{e}}} \frac{u_{\mathbf{k}}}{k_{\perp}^{2}}, \quad \text { for } \quad \mathbf{k} \in \mathscr{K} \backslash\left\{\left(0,0, \pi p / L_{z}\right), p \in \mathbb{Z}\right\},  \tag{34}\\
& \phi_{f l}[u]_{\mathbf{k}}=\phi_{0_{p}}, \quad \text { for } \quad \mathbf{k} \in\left\{\left(0,0, \pi p / L_{z}\right), p \in \mathbb{Z}\right\} \tag{35}
\end{align*}
$$

and

$$
\begin{equation*}
A_{f l}[w](\mathbf{x})=\sum_{\mathbf{k} \in \mathscr{K}} A_{f l}[w]_{\mathbf{k}} \mathrm{e}^{i \mathbf{k} \cdot \mathbf{x}} \tag{36}
\end{equation*}
$$

with

$$
\begin{equation*}
A_{f l}[w]_{\mathbf{k}}=-\sqrt{\frac{\beta_{e}}{2}} \frac{w_{\mathbf{k}}}{1+k_{\perp}^{2}}, \quad \text { for } \quad \mathbf{k} \in \mathscr{K}, \tag{37}
\end{equation*}
$$

one has, for $g \in \mathcal{G}$, the relations

$$
\begin{align*}
& \phi_{d k}[g]_{\mathbf{k}}=-\delta^{2} \sqrt{\frac{2}{\beta_{e}}} \sum_{n=0}^{+\infty} \frac{1}{\sqrt{n!}} \frac{g_{n_{\mathbf{k}}}}{k_{\perp}^{2}} \int_{-\infty}^{+\infty} d v H_{n} F_{e q} \\
& =-\delta^{2} \sqrt{\frac{2}{\beta_{e}}} \frac{g_{0_{\mathbf{k}}}}{k_{\perp}^{2}}=\phi_{f l}\left[g_{0}\right]_{\mathbf{k}}, \quad \text { for } \quad \mathbf{k} \in \mathscr{K} \backslash\left\{\left(0,0, \pi p / L_{z}\right), p \in \mathbb{Z}\right\},  \tag{38}\\
& \phi_{d k}[g]_{\mathbf{k}}=\phi_{0_{p}}=\phi_{f l}\left[g_{0}\right]_{\mathbf{k}}, \quad \text { for } \quad \mathbf{k} \in\left\{\left(0,0, \pi p / L_{z}\right), p \in \mathbb{Z}\right\},  \tag{39}\\
& A_{d k}[g]_{\mathbf{k}}=-\sqrt{\frac{\beta_{e}}{2}} \sum_{n=0}^{+\infty} \frac{1}{\sqrt{n!}} \frac{g_{n_{\mathbf{k}}}}{1+k_{\perp}^{2}} \int_{-\infty}^{+\infty} d v v H_{n} F_{e q} \\
& =-\sqrt{\frac{\beta_{e}}{2}} \frac{g_{1_{\mathbf{k}}}}{1+k_{\perp}^{2}}=A_{f l}\left[g_{1}\right]_{\mathbf{k}}, \quad \text { for } \quad \mathbf{k} \in \mathscr{K} . \tag{40}
\end{align*}
$$

where, in Eqs. (38)-(40), we made use of Eqs. (22), (25) and of the orthogonality relation 20. The electrostatic and magnetic potentials $\phi$ and $A$ can thus be expressed in terms of the zeroth and first order moments of $g$ by

$$
\begin{equation*}
\phi=\phi_{f l}\left[g_{0}\right], \quad A=A_{f l}\left[g_{1}\right] \tag{41}
\end{equation*}
$$

respectively.
The condition for $g_{0}$ of having zero spatial average on planes $z=$ constant ${ }_{230}$ comes from the fact that Eq. (3), for each $\mathbf{k} \in \mathscr{K}$, implies $-\left(k_{x}^{2}+k_{y}^{2}\right) \phi_{\mathbf{k}}=$ $\delta^{2} \sqrt{2 / \beta_{e}} g_{0_{\mathbf{k}}}$. When evaluated at $\mathbf{k}=\left(0,0, \pi p / L_{z}\right)$, for $p \in \mathbb{Z}$ and $\phi_{\left(0,0, \pi p / L_{z}\right)} \neq$ 0 , this relation implies $g_{0\left(0,0, \pi p / L_{z}\right)}=0$. The electrostatic potential $\phi$, on the other hand, is determined up to the choice of the arbitrary constants $\phi_{0_{n}}$, with $n \in \mathbb{N}$.

In order to introduce the Hamiltonian structures of the drift-kinetic and fluid models, it is also convenient to define here the functional derivatives that we will make use of, later in the paper.

Given a real functional $F: \mathcal{G} \rightarrow \mathbb{R}$, we denote its functional derivative, with respect to $g \in \mathcal{G}$, as $\delta F / \delta g$ and we define it by means of the relation

$$
\begin{equation*}
\lim _{\epsilon \rightarrow 0} \frac{1}{\epsilon}(F(g+\epsilon \delta g)-F(g))=\int_{-\infty}^{+\infty} d v \int_{\mathcal{D}} d^{3} x \delta g \frac{\delta F}{\delta g} \tag{42}
\end{equation*}
$$

for all $\delta g \in \mathcal{G}$. Functional derivatives of this type will appear in the Hamiltonian formulation of the drift-kinetic model.

With regard to the functional derivatives with respect to the moments, occurring in fluid models, given a functional $F: \mathcal{F}_{N} \rightarrow \mathbb{R}$, its functional derivative with respect to $h \in \mathcal{F}_{N}$ is denoted as $\delta F / \delta h$ and corresponds to $\delta F / \delta h=\left(\delta F / \delta h_{0}, \delta F / \delta h_{1}, \cdots, \delta F / \delta h_{N}\right)$ defined by

$$
\begin{equation*}
\lim _{\epsilon \rightarrow 0} \frac{1}{\epsilon}(F(h+\epsilon \delta h)-F(h))=\int_{\mathcal{D}} d^{3} x \delta h \cdot \frac{\delta F}{\delta h}, \tag{43}
\end{equation*}
$$

for all $\delta h=\left(\delta h_{0}, \delta h_{1}, \cdots, \delta h_{N}\right) \in \mathcal{F}_{N}$. Note that the dot on the right-hand side of Eq. 43) denotes a scalar product, so that $\delta h \cdot \delta F / \delta h=\sum_{i=0}^{N} \delta h_{i}\left(\delta F / \delta h_{i}\right)$. Moreover, by varying one $\delta h_{i}$ at the time, for $i=0,1, \cdots, N$, while keeping $\delta h_{j}=0$ for $j \neq i$, one obtains
$\lim _{\epsilon \rightarrow 0} \frac{1}{\epsilon}\left(F\left(h_{0}, \cdots, h_{i}+\epsilon \delta h_{i}, \cdots, h_{N}\right)-F\left(h_{0}, \cdots, h_{i}, \cdots, h_{N}\right)\right)=\int_{\mathcal{D}} d^{3} x \delta h_{i} \frac{\delta F}{\delta h_{i}}$, $F$ with respect to $h_{i}$, for a given $i$.

As will be reviewed in Sec. 3.2, the Poisson bracket of a Hamiltonian system acts on observables, which are functions defined on the phase space. In the case of the parent drift-kinetic model and of the reduced fluid models derived from it, the Jacobi identity for the Poisson bracket relies on the relation

$$
\begin{equation*}
\int_{\mathcal{D}} d^{3} x u[v, w]=\int_{\mathcal{D}} d^{3} x w[u, v], \tag{45}
\end{equation*}
$$

for functions $u, v, w \in \mathcal{F}$. The periodic boundary conditions satisfied by $u, v$ and $w$ imply that boundary terms vanish when integrating by parts, which is required to prove the identity (45). Because the Poisson bracket for the driftkinetic model and the reduced fluid models will contain functional derivatives of observables, at the place of $v$ and $w$, in an expression analogous to the left-hand side of Eq. 45) (see, for instance Eqs. (50) and (87)), we require the observables to be such that their functional derivatives satisfy periodic boundary conditions on $\mathcal{D}$.

We will indicate with $\bar{C}^{\infty}(\mathcal{P})$ the set of real smooth functionals over a phase space $\mathcal{P}$, such that the functional derivatives of these functionals satisfy periodic
boundary conditions on $\mathcal{D}$. So, for instance, the set of observables for the parent drift-kinetic model will be denoted as $\bar{C}^{\infty}(\mathcal{G})$, whereas the set of observables for the $(N+1)$-moment fluid models will be $\bar{C}^{\infty}\left(\mathcal{F}_{N}\right)$.

### 3.2. Hamiltonian formulation of the parent drift-kinetic model

Based on Refs. [38, 32, we review here the (noncanonical) Hamiltonian structure of the parent drift-kinetic model (2)-(4). First, we recall that (see, e.g. Refs. [30, 31, 40]) a Hamiltonian system on a phase space $\mathcal{P}$ is a dynamical system

$$
\begin{equation*}
\dot{u}(t)=J(u(t)) \partial_{u} H(u(t)) \tag{46}
\end{equation*}
$$

where the dot indicates the time derivative and $u: \mathbb{R} \rightarrow \mathcal{P}$ is a curve describing the time evolution in phase space of the dynamical variable $u$ following an initial condition $u(0)=u_{0} \in \mathcal{P}$. We denote with $\Phi(\mathcal{P})$ the set of observables of the dynamical system. On the right-hand side of Eq. 46 the symbol $J$ indicates the Poisson operator $J: T^{*} \mathcal{P} \rightarrow T \mathcal{P}$ (where $T^{*} \mathcal{P}$ and $T \mathcal{P}$ are the cotangent and tangent bundle of $\mathcal{P}$, respectively ) and $H \in \Phi(\mathcal{P})$ is the observable corresponding to the Hamiltonian of the system. In Eq. 46 we also denoted with $\partial_{u}$ a derivative with respect to $u$, which, in the infinite-dimensional case, takes the form of a functional derivative, as those defined in Eqs. (42) and 43).

A generic observable $F \in \Phi(\mathcal{P})$ of the Hamiltonian system (46) evolves according to

$$
\begin{equation*}
\dot{F}=\{F, H\} \tag{47}
\end{equation*}
$$

where $\{$,$\} is a Poisson bracket, i.e. an antisymmetric bilinear form satisfying the$ Leibniz rule and the Jacobi identity. Note that, for two observables $F, G \in \Phi(\mathcal{P})$, the Poisson operator and the Poisson bracket are related by

$$
\begin{equation*}
\{F, G\}=<\partial_{u} F, J(u) \partial_{u} G> \tag{48}
\end{equation*}
$$

where $<,>$ indicates a dual pairing.
The Hamiltonian structure of a Hamiltonian system on a given phase space is thus determined by its Hamiltonian and its Poisson bracket. Due to the
antisymmetry property of the bracket, it follows from 47 that $\dot{H}=0$, which expresses the conservation of the total energy.

As anticipated in Sec. 3.1, the phase space for the parent drift-kinetic system is given by the set $\mathcal{G}$. The Hamiltonian structure of this system 38] consists of the following Hamiltonian functional $H_{d k} \in \bar{C}^{\infty}(\mathcal{G})$ :

$$
\begin{equation*}
H_{d k}(g)=\frac{1}{2} \int_{\mathcal{D}} d^{3} x \int_{-\infty}^{+\infty} d v\left(\frac{g^{2}}{F_{e q}}-\sqrt{\frac{2}{\beta_{e}}} g(\phi-v A)\right) \tag{49}
\end{equation*}
$$

where $\phi=\phi_{d k}[g], A=A_{d k}[g]$, and of the Poisson bracket

$$
\begin{equation*}
\{F, G\}_{d k}=\int_{\mathcal{D}} d^{3} x \int_{-\infty}^{+\infty} d v\left(\sqrt{\frac{\beta_{e}}{2}} g\left[F_{g}, G_{g}\right]-v F_{e q} F_{g} \frac{\partial G_{g}}{\partial z}\right) \tag{50}
\end{equation*}
$$ bution function around a Maxwellian equilibrium (first term on the right-hand side), with the energy associated with electromagnetic fluctuations, given by the remaining two terms on the right-hand side.

We proceed at a formal level, assuming that, for a generic initial condition ${ }_{280} g(x, y, z, v, 0)=\mathrm{g}_{0}(x, y, z, v) \in \mathcal{G}$, the solution of the system (2)-(4) exists for $0 \leqslant t \leqslant T$, with $T>0$, and that, for each solution, $g(x, y, z, v, t) \in \mathcal{G}$ for $0 \leqslant t \leqslant T$. We then identify a solution $g(x, y, z, v, t)$, corresponding to a given initial condition, with a curve $g:[0, T] \rightarrow \mathcal{G}$ that associates, at each time $t$, a point $g(t) \in \mathcal{G}$ in phase space.

Concerning the Hamiltonian structure of the drift-kinetic model, we remark that

$$
\begin{equation*}
\frac{\delta H_{d k}}{\delta g}=\frac{g}{F_{e q}}-\sqrt{\frac{2}{\beta_{e}}}\left(\phi_{d k}[g]-v A_{d k}[g]\right) \tag{51}
\end{equation*}
$$

In order to derive the latter relation one makes use of the following symmetry
properties of the operators $\phi_{d k}[]$ and $A_{d k}[]$ :

$$
\begin{align*}
& \int_{\mathcal{D}} d^{3} x \int_{-\infty}^{+\infty} d v h \phi_{d k}[g]=\int_{\mathcal{D}} d^{3} x \int_{-\infty}^{+\infty} d v g \phi_{d k}[h]  \tag{52}\\
& \int_{\mathcal{D}} d^{3} x \int_{-\infty}^{+\infty} d v h A_{d k}[g]=\int_{\mathcal{D}} d^{3} x \int_{-\infty}^{+\infty} d v g A_{d k}[h] \tag{53}
\end{align*}
$$

for $g, h \in \mathcal{G}$. The properties $(52)-53)$ easily follow from the definitions (30)-(32) and from the orthogonality of the Fourier and Hermite bases.

From the expressions (48) and (50), using integration by parts, it follows that the Poisson operator associated with the drift-kinetic Poisson bracket 50 ) is given by

$$
\begin{equation*}
J_{d k}(g)=-\sqrt{\frac{\beta_{e}}{2}}[g, \cdot]-v F_{e q} \frac{\partial}{\partial z} \tag{54}
\end{equation*}
$$

with respect to the dual pairing

$$
\begin{equation*}
<f, h>=\int_{\mathcal{D}} d^{3} x \int_{-\infty}^{+\infty} d v f h \tag{55}
\end{equation*}
$$

for two functions $f, h$. Combining Eqs. (51) and (54) with the general expression (46) for a Hamiltonian system, one retrieves namely the drift-kinetic equation (2).

