

An Insight Into Ancient Aeolian Processes and Post-Noachian Aqueous Alteration in Gale Crater, Mars, Using ChemCam Geochemical Data From the Greenheugh Capping Unit

Candice Bedford, Steven Banham, John Bridges, Olivier Forni, Agnes Cousin, Donald Bowden, Stuart Turner, Roger Wiens, Patrick Gasda, Jens Frydenvang, et al.

▶ To cite this version:

Candice Bedford, Steven Banham, John Bridges, Olivier Forni, Agnes Cousin, et al.. An Insight Into Ancient Aeolian Processes and Post-Noachian Aqueous Alteration in Gale Crater, Mars, Using ChemCam Geochemical Data From the Greenheugh Capping Unit. Journal of Geophysical Research. Planets, 2022, 127 (9), pp.e2021JE007100. 10.1029/2021JE007100. insu-03853134

HAL Id: insu-03853134 https://insu.hal.science/insu-03853134v1

Submitted on 15 Nov 2022

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.



Distributed under a Creative Commons Attribution - NonCommercial - NoDerivatives 4.0 International License

An insight into ancient aeolian processes and post-Noachian aqueous alteration in 1

Gale crater, Mars, using ChemCam geochemical data from the Greenheugh capping 2 unit. 3

- Candice C. Bedford^{1,2}, Steven G. Banham³, John C. Bridges⁴, Olivier Forni⁵, Agnes 4
- Cousin⁵, Donald Bowden⁴, Stuart M. R. Turner⁶, Roger C. Wiens⁷, Patrick J. Gasda⁷, Jens 5
- 6
- Frydenvang⁸, Olivier Gasnault⁵, Kristin Rammelkamp⁵, Frances Rivera-Hernandez⁹, Elizabeth B. Rampe², Rebecca Smith¹⁰, Cherie Achilles¹¹, Erwin Dehouck¹², Alexander B. Bryk¹³, Susanne P. Schwenzer⁶, Horton Newsom¹⁴. 7
- 8
- ¹Lunar and Planetary Institute, Universities Space Research Association, Houston, Texas, USA. 9
- ²Astromaterials Research and Exploration Science Division, NASA Johnson Space Center, 10
- Houston, Texas, USA. 11
- ³Imperial College, London, UK. 12
- ⁴University of Leicester, Leicester, UK. 13
- ⁵Institut de Recherche en Astrophysique et Planétologie, Université de Toulouse, UPS, CNRS, 14
- CNES, Toulouse, France. 15
- ⁶AstrobiologyOU, School of Environment, Earth and Ecosystem Sciences, The Open 16
- University, Milton Keynes, UK. 17
- 18 ⁷Los Alamos National Laboratory, New Mexico, USA.
- ⁸Globe Institute, University of Copenhagen, Denmark. 19
- ⁹Georgia Institute of Technology, Georgia, USA. 20
- ¹⁰Stony Brook University, New York, USA. 21
- ¹¹NASA Goddard space flight center, Maryland, USA. 22
- ¹²Universite de Lyon, Lyon, France. 23
- ¹³University of Berkeley, California, USA. 24
- ¹⁴University of New Mexico, Albuquerque, New Mexico, USA. 25
- Corresponding author: Candice C. Bedford (cbedford@lpi.usra.edu) 26
- **Key Points:** 27
- ChemCam geochemical results support that the Greenheugh pediment capping unit was 28 deposited within the Stimson sand sea in the Hesperian. 29
- The Greenheugh pediment capping unit was sourced from olivine-rich sands, potentially 30 recycled from units farther up Mt Sharp. 31
- Greenheugh experienced less interaction with groundwaters than Stimson at the Emerson 32 and Naukluft plateaus, but shared diagenesis at the unconformity. 33

34

35 Abstract

36 Aeolian processes have shaped and contributed to the geological record in Gale crater, 37 Mars, long after the fluviolacustrine system existed ~3 Ga ago. Understanding these aeolian deposits, particularly those which have been lithified and show evidence for aqueous alteration, 38 can help to constrain the environment at their time of deposition and the role of liquid water later 39 in Mars' history. The NASA Curiosity rover investigated a prominent outcrop of aeolian 40 41 sandstone within the Stimson formation at the Greenheugh pediment as part of its investigation 42 of the Glen Torridon area. In this study, we use geochemical data from ChemCam to constrain the effects of aeolian sedimentary processes, sediment provenance, and diagenesis of the 43 sandstone at the Greenheugh pediment, comparing the Greenheugh data to the results from 44 45 previous Stimson localities situated 2.5 km north and >200 m lower in elevation. Our results, supported by mineralogical data from CheMin, show that the Stimson formation at the 46 47 Greenheugh pediment was likely sourced from an olivine-rich unit that may be present farther up 48 the slopes of Gale crater's central mound. Our results also suggest that the Greenheugh pediment Stimson formation was cemented by surface water runoff such as that which may have formed 49 Gediz Vallis. The lack of alteration features in the Stimson formation at the Greenheugh 50 51 pediment relative to those of the Emerson and Naukluft plateaus suggests that groundwater was not as available at this locality compared to the others. However, all sites share diagenesis at the 52

53 unconformity.

54 Plain language summary

55 We use geochemical data from the Curiosity rover's ChemCam instrument to investigate the

- 56 geological history of the Greenheugh pediment capping unit. The Greenheugh pediment is a
- 57 geologic feature that has a gently sloping top, forming a broad plain extending from the base of
- the valley, Gediz Vallis. The Greenheugh pediment is capped by a unit of broadly uniform
- 59 thickness which represents the remains of the Stimson dune field that existed <2.5 Ga (mid- to
- 60 late-Hesperian). ChemCam geochemical data shows that the sands deposited at the Greenheugh
- capping unit were sourced from a nearby olivine-rich unit. Surface waters then cemented the
- 62 windblown sand deposits, ponding at the unconformity with the underlying mudstone unit,
- 63 creating concretions towards the base. Episodes of groundwater circulation did not affect the 64 rocks at Greenheugh as much as they did at other Stimson localities with the exception of acid-
- sulfate alteration that occurred along the unconformity. These results suggest that the ancient
- 66 Stimson dune field was a dynamic environment, incorporating grains from the surrounding
- 67 geological units on Mt Sharp. Furthermore, liquid water was stable at the surface in the
- Hesperian and was available for multiple diagenetic events along bedrock weaknesses.

69 **1 Introduction**

Gale crater, where the NASA Mars Science Laboratory (MSL) Curiosity rover has been exploring since August, 2012, contains a diverse geological record which has been shaped by fluviolacustrine and later aeolian processes (Banham et al., 2021; Bridges et al., 2017; Edgar et al., 2020; Grotzinger et al., 2012, 2014, 2015; Hobbs et al., 2010; Stack et al., 2019; Thomson et

- al., 2011). After landing, the Curiosity rover traversed across the Bradbury plains to the foothills
- of Gale crater's central mound, Aeolis Mons (informally known as Mt Sharp). Along the
- 75 of Gale crater's central mound, Acons Mons (mornary known as Wt Sharp). Along the 76 traverse, the MSL mission has investigated poorly sorted conglomerates, cross-bedded
- sandstones, and thinly laminated mudstones within the Bradbury group (Fig. 1 and 2), indicative

of deposition in a fluviolacustrine sedimentary system (Anderson et al., 2015; Grotzinger et al.,

- 79 2014; Mangold et al., 2016; Treiman et al., 2016a; Vasavada et al., 2014). Since reaching Mt
- 80 Sharp, the rover has investigated over 300 m of stratigraphy within the Mt Sharp group,
- consisting mainly of laminated mudstone deposits of the Murray and Carolyn Shoemaker
 formations, which are interpreted as evidence for a long-lived, perennial lacustrine environment
- formations, which are interpreted as evidence for a long-lived, perennial lacustrine environment (Bennett et al., *this issue*: Edgar et al., 2020; Fedo et al., 2018; Grotzinger et al., 2015; Stack et
- al., 2019). This long-lived river-lake system within Gale crater likely existed soon after the crater
- formed approximately 3.6–3.8 Ga based on crater counting (Thomson et al., 2011) with aqueous
- activity in the crater evidenced beyond the analyses conducted by the MSL mission by the
- detection from orbit of hydrated minerals (Milliken et al., 2010a; Thomson et al., 2011).
- 88 However, at some point in Gale crater's geological record, the environment changed to the cold
- and dry conditions of today, in which geologic processes are dominated by aeolian activity.

The geological units preserved in Mt Sharp are hypothesized to have previously filled in 90 the crater, with aeolian erosion largely responsible for the exhumation of Mt Sharp to its current 91 topographic expression (Bennett et al., 2016; Grotzinger et al., 2015; Malin et al., 2000; Steele et 92 93 al., 2018). Aeolian processes have also resulted in the formation of modern aeolian landforms including wind-ripples, sand shadows (Blake et al., 2013), transverse aeolian ridges (Zimbelman 94 et al., 2020), and dune fields (Bridges et al., 2017, 2018; Ewing et al., 2017; Lapotre et al., 2018) 95 using the available sediment eroded from the surrounding landscape (Day et al., 2016; Meslin et 96 al., 2013), or transported into the crater from elsewhere (Hobbs et al., 2010). 97

98 The Curiosity rover has encountered lithified aeolian sandstone deposits unconformably overlying the Murray formation mudstones as part of the Stimson formation (Fig. 1 and 2) at the 99 Emerson plateau (Banham et al., 2018), Naukluft plateau (Bedford et al., 2020a), and the Murray 100 Buttes (Banham et al., 2021). Most recently, the Stimson formation has been encountered again 101 at the Greenheugh pediment, this time overlying the Carolyn Shoemaker formation mudstone 102 deposits (Fig. 1 and 2). The Stimson formation is currently the youngest in situ geological unit 103 104 investigated by the Curiosity rover in Gale crater. In this study, we use geochemical data from the Chemistry and Camera (ChemCam) instrument suite and mineralogical data from the 105 Chemistry and Mineralogy (CheMin) instrument to investigate the geological history of the 106 107 Stimson formation at the Greenheugh pediment. Thereby providing constraints on the 108 paleoenvironment of Gale crater beyond the time in which the river-lake system existed.

109 1.1 An overview of the Stimson formation and previous localities

Initially identified from orbital imagery on the basis of its high thermal inertia with 110 respect to the surrounding units (Fraeman et al., 2016), the Stimson formation is a dark grey, 111 cross-bedded sandstone within the Siccar Point group (Fig. 2) that unconformably overlies the 112 Mt Sharp group (Banham et al., 2018; Watkins et al., 2016). The unconformity between the Mt 113 Sharp group and the Stimson formation has been informally named the basal Siccar Point group 114 unconformity and is a noticeable divide between the dark grey aeolian sandstones above, and the 115 light-toned mudstones of the Mt Sharp group below (Banham et al., 2018; Watkins et al., 2016). 116 Mineral veins in the Mt Sharp group mudstones are observed to terminate at this unconformity 117 (Fig. 3) suggesting that they formed prior to the erosion of the Mt Sharp group and deposition 118 119 and lithification of the Stimson formation (Kronyak, et al., 2019a).

Based on orbital mapping, the Stimson formation is exposed over a region of 20 km² (Kronyak et al., 2019b). Previous ground-based observations of the Stimson formation showed 122 that it contains well-sorted to sub-well sorted, rounded sand grains with an average grain size of

medium sand with a coarse-grained skew (Banham et al., 2018). Meter-scale cross-bedding is

also common across all Stimson formation outcrops so far investigated by the Curiosity rover

125 (Banham et al., 2018, 2021, *this issue.*; Bedford et al., 2020a). The grain size characteristics are 126 indicative of the physical attrition of grains during transportation by saltation and surface creep

indicative of the physical attrition of grains during transportation by saltation and surface creep
 within migrating dune bedforms in a dry, aeolian environment (Banham et al., 2018). As such,

the Stimson formation is a record of a large aeolian dune field that covered the north-west flank

of Mt Sharp (Banham et al., 2018, 2021; Bedford et al., 2020a). Further evidence for this

130 interpretation is: the presence of wind ripple strata and meter-scale cross-bedding associated with

131 sediment transport on dune surfaces; and bounding surfaces associated with migration of

132 superimposed dunes.

The Emerson plateau was the first locality in which the Stimson formation was
investigated (Fig. 1). The Stimson formation at the Emerson plateau overlies the Murray

135 formation mudstones of the Mt Sharp group surrounding the Pahrump Hills locality (sols 750–

136 940) but was first investigated close-up by the rover at the Marias Pass locality (sols 992–1154).

- The Stimson formation at the Emerson plateau consists of \sim 7 m tall outcrops of blocky, dark grey sandstones, predominantly comprising simple cross sets that are usually 0.3–1.0 m thick and
- grey sandstones, predominantly comprising simple cross sets that are usually 0.3–1.0 m thick and are themselves derived of cross-laminations with a uniform thickness of 4 mm (c.f. Banham et

140 al., 2018 for a detailed description of this locality) (Fig. 3C). Bounding surfaces at the Emerson

plateau were traced laterally to a distance of ~ 30 m and are hypothesized to relate to the

interdune surfaces (Banham et al., 2018). Three drilled samples were acquired at the Emerson

plateau (Fig. 1): the Buckskin drilled sample (sol 1053) of the underlying Murray formation at
 Marias Pass, the Big Sky drilled sample (sol 1119) representing parent Stimson, and the

Greenhorn drilled sample (sol 1137) representing altered Stimson at a fracture-associated halo.

146 The Naukluft plateau is situated west of the Emerson plateau and was the second Stimson locality to be investigated by the Curiosity rover after the first Bagnold dune field campaign (Fig. 147 148 1). The Naukluft plateau is 250 km wide, with the Stimson formation at this locality varying in thickness between 0.5–10 m (Bedford et al., 2020a). The Stimson formation here is characterized 149 by trough-cross bedded sandstones with a similar sedimentary architecture to the Stimson 150 formation at the Emerson plateau (Bedford et al., 2020a) (Fig. 3D). Three drilled samples were 151 also acquired at the Naukluft plateau locality: the Okoruso sample representing Stimson bedrock 152 (sol 1332, Fig. 3d), the Lubango sample of another light-toned fracture-associated halo in the 153 154 Stimson bedrock (sol 1320), and the Oudam drilled sample taken in the Murray formation just below the Siccar Point group unconformity (sol 1361). Cross-set dip directions of the Stimson 155 formation at the Emerson plateau and Murray Buttes (Fig. 1) suggest sediment transport toward 156 157 the northeast (Banham et al., 2018, 2021), which is supported by geochemical trends relating to the physical grain segregation of different basaltic materials that also extend this transport 158 direction to the Stimson formation at the Naukluft plateau (Bedford et al., 2020a). 159

The Murray Buttes locality was the third locality in which the Stimson formation was investigated (Fig. 1). The Murray Buttes were encountered to the southwest of the Naukluft plateau between sols 1412 and 1482 and form a set of isolated buttes and mesas between 5 and 16 m high (Banham et al., 2021). The talus and steep slopes present at the buttes prevented contact science opportunities. As such, the Stimson formation here was only analyzed using Mastcam images, with no opportunities for in situ geochemical or mineralogical analysis (Banham et al., 2021). The Stimson formation at this locality is present as erosion resistant, 167 cross-bedded cap rock on the buttes and provides the thickest, three-dimensional exposure of the

aeolian sandstone deposit of all locations to date (Banham et al., 2021). Once again, the

sedimentary architecture at the Murray Buttes is similar to that of the Emerson and Naukluft

plateaus, containing meter-scale planar and trough cross-bedding, with the cross-beds consisting
 of millimeter-thick pinstripe laminations that are laterally continuous for tens of centimeters

(Banham et al., 2021). The results from the image analysis supports that the Stimson formation

represents a large aeolian sand sea, composed of aggrading compound dunes that, at this

174 location, migrated approximately to the north, downslope of Mt Sharp, which is a slight

175 deviation from the northeast-southwest direction of the Emerson and Naukluft plateaus (Banham

et al., 2021). Alteration halos were not observed at the Murray Buttes locality, but calcium

177 sulfate cement and veins and nodules were present in the Murray formation mudstone and

178 Stimson formation sandstone.

179 1.1.1. The Greenheugh pediment capping unit mini-campaign

The Greenheugh pediment was investigated by the Curiosity rover during the Glen 180 Torridon campaign (Bennett et al., this issue; Bryk et al., 2019). The Glen Torridon locality 181 comprises lacustrine mudstone of the Carolyn Shoemaker formation within the Mt Sharp group 182 that has a strong phyllosilicate-signature from orbit (Bennett et al., *this issue*), and a diverse 183 secondary mineralogy containing phyllosilicates, carbonates, and iron oxides that may be 184 indicative of lacustrine groundwater mixing (Thorpe et al., this issue). The Stimson formation of 185 the Greenheugh capping unit overlies the Mt Sharp group in the Glen Torridon area (Banham et 186 al., this issue; Bryk et al., 2019). On the approach to the Greenheugh pediment capping unit in 187 Glen Torridon, the underlying Carolyn Shoemaker formation mudstone changed in 188 geochemistry, became visually lighter-toned in color (Dehouck et al., this issue), and increased 189 in the abundance of F/MgO-rich and FeO_T/MnO-rich alteration features (Gasda et al., *this issue*: 190 Forni et al., 2021), indicative of alteration of the mudstone just below the Basal Siccar Point 191 group unconformity. The mudstone situated close to the Basal Siccar Point group unconformity 192 at the Hutton drill hole (Fig. 2) shows a change in mineralogy with a lower abundance of clay 193 minerals relative to the other drilled samples in Glen Torridon, in addition to minor cristobalite 194 195 and opal-CT (Thorpe et al., this issue; Rampe et al., 2020).