### 3.3. Hamiltonian formulation of the fluid models

Multiplying both sides of Eq. 22 by $H_{n} / \sqrt{n!}$, for $n=0,1,2, \cdots$ and integrating over $v$, one obtains the following infinite system of fluid equations

$$
\begin{align*}
& \frac{\partial g_{0}}{\partial t}+\left[\phi, g_{0}\right]-\left[A, g_{1}\right]+\frac{\partial}{\partial z}\left(g_{1}+\sqrt{\frac{2}{\beta_{e}}} A\right)=0  \tag{56}\\
& \frac{\partial g_{1}}{\partial t}+\left[\phi, g_{1}\right]-\sqrt{2}\left[A, g_{2}\right]-\left[A, g_{0}\right]+\frac{\partial}{\partial z}\left(\sqrt{2} g_{2}+g_{0}-\sqrt{\frac{2}{\beta_{e}}} \phi\right)=0  \tag{57}\\
& \frac{\partial g_{2}}{\partial t}+\left[\phi, g_{2}\right]-\sqrt{3}\left[A, g_{3}\right]-\sqrt{2}\left[A, g_{1}\right]+\frac{\partial}{\partial z}\left(\sqrt{3} g_{3}+\sqrt{2} g_{1}+\sqrt{2} \sqrt{\frac{2}{\beta_{e}}} A\right)=0 \tag{58}
\end{align*}
$$

$\frac{\partial g_{N}}{\partial t}+\left[\phi, g_{N}\right]-\sqrt{N+1}\left[A, g_{N+1}\right]-\sqrt{N}\left[A, g_{N-1}\right]+\frac{\partial}{\partial z}\left(\sqrt{N+1} g_{N+1}+\sqrt{N} g_{N-1}\right)=0$,
where

$$
\begin{align*}
& \Delta_{\perp} \phi=\delta^{2} \sqrt{\frac{2}{\beta_{e}}} g_{0}  \tag{60}\\
& \Delta_{\perp} A-A=\sqrt{\frac{\beta_{e}}{2}} g_{1} \tag{61}
\end{align*}
$$

and the fluid moments $g_{0}, g_{1}, g_{2}, \cdots$ are defined in Eq. 21. We remark that, on the basis of our definitions, $g_{0}, g_{1}$ and $g_{2}$ are proportional to the fluctuations of the electron gyrocenter density, parallel velocity and parallel temperature, respectively, where 'parallel' refers to the direction of the guide field. Note also that the first three equations (56)-58) of the hierarchy are peculiar, as they involve also derivatives, with respect to $z$, of the electromagnetic potentials $\phi$ and $A$. For $N>2$, on the other hand, the equations of the hierarchy are given by Eq. (59).

Given a fixed integer $N \geqslant 1$, the infinite hierarchy of equations (56)-59) can be truncated by imposing

$$
\begin{equation*}
g_{N+1}=0 \tag{62}
\end{equation*}
$$

The resulting closed system, given that we are assuming $g \in \mathcal{G}$ in the parent drift-kinetic system, can be written as

$$
\begin{align*}
& \frac{\partial g_{m}}{\partial t}+\left[\phi, g_{m}\right]-S_{N_{m n}}\left[A, g_{n}\right]+S_{N_{m n}} \frac{\partial g_{n}}{\partial z} \\
& +\sqrt{\frac{2}{\beta_{e}}} \frac{\partial}{\partial z}\left(\sqrt{m!}\left(\delta_{m 0}+\delta_{m 2}\right) A-\delta_{m 1} \phi\right)=0, \quad m=0,1, \cdots, N  \tag{63}\\
& \Delta_{\perp} \phi=\delta^{2} \sqrt{\frac{2}{\beta_{e}}} g_{0}  \tag{64}\\
& \Delta_{\perp} A-A=\sqrt{\frac{\beta_{e}}{2}} g_{1} \tag{65}
\end{align*}
$$

where the sum over repeated indices is understood and where $S_{N_{m n}}$ indicates
the element of row $m$ and column $n$, of the $(N+1) \times(N+1)$ matrix

$$
S_{N}=\left(\begin{array}{cccccc}
0 & 1 & 0 & 0 & \ldots & 0  \tag{66}\\
1 & 0 & \sqrt{2} & 0 & \ldots & 0 \\
0 & \sqrt{2} & 0 & \sqrt{3} & \ldots & 0 \\
0 & 0 & \sqrt{3} & 0 & \ldots & 0 \\
& \ldots & & & \ldots & \\
& \cdots & & & \ldots & \\
0 & 0 & 0 & \ldots & 0 & \sqrt{N} \\
0 & 0 & 0 & \ldots & \sqrt{N} & 0
\end{array}\right) .
$$

We specify that, in this paper, the indices of the rows and of the columns of $(N+1) \times(N+1)$ matrices run from 0 to $N$.

Because the matrix $S_{N}$ is real symmetric, it exists an orthogonal matrix $U_{N} \in$ $O(N+1)$ such that $U_{N}^{T} S_{N} U_{N}=\operatorname{diag}\left(\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}\right)$, where $\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}$ are the eigenvalues of $S_{N}$. One can then introduce the alternative variables

$$
\begin{equation*}
G_{i}=U_{N_{i m}}^{T} g_{m}, \quad i=0,1, \cdots, N \tag{67}
\end{equation*}
$$

in terms of which the system (63)-65) can be rewritten as

$$
\begin{align*}
& \frac{\partial G_{i}}{\partial t}+\left[\phi-\lambda_{i} A, G_{i}\right]+\lambda_{i} \frac{\partial G_{i}}{\partial z} \\
& +\sqrt{\frac{2}{\beta_{e}}} \frac{\partial}{\partial z}\left(\left(U_{N_{i 0}}^{T}+\sqrt{2!} U_{N_{i 2}}^{T}\right) A-U_{N_{i 1}}^{T} \phi\right)=0, \quad i=0,1, \cdots, N  \tag{68}\\
& \Delta_{\perp} \phi=\delta^{2} \sqrt{\frac{2}{\beta_{e}}} \sum_{m=0}^{N} U_{N_{0 m}} G_{m}  \tag{69}\\
& \Delta_{\perp} A-A=\sqrt{\frac{\beta_{e}}{2}} \sum_{m=0}^{N} U_{N_{1 m}} G_{m} \tag{70}
\end{align*}
$$

In Ref. [32] it was shown that the system 68-70 (which is equivalent to Eqs. (63)-(65) is Hamiltonian, with Hamiltonian functional

$$
\begin{equation*}
\mathcal{H}\left(G_{0}, G_{1}, \cdots, G_{N}\right)=\frac{1}{2} \int_{\mathcal{D}} d^{3} x\left(\sum_{n=0}^{N} G_{n}^{2}-\sqrt{\frac{2}{\beta_{e}}} U_{N_{0 i}} G_{i}\left(\phi_{f l}\left[U_{N_{0 l}} G_{l}\right]-\lambda_{i} A_{f l}\left[U_{N_{1 l}} G_{l}\right]\right)\right) \tag{71}
\end{equation*}
$$

where, due to the orthogonality of the matrix $U_{N}$, from Eq. 67), one has $U_{N_{0 l}} G_{l}=g_{0}$ and $U_{N_{1 l}} G_{l}=g_{1}$ (we recall that, also in the expression 71),
the sum over repeated indices is understood). In Eq. 71) we are considering $\mathcal{H} \in \bar{C}^{\infty}\left(\mathcal{G}_{N}\right)$, where

$$
\begin{equation*}
\mathcal{G}_{N}=\left\{\left(G_{0}, G_{1}, \cdots, G_{N}\right) \in \mathcal{F} \times \mathcal{F} \times \cdots \times \mathcal{F} \mid \sum_{l=0}^{N} U_{N_{0 l}} G_{l} \in \mathcal{F}_{0}\right\} \tag{72}
\end{equation*}
$$

The introduction of this set is required in order for the expression $\phi_{f l}\left[U_{N_{0 l}} G_{l}\right]$ to be well defined and, in turn, for the equation 6 to be solved with respect to $\phi$.

The Poisson bracket, on the other hand, is given by

$$
\begin{equation*}
\{F, K\}_{G}=\{F, K\}_{G_{\perp}}+\{F, K\}_{G_{\|}} \tag{73}
\end{equation*}
$$

where

$$
\begin{aligned}
\{F, K\}_{G_{\perp}} & =\sqrt{\frac{\beta_{e}}{2}} \sum_{i=0}^{N} \frac{1}{U_{N_{0 i}}} \int_{\mathcal{D}} d^{3} x G_{i}\left[F_{G_{i}}, K_{G_{i}}\right] \\
\{F, K\}_{G_{\|}} & =-\sum_{i=0}^{N} \lambda_{i} \int_{\mathcal{D}} d^{3} x F_{G_{i}} \frac{\partial K_{G_{i}}}{\partial z}
\end{aligned}
$$

for $F, K \in \bar{C}^{\infty}\left(\mathcal{G}_{N}\right)$. Using the relation

$$
\begin{equation*}
\frac{\delta \mathcal{H}}{\delta G_{i}}=G_{i}-\sqrt{\frac{2}{\beta_{e}}} U_{N_{0 i}}\left(\phi_{f l}\left[U_{N_{0 l}} G_{l}\right]-\lambda_{i} A_{f l}\left[U_{N_{1 l}} G_{l}\right]\right), \quad i=0,1, \cdots, N \tag{74}
\end{equation*}
$$

one can indeed obtain Eqs. (68)- 70 , from Eqs. (71) and (73), applying the expression 46.

We remark that, in order to obtain the relation $\sqrt[74)]{ }$, we made use of the symmetry properties

$$
\begin{align*}
& \int_{\mathcal{D}} d^{3} x \eta \phi_{f l}[\xi]=\int_{\mathcal{D}} d^{3} x \xi \phi_{f l}[\eta],  \tag{75}\\
& \int_{\mathcal{D}} d^{3} x u A_{f l}[w]=\int_{\mathcal{D}} d^{3} x w A_{f l}[u], \tag{76}
\end{align*}
$$

for $\eta, \xi \in \mathcal{F}_{0}$ and $u, w \in \mathcal{F}$, which are a straightforward consequence of the relations 52 and (53).

## 4. Hamiltonian structure of the family of fluid models in terms of the moments $g_{0}, g_{1}, \cdots, g_{N}$

The Hamiltonian formulation of the fluid models described in Sec. 3.3 crucially depends on the knowledge of the matrix $U_{N}$ and of the eigenvalues
$\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}$. The existence of such matrix is guaranteed by the spectral theorem and this is sufficient to show the existence of the Hamiltonian structure. However, in the absence of an explicit expression for the matrix $U_{N}$ and for the eigenvalues of $S_{N}$, the general expression for the Hamiltonian (71) and the Poisson bracket (73) cannot be determined, and one is forced to find such matrix and such eigenvalues case by case, when possible, for a given $N$ of interest. Moreover, the general expression of the transformation (67) is not known either. This expression is required in order to rewrite the Poisson bracket (73), for an arbitrary $N$, in terms of the fluid moments $g_{0}, g_{1}, \cdots, g_{N}$, which are the variables that one naturally adopts in applications.

In this Section we remedy this deficiency and provide the explicit expression for the matrix $U_{N}$ and for the eigenvalues of $S_{N}$, as well as the Hamiltonian formulation of the fluid models $(63)-(\sqrt{65})$, for arbitrary $N$, in terms of the moments $g_{0}, g_{1}, \cdots, g_{N}$.

Indeed, in Refs. [32, 33] it was not realized that the eigenvalues of the matrix $S_{N}$ correspond to the zeros of the Hermite polynomial $H_{N+1}(x)$ (which is actually a well known fact, see, e.g. Ref. 41]). From the recurrence relation $x H_{n}(x)=H_{n+1}(x)+n H_{n-1}(x)$ it follows that

$$
\begin{equation*}
x \hat{H}_{n}(x)=\sqrt{n+1} \hat{H}_{n+1}(x)+\sqrt{n} \hat{H}_{n-1}(x) \tag{77}
\end{equation*}
$$

where we defined

$$
\begin{equation*}
\hat{H}_{n}(x)=\frac{H_{n}(x)}{\sqrt{n!}}, \quad n \geqslant 0 \tag{78}
\end{equation*}
$$

Equation (77), evaluated at $x=\lambda_{i}$, for a given eigenvalue $\lambda_{i} \in\left\{\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}\right\}$, yields

$$
\begin{equation*}
\lambda_{i} \hat{H}_{n}\left(\lambda_{i}\right)=\sqrt{n+1} \hat{H}_{n+1}\left(\lambda_{i}\right)+\sqrt{n} \hat{H}_{n-1}\left(\lambda_{i}\right) \tag{79}
\end{equation*}
$$

Combining Eq. (79) with the expression of $S_{N}$, one sees that the vector $\left(\hat{H}_{0}\left(\lambda_{i}\right), \hat{H}_{1}\left(\lambda_{i}\right), \cdots, \hat{H}_{N}\left(\lambda_{i}\right)\right)^{T}$ is an eigenvector of $S_{N}$ associated with the eigenvalue $\lambda_{i}$. Because the columns of the matrix $U_{N}$ correspond to orthonormal eigenvectors of $S_{N}$, we have that a generic element of $U_{N}$ is given by

$$
\begin{equation*}
U_{N_{m n}}=\frac{\hat{H}_{m}\left(\lambda_{n}\right)}{u_{(n)}}, \quad m, n=0,1, \cdots, N \tag{80}
\end{equation*}
$$

where

$$
\begin{equation*}
u_{(n)}=\sqrt{\sum_{i=0}^{N} \hat{H}_{i}^{2}\left(\lambda_{n}\right)}, \quad n=0,1, \cdots, N \tag{81}
\end{equation*}
$$

are normalization constants. The expression for these constants can be simplified making use of the Christoffel-Darboux identity 42

$$
\begin{equation*}
\sum_{n=0}^{N} \frac{\bar{H}_{n}(x) \bar{H}_{n}(y)}{2^{n} n!}=\frac{\bar{H}_{N+1}(x) \bar{H}_{N}(y)-\bar{H}_{N}(x) \bar{H}_{N+1}(y)}{2^{N+1} N!(x-y)} \tag{82}
\end{equation*}
$$

where $\bar{H}_{n}(x)=2^{\frac{n}{2}} H_{n}(\sqrt{2} x)$ are rescaled Hermite polynomials. The simplification is obtained by first taking the limit $x \rightarrow y$, with the help of the de l'Hôpital rule, in the expression 82). Then one sets $y=\lambda_{n}$, and given that $H_{N+1}\left(\lambda_{n}\right)=0$, one can obtain the following simplified expression [43, 41]

$$
\begin{equation*}
u_{(n)}=\sqrt{\frac{N+1}{N!}}\left|H_{N}\left(\lambda_{n}\right)\right| \tag{83}
\end{equation*}
$$

Combining Eqs. 80 and 83, it follows that the explicit expression for a generic element of the matrix $U_{N}$ is given by

$$
\begin{equation*}
U_{N_{m n}}=\sqrt{\frac{N!}{N+1}} \frac{\hat{H}_{m}\left(\lambda_{n}\right)}{\left|H_{N}\left(\lambda_{n}\right)\right|}, \quad m, n=0,1, \cdots, N \tag{84}
\end{equation*}
$$

so that the generic matrix $U_{N}$ has the form

$$
U_{N}=\left(\begin{array}{ccccc}
\sqrt{\frac{N!}{N+1}} \frac{\hat{H}_{0}\left(\lambda_{0}\right)}{H_{N}\left(\lambda_{0}\right) \mid} & \sqrt{\frac{N!}{N!}} \frac{\hat{H}_{0}\left(\lambda_{1}\right)}{\left|H_{N}\left(\lambda_{1}\right)\right|} & \cdots & \cdots & \sqrt{\frac{N!}{N+1}} \frac{\hat{H}_{0}\left(\lambda_{N}\right)}{H_{N_{N}}\left(\lambda_{N}\right) \mid}  \tag{85}\\
\sqrt{\frac{N!}{N+1}} \frac{H_{1}\left(\lambda_{0}\right)}{\left|H_{N}\left(\lambda_{0}\right)\right|} & \sqrt{\frac{N!}{N+1}} \frac{\hat{H}_{1}\left(\lambda_{1}\right)}{H_{N}\left(\lambda_{1}\right) \mid} & \cdots & \cdots & \sqrt{\frac{N!}{N+1}} \frac{H_{1}\left(\lambda_{N}\right)}{H_{N}\left(\lambda_{N}\right) \mid} \\
\cdots & \cdots & \cdots & \cdots & \cdots \\
\cdots & \cdots & \cdots & \cdots & \cdots \\
\cdots & \cdots & & & \cdots \\
\cdots & \cdots & \cdots & \cdots & \cdots \\
\sqrt{\frac{N!}{N+1}} \frac{\hat{H}_{N}\left(\lambda_{0}\right)}{\left|H_{N}\left(\lambda_{0}\right)\right|} & \sqrt{\frac{N!}{N+1} \frac{\hat{H}_{N}\left(\lambda_{1}\right)}{H_{N}\left(\lambda_{1}\right) \mid}} & & & \sqrt{\frac{N!}{N+1}} \frac{\hat{H}_{N}\left(\lambda_{N}\right)}{\left|H_{N}\left(\lambda_{N}\right)\right|}
\end{array}\right) .
$$

We are now ready to provide the explicit Hamiltonian structure, in terms of the fluid moments, for an arbitrary member of the class of reduced fluid models under consideration. We formalize this result by means of the following

Proposition 4.1. For any integer $N \geqslant 1$, the system (63)-65), with $g_{0}, g_{1}, \cdots, g_{N} \in$ $\mathcal{F}_{N}$, is a Hamiltonian system with Hamiltonian $H \in \bar{C}^{\infty}\left(\mathcal{F}_{N}\right)$ given by

$$
\begin{equation*}
H\left(g_{0}, g_{1}, \cdots, g_{N}\right)=\frac{1}{2} \int_{\mathcal{D}} d^{3} x\left(\sum_{n=0}^{N} g_{n}^{2}-\sqrt{\frac{2}{\beta_{e}}}\left(g_{0} \phi_{f l}\left[g_{0}\right]-g_{1} A_{f_{l}}\left[g_{1}\right]\right)\right) \tag{86}
\end{equation*}
$$

and Poisson bracket given by

$$
\begin{align*}
& \{F, K\}_{g}=\sqrt{\frac{\beta_{e}}{2}} \sum_{l, m, n=0}^{N} W_{(N) l}^{m n} \int_{\mathcal{D}} d^{3} x g_{l}\left[F_{g_{m}}, K_{g_{n}}\right] \\
& -\sum_{m, n=0}^{N} S_{N_{m n}} \int_{\mathcal{D}} d^{3} x F_{g_{m}} \frac{\partial K_{g_{n}}}{\partial z} \tag{87}
\end{align*}
$$

for $F, K \in \bar{C}^{\infty}\left(\mathcal{F}_{N}\right)$, where

$$
\begin{equation*}
W_{(N) l}^{m n}=\frac{N!}{N+1} \sum_{i=0}^{N} \frac{H_{l}\left(\lambda_{i}\right) H_{m}\left(\lambda_{i}\right) H_{n}\left(\lambda_{i}\right)}{H_{N}^{2}\left(\lambda_{i}\right) \sqrt{l!m!n!}} \tag{88}
\end{equation*}
$$

In the expression 88), the numbers $\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}$ are the eigenvalues of the matrix $S_{N}$ given in Eq. (66), and are known to correspond to the zeros of the Hermite polynomial $H_{N+1}(v)$.

Proof. The system (63)-65) is equivalent to the system (68)-70). In particular, given that the orthogonal matrix $U_{N}$ is invertible with inverse $U_{N}^{-1}=U_{N}^{T}$, one has a linear invertible $\operatorname{map} \mathcal{U}_{N}: \mathcal{G}_{N} \rightarrow \mathcal{F}_{N}$ defined by

$$
\begin{equation*}
\mathcal{U}_{N} z=U_{N} z \tag{89}
\end{equation*}
$$

for $z \in \mathcal{G}_{N}$, which preserves the properties of a Poisson bracket. Therefore, in order to determine the Hamiltonian structure of the system (63)-65), it is sufficient to express the Hamiltonian structure of the system 68)-70), in terms of the variables $g_{0}, g_{1}, \cdots, g_{N}$, making use of the transformation 67) and of its inverse

$$
\begin{equation*}
g_{m}=U_{N_{m n}} G_{n}, \quad m=0,1, \cdots, N \tag{90}
\end{equation*}
$$

We proceed with transforming first the Hamiltonian (71). Using the orthogonality of $U_{N}$, and Eq. (67), one readily has that

$$
\begin{equation*}
\sum_{n=0}^{N} G_{n}^{2}=\sum_{n, i, j=0}^{N} U_{N_{n i}}^{T} g_{i} U_{N_{n j}}^{T} g_{j}=\sum_{n, i, j=0}^{N} g_{j} U_{N_{j n}} U_{N_{n i}}^{T} g_{i}=\sum_{i=0}^{N} g_{i}^{2} \tag{91}
\end{equation*}
$$

which permits to transform the first term on the right-hand side of Eq. (71). With regard to the remaining terms, they readily follow from the relations $U_{N_{0 l}} G_{l}=g_{0}$ and $U_{N_{1 l}} G_{l}=g_{1}$. Thus, one has straightforwardly

$$
\begin{align*}
& \mathcal{H}\left(G_{0}, G_{1}, \cdots, G_{N}\right)=\frac{1}{2} \int_{\mathcal{D}} d^{3} x\left(\sum_{n=0}^{N} g_{n}^{2}-\sqrt{\frac{2}{\beta_{e}}}\left(g_{0} \phi_{f l}\left[g_{0}\right]-g_{1} A_{f l}\left[g_{1}\right]\right)\right) \\
& =H\left(g_{0}, g_{1}, \cdots, g_{N}\right) \tag{92}
\end{align*}
$$

With regard to the Poisson bracket, in order to transform $\{,\}_{G}$, one needs to transform also the functional derivatives in terms of the variables $g_{0}, g_{1}, \cdots, g_{N}$. This is accomplished using Eqs. (44) and (90), from which one obtains

$$
\begin{equation*}
\frac{\delta \bar{F}}{\delta G_{i}}=U_{N i j}^{T} \frac{\delta F}{\delta g_{j}}, \quad i=0,1, \cdots, N \tag{93}
\end{equation*}
$$

for $\bar{F} \in \bar{C}^{\infty}\left(\mathcal{G}_{N}\right)$ and $F=\bar{F} \circ \mathcal{U}_{N}^{-1} \in \bar{C}^{\infty}\left(\mathcal{F}_{N}\right)$.
Using the relations (67) and 93 in the expression 73 , for $\bar{F}, \bar{K} \in \bar{C}^{\infty}\left(\mathcal{G}_{N}\right)$, yields

$$
\begin{align*}
& \{\bar{F}, \bar{K}\}_{G}=\sqrt{\frac{\beta_{e}}{2}} \sum_{i=0}^{N} \sum_{l, m, n=0}^{N} \frac{1}{U_{N_{0 i}}} U_{N_{i l}}^{T} U_{N_{i m}}^{T} U_{N_{i n}}^{T} \int_{\mathcal{D}} d^{3} x g_{l}\left[F_{g_{m}}, K_{g_{n}}\right]  \tag{94}\\
& -\sum_{i=0}^{N} \sum_{m, n=0}^{N} \lambda_{i} U_{N_{i m}}^{T} U_{N_{i n}}^{T} \int_{\mathcal{D}} d^{3} x F_{g_{m}} \frac{\partial K_{g_{n}}}{\partial z} \\
& =\sqrt{\frac{\beta_{e}}{2}} \sum_{l, m, n=0}^{N} \sum_{i=0}^{N} u_{(i)} \frac{\hat{H}_{l}\left(\lambda_{i}\right)}{u_{(i)}} \frac{\hat{H}_{m}\left(\lambda_{i}\right)}{u_{(i)}} \frac{\hat{H}_{n}\left(\lambda_{i}\right)}{u_{(i)}} \int_{\mathcal{D}} d^{3} x g_{l}\left[F_{g_{m}}, K_{g_{n}}\right]  \tag{95}\\
& -\sum_{i=0}^{N} \sum_{m, n=0}^{N} \lambda_{i} \frac{\hat{H}_{m}\left(\lambda_{i}\right)}{u_{(i)}} \frac{\hat{H}_{n}\left(\lambda_{i}\right)}{u_{(i)}} \int_{\mathcal{D}} d^{3} x F_{g_{m}} \frac{\partial K_{g_{n}}}{\partial z}  \tag{96}\\
& =\sqrt{\frac{\beta_{e}}{2}} \sum_{l, m, n=0}^{N} \frac{N!}{N+1} \sum_{i=0}^{N} \frac{H_{l}\left(\lambda_{i}\right) H_{m}\left(\lambda_{i}\right) H_{n}\left(\lambda_{i}\right)}{H_{N}^{2}\left(\lambda_{i}\right) \sqrt{l!m!n!}} \int_{\mathcal{D}} d^{3} x g_{l}\left[F_{g_{m}}, K_{g_{n}}\right]  \tag{97}\\
& -\frac{N!}{N+1} \sum_{m, n=0}^{N} \sum_{i=0}^{N} \lambda_{i} \frac{H_{m}\left(\lambda_{i}\right) H_{n}\left(\lambda_{i}\right)}{H_{N}^{2}\left(\lambda_{i}\right) \sqrt{m!n!}} \int_{\mathcal{D}} d^{3} x F_{g_{m}} \frac{\partial K_{g_{n}}}{\partial z} \tag{98}
\end{align*}
$$

where $F=\bar{F} \circ \mathcal{U}_{N}^{-1}, K=\bar{K} \circ \mathcal{U}_{N}^{-1}$, and where in the steps $94-98$, we made use of the expressions 83) and (84).

Given the definition (88), we see that the expression (97) is already in the desired form. We focus then on the expression (98). Making use of the formula
(77) we obtain

$$
\begin{align*}
& \frac{N!}{N+1} \sum_{i=0}^{N} \lambda_{i} \frac{H_{m}\left(\lambda_{i}\right) H_{n}\left(\lambda_{i}\right)}{H_{N}^{2}\left(\lambda_{i}\right) \sqrt{m!n!}} \\
& =\sum_{m, n=0}^{N} \sum_{i=0}^{N} \sqrt{\frac{N!}{N+1}} \frac{\left(\sqrt{m+1} \hat{H}_{m+1}\left(\lambda_{i}\right)+\sqrt{m} \hat{H}_{m-1}\left(\lambda_{i}\right)\right) \hat{H}_{n}\left(\lambda_{i}\right)}{H_{N}\left(\lambda_{i}\right)} \sqrt{\frac{N!}{N+1}} \frac{\hat{H}_{n}\left(\lambda_{i}\right)}{H_{N}\left(\lambda_{i}\right)} \\
& =\sum_{i=0}^{N}\left(\sqrt{m+1} U_{N_{m+1 i}} U_{N_{i n}}^{T}+\sqrt{m} U_{N_{m-1} i} U_{N_{i n}}^{T}\right) \\
& =\left(\sqrt{m+1} \delta_{m+1, n}+\sqrt{m} \delta_{m-1, n}\right)=S_{N_{m n}} \tag{99}
\end{align*}
$$

Using the expression (88) and the result (99) in the two final steps of (94)- (98) yields

$$
\begin{align*}
& \{\bar{F}, \bar{K}\}_{G}=\sqrt{\frac{\beta_{e}}{2}} \sum_{l, m, n=0}^{N} W_{(N) l}^{m n} \int_{\mathcal{D}} d^{3} x g_{l}\left[F_{g_{m}}, K_{g_{n}}\right] \\
& -\sum_{m, n=0}^{N} S_{N_{m n}} \int_{\mathcal{D}} d^{3} x F_{g_{m}} \frac{\partial K_{g_{n}}}{\partial z}  \tag{100}\\
& =\{F, K\}_{g}
\end{align*}
$$

which completes the proof.

Remark 4.1. The Poisson bracket (87) generalizes, up to the normalization, various Poisson brackets present in the literature. For instance, for $N=1$, it reduces to the bracket for the electron dynamics of the models of Refs. [24, 25, 44 and to the bracket of the model of Ref. [12] in the cold-ion limit. For $N=2$ one retrieves the bracket for the ion gyrofluid dynamics of the model of Ref. [26], whereas for $N=3$, in the 2D limit, the bracket pertaining to the electron dynamics with heat flux of Ref. [27] is obtained.

Remark 4.2. Comparing Eq. 87) with Eq. 73, i.e. the expressions of the Poisson bracket in terms of the fluid moments $g_{0}, g_{1}, \cdots, g_{N}$ and in terms of the variables $G_{0}, G_{1}, \cdots, G_{N}$, respectively, it emerges that the Poisson bracket takes a much simpler form in terms of the latter variables. In particular, when
considering the perpendicular component $\{F, K\}_{G_{\perp}}$, one sees that this is the direct sum of independent Poisson brackets of the form

$$
\begin{equation*}
c_{i} \int_{\mathcal{D}} d^{3} x G_{i}\left[F_{G_{i}}, K_{G_{i}}\right] \tag{101}
\end{equation*}
$$

(a) $W_{(N) l}^{m n}=0 \quad$ if $l+m+n$ is an odd number,
(b) For fixed integers $l, m, n$, one has $W_{N \sigma(l)}^{\sigma(m) \sigma(n)}=W_{(N) l}^{m n}$, where $\sigma:\{l, m, n\} \rightarrow$ $\{l, m, n\}$ is a permutation of the integers $l, m, n$.

The proof of Lemma 5.1 is provided in AppendixA.
Remark 5.1. We observe that the property $W_{(N) l}^{m n}=W_{(N) l}^{n m}$, following from Lemma 5.1 (b), is required by the antisymmetry of the Poisson bracket $\{,\}_{g}$ [20].