On sol 2695, the Curiosity rover ascended the Greenheugh pediment and reached the 196 pediment capping unit. The Stimson formation of the Greenheugh pediment capping unit can be 197 divided into three stratigraphic intervals. The Gleann Beag interval is the lowermost interval with 198 a thickness of ~3.6 m (Fig. 3B). The Gleann Beag interval is composed of grey, compound cross-199 bedded sandstones with concretions. Here, cross-bedding dip azimuths are oriented to the 200 northeast (Banham et al., this issue). In total, 13 ChemCam targets (113 observation points) were 201 measured at the Gleann Beag interval analyzing concretionary and non-concretionary bedrock. 202 The concretions of this interval range in size between 5–20 mm diameter and preserve the 203 sedimentary features of the sandstone (Banham et al., this issue). Two additional targets were 204 measured at the base of the Gleann Beag interval that also sampled light-toned Stimson bedrock 205 immediately above the unconformity with irregularly shaped, white clasts ~1 mm long (Banham 206 et al., this issue). These clasts were interpreted as rip-up clasts derived from the Carolyn 207 Shoemaker formation due to their similar visual appearance to the underlying mudstone and 208 209 similar shape to rip-up clasts identified at the Marias Pass locality (Banham et al., *this issue*;

210 Newsom et al., 2016).

Above the Gleann Beag interval is the Ladder interval (Fig. 3A and 3B). This interval consists of simple and concordant cross-sets with cross-laminations that dip to the south. Concretions are rare in this interval (Banham et al., *this issue*). Five ChemCam targets sampled the Ladder interval, although only three analyzing the uppermost Machir Bay outcrop (25 observation points) are of a sufficient data quality to be included in the study.

216 Finally, the Edinburgh interval is the uppermost stratigraphic interval that was investigated by the Curiosity rover (Fig. 3A). This interval contains trough cross-bedded 217 sandstone with cross-set dip directions to the west (Banham et al., this issue). In this interval, 218 concretions were very rare, and one drilled sample (Edinburgh, sol 2715, Fig. 3a) was delivered 219 to the CheMin and SAM instruments (Sutter et al., 2022; Rampe et al., 2020). The Edinburgh 220 drilled sample contained a high abundance of plagioclase feldspar (28.9 wt%), pyroxene (20.2 221 222 wt%), olivine (8.4 wt%) and magnetite (10.2 wt%), in addition to K-feldspar (3.4 wt%), phyllosilicates (7 wt%), trace fluorapatite and an X-ray amorphous component (20 wt%) (Thorpe 223 et al., this issue; Rampe et al., 2020). ChemCam analyzed 30 targets in this interval (259 224 observation points) in addition to 1 target of a sedimentary float rock (Cheviot Hills, sol 2696) 225 and 1 soil (Digg, sol 2710). 226

227

228

1.2.1. Constraining aeolian net sediment transport direction and provenance

1.2 A geochemical and mineralogical analysis of Gale crater's ancient dunes

Previous studies of aeolian dunes derived from basaltic material on Earth and Mars have 229 230 shown that major element chemistry can provide information on mineral sorting processes that can in turn constrain net sediment transport direction and changes in the sediment source 231 (Baratoux et al., 2011; Bedford, et al., 2020a; Cornwall et al., 2015; Sara, 2017; Mangold et al., 232 2011; Mountney et al., 2004; O'Connell-Cooper et al., 2017; Rampe et al., 2018). Basaltic sand 233 grains can be monomineralic, composed of olivine, pyroxene, and feldspar grains, derived from 234 volcanic or impact glass, or polymineralic, composed of a mixture of basaltic minerals and glass 235 (Cornwall et al., 2015; Fedo et al., 2015). The minerals within each grain determine the density, 236 shape, and resistance to physical and chemical weathering that in turn determine how they will 237 be sorted hydrodynamically with transportation in an aeolian environment (Cornwall et al., 2015; 238 Fedo et al., 2015; Jerolmack et al., 2006; Morrone et al., 2017; Willetts et al., 1982). 239

On Earth, most mature aeolian sands are dominated by quartz as this mineral is the most 240 resistant to physical and chemical weathering (Muhs, 2004; Pettijohn, 1954). Mafic minerals in 241 sedimentary environments on the Earth are uncommon because Earth's continental crust largely 242 consists of felsic materials and reactive basaltic materials are easily chemically weathered under 243 Earth's active hydrological cycle (Nesbitt et al., 1992). However, on Mars, the crust is 244 predominantly basaltic, and the cold and arid climate of the Amazonian geological period has 245 resulted in physical weathering being the main driver in the diminution of aeolian sediments. 246 Physical abrasion studies by Cornwall et al (2015) show that olivine grains are rapidly abraded to 247 a spherical shape that aides its transportation in aeolian environments. Meanwhile the cleavage 248 planes of feldspar and pyroxene act as zones of weakness in monomineralic sand grains and 249 hinder the sphericity of these grains. Glass and polymineralic basaltic grains show the most 250 resilience to physical abrasion. These properties result in olivine minerals, glass, and basaltic 251

- fragments being transported farther and concentrating in mature aeolian sands, relative to
- feldspar grains. Mineral sorting studies of basaltic aeolian sands in Iceland show that an increase
- in MgO + FeO_T with distance from the source corresponds to an increase in olivine relative to
- feldspar regardless of the presence of basaltic grains and glass (Baratoux et al., 2011; Sara, 2017;
- 256 Mangold et al., 2011). As the mineral sorting acts on feldspar minerals and olivine, it is possible
- to use the variation in MgO + FeO_T relative to $Al_2O_3 + Na_2O + K_2O$ to measure the variation in the mafic and felsic basaltic materials, respectively (Baratoux et al., 2011; Mangold et al., 2011).
- 259 In Gale crater, several studies have used the geochemical data from the ChemCam instrument to probe the geochemical variation in aeolian sediments in relation to mineralogical 260 variations (Bedford et al., 2020a; Cousin et al., 2017; Meslin et al., 2013). ChemCam 261 262 geochemical data of modern aeolian sediments from Gale crater show that locally derived sands consist of coarse feldspar grains in addition to finer mafic material (Cousin et al., 2015; Meslin et 263 al., 2013), whereas, the mature aeolian sands of the Bagnold dune field only contain feldspar 264 within polymineralic grains, or as fine monomineralic grains (Cousin et al., 2017; Rampe et al., 265 2018). A statistical geochemical analysis of ChemCam data from the Emerson and Naukluft 266 plateaus was applied to determine trends associated with mineral sorting during transportation 267 prior to deposition and lithification (Bedford et al., 2020a). The results showed that the Emerson 268 plateau had a higher abundance of MgO and FeO_T relative to the Naukluft plateau, which was 269 more felsic in chemistry and had a greater proportion of coarser feldspar grains identified in 270 ChemCam Remote Micro-Imager (RMI) images (Bedford et al., 2020a). The geochemical 271 variation of the Stimson formation at the Emerson and Naukluft plateaus indicated that net 272 sediment transport direction was from southwest-to-northeast (Bedford et al., 2020a), supporting 273 the results of Banham et al. (2018). This net sediment transport direction is opposite to that of the 274 active Bagnold dunes, and when the geochemistry of the ancient dunes and modern dunes are 275 276 directly compared to one another, they show geochemical differences indicative of different basaltic provenances (Bedford et al., 2020a). The geochemistry of the Stimson formation 277 sandstone was more similar to that of the subalkaline basaltic sediments that are a source for the 278 Bradbury group fluvial sandstones, suggesting that the two geological units either shared a 279 similar sediment source, or that the Stimson sandstones may consist of recycled Bradbury group 280 material (Bedford et al., 2020a). Due to the variation in cross-set dip directions between intervals 281 in the Stimson formation at the Greenheugh capping unit (cf Banham et al., this issue), we will 282 compare the geochemistry of each interval at the Greenheugh pediment to that of the Stimson 283 formation at the Emerson and Naukluft plateaus using the methods of Bedford et al. (2020a) to 284 determine if the geochemical variation supports changes in sediment transport direction and 285 potentially sediment provenance. 286
- 287

1.2.2. Constraining aqueous alteration of the Stimson sandstone

The alteration features of the Stimson sandstone have been of particular interest in several studies as they can provide information on the nature of the groundwater that existed after the habitable river-lake system ended, and after the erosional moat formed within Gale crater. Although the timing of Stimson deposition and lithification on the slopes of Mt Sharp is not well constrained, it is likely that the Stimson sandstones formed in the mid- to late-Hesperian (Banham et al., 2021), making the alteration features within the Stimson formation representative of some of the youngest aqueous alteration events observed in situ, at the time of this

295 publication. Previous research suggests that the Stimson formation at the Emerson and Naukluft

296 plateaus was potentially lithified by a rising capillary fringe that did not interact with the

depositional surface (Banham et al., 2018). This interaction between groundwater and mafic
 minerals in the aeolian sediments (particularly olivine) caused the formation of an iron oxide and

minerals in the aeolian sediments (particularly olivine) caused the formation of an iron oxide a
 X-ray amorphous materials as cement, isochemical to the previous sandstone deposit, and

preferentially cemented as concretions closer to the unconformity (Bedford et al., 2020a;

301 Hausrath et al., 2018).

After lithification of the Stimson sandstone, repeated episodes of groundwater 302 emplacement and desiccation formed the polygonal fractures and calcium sulfate hydrofracture 303 veins in the Stimson sandstone (Frydenvang et al., 2017; Hausrath et al., 2018; Kronyak et al., 304 2019a, 2019b; Yen et al., 2017, 2021). Multiple alteration features occur within the Stimson 305 formation of the Emerson and Naukluft plateaus such as concretions, mineral veins and the 306 alteration halos which were sampled at Greenhorn and Lubango (Fig. 4). Concretions are more 307 common towards the unconformity, similar to what is observed at the Greenheugh pediment 308 (Banham et al., 2018, this issue; Bedford et al., 2020a) (Fig. 4). An analysis of concretionary 309 versus non-concretionary Stimson bedrock at the Emerson and Naukluft plateaus by the 310 ChemCam instrument showed that there is no statistically significant geochemical variation of 311 these features to the surrounding bedrock (Bedford et al., 2020a). This supports the hypothesis 312 that the concretions at these localities likely formed through in situ isochemical alteration and 313 represent parts of the bedrock that were preferentially cemented during the lithification of the 314 sandstone (Banham et al., 2018; Bedford et al., 2020a). 315

Light-toned, silica-rich Stimson bedrock surrounding large fractures, known as silica-rich 316 alteration halos, are present in the Stimson formation at the Emerson and Naukluft plateaus 317 (Bedford, et al., 2020a; Frydenvang et al., 2017; Yen et al., 2017). Silica-rich halos are 318 geochemically characterized by an increase in SiO₂ abundances up to 80 wt% and a decrease in 319 all other major element oxides except TiO2 and K2O whose compositions are scattered across 320 observation points (Bedford et al., 2020a; Frydenvang et al., 2017; Yen et al., 2017). Two 321 hypotheses exist for the formation of the halos. Either, the halo alteration occurred through the 322 remobilization of silica from the underlying mudstones by neutral to alkaline pH fluids into the 323 324 fractures in the Stimson bedrock where they precipitated to form the halos (Frydenvang et al., 2017). Alternatively, it has been hypothesized that they formed as a result of acidic fluids 325 326 circulating through the fractures, interacting with the mafic minerals in the surrounding bedrock, remobilizing their cations and forming a silica-rich residue (Hausrath et al., 2018; Yen et al., 327 2017). The timing of the silica-enrichment that formed the light-toned halos relative to the 328 329 formation of the calcium sulfate veins and cement is currently unknown, however, previous studies agree that these represent some of the latest stages of diagenesis in the Stimson formation 330 and underlying Murray formation at these localities (Frydenvang et al., 2017; Yen et al., 2017). 331

Polygonal fractures with average widths of 7.5 m can also be identified by orbital images and from the surface in the Stimson formation at the Emerson plateau, Naukluft plateau, and to a lesser extent at the Greenheugh pediment (Kronyak et al., 2019b). These polygonal fractures are similar in appearance to the surrounding bedrock, only evident as they are slightly raised due to their greater erosion resistance (Kronyak et al., 2019b). Based on their geomorphology, the polygons likely formed as a result of multiple cycles of expansion and contraction of the bedrock caused by the recharge of near-surface groundwater preferentially cementing these zones (Kronyak et al., 2019b). The polygons are not likely related to the diagenesis that formed the

- silica-rich halos or the calcium sulfate mineral veins as these features do not consistently
- correlate spatially across the Siccar Point group (Kronyak et al., 2019b).

The small sampling size of the ChemCam laser in addition to the thorough documentation of each target by RMI images before and after sample acquisition makes it possible to isolate observation points with alteration features and compare them to the bedrock, constraining alteration processes. Here, we compare alteration features and examine variations in potential sandstone cementing agents between the different Stimson formation locations in order to further constrain the spatial scale of these late-stage aqueous alteration events in Gale crater.