We now proceed with re-expressing the Poisson bracket $\{,\}_{g}$ in a way that will facilitate its comparison with the bilinear form obtained from the truncated series, which will be derived in Sec. 5.1. In order to formulate the corresponding Proposition, it is convenient to the define, for a given integer $N \geqslant 1$, the following sets $A_{N}$ and $B_{N}$ :

$$
\begin{align*}
& A_{N}=\left\{(l, m, n) \in \mathbb{Z}_{\geqslant 0}^{3} \mid l, m, n \leqslant N, l+m+n\right. \text { is even, } \\
& m+n \geqslant l, n+l \geqslant m, l+m \geqslant n\},  \tag{102}\\
& B_{N}=\left\{(l, m, n) \in \mathbb{Z}_{\geqslant 0}^{3} \mid l, m, n \leqslant N, l+m+n\right. \text { is even, } \\
& m+n>N+1, n+l>N+1, l+m>N+1\} \tag{103}
\end{align*}
$$

The set $B_{N}$ is thus a subset of $A_{N}$.
Also, for two given positive integers $m$ and $n$, such that $m+n>N+1$, we introduce the number $r_{N_{m n}}$ defined by

$$
\begin{equation*}
r_{N_{m n}}=\min \left(R_{N_{m n}}\right), \tag{104}
\end{equation*}
$$

where

$$
\begin{equation*}
R_{N_{m n}}=\left\{r \in \mathbb{Z}_{>0}:(m+n-N-1) / 2 \leqslant r \leqslant \min (m, n)\right\} \tag{105}
\end{equation*}
$$

We can now formulate the following

Proposition 5.2. Given two functionals $F, K \in \bar{C}^{\infty}\left(\mathcal{F}_{N}\right)$, the Poisson bracket $\{,\}_{g}$ can be expressed in the following way:

$$
\begin{equation*}
\{F, K\}_{g}=\{F, K\}_{g_{\perp}}+\{F, K\}_{g_{\|}} . \tag{106}
\end{equation*}
$$

In the expression 106) one has

$$
\begin{equation*}
\{F, K\}_{g_{\perp}}=\sqrt{\frac{\beta_{e}}{2}} \sum_{l, m, n=0}^{N} W_{(N) l}^{m n} \int_{\mathcal{D}} d^{3} x g_{l}\left[F_{g_{m}}, K_{g_{n}}\right] \tag{107}
\end{equation*}
$$

where

$$
W_{(N) l}^{m n}= \begin{cases}\frac{\sqrt{l!m!n!}}{\left(\frac{l+m-n}{2}\right)!\left(\frac{n+l-m}{2}\right)!\left(\frac{m+n-l}{2}\right)!}, & i f(l, m, n) \in A_{N} \backslash B_{N},  \tag{108}\\ \frac{N!}{N+1} \sum_{i=0}^{N} \frac{H_{l}\left(\lambda_{i}\right)}{H_{N}^{2}\left(\lambda_{i}\right)} \sum_{r=0}^{r_{N} m_{n}-1} \frac{m!}{(m-r)!} \frac{n!}{(n-r)!r!} \frac{H_{m+n-2 r}\left(\lambda_{i}\right)}{\sqrt{l!m!n!}} \\ +\frac{\sqrt{l!m!n!}}{\left(\frac{l+m-n}{2}\right)!\left(\frac{n+l-m}{2}\right)!\left(\frac{m+n-l}{2}\right)!}, & i f(l, m, n) \in B_{N}, \\ 0 \quad \text { otherwise }\end{cases}
$$

and

$$
\begin{equation*}
\{F, K\}_{g_{\|}}=-\sum_{m, n=0}^{N} S_{N_{m n}} \int_{\mathcal{D}} d^{3} x F_{g_{m}} \frac{\partial K_{g_{n}}}{\partial z} . \tag{109}
\end{equation*}
$$

The proof of Proposition 5.2 is provided in AppendixB.
5.1. Poisson bracket $\{,\}_{g}$ vs. bilinear structure obtained from truncated series

The closed fluid system $\sqrt{63}$ - 65 can be obtained from the drift-kinetic system (2)-(4), assuming that $g$, at any time $t$, is a truncated Hermite series, retaining only the moments from order 0 up to order $N$. Therefore it is natural to ask whether the Hamiltonian and the Poisson bracket of the fluid model (63)-65) can be obtained from those of the drift-kinetic model, by replacing $g$, in the latter, with its truncated series, or, more precisely, by restricting the Poisson algebra of observables, to functionals of Hermite series truncated at the moment of order $N$, and then expressing the outcome of the resulting restricted algebra in terms of the fluid moment variables. A delicate point in this operation concerns the Poisson bracket. Indeed, in order for the set of functionals of the truncated Hermite series to be closed under the restricted bilinear algebra operation (descending from the drift-kinetic Poisson bracket), one has to restrict the Poisson operator of the drift-kinetic bracket, taking its composition with the projector onto the subspace of truncated Hermite series. Because this operation is not invertible, the resulting bilinear form can fail to satisfy the Jacobi identity and therefore not be a Poisson bracket.

In order to carry out this analysis, let us first introduce the operator $\mathcal{P}_{N}$ : $\mathcal{G} \rightarrow \Gamma_{N}$ defined by

$$
\begin{equation*}
\mathcal{P}_{N}(g)=\sum_{n=0}^{N} g_{n} \frac{H_{n}}{\sqrt{n!}} F_{e q} \tag{110}
\end{equation*}
$$

The operator $\mathcal{P}_{N}$ projects a generalized perturbed distribution function $g$ onto the subspace

$$
\begin{equation*}
\Gamma_{N}=\left\{g \in \mathcal{G} \mid g_{i}=0 \quad \forall i \geqslant N+1\right\} \tag{111}
\end{equation*}
$$

whose elements are the truncated Hermite series. Each $\tilde{g} \in \Gamma_{N}$ can be written as

$$
\begin{equation*}
\tilde{g}(\mathbf{x}, v)=\sum_{n=0}^{N} g_{n}(\mathbf{x}) \frac{H_{n}(v)}{\sqrt{n!}} F_{e q}(v) \tag{112}
\end{equation*}
$$

with $\left(g_{0}, g_{1}, \cdots, g_{N}\right) \in \mathcal{F}_{N}$ satisfying

$$
\begin{equation*}
g_{n}=\frac{1}{\sqrt{n!}} \int_{-\infty}^{+\infty} d v H_{n} \tilde{g}, \quad n=0,1, \cdots, N \tag{113}
\end{equation*}
$$

We want to identify the restriction of the Poisson algebra of observables of the drift-kinetic system, consisting of $\bar{C}^{\infty}(\mathcal{G})$, to the algebra $\bar{C}^{\infty}\left(\Gamma_{N}\right)$ of the truncated Hermite series, with the corresponding restricted bilinear algebra operation. In particular, we want to determine the expression of the restricted Hamiltonian functional $\mathbb{H}: \mathcal{F}_{N} \rightarrow \mathbb{R}$ defined by

$$
\begin{equation*}
\mathbb{H}\left(g_{0}, g_{1}, \cdots, g_{N}\right)=H_{d k} \circ \mathcal{P}_{N}(g)=H_{d k}(\tilde{g}) \tag{114}
\end{equation*}
$$

for $\tilde{g} \in \Gamma_{N}$ given by Eq. 112 , and of the bilinear operator $\llbracket, \rrbracket: \bar{C}^{\infty}\left(\mathcal{F}_{N}\right) \times$ $\bar{C}^{\infty}\left(\mathcal{F}_{N}\right) \rightarrow \bar{C}^{\infty}\left(\mathcal{F}_{N}\right)$ given by

$$
\begin{align*}
& \llbracket F, K \rrbracket=\left\langle\frac{\delta \bar{F} \circ \mathcal{P}_{N}(g)}{\delta g}, J_{d k} \circ \mathcal{P}_{N}(g) \frac{\delta \bar{K} \circ \mathcal{P}_{N}(g)}{\delta g}\right\rangle \\
& =\left\langle\frac{\delta \bar{F}}{\delta \tilde{g}}, J_{d k}(\tilde{g}) \frac{\delta \bar{K}}{\delta \tilde{g}}\right\rangle \tag{115}
\end{align*}
$$

where $\bar{F}, \bar{K} \in \bar{C}^{\infty}\left(\Gamma_{N}\right)$ and $F\left(g_{0}, g_{1}, \cdots, g_{N}\right)=\bar{F} \circ \mathcal{P}_{N}(g)=\bar{F}(\tilde{g}), K\left(g_{0}, g_{1}, \cdots, g_{N}\right)=$ $\bar{K} \circ \mathcal{P}_{N}(g)=\bar{K}(\tilde{g})$, with $\tilde{g} \in \Gamma_{N}$.

In Eqs. 114 and 115, the functional $H_{d k}$, the Poisson bracket $\{,\}_{d k}$ and the Poisson operator $J_{d k}$ are those of the drift-kinetic model given by Eqs. (49), (50) and (54), respectively. In Eq. (115), as above mentioned, we are
considering a bilinear form involving the composition of the Poisson operator $J_{d k}$, with the projector $\mathcal{P}_{N}$. The composition of these two operators, in general, nian (86) and the Poisson bracket 106 of the fluid model.

The wanted expressions for $\mathbb{H}$ and $\llbracket, \rrbracket$ are provided by the next two Propositions.

Proposition 5.3. The functional $H$, defined by the relation (114), coincides with the Hamiltonian functional $H$ of the fluid model given in Eq. 886) of Proposition 4.1.

Proof. From Eqs. (114), 49) and 112 one has

$$
\begin{align*}
& \mathbb{H}\left(g_{0}, g_{1}, \cdots, g_{N}\right)=H_{d k}(\tilde{g})=\frac{1}{2} \int_{\mathcal{D}} d^{3} x \int_{-\infty}^{+\infty} d v\left(\sum_{n, n^{\prime}=0}^{N} g_{n} \frac{H_{n}}{\sqrt{n!}} g_{n^{\prime}} \frac{H_{n^{\prime}}}{\sqrt{n^{\prime}!}} F_{e q}\right.  \tag{116}\\
& \left.-\sqrt{\frac{2}{\beta_{e}}} \sum_{n=0}^{N} g_{n} \frac{H_{n}}{\sqrt{n!}} F_{e q}\left(\phi_{d k}\left[\sum_{n^{\prime}=0}^{N} g_{n^{\prime}} \frac{H_{n^{\prime}}}{\sqrt{n^{\prime}!}} F_{e q}\right]-v A_{d k}\left[\sum_{n^{\prime}=0}^{N} g_{n^{\prime}} \frac{H_{n^{\prime}}}{\sqrt{n^{\prime}!}} F_{e q}\right]\right)\right)  \tag{117}\\
& =\frac{1}{2} \int_{\mathcal{D}} d^{3} x\left(\sum_{n=0}^{N} g_{n}^{2}-\sqrt{\frac{2}{\beta_{e}}}\left(g_{0} \phi_{f l}\left[g_{0}\right]-g_{1} A_{f l}\left[g_{1}\right]\right)\right)=H\left(g_{0}, g_{1}, \cdots, g_{N}\right) . \tag{118}
\end{align*}
$$

To go from the expressions 116 - 117 to the expressions 118 we made use of the orthogonality relation 20 for Hermite polynomials and of the relations 41, permitting to express electromagnetic potentials in terms of fluid moments.

Proposition 5.4. Given two functionals $F, K \in \bar{C}^{\infty}\left(\mathcal{F}_{N}\right)$, the bilinear operator
$\llbracket, \rrbracket$, defined by the relation 115), can be written as

$$
\begin{equation*}
\llbracket F, K \rrbracket=\llbracket F, K \rrbracket_{\perp}+\llbracket F, K \rrbracket_{\|}, \tag{119}
\end{equation*}
$$

where

$$
\begin{equation*}
\llbracket F, K \rrbracket_{\perp}=\sqrt{\frac{\beta_{e}}{2}} \sum_{l, m, n=0}^{N} \mathbb{W}_{(N) l}^{m n} \int_{\mathcal{D}} d^{3} x g_{l}\left[F_{g_{m}}, K_{g_{n}}\right], \tag{120}
\end{equation*}
$$

with

$$
\mathbb{W}_{(N) l}^{m n}= \begin{cases}\frac{\sqrt{l!m!n!}}{\left(\frac{l+m-n}{2}\right)!\left(\frac{n+l-m}{2}\right)!\left(\frac{m+n-l}{2}\right)!}, & \text { if }(l, m, n) \in A_{N},  \tag{121}\\ 0 \quad \text { otherwise }\end{cases}
$$

and

$$
\begin{equation*}
\llbracket F, K \rrbracket_{\|}=-\sum_{m, n=0}^{N} S_{N_{m n}} \int_{\mathcal{D}} d^{3} x F_{g_{m}} \frac{\partial K_{g_{n}}}{\partial z} . \tag{122}
\end{equation*}
$$

Proof. From Eq. 115), using Eqs. (50) and 112), we obtain

$$
\begin{align*}
& \llbracket F, K \rrbracket=\left\langle\frac{\delta \bar{F}}{\delta \tilde{g}}, J_{d k}(\tilde{g}) \frac{\delta \bar{K}}{\delta \tilde{g}}\right\rangle \\
& =\int_{\mathcal{D}} d^{3} x \int_{-\infty}^{+\infty} d v\left(\sqrt{\frac{\beta_{e}}{2}} \tilde{g}\left[\bar{F}_{\tilde{g}}, \bar{K}_{\tilde{g}}\right]-v F_{e q} \bar{F}_{\tilde{g}} \frac{\partial \bar{K}_{\tilde{g}}}{\partial z}\right) . \tag{123}
\end{align*}
$$

From the definitions of functional derivatives (42) and (44), using also the relations (112) and (113), one obtains the following chain rule for functional derivatives:

$$
\begin{equation*}
\frac{\delta \bar{F}}{\delta \tilde{g}}(\mathbf{x}, v)=\sum_{n=0}^{N} \frac{H_{n}(v)}{\sqrt{n!}} \frac{\delta F}{\delta g_{n}}(\mathbf{x}), \tag{124}
\end{equation*}
$$

for $\bar{F}(\tilde{g})=F\left(g_{0}, g_{1}, \cdots, g_{N}\right)$. Inserting the expressions 112 and 124 into Eq. (123) yields

$$
\begin{align*}
& \llbracket F, K \rrbracket=\int_{\mathcal{D}} d^{3} x \int_{-\infty}^{+\infty} d v\left(\sqrt{\frac{\beta_{e}}{2}} \sum_{l, m, n=0}^{N} F_{e q} \frac{H_{l} H_{m} H_{n}}{\sqrt{l!m!n!}} g_{l}\left[F_{g_{m}}, K_{g_{n}}\right]\right. \\
& \left.-\sum_{m, n=0}^{N} F_{e q} \frac{H_{1} H_{m} H_{n}}{\sqrt{m!n!}} F_{g_{m}} \frac{\partial K_{g_{n}}}{\partial z}\right), \tag{125}
\end{align*}
$$

where we also made use of the relation $v=H_{1}(v)$.

At this point, we can apply to the expression (125) the following identity for Hermite polynomials 45]:

$$
\begin{align*}
& \int_{-\infty}^{+\infty} d v F_{e q}(v) \frac{H_{l}(v) H_{m}(v) H_{n}(v)}{\sqrt{l!m!n!}} \\
& =\left\{\begin{array}{l}
\frac{\sqrt{l!m!n!}}{\left(\frac{l+m-n}{2}\right)!\left(\frac{n+l-m}{2}\right)!\left(\frac{m+n-l}{2}\right)!}, \quad \text { if } l+m+n \text { is even and } l+m \geqslant n, m+n \geqslant l, n+l \geqslant m \\
0 \quad \text { otherwise. }
\end{array}\right. \tag{127}
\end{align*}
$$

In particular, with regard to the expression in the second line of Eq. 125, we have that

$$
\begin{equation*}
\int_{-\infty}^{+\infty} d v F_{e q} \frac{H_{1} H_{m} H_{n}}{\sqrt{m!n!}}=\frac{\sqrt{m!n!}}{\left(\frac{1+m-n}{2}\right)!\left(\frac{m+n-1}{2}\right)!\left(\frac{n+1-m}{2}\right)!} \tag{128}
\end{equation*}
$$

when $m+n+1$ is even and the three conditions

$$
\begin{equation*}
1+m \geqslant n, \quad m+n \geqslant 1, \quad n+1 \geqslant m \tag{129}
\end{equation*}
$$

are satisfied. Otherwise,

$$
\begin{equation*}
\int_{-\infty}^{+\infty} d v F_{e q} \frac{H_{1} H_{m} H_{n}}{\sqrt{m!n!}}=0 \tag{130}
\end{equation*}
$$

The conditions 129 , together with the constraint for $m+n+1$ of being even, imply

$$
\begin{equation*}
m-1 \leqslant n \leqslant m+1 \tag{131}
\end{equation*}
$$

From the relation (131) it follows that the integral 128 can be non-zero only if $n=m-1, n=m$ or $n=m+1$. However, the case $n=m$ implies that $m+n+1=2 m+1$ is odd. Therefore, also in this case the integral is zero. Using Eq. 128, it follows that

$$
\begin{equation*}
\int_{-\infty}^{+\infty} d v F_{e q} \frac{H_{1} H_{m} H_{n}}{\sqrt{m!n!}}=\sqrt{m+1} \delta_{m+1, n}+\sqrt{m} \delta_{m-1, n}=S_{N_{m n}} \tag{132}
\end{equation*}
$$

Using Eqs. 126) (recalling that $l, m$ and $n$ go from 0 to $N$ ) and 132 into Eq.
(125), we can write

$$
\begin{align*}
& \llbracket F, K \rrbracket=\sqrt{\frac{\beta_{e}}{2}} \sum_{l, m, n=0}^{N} \mathbb{W}_{(N) l}^{m n} \int_{\mathcal{D}} d^{3} x g_{l}\left[F_{g_{m}}, K_{g_{n}}\right] \\
& -\sum_{m, n=0}^{N} S_{N_{m n}} \int_{\mathcal{D}} d^{3} x F_{g_{m}} \frac{\partial K_{g_{n}}}{\partial z} \tag{133}
\end{align*}
$$

where

$$
\mathbb{W}_{(N) l}^{m n}=\left\{\begin{array}{l}
\frac{\sqrt{l!m!n!}}{\left(\frac{l+m-n}{2}\right)!\left(\frac{n+l-m}{2}\right)!\left(\frac{m+n-l}{2}\right)!}, \quad \text { if }(l, m, n) \in A_{N}  \tag{134}\\
0 \quad \text { otherwise }
\end{array}\right.
$$

Therefore we can conclude

$$
\begin{equation*}
\llbracket F, K \rrbracket=\llbracket F, K \rrbracket_{\perp}+\llbracket F, K \rrbracket_{\|} . \tag{135}
\end{equation*}
$$

From Proposition 5.3 it follows that, for any $N$, the Hamiltonian of the fluid model can be derived from the Hamiltonian of the drift-kinetic model, by replacing, in the latter, the generalized perturbed distribution function with its truncated series of order $N$. On the other hand, Proposition 5.4 shows that the Poisson bracket of the fluid model cannot be obtained from the Poisson bracket of the drift-kinetic model, by considering a Poisson operator depending only on truncated Hermite series, and by restricting the set of observables to $\bar{C}^{\infty}\left(\mathcal{F}_{N}\right)$. More in detail, by comparing Proposition 5.2 with Proposition 5.4 one sees that

$$
\begin{equation*}
\{,\}_{g_{\|}}=\llbracket, \rrbracket_{\|} . \tag{136}
\end{equation*}
$$

However, the difference arises in the "perpendicular' components $\{,\}_{g_{\perp}}$ and $\llbracket, \rrbracket_{\perp}$, characterized by the coefficients $W_{(N) l}^{m n}$ and $\mathbb{W}_{(N) l}^{m n}$. More specifically, one sees that in general $W_{(N) l}^{m n}=\mathbb{W}_{(N) l}^{m n}$ except when $(l, m, n) \in B_{N}$, i.e. when $l$, $m$ and $n$ are all sufficiently 'large'. For most of the values of $l, m$ and $n$, the coefficients of the bilinear form, obtained from the truncated series, and the


Figure 1: Plot illustrating, in the $l m n$ space, the points where $W_{(N) l}^{m n}=\mathbb{W}_{(N) l}^{m n}=0$, the points of the set $A_{N} \backslash B_{N}$ and the points of the set $B_{N}$, for the case $N=5$. The yellow points are those for which $W_{(N) l}^{m n}=\mathbb{W}_{(N) l}^{m n}=0$. The blue points are the points of the set $A_{N} \backslash B_{N}$. At such points one has $W_{(N) l}^{m n}=\mathbb{W}_{(N) l}^{m n}$. The red color indicate the points of the set $B_{N}$. At such points $W_{(N) l}^{m n} \neq \mathbb{W}_{(N) l}^{m n}$. It is at such points that the coefficients of $\{,\}_{g_{\perp}}$ differ from those of $\llbracket, \rrbracket \perp$.
coefficients of the Poisson bracket, are nevertheless identical. It is in this sense, that we previously stated that the bilinear form $\llbracket, \rrbracket$ and the Poisson bracket $\{,\}_{g}$ are "very similar". Note also that the expression for the coefficients $\mathbb{W}_{(N) l}^{m n}$,

## 6. Example : a three-moment model

In order to exemplify the results presented in Secs. 4 and 5 we treat here in detail the case $N=2$. In this case, the resulting fluid model evolves the three moments $g_{0}, g_{1}$ and $g_{2}$, which, as anticipated in 3.1, are proportional to the fluctuations of electron gyrocenter density, parallel canonical momentum and given by Eq. 121, does not depend on $N$, whereas this is the case for $W_{(N) l}^{m n}$ when $(l, m, n) \in B_{N}$, as shown by Eq. 108 .
The distribution, in the $l m n$ space, of the points where $W_{(N) l}^{m n}=\mathbb{W}_{(N) l}^{m n}=0$, of the points of $A_{N} \backslash B_{N}\left(\right.$ where $\left.W_{(N) l}^{m n}=\mathbb{W}_{(N) \downarrow}^{m n}\right)$ and of the points of $B_{N}$ (where $\left.W_{(N) l}^{m n} \neq \mathbb{W}_{(N) l}^{m n}\right)$ is illustrated, for the case $N=5$, in Fig. 1 parallel temperature, respectively. The closure adopted in this case is $g_{3}=0$,
which amounts to set the parallel heat flux fluctuations of the electron gyrocenters equal to zero.

The matrix $S_{2}$ reads

$$
S_{2}=\left(\begin{array}{ccc}
0 & 1 & 0  \tag{137}\\
1 & 0 & \sqrt{2} \\
0 & \sqrt{2} & 0
\end{array}\right)
$$

and its eigenvalues, corresponding to the zeros of $H_{3}(x)=x^{3}-3 x$, are

$$
\begin{equation*}
\left(\lambda_{0}, \lambda_{1}, \lambda_{2}\right)=(-\sqrt{3}, 0, \sqrt{3}) \tag{138}
\end{equation*}
$$

From Eq. 86) of Proposition 4.1, we have that the Hamiltonian of the fluid model is given by

$$
\begin{equation*}
H\left(g_{0}, g_{1}, g_{2}\right)=\frac{1}{2} \int_{\mathcal{D}} d^{3} x\left(g_{0}^{2}+g_{1}^{2}+g_{2}^{2}-\sqrt{\frac{2}{\beta_{e}}}\left(g_{0} \phi_{f l}\left[g_{0}\right]-g_{1} A_{f l}\left[g_{1}\right]\right)\right) \tag{139}
\end{equation*}
$$

For the Poisson bracket, we consider the expression following from Eq. 106 of Proposition 5.2. For two functionals $F, K \in \bar{C}^{\infty}\left(\mathcal{F}_{2}\right)$ the Poisson bracket can thus be written as

$$
\begin{align*}
& \{F, K\}_{g}=\sqrt{\frac{\beta_{e}}{2}} \sum_{l, m, n=0}^{2} W_{(2) l}^{m n} \int_{\mathcal{D}} d^{3} x g_{l}\left[F_{g_{m}}, K_{g_{n}}\right] \\
& -\sum_{m, n=0}^{2} S_{2_{m n}} \int_{\mathcal{D}} d^{3} x F_{g_{m}} \frac{\partial K_{g_{n}}}{\partial z} \tag{140}
\end{align*}
$$

where

$$
\begin{gather*}
W_{(2) l}^{m n}=0, \quad \text { for } 0 \leqslant l, m, n \leqslant 2 \text { except for } \\
\\
W_{(2) 0}^{00}=1, \\
W_{(2) 0}^{11}=W_{(2) 1}^{10}=W_{(2) 1}^{01}=1,  \tag{141}\\
\\
W_{(2) 0}^{22}=W_{(2) 2}^{20}=W_{(2) 2}^{02}=1, \\
\\
W_{(2) 2}^{11}=W_{(2) 1}^{12}=W_{(2) 1}^{21}=\sqrt{2}, \\
\\
W_{(2) 2}^{22}=\frac{1}{\sqrt{2}} .
\end{gather*}
$$

In a less compact form, the Poisson bracket can be written as

$$
\begin{align*}
& \{F, K\}_{g}=\int_{\mathcal{D}} d^{3} x\left(\sqrt { \frac { \beta _ { e } } { 2 } } \left(g_{0}\left(\left[F_{g_{0}}, K_{g_{0}}\right]+\left[F_{g_{1}}, K_{g_{1}}\right]+\left[F_{g_{2}}, K_{g_{2}}\right]\right)\right.\right. \\
& +g_{1}\left(\left[F_{g_{0}}, K_{g_{1}}\right]+\left[F_{g_{1}}, K_{g_{0}}\right]+\sqrt{2}\left[F_{g_{1}}, K_{g_{2}}\right]+\sqrt{2}\left[F_{g_{2}}, K_{g_{1}}\right]\right) \\
& \left.+g_{2}\left(\left[F_{g_{0}}, K_{g_{2}}\right]+\left[F_{g_{2}}, K_{g_{0}}\right]+\sqrt{2}\left[F_{g_{1}}, K_{g_{1}}\right]+\frac{1}{\sqrt{2}}\left[F_{g_{2}}, K_{g_{2}}\right]\right)\right)  \tag{142}\\
& \left.-F_{g_{0}} \frac{\partial K_{g_{1}}}{\partial z}-F_{g_{1}} \frac{\partial K_{g_{0}}}{\partial z}-\sqrt{2} F_{g_{1}} \frac{\partial K_{g_{2}}}{\partial z}-\sqrt{2} F_{g_{2}} \frac{\partial K_{g_{1}}}{\partial z}\right) .
\end{align*}
$$

The Hamiltonian 139 and the Poisson bracket 142 yield the following equations of motion:

$$
\begin{align*}
& \frac{\partial g_{0}}{\partial t}+\left[\phi, g_{0}\right]-\left[A, g_{1}\right]+\frac{\partial g_{1}}{\partial z}+\sqrt{\frac{2}{\beta_{e}}} \frac{\partial A}{\partial z}=0 \\
& \frac{\partial g_{1}}{\partial t}+\left[\phi, g_{1}\right]-\left[A, g_{0}\right]-\sqrt{2}\left[A, g_{2}\right]+\frac{\partial g_{0}}{\partial z}+\sqrt{2} \frac{\partial g_{2}}{\partial z}-\sqrt{\frac{2}{\beta_{e}}} \frac{\partial \phi}{\partial z}=0  \tag{143}\\
& \frac{\partial g_{2}}{\partial t}+\left[\phi, g_{2}\right]-\sqrt{2}\left[A, g_{1}\right]+\sqrt{2} \frac{\partial g_{1}}{\partial z}+\sqrt{2} \sqrt{\frac{2}{\beta_{e}}} \frac{\partial A}{\partial z}=0
\end{align*}
$$

where $\phi=\phi_{f l}\left[g_{0}\right]$ and $A=A_{f l}\left[g_{1}\right]$. The system 143 coincides namely with the fluid system $\sqrt{63}$ in the case $N=2$, obtained by truncating the hierarchy of equations $\sqrt{56})-(59)$ with $g_{3}=0$.

Also, from Eq. (85) one obtains

$$
U_{2}=\left(\begin{array}{ccc}
\frac{1}{\sqrt{6}} & \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{6}}  \tag{144}\\
-\frac{1}{\sqrt{2}} & 0 & \frac{1}{\sqrt{2}} \\
\frac{1}{\sqrt{3}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{3}}
\end{array}\right)
$$

(note a difference with respect to Eq. (87) of Ref. [33], due to a different numbering of the eigenvalues and a different choice in the normalization constant). From the matrix $U_{2}$ one can, by means of Eq. (67), obtain the expressions for the variables $\left(G_{0}, G_{1}, G_{2}\right) \in \mathcal{G}_{2}$, which read

$$
\begin{align*}
G_{0} & =\frac{g_{0}}{\sqrt{6}}-\frac{g_{1}}{\sqrt{2}}+\frac{g_{2}}{\sqrt{3}}, \\
G_{1} & =\sqrt{\frac{2}{3}} g_{0}-\frac{g_{2}}{\sqrt{3}},  \tag{145}\\
G_{2} & =\frac{g_{0}}{\sqrt{6}}+\frac{g_{1}}{\sqrt{2}}+\frac{g_{2}}{\sqrt{3}} .
\end{align*}
$$

In terms of these variables, following Eqs. (68)-(70), the system (143) can be rewritten as

$$
\begin{align*}
& \frac{\partial G_{0}}{\partial t}+\left[\phi+\sqrt{3} A, G_{0}\right]-\sqrt{3} \frac{\partial G_{0}}{\partial z}+\sqrt{\frac{2}{\beta_{e}}} \frac{\partial}{\partial z}\left(\sqrt{\frac{3}{2}} A+\frac{\phi}{\sqrt{2}}\right)=0 \\
& \frac{\partial G_{1}}{\partial t}+\left[\phi, G_{1}\right]=0  \tag{146}\\
& \frac{\partial G_{2}}{\partial t}+\left[\phi-\sqrt{3} A, G_{2}\right]+\sqrt{3} \frac{\partial G_{2}}{\partial z}+\sqrt{\frac{2}{\beta_{e}}} \frac{\partial}{\partial z}\left(\sqrt{\frac{3}{2}} A-\frac{\phi}{\sqrt{2}}\right)=0
\end{align*}
$$

where $\phi=\phi_{f l}\left[U_{2_{0 l}} G_{l}\right]$ and $A=A_{f l}\left[U_{2_{1 l}} G_{l}\right]$.
The formulation 146 puts in evidence the existence of a Lagrangian invariant, corresponding to $G_{1}$, which is simply transported by the velocity field

$$
\begin{equation*}
\mathbf{v}_{\mathbf{E} \times \mathbf{B}}=\hat{z} \times \nabla \phi, \tag{147}
\end{equation*}
$$

which identifies with the so-called $\mathbf{E} \times \mathbf{B}$ velocity field. We note that, on the basis of our definitions of the fluid moments, $G_{1}$ is proportional to $n_{e}-T_{\| e} / 2$, where $n_{e}$ and $T_{\| e}$ are normalized fluctuations of electron gyrocenter density and parallel temperature, respectively. We observe that the relation $T_{\| e}=$ $2 n_{e}$ (implying $G_{1}=0$ ) expresses an adiabatic law for the parallel electron temperature. Therefore, if $T_{\| e}=2 n_{e}$ at $t=0, G_{1}=0$ is a solution of Eq. 146, which expresses the adiabatic relation between parallel electron temperature and density at all times. If $T_{\| e} \neq 2 n_{e}$ at $t=0$, on the other hand, such initial departure from adiabaticity is conserved along the flow of the $\mathbf{E} \times \mathbf{B}$ velocity at all times. From Eqs. 146 it can also be seen that, in the 2D limit in which the fields $G_{0}, G_{1}$ and $G_{2}$ do not depend on the $z$ coordinate, all these three fields become Lagrangian invariants, as already pointed out in Refs. [32, 33].