347 348

349 2 Materials and Methods

350 2.1 The ChemCam instrument

The ChemCam instrument suite utilizes Laser-Induced Breakdown Spectroscopy (LIBS) 351 to obtain elemental analyses of a geological target situated over ~ 2 m from the rover mast, with 352 most of the targets selected at a distance less than 5 m from the rover mast (Maurice et al., 2012, 353 2016; Wiens et al., 2012). LIBS spectra are pre-processed to remove ambient light background, 354 continuum background, and noise (Wiens et al., 2013). The quantification of elemental 355 abundances is obtained for major elements (Si, Ti, Al, Fe, Mg, Ca, Na, K) through a combination 356 of partial least squares regression sub-models and independent component analysis using a large 357 spectral training set (Anderson et al., 2017; Clegg et al., 2017; Forni et al., 2013). Minor and 358 trace elements are quantified individually, often using the same spectral library. Relevant to this 359 work, MnO quantification is described in Gasda et al. (2021). Hydrogen quantification has been 360 shown to require special attention to the matrix (Rapin et al., 2017a; Rapin et al., 2017b; 361 Schröder et al., 2015; Thomas et al., 2018); nevertheless it has proven useful in some specific 362 cases (e.g., Rapin et al., 2016), and rough quantification can be used for large-scale studies 363 (Thomas et al., 2020). The ChemCam team discovered that while F atomic emission lines are not 364 generally detectable, F enrichments can be observed and quantified via the CaF molecular 365 emission band, which has indicated the presence of apatite, fluorite, and mica grains (Forni et al., 366 2015). 367

The uncertainty of major elemental abundances based on ChemCam data has been 368 presented as accuracy and precision to 1σ (Anderson et al., 2017; Clegg et al., 2017; Wiens et al., 369 2012) and is provided in the Supplementary information (Table S.1). Accuracy relates to the root 370 mean square error of prediction (RMSEP) for representative geological samples that share 371 abundance ranges similar to those in the calibration regression models (Clegg et al., 2017). The 372 precision of the predictions is calculated as the variation observed across the 25 shots that derive 373 the average spectrum for each observation point (Blaney et al., 2014), or across targets from a 374 375 uniform unit (Mangold et al., 2016).

ChemCam targets are taken in a linear or grid raster configuration i.e., 3×3 or 1×10 (Maurice et al., 2016). Each point of the raster is known as an observation point, where each observation point consists of between 30–50 laser shots. Spectra from the first 5 laser shots are not included in the average spectrum and serve to minimize surface dust contamination (Clegg et al., 2017). The ChemCam instrument suite also contains a Remote Micro-Imager (RMI) which takes high-resolution (60–100 μrad) images of targets pre- and post-acquisition of the LIBS
 spectral data in order to provide geological context for the targets in question (Le Mouélic et al.,
 2015).

2.2 The CheMin instrument suite

CheMin housed within the rover body and generates X-ray diffraction (XRD) patterns of drilled 385 rock or scooped sediment samples (Blake et al., 2012; Morrison et al., 2018a). Rietveld 386 refinement and full-pattern fitting techniques are used to quantify abundances of minerals and X-387 ray amorphous material and to determine unit-cell parameters of the major crystalline phases 388 (Blake et al., 2012, Morrison et al., 2018a). Mass-balance calculations using mineral abundances 389 and crystal chemistry from CheMin and geochemical data from APXS provide estimates of the 390 composition of the amorphous component (e.g., Smith et al., 2018). Evolved gas analyses from 391 the Sample Analysis at Mars (SAM) instrument suite help identify clay minerals and amorphous 392 393 materials (e.g., Sutter et al., 2017). In this study, we use the CheMin-derived mineral abundances of the Stimson formation sandstone at the Emerson plateau (Big Sky, sol 1119; Morrison et al., 394 2018; Yen et al., 2017), Naukluft plateau (Okoruso, sol 1332; Morrison et al., 2018b; Yen et al., 395 396 2017), and Greenheugh pediment capping unit (Edinburgh, sol 2711; Rampe et al., 2020) to place the geochemical compositions from the ChemCam instrument into mineralogical context. 397

2.3 The on-board science cameras and target classifications

In this study, we have individually classified each ChemCam observation point using the 399 RMI images according to whether the observation point has analyzed bedrock, secondary 400 diagenetic features, unconsolidated sediments, or float. Where available, images from the 401 402 Curiosity rover's other on-board science cameras, Mastcam (Mast camera) and MAHLI (Mars Hand Lens Imager), were used to support the classification at the scale of the outcrop and target, 403 respectively. Mastcam includes two, fixed-focal length, multispectral, color CCD imagers 404 situated on Curiosity's mast (Malin et al., 2017). The two Mastcam imagers are separated by 405 24.2 cm and positioned approximately 2 m from the surface (Malin et al., 2017). The left 406 Mastcam (M-34) has a focal length of 34 mm, a field of view (FOV) of $20^{\circ} \times 15^{\circ}$, and a pixel 407 scale of 0.22 mrad/pixel (Malin et al., 2017). Meanwhile, the right Mastcam (M-100) has a 100 408 mm focal length, a FOV of 6.8°×5.1°, and a 0.074 mrad/pixel scale of sampling (Malin et al., 409 2017). Data from this instrument were used to identify large-scale features, such as cross-410 411 bedding or alteration halos, and to locate the ChemCam targets with respect to the stratigraphic units. MAHLI is situated on the turret at the end of the Curiosity rover's arm and uses a 2-mega-412 pixel color CCD camera to analyze the textures, structures, and morphologies present at the 413 finest grain size possible for Curiosity's cameras, 0.0042 mm (silt) (Edgett et al., 2012; Mangold 414 et al., 2017). Data from this instrument were used to provide additional, high resolution, color 415 data to support the classifications of the black and white images from the ChemCam RMI. 416

Bedrock is defined here as an indurated, lithified unit, possibly containing sedimentary structures such as bedding or laminations, and secondary structural features such as bedrock fractures. As the sedimentary units of the Bradbury group, Mt Sharp group, and Siccar Point group (Stimson formation) have undergone compaction and cementation during diagenesis, it is important to note that these analyses are not completely free of alteration as they include diagenetic cement. However, provided that the diagenesis which formed the sandstone cement was relatively isochemical, the geochemistry can be used to identify mineral sorting trends

before the units became lithified, in addition to sediment source rock characteristics. Concretions 424 are abundant in the Stimson formation which likely formed through the preferential cementation 425 of the Stimson sands (Banham et al., 2018; Bedford et al., 2020a; Potter-McIntyre et al., 2014; 426 Potter et al., 2011). A previous analysis of concretions in the Stimson formation sandstone by 427 both the Alpha Particle X-ray Spectrometer (APXS; e.g., Gellert & Clark, 2015) and ChemCam 428 instruments showed that they have the same geochemical composition to non-concretionary 429 Stimson (Bedford et al., 2020a; Siebach et al., 2017) and are hypothesized to relate to closed-430 system diagenesis that preserves the bulk rock geochemistry (Banham et al., 2018; Bedford et al., 431 2020a; Siebach et al., 2017). This is also true for the Greenheugh pediment capping unit for 432 concretions smaller than 5 mm (Appendix A); as such they are included in the bedrock dataset. 433 However, larger concretions that are greater than 5 mm (Fig. 4), with multiple ChemCam 434 observation points directly on the concretions, and which occur closer to the unconformity have 435 been isolated from the bedrock dataset of all Stimson localities for a more in-depth analysis on 436 the diagenesis that formed these features. As the Stimson formation across all localities has a 437 relatively uniform grain size (Banham et al., 2018; 2021; this issue), we have not classified 438 bedrock targets according to grain size. Bedrock targets do not include analyses of drill tailings 439 and dump piles as these samples have been modified by the rover. Furthermore, bedrock 440 analyses have been restricted to only include observation point data with oxides that sum to 441 between 90–105 wt% to minimize the influence of volatile elements (S, Cl, and H), or elements 442 443 such as Ni which are not nominally quantified in ChemCam data. Removing the potential impact of volatile elements is important for our bedrock dataset as these elements may have been 444 emplaced due to open-system alteration. Restricting the total sum of oxides can further remove 445 any targets that have been contaminated with Martian soil, as these deposits also have higher 446 concentrations of volatiles (Bish et al., 2013; Ehlmann et al., 2017; Lasue et al., 2018; Meslin et 447 al., 2013; O'Connell-Cooper et al., 2017). 448

ChemCam analyses that have targeted secondary diagenetic features in the Stimson 449 sandstone have been placed into an alteration feature dataset and subclassified according to 450 "vein", and "halo" categories. Vein alteration targets relate to light-toned mineral veins that in-451 fill hydrofractures in the Stimson formation (Bedford et al., 2020a; Kronyak et al., 2019a). Halos 452 relate to the light-toned halos that extend into the Stimson bedrock ~20 cm either side of large 453 fractures (Fig. 4) (Bedford et al., 2020a; Frydenvang et al., 2017; Yen et al., 2017). These 454 secondary diagenetic features have been separated from the bulk rock dataset to isolate the 455 effects of later diagenetic alteration from geochemical trends that may relate to the sediment 456 source characteristics of the Stimson formation across all localities, in addition to pre-457 lithification mineral sorting processes, and the formation of the sandstone cement. 458

Other classifications include float and soils. Float targets in Gale crater include pebble to 459 boulder grain size fragments of rock that can be igneous, sedimentary, or meteoritic in nature. 460 Soil targets are classified as unlithified deposits, with a lack of obvious cement that derive fine-461 grained accumulations of sand- to silt-grained particles (Bish et al., 2013; Cousin et al., 2015; 462 Meslin et al., 2013). The Curiosity rover also encountered two large, active, aeolian dune fields, 463 and investigated them in two phases of the Bagnold dune campaigns (Bridges et al., 2018; 464 Lapotre et al., 2018). The aeolian dunes occur as large bedforms, ~ 15 m across, which are dark 465 in color and can be considered a modern analog of the ancient Stimson formation (Achilles et al., 466 2017; Bedford et al., 2020a; Hausrath et al., 2018). Due to the similarities in characteristics of 467

the dune and soil deposits, ChemCam analyses of these deposits have been placed into the soilcategory.

470 *2.4 Statistical methods and data interpretation*

The ChemCam LIBS laser sampling footprint is 350–750 µm for distances 3–7 m from 471 the rover mast (Maurice et al., 2012), which is a similar size to the average grain size of the 472 Stimson formation (336–634 um at the Greenheugh pediment, Banham et al., *this issue*). As 473 such, target analyses are not likely to be representative of whole rock compositions as individual 474 grains are more likely to be sampled (Cousin et al., 2017). However, the bedrock Stimson 475 formation ChemCam dataset to date (August, 2021) consists of 741 data points which permits 476 the use of a statistical approach to both estimate a bulk rock composition and detect geochemical 477 trends that may relate to mineralogical variations in the relatively alteration-free bedrock dataset. 478 For each locality and stratigraphic interval of the Stimson formation, we give the basic statistics 479 480 (mean, median, range, and standard deviations), followed by density contour plots which will be used to estimate the bulk rock composition. Density contours display the data across two 481 variables and represent the density of smoothed data within each pixel on the bivariate plot 482 483 defined according to the bin size (Bedford et al., 2019a; Eilers et al., 2004). Data densities are smoothed according to Eilers and Goeman (2004), and contours are generated using MATLAB 484 (2019a). This method is used to estimate the bulk rock composition which is in turn used as a 485 comparison to the geochemical variation associated with the alteration feature datasets. By 486 comparing the estimated whole rock composition to alteration features, we can constrain how 487 these features formed. 488

489 After the density contour analysis, sandstone components (the minerals and amorphous materials that comprise the sand and cement) are delineated using a hierarchical, agglomerative 490 cluster algorithm on the bedrock dataset with Minitab 19 statistical software (2020), following 491 the methods of Bedford et al. (2020a) and Bedford (2019a). In this study, we use the complete 492 493 linkage measure of association for Euclidean distances with all major oxides, now including MnO, and the total sum of oxides, as the variables. Complete linkages and Euclidean distances 494 were used as these methods provided the best fit for the dataset showing the most defined 495 clusters and was successfully used in a previous investigation of the Stimson formation (Bedford 496 497 et al., 2020a). Clustering methods have also been successfully employed in other studies of Gale crater sediments (Rammelkamp et al., 2021; Gasnault et al., 2019, 2013). Variables were 498 standardized in order to minimize the effect of scale differences. 499

500 **3 Results**

501 *3.1. Distribution of the classified data across the Stimson formation*

In total, 430 ChemCam observation points were acquired during the Greenheugh 502 pediment mini campaign that were analyzed <4 m from the rover and were of good data quality. 503 504 This included 324 bedrock analyses, two analyses of calcium sulfate veins, 10 soil analyses, 10 float analyses, 20 analyses that targeted large (>5 mm) nodules, and 15 analyses of the 505 Edinburgh drill hole wall and tailings (Fig. 5). This is in contrast to the distribution of 506 observation point classifications for the Emerson and Naukluft plateaus which has a much larger 507 proportion of observation points (50% at Emerson plateau, 65% at Naukluft plateaus) that 508 targeted alteration features such as halos, concretions, and calcium sulfate veins (Fig. 5). These 509

proportions are not a result of the true distribution of these features relative to bedrock as 510

- 511 targeting bias to investigate certain features has played a role. Regardless, the diversity of
- alteration at the Greenheugh capping unit is significantly less than the diversity of the features 512
- seen previously, with no alteration halos were identified at the Greenheugh pediment capping 513
- unit, unlike at the Emerson and Naukluft plateaus (Fig. 5) or any obvious raised ridge features in 514
- comparison to the 15 raised ridge analyses at the Emerson plateau spanning 2 targets (Tinder 515
- Box and Duperow), and the 7 raised ridge analyses of the Uubvley and Bloedkoppie targets at 516
- the Naukluft plateau. 517

Calcium sulfate mineral veins are rare at the Greenheugh pediment capping unit with 518 only two observation points targeting a very thin mineral vein (Fig. 4) and rare occurrences of 519 520 veins identified in the Mastcam images. Calcium sulfate veins were also rare at the Emerson and Naukluft plateaus with 15 and 27 ChemCam targets of veins respectively, but where they did 521 522 occur the veins were larger in size and extent than the vein investigated at Greenheugh (Fig. 4). Concretions are also present across all sites. All the large and small concretions at the 523 Greenheugh locality were observed in the lowermost Gleann Beag interval in the ChemCam data 524 and Mastcam images (Fig. 3B and 4). Consequently, there is a similar number of large 525 concretion analyses in the Stimson formation at the Greenheugh pediment (20 points over 4 526 targets) and the Naukluft plateau (29 points over 6 targets). Only three observation points in the 527 targets Dakota and Heath sampled large concretions in the Stimson formation at the Emerson 528 plateau. 529

530 Of the 324 bedrock analyses, 220 constituted data within the 90–105 wt% total sum of oxide range. This suggests that 104 observation points on the Greenheugh capping unit contains 531 a higher abundance of volatile elements (S, H, or Cl) or other elements which are not routinely 532 quantified, 82% of which are within the Edinburgh interval. Due to a distinct geochemistry in 533 two ChemCam targets situated just above the Siccar Point group unconformity, 20 observation 534 points were placed into their own dataset for an analysis of the unconformity relative to the bulk 535 Greenheugh pediment capping unit. 536

537

3.2. Geochemistry of the Greenheugh pediment capping unit compared to the Stimson formation at the Emerson and Naukluft plateaus 538

The geochemistry of the Greenheugh pediment capping unit is similar to that of the 539 Stimson formation analyzed previously, with geochemical distributions, mean compositions, and 540 standard deviations within family to both previous Stimson formation localities for TiO₂, FeO_T, 541 and K₂O (Fig. 6 and Table 1). For all localities, TiO₂, FeO_T, and K₂O have unimodal 542 distributions, with mean compositions and standard deviations of 0.9 ± 0.2 wt%, 19.5 ± 2.8 wt%, 543 and 0.6 ± 0.4 wt%, respectively (Fig. 6). Both distributions for TiO₂ and K₂O have a heavy right 544 tail, which is the most prominent for the Emerson plateau, though these heavy right tails do not 545 significantly skew the median compositions from the mean. Two-way equivalence tests also 546 support that these major element oxides are equivalent across the three Stimson localities within 547 548 the upper and lower limits defined by ChemCam precision (Supplementary material) and they suggest that Na₂O can be claimed as equivalent across the three sites, despite the slight variation 549 in distribution (Table S.2). For all three sites, Na₂O has a mean composition and standard 550 deviation of 3.0 ± 0.8 wt%, with multimodal distributions presenting peaks at 1.8 wt%, 2.7 wt%, 551 3.1 wt%, and a heavy tail to the right. However, the distribution in data between these peaks 552 differ across all sites (Fig. 6). The dominant peak for the Emerson plateau and Greenheugh 553

capping unit is at 2.7 wt%, with both of these localities showing fewer data extending to higher

Na₂O concentrations. In comparison, the Naukluft plateau's dominant peak is at 3.1 wt%, with a greater proportion of data points situated over 4 wt% Na₂O.