We now apply, to the case $N=2$, the approach based on the truncated series. From Proposition 5.3 we obtain that the truncated Hermite series $\tilde{g} \in \Gamma_{2}$, given by

$$
\begin{equation*}
\tilde{g}(\mathbf{x}, v)=\sum_{n=0}^{2} g_{n}(\mathbf{x}) \frac{H_{n}(v)}{\sqrt{n!}} F_{e q}(v) \tag{148}
\end{equation*}
$$

when inserted into the drift-kinetic Hamiltonian $H_{d k}$, yields the functional $\mathbb{H}\left(g_{0}, g_{1}, g_{2}\right)=H\left(g_{0}, g_{1}, g_{2}\right)$, where $H$ is the fluid Hamiltonian 139). On
the other hand, the expression of the bilinear form $\llbracket, \rrbracket$, for two functionals $F, K \in \bar{C}^{\infty}\left(\mathcal{F}_{2}\right)$, follows from Proposition 5.4 and is given by

$$
\begin{align*}
& \llbracket F, K \rrbracket=\sqrt{\frac{\beta_{e}}{2}} \sum_{l, m, n=0}^{2} \mathbb{W}_{(2) l}^{m n} \int_{\mathcal{D}} d^{3} x g_{l}\left[F_{g_{m}}, K_{g_{n}}\right] \\
& -\sum_{m, n=0}^{2} S_{2_{m n}} \int_{\mathcal{D}} d^{3} x F_{g_{m}} \frac{\partial K_{g_{n}}}{\partial z} \tag{149}
\end{align*}
$$

where

$$
\begin{align*}
& \mathbb{W}_{(2) l}^{m n}= 0, \quad \text { for } 0 \leqslant l, m, n \leqslant 2 \text { except for } \\
& \mathbb{W}_{(2) 0}^{00}=1, \\
& \mathbb{W}_{(2) 0}^{11}=\mathbb{W}_{(2) 1}^{10}=\mathbb{W}_{(2) 1}^{01}=1 \\
& \mathbb{W}_{(2) 0}^{22}=\mathbb{W}_{(2) 2}^{20}=\mathbb{W}_{(2) 2}^{02}=1  \tag{150}\\
& \mathbb{W}_{(2) 2}^{11}=\mathbb{W}_{(2) 1}^{12}=\mathbb{W}_{(2) 1}^{21}=\sqrt{2} \\
& \mathbb{W}_{(2) 2}^{22}=2 \sqrt{2}
\end{align*}
$$

Comparing Eq. 142 with Eq. 149 and Eq. 141 with Eq. 150 , it emerges that the bilinear form $\llbracket, \rrbracket$ differs from the Poisson bracket $\{,\}_{g}$ only by the coefficient $\mathbb{W}_{(2) 2}^{22}=2 \sqrt{2}$, which is not equal to $W_{(2) 2}^{22}=1 / \sqrt{2}$. Indeed, the set $B_{2}$ is given by $B_{2}=\{(2,2,2)\}$ and, from Eq. 108) of Proposition 5.2, it follows that only the expression for $W_{(2) 2}^{22}=1 / \sqrt{2}$ differs from the expression of the corresponding coefficient of $\llbracket$, 】.

Remark 6.1. It turns out that the bilinear form 149 is antisymmetric, satisfies the Leibniz rule but it does not satisfy the Jacobi identity. Therefore, the form (149) is not a Poisson bracket. In order to see this, we first remark that the Jacobi identity for the bilinear form $\llbracket, \rrbracket$ can be written as

$$
\begin{equation*}
\llbracket \llbracket F, K \rrbracket_{\perp}, E \rrbracket_{\perp}+\llbracket \llbracket F, K \rrbracket_{\perp}, E \rrbracket_{\|}+\llbracket \llbracket F, K \rrbracket_{\|}, E \rrbracket_{\perp}+\llbracket \llbracket F, K \rrbracket_{\|}, E \rrbracket_{\|}+\circlearrowleft=0, \tag{151}
\end{equation*}
$$

where the symbol $\circlearrowleft$ indicates the additional terms obtained by cyclic permutation of $F, K$ and $E$. The identity 151 must be valid for all functionals $F, K, E \in \bar{C}^{\infty}\left(\mathcal{F}_{2}\right)$.

In order to investigate the relation we can make use of a result presented in Sec. 3.2 of Ref. [23]. According to such result, one has that

$$
\begin{equation*}
\llbracket \llbracket F, K \rrbracket_{\|}, E \rrbracket_{\perp}+\llbracket \llbracket F, K \rrbracket_{\|}, E \rrbracket_{\|}+\circlearrowleft=0, \tag{152}
\end{equation*}
$$

which follows from the fact that $\llbracket, \rrbracket$ || is itself a Poisson bracket with constant Poisson operator. With regard to the expression

$$
\begin{equation*}
\llbracket \llbracket F, K \rrbracket_{\perp}, E \rrbracket_{\|}+\circlearrowleft \tag{153}
\end{equation*}
$$

it vanishes for all $F, K, E \in \bar{C}^{\infty}\left(\mathcal{F}_{2}\right)$ if and only if the condition

$$
\begin{equation*}
S_{2_{i n}} \mathbb{W}_{(2) i}^{j m}=S_{2_{i j}} \mathbb{W}_{(2) i}^{m n}=S_{2_{i m}} \mathbb{W}_{(2) i}^{n j}, \quad \text { for } 0 \leqslant j, m, n \leqslant 2 \tag{154}
\end{equation*}
$$

(where the sum over the repeated index $i$ is understood) is satisfied. However, if we consider the case $j=2, m=2, n=1$, from Eqs. 137) and 150, we obtain

$$
\begin{equation*}
S_{2_{i 1}} \mathbb{W}_{(2) i}^{22}=5, \quad S_{2_{i 2}} \mathbb{W}_{(2) i}^{21}=2 \tag{155}
\end{equation*}
$$

Therefore $S_{2_{i 1}} \mathbb{W}_{(2) i}^{22} \neq S_{2_{i 2}} \mathbb{W}_{(2) i}^{21}$ and the condition 154 is not satisfied. Thus, the expression 153 does not vanish for all $F, K, E \in \bar{C}^{\infty}\left(\mathcal{F}_{2}\right)$. Similarly, one can show that the

$$
\begin{equation*}
\llbracket \llbracket F, K \rrbracket_{\perp}, E \rrbracket_{\perp}+\circlearrowleft=0, \tag{156}
\end{equation*}
$$

does not hold for all $F, K, E \in \bar{C}^{\infty}\left(\mathcal{F}_{2}\right)$. An effective way to see this is to take advantage from a result in Sec. 2.3 of Ref. [20] (actually, the result of Ref. [20] refers to a purely 2D domain, but this difference is irrelevant for the present purpose). According to this result, for a given $N$, antisymmetric bilinear operations of the form satisfy the Jacobi identity if and only if all the $N+1$ matrices $\mathbb{W}^{(k)}$, whose elements of row $i$ and column $j$ are defined by

$$
\begin{equation*}
\mathbb{W}_{i j}^{(k)}=\mathbb{W}_{(N) i}^{j k}, \quad i, j, k=0, \cdots, N \tag{157}
\end{equation*}
$$

pairwise commute.
In the present case $N=2$, from Eq. (150, one obtains that the three
matrices $\mathbb{W}^{(0)}, \mathbb{W}^{(1)}$ and $\mathbb{W}^{(2)}$ are given by
$\mathbb{W}^{(0)}=\left(\begin{array}{ccc}1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1\end{array}\right), \quad \mathbb{W}^{(1)}=\left(\begin{array}{ccc}0 & 1 & 0 \\ 1 & 0 & \sqrt{2} \\ 0 & \sqrt{2} & 0\end{array}\right), \quad \mathbb{W}^{(2)}=\left(\begin{array}{ccc}0 & 0 & 1 \\ 0 & \sqrt{2} & 0 \\ 1 & 0 & 2 \sqrt{2}\end{array}\right)$.

Clearly, both $\mathbb{W}^{(1)}$ and $\mathbb{W}^{(2)}$ commute with $\mathbb{W}^{(0)}$, which is the identity matrix. However, one has

$$
\mathbb{W}^{(1)} \mathbb{W}^{(2)}=\left(\begin{array}{ccc}
0 & \sqrt{2} & 0  \tag{159}\\
\sqrt{2} & 0 & 5 \\
0 & 2 & 0
\end{array}\right), \quad \mathbb{W}^{(2)} \mathbb{W}^{(1)}=\left(\begin{array}{ccc}
0 & \sqrt{2} & 0 \\
\sqrt{2} & 0 & 2 \\
0 & 5 & 0
\end{array}\right)
$$

Therefore, $\mathbb{W}^{(1)}$ and $\mathbb{W}^{(2)}$ do not commute and, consequently, the condition 156 Thus, the Jacobi identity would have been satisfied, which confirms that $\{,\}_{g}$ is a valid Poisson bracket.

The additional term, with respect to the expression (121), present in Eq. $\left(108\right.$ for $(l, m, n) \in B_{N}$, is what 'corrects' the coefficient $\mathbb{W}_{(N) l}^{m n}$ in order to modify the form 【, 】 into a Poisson bracket. In particular, it modifies the perpendicular form $\llbracket, \rrbracket_{\perp}$ turning it into $\{,\}_{g_{\perp}}$, which is a Poisson bracket obtained by extension of a Lie-Poisson bracket [20].

Finally, we note that, had one naively (and not correctly) used the bilinear form (149) as a Poisson bracket, and derived a set of equations using the formula $\dot{g}_{i}=\llbracket g_{i}, H \rrbracket$, for $i=0,1,2$, with $H$ given by Eq. 139), one would have nevertheless obtained the correct equations of motion 143. Indeed, the coefficient $\mathbb{W}_{(2) 2}^{22}$, which spoils the Jacobi identity, would have produced no finite terms in the equations of motion. Therefore, this is one of the examples that points out subtleties existing in the identification of Poisson brackets among 'almost

## 7. Physical interpretations of the variables $G_{0}, G_{1}, \cdots, G_{N}$. Analogy with the problem of the quantum harmonic oscillator

The set of variables $G_{0}, G_{1}, \cdots, G_{N}$ introduced with Eq. 67), proved to be useful [32] in order to find the Poisson bracket for the fluid models which, in terms of these variables, takes the simple form 73 . Such variables are related to the Casimir invariants of the perpendicular Poisson bracket $\{,\}_{g_{\perp}}$, i.e. to observables $C \in \bar{C}^{\infty}\left(\mathcal{G}_{N}\right)$ such that

$$
\begin{equation*}
\{C, F\}_{G_{\perp}}=0, \quad \forall F \in \bar{C}^{\infty}\left(\mathcal{G}_{N}\right) \tag{160}
\end{equation*}
$$

More precisely, one has that the observables $C_{i} \in \bar{C}^{\infty}\left(\mathcal{G}_{N}\right)$, with $i=0,1, \cdots, N$, and defined by

$$
\begin{equation*}
C_{i}=\int_{\mathcal{D}} d^{3} x \mathcal{C}_{i}\left(G_{i}\right), \quad \text { for } i=0,1, \cdots, N \tag{161}
\end{equation*}
$$

with $\mathcal{C}_{1}, \mathcal{C}_{2}, \cdots, \mathcal{C}_{N}$ arbitrary smooth functions, are infinite families of Casimir invariants for $\{,\}_{G_{\perp}}$. Casimir invariants of this form have been found in the 2D limit of a number of reduced fluid models, as for instance in those of Refs. [46, 24, 25, 26, 27. The associated variables $G_{0}, G_{1}, \cdots, G_{N}$ have also been used to investigate simulations of collisionless magnetic reconnection [11, 46,

27, 10, 47] or directions of spectral cascades in 2D plasma turbulence 48. In the 2D limit, such Casimir invariants are analogous to the Casimir invariants of the 2D Euler equation for an incompressible fluid [30. However, in spite of their frequent occurrence in reduced fluid models for plasmas, a complete physical interpretation of the variables $G_{0}, G_{1}, \cdots, G_{N}$, associated with such Casimir invariants, still appears to be lacking. Here we provide an attempt to remedy this gap.

We begin by recalling, as stated at the beginning of Sec. 5.1, that the $N+1-$ moment fluid model (63)-65) can be obtained from the parent drift-kinetic system (2)-4) by replacing $g$ with its truncated Hermite series

$$
\begin{equation*}
\tilde{g}(\mathbf{x}, v, t)=\sum_{m=0}^{N} g_{m}(\mathbf{x}, t) \frac{H_{m}(v)}{\sqrt{m!}} F_{e q}(v) \tag{162}
\end{equation*}
$$

Using the relation $g_{m}=U_{N_{m n}} G_{n}$, Eq. 162 can be rewritten as

$$
\begin{equation*}
\tilde{g}(\mathbf{x}, v, t)=\sum_{m, n=0}^{N} U_{N_{m n}} G_{n}(\mathbf{x}, t) \frac{H_{m}(v)}{\sqrt{m!}} F_{e q}(v) \tag{163}
\end{equation*}
$$

Evaluating Eq. 163 at $v=\lambda_{i}$, for $i=0,1, \cdots, N$ and using the relation

$$
\begin{equation*}
\frac{N!}{N+1} \sum_{i=0}^{N} \frac{H_{i}\left(\lambda_{j}\right) H_{i}\left(\lambda_{k}\right)}{\left|H_{N}\left(\lambda_{j}\right)\right|\left|H_{N}\left(\lambda_{k}\right)\right| i!}=\delta_{j, k}, \tag{164}
\end{equation*}
$$

which, analogously to Eq. B.3), follows from the orthogonality between $U_{N}$ and $U_{N}^{T}$, one obtains the relation

$$
\begin{equation*}
\tilde{g}\left(\mathbf{x}, \lambda_{i}, t\right)=\sqrt{\frac{N+1}{N!}}\left|H_{N}\left(\lambda_{i}\right)\right| G_{i}(\mathbf{x}, t) F_{e q}\left(\lambda_{i}\right), \quad \text { for } i=0,1, \cdots, N \tag{165}
\end{equation*}
$$

This relation can alternatively be written as

$$
\begin{equation*}
G_{i}(\mathbf{x}, t)=\alpha_{N, \lambda_{i}} \tilde{g}\left(\mathbf{x}, \lambda_{i}, t\right), \quad \text { for } i=0,1, \cdots, N \tag{166}
\end{equation*}
$$

where

$$
\begin{equation*}
\alpha_{N, \lambda_{i}}=\sqrt{\frac{N!}{N+1}} \frac{1}{\left|H_{N}\left(\lambda_{i}\right)\right| F_{e q}\left(\lambda_{i}\right)} . \tag{167}
\end{equation*}
$$

From the relation 166 one can then interpret each $G_{i}$ as a field providing, up to the multiplicative constant factor $\alpha_{N, \lambda_{i}}$, the spatial and temporal evolution $v=\lambda_{i}$.

Using Eq. 80, the evolution equation for a given $G_{i}$, can be written as

$$
\begin{equation*}
\frac{\partial G_{i}}{\partial t}+\left[\phi-\lambda_{i} A, G_{i}\right]+\lambda_{i} \frac{\partial G_{i}}{\partial z}+\sqrt{\frac{2}{\beta_{e}}} \sqrt{\frac{N!}{N+1}} \frac{\lambda_{i}}{\left|H_{N}\left(\lambda_{i}\right)\right|} \frac{\partial}{\partial z}\left(\lambda_{i} A-\phi\right)=0 \tag{168}
\end{equation*}
$$

where $\phi=\phi_{f l}\left[U_{N_{0 l}} G_{l}\right]$ and $A=A_{f l}\left[U_{N_{1 l}} G_{l}\right]$. Upon defining

$$
\begin{equation*}
\tilde{\mathrm{g}}_{\lambda_{i}}(\mathbf{x}, t)=\tilde{g}\left(\mathbf{x}, \lambda_{i}, t\right), \tag{169}
\end{equation*}
$$

one obtains, through Eq. 165, the evolution equation for $\tilde{\mathrm{g}}_{\lambda_{i}}$, which reads

$$
\begin{equation*}
\frac{\partial \tilde{\mathrm{g}}_{\lambda_{i}}}{\partial t}+\left[\phi-\lambda_{i} A, \tilde{\mathrm{~g}}_{\lambda_{i}}\right]+\lambda_{i} \frac{\partial \tilde{\mathrm{~g}}_{\lambda_{i}}}{\partial z}+\sqrt{\frac{2}{\beta_{e}}} F_{e q}\left(\lambda_{i}\right) \lambda_{i} \frac{\partial}{\partial z}\left(\lambda_{i} A-\phi\right)=0 . \tag{170}
\end{equation*}
$$

We remark that analogies between the dynamics of variables $G_{i}$ in fluid models (for $N=1$ ) and the dynamics of the electron drift-kinetic distribution function was observed by means of numerical simulations of magnetic reconnection in 0 Refs. 34, 35, 36.

From the relation 165 it follows that, knowing the location of $\lambda_{i}$ (i.e. of the zeros of $\left.H_{N+1}(x)\right)$, tells what are the particular values of $v$ for which the truncated generalized perturbed distribution function is proportional to $G_{i}$. From the properties of the zeros of Hermite polynomials, some information about the location of the eigenvalues $\lambda_{i}$ can actually be inferred.

We proceed by recalling some known properties about zeros of Hermite polynomials, from which we also draw some conclusions on the dynamics described by the reduced fluid models, and their relation with the parent drift-kinetic model.

First, for a given $N$, the eigenvalues $\lambda_{i}$ are distinct 42. Moreover, as already noticed in Sec. 5, they are symmetrically distributed around $x=0$, and, when $N$ is even, there exists one $m$ such that $\lambda_{m}=0$. Therefore, for fluid models with even $N$ (i.e. evolving an odd number of moments), one has

$$
\begin{equation*}
\frac{\partial G_{m}}{\partial t}+\left[\phi, G_{m}\right]=0 \tag{171}
\end{equation*}
$$

i.e., there exists one, and only one field $G_{m}$, which is purely advected by the $\mathbf{E} \times \mathbf{B}$ velocity field (this was the case, for instance, of $G_{1}$ in Eq. 146). This reflects the fact that, when $N$ is even, the Poisson bracket 73 can be written as

$$
\begin{align*}
& \{F, K\}_{G}=\sum_{\substack{i=0 \\
i \neq m}}^{N}\left(\sqrt{\frac{\beta_{e}}{2}} \frac{1}{U_{N_{0 i}}} \int_{\mathcal{D}} d^{3} x G_{i}\left[F_{G_{i}}, K_{G_{i}}\right]-\lambda_{i} \int_{\mathcal{D}} d^{3} x F_{G_{i}} \frac{\partial K_{G_{i}}}{\partial z}\right) \\
& +\sqrt{\frac{\beta_{e}}{2}} \frac{1}{U_{N_{0 m}}} \int_{\mathcal{D}} d^{3} x G_{m}\left[F_{G_{m}}, K_{G_{m}}\right] . \tag{172}
\end{align*}
$$

Casimir invariants of the bracket 172 are given by

$$
\begin{align*}
& \mathfrak{C}_{i}=\int_{\mathcal{D}} d^{3} x G_{i}, \quad i=0, \cdots, N, \quad i \neq m  \tag{173}\\
& \mathfrak{C}_{m}=\int_{\mathcal{D}} d^{3} x \mathcal{C}_{m}\left(G_{m}\right) \tag{174}
\end{align*}
$$

where $\mathcal{C}_{m}$ is an arbitrary smooth function. Therefore, when the number of retained moments is odd, the system admits, even in 3D, an infinite number of Casimir invariants, corresponding to the family $\mathfrak{C}_{m}$, in addition to the finite number of Casimir invariants $\mathfrak{C}_{i}$, with $i=0, \cdots, N$ and $i \neq m$. A similar feature was encountered also in the Casimir invariants of the model described in Ref. [28].

A different situation occurs when the number of moments is even (i.e. $N$ is odd). In this case, $\lambda_{i} \neq 0$ for all $i$ and the Poisson bracket reads

$$
\begin{equation*}
\{F, K\}_{G}=\sum_{i=0}^{N}\left(\sqrt{\frac{\beta_{e}}{2}} \frac{1}{U_{N_{0 i}}} \int_{\mathcal{D}} d^{3} x G_{i}\left[F_{G_{i}}, K_{G_{i}}\right]-\lambda_{i} \int_{\mathcal{D}} d^{3} x F_{G_{i}} \frac{\partial K_{G_{i}}}{\partial z}\right) . \tag{175}
\end{equation*}
$$

Casimir invariants thus reduce to

$$
\begin{equation*}
\mathfrak{C}_{i}=\int_{\mathcal{D}} d^{3} x G_{i}, \quad i=0, \cdots, N \tag{176}
\end{equation*}
$$

and remain in a finite number.
The explicit form of the matrix $U_{N}$ provided in the present paper, makes it also now possible to easily express Casimir invariants in terms of the fluid moments, for arbitrary $N$. This becomes particularly relevant in the 2D case, advection equations for the Lagrangian invariants $G_{0}, G_{1}, \cdots, G_{N}$.

More in general, as it emerges from Eq. (68), the eigenvalues $\lambda_{i}$ express the weight of the magnetic contribution, relative to the $\mathbf{E} \times \mathbf{B}$ contribution, of the generalized velocity fields

$$
\begin{equation*}
\mathbf{v}_{i}=\hat{z} \times \nabla\left(\phi+\lambda_{i} A\right), \quad i=0,1, \cdots, N \tag{177}
\end{equation*}
$$

The incompressible velocity fields $\mathbf{v}_{i}$ are those that advect the fields $G_{i}$ in the plane perpendicular to the guide field.

Another property of zeros of Hermite polynomials is that they interlace (see, e.g. Ref. 49]). This means that, if $\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}$ are the zeros of $H_{N+1}$ and $\lambda_{0}^{\prime}, \lambda_{1}^{\prime}, \cdots, \lambda_{N+1}^{\prime}$ are the zeros of $H_{N+2}$, one has

$$
\begin{equation*}
\lambda_{0}^{\prime}<\lambda_{0}<\lambda_{1}^{\prime}<\cdots<\lambda_{N}<\lambda_{N+1}^{\prime} \tag{178}
\end{equation*}
$$

Therefore, between two consecutive eigenvalues $\lambda_{i}$ and $\lambda_{i+1}$ of $S_{N}$, there will in the approximation of the drift-kinetic dynamics with that of the reduced fluid models. Indeed, reduced fluid models with the adopted Hamiltonian closure, as mentioned at the beginning of Sec. 5.1 replace the actual dynamics of $g$ with that of a truncated series. Through the relation 166 , one has a direct always be one eigenvalue $\lambda_{i+1}^{\prime}$ of $S_{N+1}$. Also, for a given interval $v_{1} \leqslant v \leqslant v_{2}$, one can always find an eigenvalue that belongs to that interval, provided $N$ is large enough. This suggests how the relative weight of magnetic vs. $\mathbf{E} \times \mathbf{B}$ contributions, in the generalized velocity fields $\mathbf{v}_{i}$, evolves, as $N$ increases. We point out that the arguments discussed in this Section, about the eigenvalues $\lambda_{i}$, hold for $N$ arbitrarily large but finite. In particular, although, as just stated, for sufficiently large $N$, one can find an eigenvalue $\lambda_{i}$ arbitrarily close to a given value of $v$, not all the real values of $v$ are eigenvalues of $S_{N}$, even for $N$ arbitrarily large. Indeed, because the eigenvalues are zeros of polynomials with rational coefficients, there exists no $N$, for which $\lambda_{i}=v$, for some $i \leqslant N$, when $v$ is a transcendental number (see also Ref. [43]). From this, we can infer a limitation correspondence between the variables $G_{i}$ of the fluid model, and a truncated
series of $g$, for a discrete set of $v \in\left\{\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}\right\}$. Thus, the dynamics of the fields $G_{i}$ can be seen as an approximation of the dynamics of the actual function $g$, for a discrete set of values of $v$. However, because of the above remark on the non-transcendental character of the eigenvalues $\lambda_{i}$, we can conclude that there values of $v$, no matter how large $N$ is.

We remark that the infinite hierarchy of equations (56)-59), obtained from the drift-kinetic system without truncations, can be written as

$$
\begin{align*}
& \frac{\partial g_{m}}{\partial t}+\left[\phi, g_{m}\right]-S_{m n}\left[A, g_{n}\right]+S_{m n} \frac{\partial g_{n}}{\partial z} \\
& +\sqrt{\frac{2}{\beta_{e}}} \frac{\partial}{\partial z}\left(\sqrt{m!}\left(\delta_{m 0}+\delta_{m 2}\right) A-\delta_{m 1} \phi\right)=0, \quad m \in \mathbb{Z}_{\geqslant 0} \tag{179}
\end{align*}
$$

where $S_{m n}=\sqrt{m} \delta_{m, n+1}+\sqrt{m+1} \delta_{m, n-1}$ are the elements of the infinite matrix

$$
S=\left(\begin{array}{ccccccc}
0 & 1 & 0 & 0 & \ldots & 0 & \ldots  \tag{180}\\
1 & 0 & \sqrt{2} & 0 & \ldots & 0 & \ldots \\
0 & \sqrt{2} & 0 & \sqrt{3} & \ldots & 0 & \ldots \\
0 & 0 & \sqrt{3} & 0 & \ldots & 0 & \ldots \\
& \ldots & & & \ldots & & \ldots \\
& \ldots & & & \ldots & & \ldots \\
0 & 0 & 0 & \ldots & 0 & \sqrt{N} & \ldots \\
0 & 0 & 0 & \ldots & \sqrt{N} & 0 & \ldots \\
\ldots & \ldots & \ldots & \ldots & \ldots & \ldots & \ldots
\end{array}\right) .
$$

One can recognize, up to multiplicative constant factors, the matrix $S$ as the Jacobi matrix of the position operator $\hat{x}$ for a quantum harmonic oscillator in the orthonormal basis consisting of eigenvectors of the number operator $\hat{N}$. The matrix $S$, in particular, can be written as $S=a^{T}+a$, where $a^{T}$ and $a$ are matrices of elements $a_{m n}^{T}=\sqrt{m} \delta_{m, n+1}$ and $a_{m n}=\sqrt{m+1} \delta_{m, n-1}$. The matrices $a^{T}$ and $a$ are associated with the ladder operators $\hat{a}^{\dagger}$ and $\hat{a}$, respectively. In terms of this analogy, one could interpret, from the first line Eq. (179), the temporal variation of the moment $g_{m}$ as influenced by the action of such two 5 operators. One of them (analogous to the "creation" operator $\hat{a}^{\dagger}$ ) corresponds
to $\sqrt{m+1} g_{m+1}$ and the second one (analogous to the "annihilation" operator $\hat{a})$ corresponds to $\sqrt{m} g_{m-1}$. We point out that a connection between creationannihilation operators and the dynamics of moments obtained from the Vlasov equation was investigated in Ref. [50. In this context, we also find it interesting to mention that our approach for closing the infinite hierarchy 179 , by replacing the infinite matrix $S$ with a finite matrix $S_{N}$, is analogous to the procedure adopted in quantum mechanics for approximating the behavior of a quantum harmonic oscillator by a truncated quantum harmonic oscillator 43, 41. Indeed, the matrices $S_{N}$ correspond to those associated with the position operators of truncated harmonic oscillators. The analogy with the quantum harmonic oscillator relies on the choice of the Hermite polynomials as basis for representing the generalized perturbed distribution function $g$. Therefore, we believe that the analogy would fail if a different basis were chosen for representing $g$.

## 8. Conclusions

We presented new results concerning an infinite class of Hamiltonian nonlinear reduced fluid models describing the dynamics of plasma and electromagnetic fields in the presence of a strong magnetic guide field. The Hamiltonian structure of all these models is now available in an explicit form, thus completing the results of Refs. 32, 33 about the existence of such structure. Although the Hamiltonian reduced fluid models can be obtained from the parent drift-kinetic model, by truncating the Hermite series expansion of the generalized perturbed distribution function, we showed that this approach cannot be applied in order to derive the Hamiltonian structure of the fluid models from that of the parent model. Indeed, by this approach, one does not retrieve the Poisson bracket of the fluid models, but a different (although, interestingly, "very similar") bilinear operator which we showed, with a counterexample, not to satisfy, in general, the Jacobi identity. Such truncations are thus shown not to preserve the Hamiltonian structure. In order to derive the Hamiltonian structure of fluid models from that of a parent kinetic model, alternative approaches should be followed.

Examples of fluid reductions from kinetic systems, that preserve a Hamiltonian structure, are provided in Refs. [51, 52, 53, 54, 55, 56.

From a more physical perspective, we showed the existence of a relation (so far unknown, to the best of our knowledge), between the variables $G_{0}, G_{1}, \cdots, G_{N}$ and the truncated generalized perturbed distribution function. Also, we put in evidence some features of the dynamics of the reduced fluid models, inferred from properties of the zeros of Hermite polynomials. A limitation in the capability of a reduced fluid model to approximate the dynamics of the drift-kinetic model was also identified. Furthermore, we pointed out an analogy between the hierarchy of fluid equations and the problem of the quantum harmonic oscillator. In particular, the closure problem in the plasma physics context, shares similarities with the problem of the truncated harmonic oscillator in quantum mechanics.

In our opinion, the present paper motivates further research in various directions. On one hand, given the above mentioned failure in deriving the Poisson brackets of the reduced fluid models, by the truncated series approach, the problem of the derivation of such Poisson brackets remains open. The identification, carried out in Sec. 5, of the terms that 'correct' the coefficients $\mathbb{W}_{(N) l}^{m n}$, turning them into the coefficients $W_{(N) l}^{m n}$ of the Poisson bracket, might give some hint on how the Lie algebra underlying the Poisson bracket $\{,\}_{g}$ descends from that of the parent Poisson bracket $\{,\}_{d k}$.

A further natural direction of investigation, potentially leading to a number of applications in terms of modelling plasmas with strong anisotropies, concerns the identification of Hamiltonian closures for reduced fluid models accounting, in addition to the evolution of moments involving the coordinate $v$, also moments with respect to the perpendicular velocity coordinate. In Ref. 33] finite Larmor radius effects involving the perpendicular velocity (or, equivalently, the magnetic moment) coordinate, were taken into account. However, no general Hamiltonian closure was found for models evolving also moments with respect to such coordinate.