- 557
- 558

The mean compositions of SiO₂, Al₂O₃, MgO, CaO, and MnO across the Stimson 559 formation localities are all situated within 1 standard deviation of each other (Table 1), however, 560 variations in the distribution of the data suggests that there may be statistically significant 561 differences, supported by equivalence test results (Table S.2). The Emerson plateau and 562 Greenheugh pediment capping unit both share unimodal SiO₂ distributions with peaks at 45 wt% 563 (Fig. 6). However, the distribution of SiO₂ at the Emerson plateau shows a right skew to higher 564 SiO₂ concentrations up to 57.1 wt%, whereas the Greenheugh pediment capping unit shows a 565 skew to the left with a greater proportion of data points <44 wt% SiO₂ causing the non-566 equivalence between the Emerson plateau and Greenheugh pediment capping unit. The Naukluft 567 plateau has a distinct bimodal distribution for SiO₂ sharing the peak at 45 wt% with an 568 additional, larger peak at 48 wt%. This second larger peak for the Naukluft plateau, in addition to 569 values up to 61.7 wt% SiO₂, results in a lack of equivalence for SiO₂ between the Naukluft 570 plateau and other Stimson localities. The high SiO2 observation points of the Emerson and 571 572 Naukluft plateaus were cross-checked with the locations of the silica-rich alteration halos and the RMI images, and either correlate with large, light-toned mineral grains (Fig. 3d), or they were 573 randomly interspersed within the target raster which is not consistent with the 10s of cms-scale 574 alteration halo features (Bedford et al., 2020a; Frydenvang et al., 2017). Instead, the correlation 575 of these high SiO₂ values with high Na₂O and/or K₂O, and low MgO and FeO_T suggest that they 576 may relate to ChemCam analyses that have targeted a greater proportion of felsic grains (Bedford 577 et al., 2020a). Thus, the statistically significant differences may be indicative of different 578 proportions of minerals comprising the sand (e.g., olivine, basaltic glass, and feldspar) or 579 differences in cement composition at each locality. 580

The Greenheugh capping unit shows statistically significant geochemical variation from 581 the Emerson and Naukluft plateaus for Al₂O₃, and CaO, which is on average lower at the 582 Greenheugh capping unit, and MnO and MgO which is higher in mean abundance (Fig. 6). Al₂O₃ 583 compositions show a bimodal distribution for the Emerson and Naukluft plateaus, with the 584 histograms showing peak abundances at 10 wt% and 17 wt% for the Emerson plateau, and peaks 585 at 12 wt% and 17 wt% at the Naukluft plateau. Meanwhile, the Greenheugh pediment capping 586 unit has a multimodal distribution with peaks at 8 wt%, 10 wt%, and 16 wt%. The additional 587 peak of the Greenheugh capping unit at 8 wt% has resulted in lower mean (10.3 ± 0.2 wt%) and 588 median (9.6 wt%) compositions compared to the rest of the Stimson formation localities with 589 average compositions of 11.3 ± 0.2 wt% and 12.1 ± 0.2 wt% for the Emerson and Naukluft 590 plateaus, respectively. 591

Both the Emerson and Naukluft plateaus have a unimodal distribution for MgO, with a peak at ~7.5 wt%, while the Greenheugh pediment capping unit has a bimodal distribution with peaks at 7 wt% and 9 wt%. All localities have a right tail to MgO concentrations of >17 wt%, generating a slight right skew in the datasets (Fig. 6).

596 The distribution in CaO is also the same between the Emerson and Naukluft plateaus in 597 that both are unimodal with a dominant peak at 6.4 wt%. Meanwhile, the Greenheugh pediment capping unit has a bimodal distribution with the 6.4 wt% peak shared with the other localities, inaddition to a different, more dominant peak at 4.0 wt%.

MnO has a unimodal distribution for all localities with a strong peak at 0.2 wt%, however, outliers exist to high concentrations at the Emerson plateau (up to 2.6 wt%) and Greenheugh capping unit (4.0 wt% in target Aegis_post_2730). The outlying MnO values at the Greenheugh pediment capping unit are more numerous than those of the Emerson plateau, and all occur within the Edinburgh interval bedrock targets, generating a mean concentration that is ~1.6x that of the Emerson and Naukluft plateaus.

To summarize, the Stimson formation at the Greenheugh pediment capping unit is 606 situated within the compositional range of earlier Stimson units, according to the standard 607 deviation of the dataset. However, statistically significant variations in the geochemical 608 distribution of the dataset exist, as is evident in the mean, equivalence test results, and the 609 histograms of major element oxides. The Greenheugh pediment capping unit has a higher 610 abundance of MgO and MnO, and lower abundances of SiO₂, Al₂O₃, and CaO relative to the 611 Stimson formation at the Emerson and Naukluft plateaus that may relate to different abundances 612 of sandstone components such as mafic minerals, felsic minerals, or sandstone cement. 613

3.3 Geochemical variation associated with diagenetic features and the unconformity at the Greenheugh capping unit.

Calcium sulfate veins across the Stimson formation show a decrease in all major oxide 616 compositions and total sum of oxides, except for CaO that can reach values up to 50 wt% 617 (Bedford et al., 2020a). The two calcium sulfate veins targeted on the Greenheugh pediment 618 capping unit were both situated in the Edinburgh interval and as a result of both veins being very 619 thin (<1 mm wide), the analyses of these veins only show a slight increase in CaO compared to 620 the average of the Edinburgh interval. Calcium sulfate can also be a cement component. For 621 Stimson analyses across all localities with total sum of oxides less than 90 wt%, 23 analyses out 622 of 191 share a similar geochemical trend to the calcium sulfate mineral veins, suggesting that 623 calcium sulfate cement is present in these ChemCam observation points. For the Greenheugh 624 pediment capping unit, 11 observation points that targeted bedrock have geochemical signatures 625 of calcium sulfate cement, suggesting that it is a small component of the sandstone and is 626 variable in its distribution. 627

Large concretions (>2 cm) in the Stimson formation preserve sedimentological features 628 such as laminations and therefore likely represent parts of the sandstone that have been 629 preferentially cemented (Banham et al., 2018). To conduct a closer investigation of how the 630 sandstone cement of the Stimson formation varies across the Greenheugh capping unit, we have 631 isolated the compositions of these large concretions and plotted them alongside the geochemical 632 633 abundances of the non-concretionary bedrock and small concretionary bedrock (S.5). Similar to the results of Bedford et al. (2020) for the Emerson and Naukluft plateaus, no statistically 634 635 significant geochemical variation is present across concretions (Fig. S.5).

To investigate how the geochemistry of the Greenheugh pediment capping unit varies with respect to the basal Siccar Point Group unconformity, and in turn how it may vary for the immediately underlying Carolyn Shoemaker formation, we have classed the bedrock data in relation to the stratigraphic levels just above (Stimson Unconf.), and just below (CS Unconf.) the unconformity (Fig. 7 and 8). In addition to the targets closest to the unconformity, we also 641 compare the results of analyses acquired at the base of the pediment around the Hutton drill hole

642 (CS Hutton) and the analyses acquired farther above the unconformity in the Stimson formation

643 stratigraphic intervals (Glean Beag interval, Ladder interval, and Edinburgh interval). As the 644 previous literature has suggested that the unconformity may have been a conduit for diagenetic

fluids to travel and alter the underlying Carolyn Shoemaker formation (Achilles et al., 2020;

Fraeman et al., 2020; Horgan et al., 2020; Rampe et al., 2020), we have not restricted the

bedrock totals for this analysis, to allow for geochemical variation associated with secondary
 phases that formed through water-rock interactions to be more apparent. The mean compositions

and their associated standard errors of the Stimson formation and Carolyn Shoemaker formation are also shown for comparison (Fig. 7). Figure 7 shows that the Carolyn Shoemaker formation

leading up to the unconformity is depleted in SiO_2 and H but enriched in CaO relative to the mean composition of the Carolyn Shoemaker formation. The Stimson formation just above the

unconformity (Stimson Unconf.) has a unique geochemistry that is low in MgO, Na₂O, and

MnO, low in total sum of oxides, but high in CaO and H relative to the mean composition of the

655 Stimson formation. These geochemical characteristics of the base of the Stimson formation

656 indicate that it may have either been affected by diagenetic fluids, been derived from a different 657 sediment provenance, or represents a complex mixture of both processes.

658 *3.4 The distribution of halogens across the Stimson formation.*

For the bulk rock dataset, all Stimson formation localities have a unimodal distribution of 659 H peak area about 0.004, however, the Stimson formation at the Greenheugh pediment has a 660 bimodal distribution with a second modal peak of H peak abundances at a value of 0.008 (Fig. 661 6). This contrasts significantly with the right-skewed unimodal distributions of the Emerson and 662 Naukluft plateaus. Looking at the distribution of the H data in relation to the stratigraphic 663 intervals at the Greenheugh pediment (Fig. 8A), these peak area values are divided between the 664 Gleann Beag interval that contains concretions and is situated close to the unconformity, and the 665 Ladder and Edinburgh intervals that do not contain notable concretions and are higher in the 666 stratigraphy (Fig. 8). The Gleann Beag interval shares the modal H peak area with that of the 667 Stimson formation at the Emerson and Naukluft plateaus, whereas the Ladder and Edinburgh 668 intervals exclusively have the higher H peak area. ChemCam analyses of the non-concretionary 669 bedrock have the highest H peak area (0.0052) compared to small concretions (0.0046) and large 670 concretions (0.0039), however the change in surface roughness between these classifications 671 likely impacts the H signal (Rapin et al., 2017a; Schröder et al., 2015). As such, we do not 672 consider the H peak area to differ significantly across the concretions. H peak area is also 673 variable among the other diagenetic features, with the calcium sulfate veins having the highest H 674 675 abundances (mean of 0.01) of all features and bedrock analyzed by ChemCam. Silica-rich halos also had a slightly higher mean H abundance (0.007) than that of the average Stimson, unaltered 676 bedrock (0.005). Only MnO has a possible positive correlation with H abundance for targets on 677 the Edinburgh interval with MnO > 1 wt% (r^2 49%, p = 0.011). 678

F detections across the Stimson formation are rare. In total, there are 20 bedrock observation points (16 non-concretionary bedrock, four small concretion-rich bedrock), and two halo analyses across all Stimson localities that have F above the limit of detection. Only one observation point in the Edinburgh interval of the Greenheugh pediment capping unit has F, contrasting with the high F targets in the underlying Carolyn Shoemaker formation of the Hutton zone (Dehouck et al., *this issue*; Gasda et al., *this issue*). For the bedrock targets, F has a potential positive correlation with CaO (r^2 48%, p = 0.001; S.3) and shares no correlations with any other major element oxide. F has the largest range in abundances for targets close to the
unconformity (0.7 wt% F), compared to those situated over 1.5 m above (0.2 wt% range in F;
Fig. S.4).

000 1 lg. 5.+).

689 *3.5 Stimson bedrock cluster analysis results*

The results of the cluster analysis show that the Stimson formation bedrock dataset can be 690 divided into 5 clusters (Fig. 9). Cluster 1 is defined as the cluster with the highest mean MgO 691 content (14.7 \pm 1.1 wt%), higher than average MnO (0.7 \pm 0.2 wt%) and the lowest Al₂O₃ (6.9 \pm 692 0.4 wt%), CaO (3.6 ± 0.4 wt%), Na₂O (2.3 ± 0.1 wt%), and K₂O (0.3 ± 0.1 wt%) of all clusters 693 (Table 2). Nineteen data points have a Cluster 1 membership from across the Stimson formation 694 with eight at the Greenheugh pediment (Fig. 9, 10). Cluster 1 is most similar to Cluster 2 and is 695 differentiated by a similarity level of 36.3 (see Bedford (2019) for the methods to calculate 696 similarity level). Cluster 2 has a moderately high abundance of MgO (8.6 ± 0.1 wt%) with the 697 698 greatest variation in TiO₂ (standard deviation of 0.21 wt%) and the greatest maximum TiO₂ (2.18 wt%) of all clusters. Cluster 2 differs from Cluster 1 by a slightly greater mean abundance of 699 Al_2O_3 (9.7 ± 0.1 wt%) and CaO (5.6 ± 0.1 wt%), and a lower abundance of MgO (Table 2). In 700 total, 372 data points have a Cluster 2 membership across all localities. It is the largest cluster, 701 comparable in size to Cluster 3 and is the most representative of the Greenheugh pediment 702 average composition as it is situated at the focal composition for this locality (Fig. 10, 12). 703

The next closest cluster is Cluster 3 with a similarity to Clusters 1 and 2 of 28.3. Cluster 3 704 705 is differentiated by Clusters 1 and 2 by having on average higher abundances of SiO_2 (48.5 ± 0.2) wt%), Al₂O₃ (13.1 \pm 0.1 wt%), CaO (6.0 \pm 1.8 wt%), and lower abundances of FeO_T (18.7 \pm 0.1 706 wt%) and MgO (6.5 ± 0.1 wt%). Cluster 3 has a total of 330 observation points with the majority 707 situated on the Naukluft plateau (Fig. 10). Clusters 2 and 3 are the largest clusters in the dataset 708 and as such are situated within the density contours of the Stimson bulk rock (Fig. 12). Similar to 709 the results of Bedford et al. (2020a) for the main clusters 2 and 3, the standard deviations are 710 711 higher than for the other clusters, which could indicate that some of the data points are not well bound to these clusters. However, the bimodal distribution of SiO2 and Na2O of the Naukluft 712 plateau (Fig. 6), which contains the majority of Cluster 3 memberships, shows peak abundances 713 714 of 45 wt% and 48 wt% SiO₂, and 2.7 wt% and 3.4 wt% Na₂O that align with the Clusters 2 and 3 715 averages, respectively. The variation of Al₂O₃ between the Emerson plateau and Greenheugh pediment localities that share a peak of 10 wt% Al₂O₃ also differs from the main peak 716 composition for Al₂O₃ of the more Cluster 3-rich Naukluft plateau (13 wt%), suggesting a more 717 bimodal distribution between all Stimson localities for this major element oxide. As such, this 718 indicates that Clusters 2 and 3 are an inherent part of the dataset and accurately represent the 719 most common sandstone components across all localities, despite the overlap of some points. 720 Furthermore, as Clusters 2 and 3 derive from a total of 702 points, the overlap of <10 % of 721 datapoints does not significantly skew the analysis of these two clusters. 722

The two most distinctive clusters are Clusters 4 and 5 (Fig. 9). Cluster 4 consists of 15 data points, distributed relatively equally between localities (four at Emerson, five at Naukluft, and six at Greenheugh, Fig. 10). Cluster 4 is defined as having the highest abundances of SiO₂ ($54.50 \pm 0.84 \text{ wt\%}$), Al₂O₃ ($17.97 \pm 0.61 \text{ wt\%}$), Na₂O ($5.74 \pm 0.34 \text{ wt\%}$), and higher than average K₂O for the Stimson formation ($0.97 \pm 0.16 \text{ wt\%}$). Cluster 5 only has five data points, one at the Emerson plateau and 4 in the Edinburgh interval of the Greenheugh capping unit and is defined by its high MnO concentration ($2.92 \pm 0.19 \text{ wt\%}$) and high K₂O concentration similar

- to Cluster 4 (1.0 ± 0.3 wt%). Aside from the MnO and K₂O concentration of Cluster 5, it is very
- similar in composition to the MgO, FeO_T , and TiO_2 -rich Clusters 1 and 2.

732 **4 Discussion**

733 *4.1 Geochemical variation associated with mineral sorting or sediment source*

The sedimentary architecture of the Stimson formation is interpreted as resembling dry 734 aeolian sand dunes (Banham et al., 2018, 2021, this issue). In a dry aeolian environment, the 735 coarsest and densest grains are rolled across the surface as they are transported by surface creep, 736 while the majority of sand-sized particles are temporarily lifted across the surface and 737 redeposited elsewhere as they travel via saltation. Meanwhile, grains finer than sand are 738 suspended in the wind. As such, mud to silt-sized grains and platy minerals such as 739 phyllosilicates are not likely to be deposited in a dry aeolian bedform (Kocurek et al., 2012), 740 supporting that the phyllosilicates at the Greenheugh capping unit are likely to be authigenic. 741 Sediments that are transported and deposited in dry aeolian environments are also subjected to 742 physical abrasion and less chemical weathering (Thomas et al., 2005). As such, sediment grains 743 that are less dense with a shape that is efficient at transportation by saltation or surface creep, and 744 that are resistant to physical weathering are more likely to be transported farther in a dry aeolian 745 environment and concentrated in distal deposits relative to those which do not meet these 746 criteria. 747

The geochemical trends of the majority of ChemCam data between mafic and felsic 748 749 minerals, and the similarity in composition between concretionary and non-concretionary bedrock, also suggests that alteration was likely minimal away from the unconformity for the 750 Greenheugh pediment capping unit. Minimal alteration at Greenheugh is further supported by the 751 presence of 8 wt% olivine in the Edinburgh drilled sample (Rampe et al., 2020), which is only 752 slightly less than the olivine abundances for the active aeolian sands at the Bagnold Dunes (e.g., 753 9.6 wt% at Ogunquit Beach) (Rampe et al., 2018). This, in addition to the low degree of 754 chemical weathering during transportation for dry aeolian systems (Jiang et al., 2019; Zhu et al., 755 2009), suggests that the geochemistry of the Stimson formation at the Greenheugh pediment 756 capping unit is likely to be representative of the mineralogy of the dune sands at the time of 757 deposition. 758

759

4.1.1. The relationship of the CheMin mineralogy to the ChemCam derived clusters

760 The crystal chemistries of the Edinburgh sample are not yet available, so we will compare our geochemical results to the crystal chemistries of the Big Sky drilled sample which 761 is considered the least affected by alteration of the Emerson and Naukluft bedrock samples (Yen 762 et al., 2017). As olivine is no longer present at the Emerson and Naukluft plateaus due to 763 diagenesis (Bedford et al., 2020a; Hausrath et al., 2018; Siebach et al., 2017), we use the olivine 764 crystal chemistry of the Gobabeb sample from the Bagnold dunes. (Achilles et al., 2017; 765 Morrison et al., 2018b). The primary igneous minerals present in the non-halo related drilled 766 samples Big Sky and Okoruso drilled at the Emerson and Naukluft plateaus respectively are 767 plagioclase feldspar (36.5 ± 1.8 wt% Big Sky, 27.2 ± 1.7 wt% Okoruso), K-feldspar (1.1 ± 0.6 768 wt% Big Sky, 1.9 ± 0.7 wt% Okuroso), and pyroxene (25.4 ± 2.3 wt% Big Sky, 20.7 ± 1.1 wt% 769 Okoruso) (Yen et al., 2017). The primary igneous minerals of the Edinburgh drilled sample from 770

the Greenheugh pediment capping unit also contains plagioclase feldspar (28.9 wt%), K-feldspar

(3.4 wt%), and pyroxene (20.2 wt%), in addition to olivine (8.4 wt%) (Rampe et al., 2020). On a 772 773 plot of $Al_2O_3 + Na_2O + K_2O$ versus MgO + FeO_T, all five clusters are situated along a trendline that extends from Stimson feldspar mineral endmembers (plagioclase feldspar and K-feldspar) in 774 775 the top left of the plot, to mafic mineral endmembers (pyroxene and olivine) situated in the bottom right (Fig. 12A). Clusters 1 and 2, in addition to Cluster 5 are situated closer to the mafic 776 mineral endmember of the plot. The Greenheugh pediment average composition is most similar 777 to Cluster 2 (Fig. 11A). The geochemical composition of Cluster 2 also relates strongly to the 778 lower SiO₂, Al₂O₃, Na₂O, and K₂O, and higher MgO dominant peak compositions of the 779 Emerson plateau and Greenheugh capping unit compositions, and the lesser of the bimodal peak 780 compositions for the Naukluft plateau (Fig. 6). This geochemical relationship of Cluster 2 with 781 one of the bimodal peak compositions and greater chemical affinity to mafic minerals suggests a 782 high proportion of mafic minerals and their diagenetic products being targeted by the ChemCam 783 laser. Cluster 2 is most prevalent in the Greenheugh capping unit (55% total datapoints) followed 784 by the Emerson plateau (46.2% total datapoints), with the fewest data at the Naukluft plateau 785 (27.5% total data points; Fig. 10), supporting the greater influence of this cluster on the 786 geochemical distribution present in the histograms. 787

Cluster 1 plots closer to the mafic mineral endmembers than Cluster 2 and likely 788 represents a mixture of mafic minerals (Fig. 11A). Few observation points (n = 17) have Cluster 789 1 memberships suggesting that it is rare for the mafic end members to be targeted on their own, 790 though it is present at each Stimson locality in relatively equal proportions (6 points at Emerson, 791 5 points at Naukluft, and 6 points at Greenheugh). An image analysis of Cluster 1 and 2 792 memberships (Fig. 13) show that Cluster 1 and 2 points likely relate to ChemCam targets that 793 794 have sampled dark-toned and fine-grained parts of the sandstone. As such, the fine grain size and association of mafic minerals with the formation of the sandstone cement may have decreased 795 the likelihood of a mafic endmember being targeted with the ChemCam laser, instead preferring 796 the more mixed composition of Cluster 2. 797

The five ChemCam observation points with a Cluster 5 membership also show a 798 geochemical composition which is closer towards the mafic mineral endmembers (Fig. 12A), but 799 with high abundances of MnO (Fig. 11B). MnO can be present as a trace element in mafic 800 minerals such as olivine and pyroxene, though not at the high abundances present in Cluster 5, 801 suggesting that this cluster may relate to the local remobilization of Mn and Fe from mafic 802 minerals that contain trace amounts of MnO in their crystal structure. As such, we interpret 803 Clusters 1 and 2 as being indicative of the relative abundance of mafic minerals initially present 804 within the deposited aeolian sands, a proportion of which was later altered to sandstone cement 805 during early diagenesis, creating Cluster 5. 806

Clusters 3 and 4 are, in turn, situated closer to the feldspar mineralogical endmember, 807 with Cluster 3 more representative of the geochemistries of the Emerson and Naukluft plateaus 808 (Fig. 12A). Our Clusters 3 and 4 are geochemically similar to Clusters 3 and 5 of Bedford et al. 809 (2020a) that were interpreted to represent more felsic minerals relative to mafic minerals and 810 cement targeted by the ChemCam laser. The ChemCam RMIs show that the LIBS spots with 811 Cluster 3 and 4 memberships have preferentially targeted light-toned mineral grains in the 812 sandstone, further supporting a felsic nature for these clusters (Fig. 12B). Cluster 3 more strongly 813 influences the bulk composition of the Emerson and Naukluft Stimson formation, skewing the 814 focal composition away from Cluster 2 and causing a greater number of contour intervals to 815 extend around the Cluster 3 composition (Fig. 11A). Looking at how the Cluster memberships 816

- vary with location (Fig. 10), Cluster 3 is most abundant at the Naukluft plateau (n = 173),
- followed by the Emerson plateau (n = 101), with the lowest abundance of Cluster 3 present at the

Greenheugh pediment capping unit (n = 56). The average composition of Cluster 3 also

scorrelates to the peak composition of the Stimson formation that is situated at higher SiO_2 ,

Al₂O₃, and Na₂O, and lower MgO and FeO_T, particularly at the Emerson plateau where Cluster 3 is most abundant (Fig. 6). As such, we interpret Clusters 3 and 4 as being representative of the

is most abundant (Fig. 6). As such, we interpret Clusters 3 and 4 as being representative of the proportion of felsic minerals that were present in the initially deposited Stimson aeolian sand

- 824 dunes.
- 825

4.1.2. Implications for sediment sorting and provenance of the Stimson dune sands

Basaltic sediments transported in aeolian environments on Earth in localities such as the 826 Dyngjusandur sand sheet, Iceland, are largely derived of volcanic glass, olivine, pyroxene, 827 feldspar, and lithic grains (Baratoux et al., 2011; Cornwall et al., 2015; Mangold et al., 2011). 828 829 These basaltic components are also present in the active aeolian sands on Mars, including those of the active Bagnold dune fields in Gale crater (Achilles et al., 2017; Lapotre et al., 2018; 830 Rampe et al., 2018). Of these basaltic components, feldspar and augite are the least likely to be 831 832 transported far as their cleavage planes and more tabular/elongate shapes make them both vulnerable to physical abrasion and less efficient at transportation by saltation and surface creep 833 in aeolian environments (Baratoux et al., 2011; Cornwall et al., 2015; Mangold et al., 2011; 834 Mountney et al., 2004). Olivine has a higher density (from forsterite at 3.29 g/cm^3 to favalite at 835 4.39 g/cm³) than most basaltic materials which could decrease the efficacy in which it can be 836 837 transported far in aeolian settings, particularly at low wind speeds (Fedo et al., 2015). However, its similar Mohs hardness to quartz and concoidal fractures promoting rapid rounding of the 838 olivine grains can make it more likely to persist in the aeolian transport of basaltic sands relative 839 to feldspar and pyroxene, provided the sands are heterogeneous (Cornwall et al., 2015). The most 840 resilient basaltic components in aeolian sediments are volcanic glass and polycrystalline basalt 841 grains. Hence, if pyroxene and feldspar are present in distal aeolian sands they are likely 842 entrained within polycrystalline basaltic grains (Cornwall et al., 2015). Geochemical trends 843 relating to an increase in MgO downwind in the Dyngusandur sand sheet, Iceland, supports an 844 845 increase in olivine minerals (Baratoux et al., 2011; Sara, 2017; Mangold et al., 2011).

The previous cluster analysis by Bedford et al. (2020a) hypothesized that physical grain 846 segregation was the main cause of geochemical variation between the Emerson and Naukluft 847 plateaus. In addition to the geochemical variation that related to mineralogical variation, a more 848 local sand source for the Naukluft plateau was further supported by the presence of coarse 849 feldspar grains which are less abundant at the Emerson plateau (Bedford et al., 2020a; Fig. 3d), 850 and the geochemical results supported the interpretation of a southwest-northeast net sediment 851 transport direction determined through cross-bed dip directions (Banham et al., 2018). The 852 Greenheugh pediment capping unit has more MgO relative to the other Stimson localities (Fig. 6, 853 11), in addition to a greater abundance of Clusters 2 and 5 (Fig. 11), indicative of higher 854 abundances of mafic minerals. This is supported by the mineralogical results from CheMin that 855 show a relatively high concentration of olivine (Rampe et al., 2020), indicating that the 856 Greenheugh capping unit may be enriched in mafic components as a result of mineral sorting 857 during transportation assuming that the sands were derived from the same sediment source. If 858 this is the case, the Greenheugh capping unit would have been deposited farther from the source 859 than the Stimson formation at the Emerson and Naukluft plateaus that respectively show an 860 increase in the felsic Clusters 3 and 4 (Fig. 10). Cross-set dip directions of the Stimson formation 861

at the Gleann Beag interval of the Greenheugh pediment shares the northeast dip direction of the

cross-sets examined at the Emerson and Naukluft plateaus. However, the cross-set dip directions

of the olivine-rich Edinburgh interval are indicative of an east-west net sediment transport

direction, different to the southwest-northeast transport direction of the earlier Stimson localities and Gleann Beag interval (Banham et al., *this issue*). Cross-set dip directions in the Ladder

and Gleann Beag interval (Banham et al., *this issue*). Cross-set dip directions in the Ladder interval show dip directions to the south, which indicates a sediment transport direction to the

south. These variable dip directions show that the wind regime driving transport and the time

- of deposition of the Greenheugh pediment was also variable (Banham et al., *this issue*), making
- the possibility of detecting geochemical variation relating to physical grain segregation with
- transportation more complex.

872 It is possible that sediment had accumulated down slope of the present-day pediment and was then remobilized after the sediment transport ratio changed, resulting in the reversal of the 873 dune at the Ladder interval and the change in net sediment transport direction for the Edinburgh 874 interval. This would maintain that the sediments at the pediment were transported farther from 875 their source than those at the Emerson and Naukluft plateaus. Despite the variations in net 876 sediment transport directions across the Greenheugh pediment intervals, the ChemCam data do 877 not show a statistically significant difference in geochemistry for the major element oxides in 878 these different stratigraphic intervals of the Greenheugh pediment capping unit away from the 879 unconformity (Fig. 7). However, the grain size of the Stimson sandstone at the Greenheugh 880 pediment capping unit is approximately the same as that of the Emerson plateau (Banham et al., 881 this issue) which is not anticipated for a more distal deposit. Furthermore, the presence of 882 pyroxene and feldspar within the Stimson sandstone of the Greenheugh capping unit suggests a 883 relatively immature aeolian sandstone (Cornwall et al., 2015), supporting that the higher 884 abundance of olivine at the Greenheugh pediment likely relates to an olivine-rich local sediment 885 source, not transportation. In the ~10 km long Dyngjusandur sand sheet of Iceland, the 886 mineralogy of the aeolian sediments show strong contributions from the source rocks of the 887 surrounding volcanoes (Baratoux et al., 2011; Sara, 2017; Mountney et al., 2004). Feldspar 888 grains are diverse in this sand sheet, representing several volcanic source rocks (Sara, 2017), and 889 the incorporation of pumice into the northeast portion of the sand sheet from the 1875 eruptives 890 of the nearby Askja volcano shows how readily aeolian processes can incorporate local bedrock 891 materials (Baratoux et al., 2011; Sara, 2017; Mountney et al., 2004). As such, it is possible that 892 the ~2.5 km distance that separates the Greenheugh pediment from the Emerson and Naukluft 893 plateaus and the different wind regimes at the Greenheugh pediment were sufficient to result in a 894 change in the main sediment source at that locality. We favor the hypothesis that a more olivine-895 rich sediment source is responsible for the geochemical difference of the Greenheugh pediment 896 capping unit from that of the Emerson and Naukluft plateaus. 897

It is widely accepted that the present topographic expression of Mt Sharp formed as a 898 result of aeolian erosion after the deposition and lithification of the Mt Sharp group (Day et al., 899 2016; Grotzinger et al., 2015; Steele et al., 2018). Orbital investigations of the units farther up 900 the slopes of Mt Sharp show several olivine-bearing marker beds to the south and southeast of 901 the Greenheugh capping unit (Rudolph et al., 2019). These marker beds could provide an 902 adequate source for the olivine-rich sandstones preserved in the Greenheugh pediment capping 903 unit, with the Stimson formation at the Emerson and Naukluft plateaus likely derived more from 904 a Bradbury sandstone source (Bedford et al., 2020a). 905