Finally, we believe it could be useful to deepen the investigation of the anal-
ogy between the present hierarchy of fluid models and the quantum harmonic oscillator. In particular, it might be interesting to see whether the techniques adopted to approximate the quantum harmonic oscillator by a truncated oscillator, as done in Ref. [41, could be transferred to the problem of approximating ${ }_{630}$ a drift-kinetic system by a reduced fluid model.

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## AppendixA. Proof of Lemma 5.1

Proof. In order to prove the statement (a) it is convenient to introduce the function

$$
\begin{equation*}
\theta_{N_{l m n}}(x)=\frac{N!}{N+1} \frac{H_{l}(x) H_{m}(x) H_{n}(x)}{H_{N}^{2}(x) \sqrt{l!m!n!}} \tag{A.1}
\end{equation*}
$$

so that

$$
\begin{equation*}
W_{(N) l}^{m n}=\sum_{i=0}^{N} \theta_{N_{l m n}}\left(\lambda_{i}\right) \tag{A.2}
\end{equation*}
$$

Because an Hermite polynomial $H_{n}(x)$ is an even (odd) function of $x$ if $n$ is an even (odd) number, it follows from the definition A.1 that, if $l+m+n$ is odd, the function $\theta_{N_{l m n}}$ is odd. Moreover, the elements $\lambda_{0}, \lambda_{1}, \cdots, \lambda_{N}$, which are the zeros of the Hermite polynomial $H_{N+1}(x)$, are symmetrically distributed around $x=0$ on the real axis, so that if $\lambda_{i}$ is a zero, also $-\lambda_{i}$ is. In particular one has $\lambda_{i}=0$, for a certain $i$, when $H_{N+1}$ is an odd function, i.e. when $N$ is even.

We first consider the case when $N$ is odd. In this case $\lambda_{i} \neq 0$ for $i=$ $0,1, \cdots, N$. We sort the eigenvalues $\lambda_{i}$ in increasing order so that $\lambda_{0}<\lambda_{1}<$ $\cdots<\lambda_{N}$. Due to the above mentioned symmetry property around $x=0$, we have that

$$
\begin{align*}
& \left\{\lambda_{0}, \lambda_{1}, \cdots, \lambda_{\frac{N-1}{2}}, \lambda_{\frac{N+1}{2}}, \cdots, \lambda_{N-1}, \lambda_{N}\right\} \\
& =\left\{-\lambda_{N},-\lambda_{N-1}, \cdots,-\lambda_{\frac{N+1}{2}}, \lambda_{\frac{N+1}{2}}, \cdots, \lambda_{N-1}, \lambda_{N}\right\} . \tag{A.3}
\end{align*}
$$

Thus, from Eq. A.2 , considering that $\theta_{N_{l m n}}$ is an odd function when $l+m+n$ is odd, one has

$$
\begin{align*}
& W_{(N) l}^{m n}=\sum_{i=0}^{\frac{N-1}{2}} \theta_{N_{l m n}}\left(\lambda_{i}\right)+\sum_{i=\frac{N+1}{2}}^{N} \theta_{N_{l m n}}\left(\lambda_{i}\right) \\
& =\sum_{i=\frac{N+1}{2}}^{N}\left(\theta_{N_{l m n}}\left(-\lambda_{i}\right)+\theta_{N_{l m n}}\left(\lambda_{i}\right)\right)=\sum_{i=\frac{N+1}{2}}^{N}\left(-\theta_{N_{l m n}}\left(\lambda_{i}\right)+\theta_{N_{l m n}}\left(\lambda_{i}\right)\right)=0 . \tag{A.4}
\end{align*}
$$

The case when $N$ is even proceeds in a similar way, with the difference that the eigenvalue $\lambda_{i}=0$ must be taken into account. Therefore one has

$$
\begin{align*}
& \left\{\lambda_{0}, \lambda_{1}, \cdots, \lambda_{\frac{N}{2}-1}, \lambda_{\frac{N}{2}}, \lambda_{\frac{N}{2}+1}, \cdots, \lambda_{N-1}, \lambda_{N}\right\} \\
& =\left\{-\lambda_{N},-\lambda_{N-1}, \cdots,-\lambda_{\frac{N}{2}+1}, 0, \lambda_{\frac{N}{2}+1}, \cdots, \lambda_{N-1}, \lambda_{N}\right\} . \tag{A.5}
\end{align*}
$$

Analogously to Eq. A.4, the expression for the coefficients $W_{(N) l}^{m n}$ is given by

$$
\begin{align*}
& W_{(N) l}^{m n}=\sum_{i=0}^{\frac{N}{2}-1} \theta_{N_{l m n}}\left(\lambda_{i}\right)+\sum_{i=\frac{N}{2}+1}^{N} \theta_{N_{l m n}}\left(\lambda_{i}\right)+\theta_{N_{l m n}}\left(\lambda_{\frac{N}{2}}\right) \\
& =\sum_{i=\frac{N}{2}+1}^{N}\left(\theta_{N_{l m n}}\left(-\lambda_{i}\right)+\theta_{N_{l m n}}\left(\lambda_{i}\right)\right)+\theta_{N_{l m n}}(0)=\sum_{i=\frac{N+1}{2}}^{N}\left(-\theta_{N_{l m n}}\left(\lambda_{i}\right)+\theta_{N_{l m n}}\left(\lambda_{i}\right)\right)=0, \tag{A.6}
\end{align*}
$$

where $\theta_{N_{l m n}}(0)=0$ because $\theta_{N_{l m n}}$ is an odd function.
With regard to the statement (b), it follows straightforwardly from commutativity under multiplication, which implies $\hat{H}_{\sigma(l)} \hat{H}_{\sigma(m)} \hat{H}_{\sigma(n)}=\hat{H}_{l} \hat{H}_{m} \hat{H}_{n}$. Consequently, from the expression (88), we obtain

$$
\begin{align*}
& W_{(N) \sigma(l)}^{\sigma(m) \sigma(n)}=\frac{N!}{N+1} \sum_{i=0}^{N} \frac{\hat{H}_{\sigma(l)}\left(\lambda_{i}\right) \hat{H}_{\sigma(m)}\left(\lambda_{i}\right) \hat{H}_{\sigma(n)}\left(\lambda_{i}\right)}{H_{N}^{2}\left(\lambda_{i}\right)} \\
& =\frac{N!}{N+1} \sum_{i=0}^{N} \frac{\hat{H}_{l}\left(\lambda_{i}\right) \hat{H}_{m}\left(\lambda_{i}\right) \hat{H}_{n}\left(\lambda_{i}\right)}{H_{N}^{2}\left(\lambda_{i}\right)}=W_{(N) l}^{m n} . \tag{A.7}
\end{align*}
$$

## AppendixB. Proof of Proposition 5.2

645 Proof. Comparing the expressions (106), 107) and 109) with the expression (87), it immediately follows that, in order to prove the Proposition, it suffices to prove the relation 108). Also, because of Lemma 5.1 (a), we only need to consider 3 -ples $(l, m, n)$ such that $l+m+n$ is even. Moreover, by virtue of Lemma 5.1 b), once we obtain an expression for $W_{(N) l}^{m n}$, we automatically obtain the expressions for all the coefficients obtained by permutations of $l, m$ and $n$.

To obtain the required expression for the coefficients $W_{(N) l}^{m n}$ we first recall the following identity for Hermite polynomials (which can be obtained from Eq. (2.01) of Ref. [57]):

$$
\begin{equation*}
H_{m}(x) H_{n}(x)=\sum_{r=0}^{\min (m, n)} r!\binom{m}{r}\binom{n}{r} H_{m+n-2 r}(x) \tag{B.1}
\end{equation*}
$$

From Eq. 88, it follows that the coefficients $W_{(N) l}^{m n}$ can then be rewritten as

$$
\begin{equation*}
W_{(N) l}^{m n}=\frac{N!}{N+1} \sum_{i=0}^{N} \frac{H_{l}\left(\lambda_{i}\right)}{H_{N}^{2}\left(\lambda_{i}\right)} \sum_{r=0}^{\min (m, n)} \frac{m!}{(m-r)!} \frac{n!}{(n-r)!r!} \frac{H_{m+n-2 r}\left(\lambda_{i}\right)}{\sqrt{l!m!n!}} \tag{B.2}
\end{equation*}
$$

Due to the orthogonality of the matrices $U$ and $U^{T}$, from the expression 85), one obtains the following relation:

$$
\begin{equation*}
\frac{N!}{N+1} \sum_{i=0}^{N} \frac{H_{j}\left(\lambda_{i}\right) H_{k}\left(\lambda_{i}\right)}{H_{N}^{2}\left(\lambda_{i}\right) \sqrt{j!k!}}=\delta_{j, k} \tag{B.3}
\end{equation*}
$$

Now we consider a coefficient $W_{(N) l}^{m n}$ for a fixed 3-ple $(l, m, n)$ and proceed by separating the analysis in two cases.

Case I: In a given 3-ple (l,m,n), the elements $l, m, n$ are such that $l+m+n$ is even and there exist at least two elements such that their sum is less than or equal to $N+1$

Let us suppose that $m$ and $n$ are such that $m+n \leqslant N+1$. Due to the orthogonality relation (B.3), when the condition $m+n-2 r=l$ (with $0 \leqslant$
$r \leqslant \min (m, n))$ is satisfied, the right-hand side of Eq. B.2 yields a finite contribution. Note also that, if $m+n=N+1$, the contribution coming from $H_{m+n}\left(\lambda_{i}\right)=H_{N+1}\left(\lambda_{i}\right)$, corresponding to $r=0$, vanishes. This contribution, actually, is not determined by the relation (B.3), because such relation involves only Hermite polynomials up to order $N$. However, one has that $H_{N+1}\left(\lambda_{i}\right)=0$, because, as recalled before Eq. (77), $\lambda_{i}$ is a zero of $H_{N+1}$ for all $i=0,1, \cdots, N$. Therefore, the case $m+n=N+1$ provides at most only one finite contribution to $W_{(N) l}^{m n}$, as does the case $m+n<N+1$. The non-zero contribution occurs for

$$
\begin{equation*}
r=\frac{m+n-l}{2} \tag{B.4}
\end{equation*}
$$

(recall that we are considering $l+m+n$ even, so that $m+n-l$ is also even), provided that

$$
\begin{equation*}
\frac{m+n-l}{2} \geqslant 0, \quad m \geqslant \frac{m+n-l}{2}, \quad n \geqslant \frac{m+n-l}{2} . \tag{B.5}
\end{equation*}
$$

The conditions (B.5) can be reformulated as

$$
\begin{equation*}
m+n \geqslant l, \quad l+m \geqslant n, \quad n+l \geqslant m \tag{B.6}
\end{equation*}
$$

If these conditions are fulfilled, from Eq. (B.2), using Eq. (B.3), one obtains

$$
\begin{aligned}
& W_{(N) l}^{m n}=\frac{\sqrt{l!m!n!}}{\left(m-\frac{m+n-l}{2}\right)!\left(n-\frac{m+n-l}{2}\right)!\left(\frac{m+n-l}{2}\right)!} \\
& =\frac{\sqrt{l!m!n!}}{\left(\frac{l+m-n}{2}\right)!\left(\frac{n+l-m}{2}\right)!\left(\frac{m+n-l}{2}\right)!} .
\end{aligned}
$$

If any of the three conditions $\overline{\mathrm{B} .6}$ is not satisfied, then $W_{(N) l}^{m n}=0$, because there would be no $r$, with $0 \leqslant r \leqslant \min (m, n)$, such to provide a non-zero contribution in the right-hand side of Eq. B.2, due to Eq. B.3. The 3-ples $(l, m, n)$ belonging to Case I and yielding $W_{(N) l}^{m n} \neq 0$ are thus those given by $(l, m, n) \in A_{N} \backslash B_{N}$.

Note that, due to the invariance of the coefficients $W_{(N) l}^{m n}$ under permuta65 tion, one can easily determine the coefficients in Case I also when the sum of two indices is greater than $N+1$. For instance, if one is in Case I and wants to compute $W_{(N) m}^{n l}$, with $n+l>N+1, l+m>N+1$ and $m+n \leqslant N+1$, it
suffices to permute the indices and, due to $W_{(N) m}^{n l}=W_{(N) l}^{m n}$, one can follow the above procedure carried out for $W_{(N) l}^{m n}$.

Case II : In a given 3-ple $(l, m, n)$, the elements $l$, $m, n$ are such that $l+m+n$ is even and $l+m>N+1, \quad m+n>N+1, \quad n+l>N+1$

We are thus referring to the case $(l, m, n) \in B_{N}$. The sum on $r$ on the right-hand side of Eq. (B.2) involves the Hermite polynomials $H_{m+n}, H_{m+n-2}$, $H_{m+n-4}, \cdots, H_{m+n-2 \min (m, n)}$. We denote with $r_{N_{m n}}$ the smallest integer $r$, with $0<r \leqslant \min (m, n)$, such that $m+n-2 r_{N_{m n}} \leqslant N+1$. This corresponds to the definition (104). We can then split the sum on $r$, in Eq. (B.2), into two parts, in the following way:

$$
\begin{align*}
& W_{(N) l}^{m n}=\frac{N!}{N+1} \sum_{i=0}^{N} \frac{H_{l}\left(\lambda_{i}\right)}{H_{N}^{2}\left(\lambda_{i}\right)}\left(\sum_{r=0}^{r_{N m n}-1} \frac{m!}{(m-r)!} \frac{n!}{(n-r)!r!} \frac{H_{m+n-2 r}\left(\lambda_{i}\right)}{\sqrt{l!m!n!}}\right.  \tag{B.7}\\
& \left.+\sum_{r=r_{N_{m n}}}^{\min (m, n)} \frac{m!}{(m-r)!} \frac{n!}{(n-r)!r!} \frac{H_{m+n-2 r}\left(\lambda_{i}\right)}{\sqrt{l!m!n!}}\right) \tag{B.8}
\end{align*}
$$

The sum from $r=r_{N_{m n}}$ to $r=\min (m, n)$, in the expression B.8), involves only Hermite polynomials of order at most equal to $N+1$. Therefore, although $m+n>N+1$, this expression can be treated in the same way as Case I, using the relation (B.3). On the other hand, the sum from $r=0$ to $r=r_{N_{m n}}-1$, in the expression (B.7), involves only Hermite polynomials of order greater than $N+1$ (and thus greater than $l$ ). For such polynomials, the orthogonality relation (B.3) does not apply. However, the terms in the expression B.7), in general, can provide additional finite contributions to $W_{(N) l}^{m n}$. It follows that, from Eq. (B.7)-( $\overline{\mathrm{B} .8}$ ), the expression for $W_{(N) l}^{m n}$ in Case II can be written as

$$
\begin{align*}
& W_{(N) l}^{m n}=\frac{N!}{N+1} \sum_{i=0}^{N} \frac{H_{l}\left(\lambda_{i}\right)}{H_{N}^{2}\left(\lambda_{i}\right)} \sum_{r=0}^{r_{N_{m n}}-1} \frac{m!}{(m-r)!} \frac{n!}{(n-r)!r!} \frac{H_{m+n-2 r}\left(\lambda_{i}\right)}{\sqrt{l!m!n!}} \\
& +\frac{\sqrt{l!m!n!}}{\left(\frac{l+m-n}{2}\right)!\left(\frac{n+l-m}{2}\right)!\left(\frac{m+n-l}{2}\right)!} . \tag{B.9}
\end{align*}
$$

Equation B.9 thus yields the required expression for $W_{(N) l}^{m n}$ when $(l, m, n) \in$

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