4.2 The contribution of sediment recycling to the Stimson formation at the Greenheughpediment

Sedimentological evidence for sediment recycling has been identified at several locations 908 in Gale crater (Edgett et al., 2020) including the basal Stimson formation at the Marias Pass 909 locality (Newsom et al., 2018). Several angular clasts are present in the basal Stimson formation 910 911 at Marias Pass <4 mm wide, and decrease in abundance with stratigraphic height in the outcrop (Newsom et al., 2018). These angular clasts were interpreted as either a mixture of Murray 912 mudstone intraclasts and calcium sulfate vein fragments or Murray clasts coated in calcium 913 sulfate according to their largely CaO-rich geochemistry and light tonation (Newsom et al., 914 2018). Away from the unconformity, the geochemistry of the Stimson sandstone shared 915 916 similarities to that of the basaltic sandstone from the Bradbury group, suggesting that they may 917 either have been derived from a similar regional sediment source, or have been recycled from Bradbury group sandstones to the southwest of the Naukluft plateau (Bedford et al., 2020a). 918

Like the basal Stimson Missoula outcrop at Marias Pass, the basal Stimson targets of the 919 Greenheugh pediment capping unit have a relatively lighter color compared to the overlying dark 920 grey Stimson sandstone, and they contain some intraclasts (Fig. 13). However, there are some 921 important differences in the physical and chemical properties between these units. First, the 922 intraclasts are not as bright as those of the Missoula outcrop of Marias Pass, indicating that they 923 924 may not have been overprinted by calcium sulfates to the same extent (Fig. 13D). Second, the underlying Carolyn Shoemaker formation is actually SiO₂-poor relative to average (Fig. 8) 925 indicative of silica-remobilization, not precipitation like that which formed the silica-rich halos 926 of the Emerson and Naukluft plateaus. 927

928 Although ChemCam did not directly analyze a mudstone intraclast at the Greenheugh pediment, it is possible that smaller grains of the Carolyn Shoemaker formation may have been 929 recycled into the aeolian Stimson deposits (Fig. 13D; Banham et al., this issue), in addition to 930 any diagenesis that could be the cause of the lighter color of the basal Stimson targets, 931 complicating the geochemistry (Fig. 13A and 13B). To delineate the geochemical trends of 932 diagenesis from those of sediment recycling of the underlying Murray formation into the aeolian 933 deposits preserved at the Greenheugh pediment, we have generated mixing lines between the 934 Carolyn Shoemaker formation and Stimson formation. Results show that at least four 935 observation points from basal Stimson targets are situated within 1 standard deviation of the 936 trendline indicative of mixing with between 50% to 90% of Carolyn Shoemaker bedrock (Fig. 937 13A and 13B). Looking at the post-ChemCam LIBS MAHLI image, the points with a more 938 Shoemaker-like chemistry appear to have hit a greater proportion of light-toned angular grains. 939 Four basal Stimson observation points have pure Stimson chemistries, and the rest of the basal 940 Stimson observation points show geochemical signatures indicative of calcium sulfates. For the 941 observation points that largely lie along the mixing line, they do show some depletions in Na₂O 942 and K₂O relative to the trendline indicating that these points are not completely free of diagenetic 943 overprint (see supplementary material S6). H is also elevated in these unconformity targets but is 944 present at average Stimson values for the other Gleann Beag interval ChemCam targets above 945 the basal Stimson targets (Fig. 9). This indicates that Carolyn Shoemaker grains were 946 incorporated into the basal Stimson formation at the Greenheugh pediment which was later 947 subjected to diagenetic fluids that remobilized the alkali elements, and precipitated calcium 948 sulfate cement and iron oxides. 949

Mt Sharp group mudstone intraclasts preserved at the base of the Stimson formation at 950 951 Marias Pass and the Greenheugh pediment supports that sediment recycling of the immediately underlying bedrock occurred. Due to grains finer than fine sand being winnowed away in dry 952 953 aeolian environments and transported in suspension, it is possible that the lacustrine mudstone of the Mt Sharp group was rapidly abraded and transported away from the Stimson sand dunes 954 which is why geochemical signatures of the Mt Sharp group are not preserved far from the 955 unconformity. The sorting of fine-grained volcaniclastic material from basaltic sand grains in 956 aeolian bedforms is seen in fluvio-aeolian sedimentary systems in Iceland, where fine-grained 957 palagonite tuff is not preserved in the well-sorted aeolian ripples, but is present in the proximal 958 fluvial deposits (Bedford et al., 2020b), supporting the hypothesis that grain size is likely a factor 959 in the Mt Sharp group mudstone's preservation in the Stimson formation. 960

4.3 Aqueous alteration and cementation of the Stimson formation at the Greenheughpediment.

963 4.3.1 Lithification of the Stimson formation and early diagenesis

The Stimson sand sea was likely a dry aeolian deposit (Banham et al., 2021), so the first measurable instance of water-sediment interaction would have occurred after deposition of the dune sands, during the interaction of the sediments with the cementing fluid which lithified the formation in what we define as the early stage diagenesis. In this section, we will discuss whether this cement-forming diagenesis occurred at the same time under the same conditions across all localities through a comparison of the mineralogy of Thorpe et al. (*this issue*) and ChemCam geochemistry.

The secondary minerals present in the non-halo Big Sky and Okoruso drill samples, 971 drilled at the Emerson and Naukluft plateaus respectively, include iron oxides magnetite (10.3 \pm 972 0.6 wt% in Big Sky, 11.2 ± 0.7 wt% in Okoruso) and hematite (3.0 ± 0.5 wt% in Big Sky, $0.7 \pm$ 973 0.4 wt% in Okoruso), and minor (<2 wt%) abundances of calcium sulfates, quartz, and 974 fluorapatite (Morrison et al., 2018; Yen et al., 2017). Both the Emerson and Naukluft plateaus 975 also contain 15-20 wt% of an X-ray amorphous component with a moderate SiO₂ and FeO_T 976 composition for Gale crater which likely formed from the aqueous alteration of basaltic 977 sediments (Smith et al., 2021). Secondary minerals at the Greenheugh pediment capping unit 978 include magnetite (10 wt%), hematite (0.4 wt%) and phyllosilicates (7 ± 3 wt%), and 20 wt% 979 amorphous component (Rampe et al., 2020). When comparing the mineralogy of these 980 sandstones to the ChemCam data, the secondary minerals and amorphous component that derive 981 the diagenetic cement were hypothesized in Bedford et al. (2020a) to have contributed to the 982 FeO_T- and MgO-rich Clusters 1 and 2, and the TiO₂-rich Cluster 4 of that study. As the MgO 983 geochemistry of the Stimson formation at these sites is similar to that of the olivine-bearing 984 985 active Bagnold dune sands and the chemistry of the bedrock comprised of small concretions is indistinguishable from non-concretionary bedrock, it was hypothesized that the cement forming 986 diagenesis at these Stimson localities was extensive enough to convert all olivine to diagenetic 987 cement, but was relatively isochemical, preserving the geochemical trends that related to mineral 988 sorting in the ancient dune field and sediment source characteristics (Fig. 11; Bedford et al., 989 2020a). Similar to the results of Bedford et al. (2020a), our cluster analysis, which now includes 990 the Greenheugh capping unit and MnO data, shows MgO- and FeO_T-rich Clusters 1 and 2, 991 respectively. These clusters are compositionally equivalent to Clusters 1 and 2 of Bedford et al. 992 (2020a) with the exception of SiO₂ that is on-average ~4 wt% lower and Al₂O₃ which is ~2 wt% 993

lower in the new clusters, likely a result of the higher abundance of mafic minerals at the

995 Stimson formation at the Greenheugh capping unit, but not at the Emerson and Naukluft

996 plateaus. This similarity in composition of the new Stimson Clusters 1 and 2 to the old ones

997 maintains that Clusters 1 and 2 are compositionally representative of the primary igneous 998 mineralogy of the pre-lithified sands that is present across all localities, even though the

999 secondary alteration minerals and amorphous component composition varies.

1000 Hydrous alteration modelling of the Stimson formation at the Emerson and Naukluft 1001 plateaus suggests that the sandstone cement which contains the secondary minerals and amorphous component primarily formed through the dissolution of olivine under slightly acidic 1002 to near neutral pH conditions (pH 6-8) during burial diagenesis (Hausrath et al., 2018). The pH 1003 1004 was largely constrained by the presence of abundant magnetite (10.3 wt%) as a lower pH would inhibit the precipitation of authigenic magnetite (Hausrath et al., 2018). The similar abundance of 1005 1006 magnetite to the Emerson and Naukluft plateaus further suggests that olivine at the Greenheugh capping unit also underwent dissolution at a pH > 6 to precipitate magnetite, which given that 8 1007 wt% olivine is still present at the Greenheugh pediment, indicates a much higher abundance of 1008 olivine in the originally deposited aeolian sands (Rampe et al., 2020). Furthermore, as olivine is 1009 1010 one of the most reactive basaltic minerals, particularly in aqueous fluids with a pH <8 (Babechuk et al., 2014; Eggleton et al., 1987; Nesbitt et al., 1992; Pokrovsky et al., 2000), the high 1011 1012 abundance of olivine in the sandstone of the Greenheugh pediment capping suggests that this 1013 sandstone has undergone less interaction with diagenetic fluids than the Stimson formation at the Emerson and Naukluft plateaus. 1014

The presence of 7 wt% phyllosilicates at the Greenheugh pediment and none detected at 1015 the Emerson and Naukluft plateaus further suggests that the early diagenetic conditions may 1016 have been slightly different at this locality even though the clusters are compositionally 1017 equivalent. Magnetite (4.3–11.3 wt%), phyllosilicates (10–22 wt%), and olivine (1.8–4.9 wt%) 1018 have been detected in the Bradbury group basaltic fluviolacustrine deposits, the John Klein and 1019 Cumberland drilled samples of the Yellowknife Bay mudstones (Vaniman et al., 2014), and the 1020 Windjana drilled sample of the Kimberley formation (Treiman et al., 2016). Hydrous alteration 1021 modelling of the formation conditions of the phyllosilicates at Yellowknife Bay by Bridges et al. 1022 (2015) suggests that these phyllosilicates formed through an inhomogeneous dissolution of \sim 70 1023 wt% amorphous component, ~20 wt% olivine, and 10 wt% host rock in an open-system with a 1024 neutral to alkaline fluid at a pH 7.5-12. A comparison of the Stimson formation sandstone 1025 1026 geochemistry to that of the Yellowknife Bay mudstone and sandstone showed that they are both basaltic and likely derived from a similar sediment provenance (Bedford et al., 2020a), as such, 1027 the modelling results of Bridges et al. (2015) are applicable to the Stimson formation. Thereby, 1028 1029 the formation of the phyllosilicates at the Greenheugh capping unit likely occurred under more alkaline conditions (pH 7.5–12) compared to the early diagenetic conditions of the Emerson and 1030 Naukluft plateaus (pH 6–8; Hausrath et al., 2018). 1031

In the ChemCam data, H can be used to trace the abundance of hydrated phases such as phyllosilicates across sedimentary units (Thomas et al., 2020). H at the Greenheugh capping unit is notably bimodal at H peak abundances of 0.003 and 0.005 (Fig. 6, 7). The Edinburgh and Ladder intervals of the Greenheugh pediment capping derive the higher H values (0.005), supporting the presence of phyllosilicates in these stratigraphic units. Meanwhile, the lower H value at the Greenheugh capping unit consists of data solely within the Glean Beag interval and is equivalent to the peak H value at the Emerson and Naukluft plateaus (0.003), suggesting minimal phyllosilicates. This indicates that the phyllosilicates are stratigraphically confined to
the Edinburgh and Ladder intervals (Fig. 8) which could result from spatial diagenetic variability
in the Stimson formation at the Greenheugh capping unit relative to the unconformity. This
spatial variability in phyllosilicate abundances with the unconformity is also present in the
underlying Carolyn Shoemaker formation, which has less phyllosilicates in the Hutton drilled
sample compared to other drill holes away from the unconformity (Thorpe et al., *this issue*).

1045 Evidence to support spatial diagenetic variability in the Stimson formation is present with the distribution of the concretions. On average, concretions are more abundant and greater in size 1046 with proximity to the unconformity for the Stimson formation at the Greenheugh pediment (Fig. 1047 5), and elsewhere (Banham et al., 2018, this issue). This relationship of the concretions with the 1048 1049 unconformity suggests that their formation may relate to the presence of a perched aquifer at the base of the Stimson formation, or that it could indicate a greater abundance of minerals that acted 1050 1051 as a nucleus for concretion formation (Banham et al., 2018; Chan et al., 2012; Potter-McIntyre et 1052 al., 2014). With the exception of the basal Stimson formation that includes mixing of mudstone intraclasts and alteration by a later diagenetic fluid, the Stimson formation is relatively 1053 geochemically uniform (Bedford et al., 2020a; Siebach et al., 2017; Thompson et al., this issue), 1054 1055 with large-scale variations only relating to late-stage diagenetic features (Bedford et al., 2020a; Yen et al., 2017). As such, a greater concentration of minerals at the unconformity for which 1056 1057 these concretions could have nucleated seems unlikely. However, the Mt Sharp group units that 1058 underly the Stimson formation are well-cemented, phyllosilicate-bearing mudstones which makes them more likely to act as an aquiclude and promote the formation of a perched aquifer 1059 which in turn promoted the formation of the concretions. If the perched aquifer is the cause of 1060 1061 the concretion formation, this would in turn support that the diagenetic fluids responsible for cementing the Stimson formation were derived from near-surface waters that accumulated 1062 towards the unconformity. The Greenheugh pediment is fan shaped and extends from Gediz 1063 Vallis whose morphology suggests that it may have formed from fluvial processes (Anderson et 1064 al., 2010; Thomson et al., 2011; Wiens et al., 2020), providing a potential source for the surficial 1065 water that lithified the Stimson formation at the Greenheugh pediment and potentially further 1066 1067 downslope.

1068 With the addition of MnO data to the Stimson formation and, therefore, the cluster analysis, there is now a MnO-rich Cluster 5 that consists of high-MnO observation points (mean 1069 of 2.9 wt%) from three host rock ChemCam targets from the Edinburgh interval of the 1070 1071 Greenheugh pediment capping unit (Fig. 13B). Additional MnO-rich host rock targets in the Edinburgh interval are also present for observation points that have ChemCam totals <90 wt% 1072 that are excluded from the cluster analysis. The ChemCam observation points with high-MnO 1073 1074 concentrations targeted fine grained, dark-toned sandstone that is not associated with any obvious feature and are sporadically distributed spatially across the sandstone. MnO 1075 concentrations are >1 wt% for these data points and there are no correlations with other major 1076 element oxides. Similar observations of sporadic enrichments in MnO were detected in the 1077 Glenelg and Gillespie Lake sandstones of the aforementioned Yellowknife Bay formation (Lanza 1078 1079 et al., 2014); note that these Mn abundances were later recalibrated (Gasda et al., 2021). The cause of these high MnO-points were hypothesized to be either detrital, from a sediment source 1080 that experienced oxidizing aqueous conditions, or diagenetic, either forming during rapidly 1081 oxidizing, but localized conditions during lithification of the sandstones, or from later diagenetic 1082 fluids moving through the sandstones (Lanza et al., 2014). As the Greenheugh capping unit has 1083

1084 minimal evidence for late diagenesis, unlike the older sedimentary units of Yellowknife Bay, it is 1085 most likely to be detrital or formed during cementation. Mn is a common trace element in 1086 olivine, substituting for Fe^{2+} . As such, the dissolution of olivine and formation of magnetite 1087 under isochemical conditions may have precipitated a late-stage Mn-bearing mineral or 1088 amorphous component if aqueous fluids were minimal and the system was closed at the scale of 1089 the stratigraphic interval.

1090 4.3.2. The formation of alteration features across the Stimson formation

Based on the Mastcam and Navcam images of the Stimson formation at the Greenheugh 1091 capping unit (Fig. 3, 4), in addition to the plots of data classifications across all Stimson 1092 localities (Fig. 5), it is evident that the Stimson formation at the Greenheugh pediment has not 1093 experienced the same later diagenetic events as those at the Emerson and Naukluft plateaus. The 1094 Greenheugh pediment lacks silica-rich halos, raised ridges, and has very few occurrences of 1095 1096 calcium sulfate mineral veins which are much thinner than those sampled and studied at the Emerson and Naukluft plateaus (Fig. 4). Erosionally resistant polygons - polygonal diagenetic 1097 features associated with preferential cementation around fractures as a result of groundwater 1098 1099 recharge - are also present to a lesser degree at the Greenheugh capping unit compared to the other Stimson localities and with a different morphology (Kronyak et al., 2019b). 1100

The lack of fractures and their associated diagenetic features, in addition to the different 1101 morphology of erosion resistant polygonal ridges at the Greenheugh pediment capping unit 1102 1103 suggests different potential scenarios: 1) that at this location, groundwater flow was not as pervasive (Kronyak et al., 2019b), 2) that the Stimson formation at the Greenheugh capping unit 1104 1105 experienced less compaction compared to that at the Emerson and Naukluft plateaus, or 3) that the Greenheugh pediment exhibits different material properties such as rock hardness and grain 1106 size. The Greenheugh pediment is situated at an elevation of -4150 m, which is 200 m higher in 1107 elevation than the Stimson formation at the Emerson and Naukluft plateaus (Fig. 1,2). As water 1108 1109 is most likely to collect at topographical lows, it is possible that there was less groundwater available to the Greenheugh pediment to either create the well-formed, large polygons of the 1110 1111 Emerson and Naukluft plateaus, or to generate sufficient fluid overpressure and hydrofracturing, providing support to scenario 1. 1112

Another factor which may have impacted the amount of groundwater available to the 1113 1114 Greenheugh pediment capping unit is the time in which it was deposited and lithified. Although the Stimson formation at the pediment shares the common sedimentary architecture of the 1115 1116 Stimson formation investigated earlier in the mission, and all outcrops are situated above the basal Siccar Point group unconformity, it is not clear whether these outcrops were lithified at the 1117 same time. While the sedimentary architecture indicates sediment accumulation under arid 1118 conditions (hence the absence of damp interdunes), water would have percolated through the 1119 sediments post-deposition to cement the strata. This could have happened at depth where 1120 subsurface water could not interact with the depositional surface (sensu Kocurek and Haveholm. 1121 1122 1993) or could have happened after the terminal phase of the dune field's evolution. The sources of cementing fluid are likely to be different between the earlier Stimson localities and the 1123 1124 Greenheugh pediment based on the different secondary minerals. The fan-shaped morphology of 1125 the Greenheugh pediment itself at the base of Gediz Vallis suggests that it may relate to surface waters flowing down and eroding the slope of Mt Sharp (Bryk et al., 2019). Meanwhile, the 1126 Stimson formation at the Emerson and Naukluft plateaus may have been cemented by rising 1127

groundwaters (Banham et al., 2018). Hence, the Greenheugh pediment capping unit may have

been cemented at a different time, potentially one in which there was less groundwateravailability in Gale crater.

Alternatively, the difference in elevation could have resulted in less overburden to 1131 compact the sediments at the Greenheugh pediment capping unit, supporting scenario 2, 1132 however, it is not known to what extent the Stimson formation was buried at either location. 1133 1134 Several sedimentary float are present on top of and close the Stimson formation at the Emerson and Naukluft plateaus (Wiens et al., 2020). Some of these float potentially include clasts of 1135 Stimson bedrock and their location on top of the Stimson formation itself suggests that additional 1136 sedimentary units did at some time overlie the Stimson formation at these localities in the past 1137 1138 (Wiens et al., 2020). But as these deposits exist as float and not as a coherent outcrop, we cannot constrain with any certainty how much burial, and therefore compaction, occurred for the 1139 1140 Stimson formation at the Emerson and Naukluft plateaus relative to the Greenheugh capping 1141 unit.

1142 4.3.3. Diagenesis at the unconformity

Changes in geochemistry and mineralogy in the underlying Mt Sharp group and 1143 immediately overlying Stimson formation have been detected along the basal Siccar Point group 1144 unconformity indicating that this geological boundary may have provided a conduit for later 1145 diagenetic fluids (Achilles et al., 2020; Horgan et al., 2020; Newsom et al., 2018; Rudolph et al., 1146 1147 this issue; Thomson et al., this issue; Yen et al., 2021). At the Greenheugh pediment, Stimson and Carolyn Shoemaker bedrock close to the unconformity exhibit differences in geochemistry, 1148 and for the Carolyn Shoemaker formation mineralogy (Dehouck et al., this issue; Thompson et 1149 al., this issue; Thorpe et al., this issue). The basal Stimson formation is lower in MgO, Na₂O, and 1150 MnO relative to the mean composition of the Stimson formation, low in total sum of oxides, but 1151 high in CaO and H relative to the average Stimson formation (Fig. 7). The high CaO, H, and low 1152 1153 totals are indicative of calcium sulfate addition, which is also supported by the mixing plots (Fig. 13, S.6), and the high abundance of sulfates detected by APXS (Thompson et al., this issue). As 1154 1155 this unique Stimson geochemistry is only observed for targets that are approximately 10 cm and 50 cm from the unconformity surface, it suggests that this diagenesis did not permeate far into 1156 1157 the overlying Stimson formation, indicating that it was lithified prior to the alteration.

Calcium sulfate cements were also identified in the basal Stimson outcrop of the Marias 1158 Pass locality according to a lighter coloration of the bedrock and geochemical compositions that 1159 1160 are high in CaO and low for all other major element oxides and totals (Newsom et al., 2018). At the Naukluft plateau, rover images also showed evidence of either silica-halo alteration or 1161 calcium sulfates concentrated at the Murray formation immediately below the unconformity (Fig. 1162 14). This indicates that the diagenesis that occurred at the unconformity between the Greenheugh 1163 pediment capping unit was widespread and also occurred along the unconformity at other 1164 Stimson localities. 1165

A possible scenario that could result in the precipitation of calcium sulfates and a decrease in MgO and alkali abundances includes acid-sulfate alteration, such as that which is hypothesized for the Murray formation of Marias Pass, the Oudam drilled sample in the Murray formation at the Naukluft plateau (Achilles et al., 2020), the alteration halos of the Emerson and Naukluft Plateaus (Hausrath et al., 2018; Yen et al., 2017, 2021), and in some places on Vera Rubin ridge (Rampe et al., 2020b). Acid-sulfate alteration can preferentially alter phenocrysts
such as olivine, pyroxene, and feldspars over volcanic glass, leaving behind a silica residue, iron
oxides, and sulfates and remobilizing MgO and the alkali elements (McCollom, et al., 2013a;
McCollom et al., 2013b; Tosca et al., 2004). However, unlike the potentially acid-sulfate altered

- 1175 mudstones and sandstones of Marias Pass, the basal Stimson formation at the Greenheugh
- 1176 capping unit does not have an enrichment in SiO_2 and TiO_2 (Fig. 7), indicating that any acid-
- sulfate alteration at this locality may have been less extensive than that which modified the
- 1178 geochemistry and mineralogy of the rocks at the Emerson and Naukluft plateaus.

1179 vThe alteration of Stimson sandstones at the unconformity suggests that either a weakly oxic, sulfate-bearing fluid or fewer resurgences of late diagenetic sulfate-bearing groundwater 1180 1181 fluids may have migrated along the unconformity, permeating the lithified sandstone directly above and the mudstone below. These fluids may have not been able to permeate as extensively 1182 1183 into the upper Stimson formation at the Greenheugh capping unit and create polygonal fractures due to the greater elevation resulting in less groundwater availability. The difference in 1184 sandstone cement between the Emerson and Naukluft plateaus and the Greenheugh pediment 1185 capping unit may have also been a factor in impeding the formation of polygons and decreasing 1186 1187 the permeability of the Stimson sandstone at the Greenheugh capping unit. The Big Sky and Okoruso drilled samples of the Emerson and Naukluft plateaus both contain iron oxides 1188 (magnetite and hematite), some calcium sulfates (bassanite), and an amorphous component 1189 1190 within the sandstone cement (Yen et al., 2017). The Edinburgh drilled sample from the Edinburgh interval of the Greenheugh pediment capping unit contains phyllosilicates in addition 1191 to magnetite and an amorphous component (Rampe et al., 2020). Phyllosilicates decrease the 1192 1193 permeability of a sandstone, which may have further impacted fluid migration at this locality and the material properties of the Stimson formation at the Edinburgh interval compared to that at the 1194 1195 Emerson and Naukluft plateaus.

The Carolyn Shoemaker formation that underlies the unconformity at the Greenheugh 1196 1197 pediment shows geochemical differences relative to the average Carolyn Shoemaker formation that are indicative of alteration (Fig. 7, 13; Dehouck et al., this issue; Gasda et al., this issue). 1198 Furthermore, the Hutton drill location in the Carolyn Shoemaker formation contains 8.7 wt% 1199 iron oxides, 4.7 wt% cristobalite and 5.2 wt% Opal-CT, silica polymorphs that can form as a 1200 result of the dissolution of mafic minerals (Thorpe et al., this issue). The immediately underlying 1201 Carolyn Shoemaker bedrock around the Hutton drill site is lower in oxide totals (94.9 wt% 1202 1203 compared to 97.3 wt%), and higher in CaO compared to average Carolyn Shoemaker bedrock (Fig. 7). This can be indicative of calcium sulfate cement which has been largely detected by the 1204 ChemCam instrument and correlated to light-toned cements in MAHLI images acquired in the 1205 traverse leading up to the Greenheugh pediment (Gasda et al., this issue; Nellesson et al., 2021). 1206 However, the mineralogy of the Hutton drilled sample acquired below the capping unit only 1207 contains 0.6 wt% anhydrite (Thorpe et al., *this issue*), which is some of the lowest abundances of 1208 crystalline calcium sulfate detected in the Carolyn Shoemaker formation since the drilled 1209 samples of the Pahrump Hills (Rampe et al., 2017), and sulfur is not detected here with APXS 1210 (Thompson et al., this issue). F abundances also increase in the Carolyn Shoemaker formation 1211 below the unconformity at the Hutton drill site (Dehouck et al., this issue; Fig. 8B), including for 1212 alteration features such as veins and raised ridges (Gasda et al., this issue). This correlated with 1213 the increase in CaO concentrations suggesting the presence of fluorapatite which is confirmed in 1214 the CheMin results of the Hutton drilled sample (Thorpe et al., *this issue*). 1215

The current hypothesis for the distinctive geochemistry and mineralogy of the Hutton 1216 1217 drill hole and surrounding sediments is that warm or acidic diagenetic fluids circulated here and concentrated below the capping unit (Dehouck et al., this issue; Gasda et al., this issue; Forni et 1218 1219 al., 2021; 2016). However, these F-rich diagenetic fluids do not appear to have permeated into the overlying Stimson formation. Fluorine is a fluid-mobile element which is in high abundance 1220 in the Murray formation bedrock and diagenetic features below the Stimson formation at the 1221 Marias Pass locality and in the Carolyn Shoemaker formation's Hutton Zone of Glen Torridon 1222 1223 (Forni et al., 2021; 2016). However, above the unconformity, fluorine detections are sparse in the Stimson formation, with only one detection in the pediment capping unit (Fig. 8B; Forni et al., 1224 2021). Meanwhile, the rare detections of fluorine in the Stimson sandstone instead suggests that 1225 these may instead relate to detrital phases, similar to what is hypothesized for the fluorine phases 1226 detected in the Bradbury group sandstones (Forni et al., 2013). The lack of fluorine detected in 1227 the basal Stimson that also contains intraclasts of the underlying mudstones further supports that 1228 the alteration of the Carolyn Shoemaker formation around the Hutton drill hole occurred after the 1229 formation of the overlying Stimson formation. 1230

1231 **5** Conclusions

The ChemCam geochemical data of the Greenheugh pediment capping unit suggests that 1232 this unit is a part of the Stimson formation investigated at the Emerson plateau, Naukluft plateau, 1233 1234 and Murray Buttes localities. However, minor but statistically significant geochemical variations exist for SiO₂, MgO, MnO, Al₂O₃, and Na₂O. These slight geochemical differences are indicative 1235 of higher abundances of mafic minerals and their diagenetic products at the Stimson formation of 1236 1237 the Greenheugh capping unit, which is supported by mineralogical data from CheMin. The higher abundance of mafic minerals preserved at the Greenheugh pediment capping unit is 1238 unlikely to relate to physical grain segregation with transportation across all Stimson localities 1239 1240 due to variable cross-set dip directions indicative of seasonal wind variations which complicate the geochemical record of net sediment transport at this locality. Similarities in grain size and the 1241 1242 high abundances of pyroxene and plagioclase feldspar in addition to olivine is also indicative that this aeolian sandstone has not undergone significant transport from its source (Cornwall et al., 1243 1244 2015). Instead, we hypothesize that the sediments of the Stimson formation at the Greenheugh pediment capping unit are predominately sourced from a nearby olivine-rich unit which may 1245 1246 relate to the olivine-bearing marker beds that have been detected close to the pediment, farther up the slopes of Mt Sharp (Rudolph et al., 2019). 1247

While the Stimson aeolian sand dunes migrated in the Hesperian, they were also able to 1248 erode and incorporate clasts from the underlying Mt Sharp group into the base of the sand dunes, 1249 1250 evidenced by the ChemCam geochemical data and images of the basal Stimson unit at the Greenheugh pediment. However, due to the fine-grained nature of the Mt Sharp group 1251 mudstones, it is unlikely that mudstones were preserved for long in the Stimson sand dunes and 1252 1253 were instead winnowed away into suspension during dune migration. Alternatively, the intraclasts within the basal Stimson formation may relate to the onset of aeolian deposition as a 1254 buried lag deposit, similar to what is hypothesized for the Murray intraclasts preserved in the 1255 Stimson formation at Marias Pass (Newsom et al., 2018). Either scenario provides additional 1256 evidence for sediment recycling in Gale crater by aeolian processes. 1257

1258 After the deposition of the Stimson sand dunes, their sediments interacted with surface 1259 waters forming a sandstone cement through the dissolution of the most reactive components, 1260 likely volcanic glass and olivine. These fluids formed a perched acquifer at the base of the

- aeolian sandstone at all localities, forming the concretionary bedrock. Geochemically, the
- 1262 sandstone cement of the Stimson formation is difficult to distinguish from that of the Emerson
- and Naukluft plateaus with the exception of MnO, however, differences in secondary mineralogy and the persistence of olivine in the Greenheugh pediment capping unit suggest that the early
- 1264 and the persistence of olivine in the Greenheugh pediment capping unit suggest that the early 1265 diagenetic fluid was more alkaline than that which cemented the Stimson formation at other
- 1265 localities. This difference in early diagenetic fluid chemistry could relate to the Greenheugh
- pediment capping unit having a different fluid source, and/or being cemented at a different time
- 1268 in the geological record of the Stimson sand sea.

Once the Stimson formation was lithified, several diagenetic events likely relating to 1269 1270 upwelling groundwater occurred to form erosion resistant polygons, silica-rich halos, calcium sulfate veins, and alteration along the unconformity. These later diagenetic events were more 1271 1272 prevalent in the Stimson formation preserved at the Emerson and Naukluft plateaus, suggesting a greater availability of groundwater fluids at these sites than the Greenheugh pediment. This 1273 greater availability of groundwater at the Emerson and Naukluft plateaus could relate to their 1274 lower elevation in Gale crater, greater burial, or a difference in age between these Stimson 1275 1276 deposits (i.e., Emerson and Naukluft Stimson are older than the Greenheugh capping unit and were cemented at a time when there was more groundwater present in Gale crater). One of the 1277 last episodes of diagenesis to occur likely relates to the unconformity, which is shared across all 1278 1279 localities. The diagenesis below the unconformity involved fluorine-rich, potentially warm fluids that did not permeate into the overlying Stimson formation. The diagenesis along the 1280 unconformity suggests that acidic, sulfate-bearing fluids migrated along the unconformity and 1281 altered the basal Stimson formation, supporting the hypotheses of previous studies. 1282

Overall, the ancient aeolian environment under which these Stimson sandstones were deposited was dynamic, recycling nearby material as the dunes migrated across the surface. These results support that liquid water was stable at the surface in the Hesperian and was available as groundwater for multiple diagenetic events after deposition along bedrock weaknesses and the basal Stimson formation in Gale crater.

1288

1289 Acknowledgments, Samples, and Data

1290 CCB acknowledges support from the LPI. JCB, SPS, and SMRT acknowledge funding from
1291 UKSA. JF acknowledges the support from the Carlsberg Foundation. The MSL engineering,

1292 ChemCam, CheMin, and science teams are gratefully acknowledged for the acquisition of the

data used in this paper. Support for this work in the US is provided by the NASA Mars

- 1294 Exploration Program and in France by CNES. Mastcam mosaics were processed by the Mastcam
- 1295 team at Malin Space Science Systems.
- 1296 The ChemCam major element oxide data and RMI data presented in this paper are archived in

1297 the Planetary Data System (PDS). The geochemical data can be found at Wiens (2013a). RMI

1298 mosaics can be found at Wiens (2013b). The derived data which includes the ChemCam targets

used in the study, target classifications, and clustering results, can be found at Bedford (2022).

1300

1301 References

- Achilles, C. N., Downs, R. T., Ming, D. W., Rampe, E. B., Morris, R. V., Treiman, A. H., et al.
 (2017). Mineralogy of an active eolian sediment from the Namib dune, Gale crater, Mars. *Journal of Geophysical Research: Planets*, *122*(11), 2344–2361.
 https://doi.org/10.1002/2017JE005262
- Achilles, C. N., Rampe, E. B., Downs, R. T., Bristow, T. F., Ming, D. W., Morris, R. V., et al.
 (2020). Evidence for Multiple Diagenetic Episodes in Ancient Fluvial-Lacustrine
 Sedimentary Rocks in Gale Crater, Mars. *Journal of Geophysical Research: Planets*, *125*(8). https://doi.org/10.1029/2019JE006295
- Anderson, R. B., & Bell III, J. F. (2010). Geologic mapping and characterization of Gale crater
 and implications for its potential as a Mars Science Laboratory landing site. *The Mars Journal*, 5, 76–128. https://doi.org/10.1555/mars.2010.0004
- Anderson, R. B., Bridges, J. C., Williams, A., Edgar, L., Ollila, A., Williams, J., et al. (2015).
 ChemCam results from the Shaler outcrop in Gale crater, Mars. *Icarus*, 249, 2–21.
 https://doi.org/10.1016/j.icarus.2014.07.025
- Anderson, R. B., Clegg, S. M., Frydenvang, J., Wiens, R. C., McLennan, S., Morris, R. V., et al.
 (2017). Improved accuracy in quantitative laser-induced breakdown spectroscopy using
 sub-models. *Spectrochimica Acta Part B: Atomic Spectroscopy*, *129*, 49–57.
 https://doi.org/10.1016/J.SAB.2016.12.002
- Anderson, R. B., Edgar, L. A., Rubin, D. M., Lewis, K. W., & Newman, C. (2018). Complex
 bedding geometry in the upper portion of Aeolis Mons, Gale crater, Mars. *Icarus*, *314*, 246–
 264. https://doi.org/10.1016/j.icarus.2018.06.009
- Babechuk, M. G., Widdowson, M., & Kamber, B. S. (2014). Quantifying chemical weathering
 intensity and trace element release from two contrasting basalt profiles, Deccan Traps,
 India. *Chemical Geology*, *363*, 56–75.
- 1326 https://doi.org/https://doi.org/10.1016/j.chemgeo.2013.10.027
- Banham, S. G., Gupta, S., Rubin, D. M., Edgett, K. S., Bryk, A. B., Dietrich, W. E., et al. (this
 issue). Reconstructing ancient sedimentary environments of the Greenheugh pediment
 capping unit (Stimson formation) and its implications for Gale crater evolution. *JGR: Planets*.
- Banham, S. G., Gupta, S., Rubin, D. M., Watkins, J. A., Sumner, D. Y., Edgett, K. S., et al.
 (2018). Ancient Martian aeolian processes and palaeomorphology reconstructed from the
 Stimson formation on the lower slope of Aeolis Mons, Gale crater, Mars. *Sedimentology*,
 65(4), 993–1042. https://doi.org/10.1111/sed.12469
- Banham, S. G., Gupta, S., Rubin, D. M., Edgett, K. S., Barnes, R., Van Beek, J., et al. (2021). A
 Rock Record of Complex Aeolian Bedforms in a Hesperian Desert Landscape: the Stimson
 formation as exposed in the Murray buttes, Gale crater, Mars. *Journal of Geophysical Research: Planets*, *126*(4), e2020JE006554. https://doi.org/10.1029/2020je006554
- Baratoux, D., Mangold, N., Arnalds, O., Bardintzeff, J.-M., Platevoët, B., Grégoire, M., & Pinet,
 P. (2011). Volcanic sands of Iceland Diverse origins of aeolian sand deposits revealed at
 Dyngjusandur and Lambahraun. *Earth Surface Processes and Landforms*, *36*(13), 1789–
 1808. https://doi.org/10.1002/esp.2201

- Bedford, C. C. (2019, July 28). Distinguishing The Geochemical Effects Of Sedimentary
 Processes And Source Region Characteristics In Gale Crater, Mars.
- 1345 Bedford, C. C., Bridges, J. C., Schwenzer, S. P., Wiens, R. C., Rampe, E. B., Frydenvang, J., &
- Gasda, P. J. (2019). Alteration trends and geochemical source region characteristics
 preserved in the fluviolacustrine sedimentary record of Gale crater, Mars. *Geochimica et*
- 1348 *Cosmochimica Acta*, 246. https://doi.org/10.1016/j.gca.2018.11.031
- Bedford, C. C., Schwenzer, S. P., Bridges, J. C., Banham, S. G., Wiens, R. C., Frydenvang, J., et
 al. (2019). Using ChemCam derived geochemistry to identify the paleonet sediment
 transport direction and source region characteristics of the Stimson formation in Gale crater,
 Mars. Retrieved from
- 1353 https%3A%2F%2Fwww.hou.usra.edu%2Fmeetings%2Flpsc2019%2F
- Bedford, C. C., Schwenzer, S. P., Bridges, J. C., Banham, S., Wiens, R. C., Gasnault, O., et al.
 (2020a). Geochemical variation in the Stimson formation of Gale crater: Provenance,
 mineral sorting, and a comparison with modern Martian dunes. *Icarus*, *341*, 113622.
- 1357 https://doi.org/10.1016/j.icarus.2020.113622
- Bedford, C. C., Rampe, E. B., Thorpe, M., Ewing, R. C., Horgan, B., Nachon, M., et al. (2020b).
 Identifying the products of volcano-ice interaction in basaltic sediments in Iceland and their
 implications for Mars. In *51st Lunar and Planetary Science Conference* (p. 2478). Houston.
- Bedford, C. (2022). Greenheugh capping unit ChemCam derived data products [Data set].
 Zenodo. https://doi.org/10.5281/ZENODO.5889318
- Bennett, K. A., & Bell, J. F. (2016). A global survey of martian central mounds: Central mounds
 as remnants of previously more extensive large-scale sedimentary deposits. *Icarus*, 264,
 331–341. https://doi.org/10.1016/j.icarus.2015.09.041
- Bennett, K. A., Fox, V. Bryk, A. Dietrich, W. Fedo, C. Dehouck, E. Cousin, A. Thorpe, M.
 Rudolph, A. Horgan, B. Gasda, P. Bristow, T. Sutter, B. McAdam, A. O'Connell-Cooper, C.
 Thompson, L. Fraeman, A. Vasavada, A. Wong, G. Banham, S. Arvidson, R. Stack, K.
 Frydenvang, J. Edgar, L. Treiman, A. Eigenbrode, J. Grotzinger, J. (this issue). An
 overview of the Curiosity rover's Campaign in Glan Torridon, Gale crater, Mars. *JGR: Planets*.
- Bish, D. L., Blake, D. F., Vaniman, D. T., Chipera, S. J., Morris, R. V., Ming, D. W., et al.
 (2013). X-ray Diffraction Results from Mars Science Laboratory: Mineralogy of Rocknest at Gale Crater. *Science*, *341*(6153), 1238932–1238932.
- 1375 https://doi.org/10.1126/science.1238932
- Blake, D., Vaniman, D., Achilles, C., Anderson, R., Bish, D., Bristow, T., et al. (2012).
 Characterization and Calibration of the CheMin Mineralogical Instrument on Mars Science
 Laboratory. *Space Science Reviews*, *170*(1–4), 341–399. https://doi.org/10.1007/s11214012-9905-1
- Blake, D. F., Morris, R. V., Kocurek, G., Morrison, S. M., Downs, R. T., Bish, D., et al. (2013).
 Curiosity at Gale Crater, Mars: Characterization and Analysis of the Rocknest Sand
 Shadow. *Science*, *341*(6153), 1239505–1239505. https://doi.org/10.1126/science.1239505
- Blaney, D. L., Wiens, R. C., Maurice, S., Clegg, S. M., Anderson, R. B., Kah, L. C., et al.
 (2014). Chemistry and texture of the rocks at Rocknest, Gale Crater: Evidence for

- sedimentary origin and diagenetic alteration. *Journal of Geophysical Research: Planets*,
 119(9), 2109–2131. https://doi.org/10.1002/2013JE004590
- Bridges, J. C., Schwenzer, S. P., Leveille, R., Westall, F., Wiens, R. C., Mangold, N., et al.
 (2015). Diagenesis and clay mineral formation at Gale Crater, Mars. *Journal of Geophysical Research: Planets*, *120*(1), 1–19. https://doi.org/10.1002/2014JE004757
- Bridges, N. T., & Ehlmann, B. L. (2018). The Mars Science Laboratory (MSL) Bagnold Dunes
 Campaign, Phase I: Overview and introduction to the special issue. *Journal of Geophysical Research: Planets*, 123(1), 3–19. https://doi.org/10.1002/2017JE005401
- Bridges, N. T., Geissler, P. E., McEwen, A. S., Thomson, B. J., Chuang, F. C., Herkenhoff, K.
 E., et al. (2007). Windy Mars: A dynamic planet as seen by the HiRISE camera. *Geophysical Research Letters*, 34(23).
 https://doi.org/https://doi.org/10.1029/2007GL 031445
- 1396 https://doi.org/https://doi.org/10.1029/2007GL031445
- Bridges, N. T., Sullivan, R., Newman, C. E., Navarro, S., van Beek, J., Ewing, R. C., et al.
 (2017). Martian aeolian activity at the Bagnold Dunes, Gale Crater: The view from the
 surface and orbit. *Journal of Geophysical Research: Planets*, *122*(10), 2077–2110.
 https://doi.org/10.1002/2017JE005263
- Bristow, T. F., Rampe, E. B., Achilles, C. N., Blake, D. F., Chipera, S. J., Craig, P., et al. (2018).
 Clay mineral diversity and abundance in sedimentary rocks of Gale crater, Mars. *Science Advances*, 4(6), eaar3330. https://doi.org/10.1126/sciadv.aar3330
- Bryk, A. A., Dientrich, W. E., Lamb, M. P., Grotzinger, J. P., Vasavada, A. R., Stack, K. M.,
 Arvidson, A. (2019). What was the original extent of the Greenheugh pediment and Gediz
 Vallis ridge deposits in Gale crater, Mars. In *Ninth International Conference on Mars*.
 vol:2089, p.6296.
- 1408 Carr, M. H., & Head, J. W. (2010). Geologic history of Mars. *Earth and Planetary Science* 1409 *Letters*, 294(3–4), 185–203. https://doi.org/10.1016/J.EPSL.2009.06.042
- Chan, M., Potter, S., Peterson, E., Parry, W., Bowman, J., Grotzinger, J., & Milliken, R. (2012).
 Characteristics of terrestrial ferric oxide concretions and implications for Mars. *Sedimentary Geology of Mars*, *102*, 253–270. Retrieved from
- 1413 https://pdfs.semanticscholar.org/b238/8ccb9b302d832316a3f17cb561868cd9391c.pdf
- Clegg, S. M., Wiens, R. C., Anderson, R., Forni, O., Frydenvang, J., Lasue, J., et al. (2017).
 Recalibration of the Mars Science Laboratory ChemCam instrument with an expanded
 geochemical database. *Spectrochimica Acta Part B: Atomic Spectroscopy*, *129*, 64–85.
 https://doi.org/10.1016/J.SAB.2016.12.003
- Cornwall, C., Bandfield, J. L., Titus, T. N., Schreiber, B. C., & Montgomery, D. R. (2015).
 Physical abrasion of mafic minerals and basalt grains: Application to martian aeolian deposits. *Icarus*, *256*, 13–21. https://doi.org/10.1016/j.icarus.2015.04.020
- Cousin, A., Meslin, P. Y., Wiens, R. C., Rapin, W., Mangold, N., Fabre, C., et al. (2015).
 Compositions of coarse and fine particles in martian soils at gale: A window into the
 production of soils. *Icarus*, *249*, 22–42. https://doi.org/10.1016/J.ICARUS.2014.04.052
- Cousin, A., Dehouck, E., Meslin, P. -Y., Forni, O., Williams, A. J., Stein, N., et al. (2017).
 Geochemistry of the Bagnold dune field as observed by ChemCam and comparison with

- 1426other aeolian deposits at Gale Crater. Journal of Geophysical Research: Planets, 122(10),14272144–2162. https://doi.org/10.1002/2017JE005261
- Czarnecki, S., Hardgrove, C., Gasda, P. J., Gabriel, T. S. J., Starr, M., Rice, M. S., et al. (2020).
 Identification and Description of a Silicic Volcaniclastic Layer in Gale Crater, Mars, Using
 Active Neutron Interrogation. *Journal of Geophysical Research: Planets*, *125*(3), 2019–
 006180. https://doi.org/10.1029/2019JE006180
- Day, M., Anderson, W., Kocurek, G., & Mohrig, D. (2016). Carving intracrater layered deposits
 with wind on Mars. *Geophysical Research Letters*, 43(6), 2473–2479.
 https://doi.org/10.1002/2016GL068011@10.1002/(ISSN)1944-8007.2016GRLEDHIGH
- Day, Mackenzie, & Kocurek, G. (2016). Observations of an aeolian landscape: From surface to
 orbit in Gale Crater. *Icarus*, 280, 37–71. https://doi.org/10.1016/j.icarus.2015.09.042
- 1437 Dehouck, E., Cousin, A., Mangold, N., Frydenvang, J., Gasnault, O., Forni, O., Rapin, W.,
- 1438 Gasda, P. J., Bedford, C. C., Caravaca, G., David, G., Lasue, J., Meslin, P.-Y.,
- 1439 Rammelkamp, K., Le Mouelic, S., Fox, V. K., Bennett, K. A., Bryk, A. A., Lanza, N. L.,
- Maurice, S., Wiens, R. (this issue). In situ geochemical characterization of the clay-bearing Glen Torridon region of Gale crater, Mars, using the ChemCam instrument. *JGR: Planets*.
- Edgar, L. A., Fedo, C. M., Gupta, S., Banham, S., Fraeman, A. A., Grotzinger, J., et al. (2020). A
 lacustrine paleoenvironment recorded at Vera Rubin ridge, Gale crater: Overview of the
 sedimentology and stratigraphy observed by the Mars Science Laboratory Curiosity rover. *Earth and Space Science Open Archive*. https://doi.org/10.1002/essoar.10501350.1
- Edgett, K. S., Yingst, R. A., Ravine, M. A., Caplinger, M. A., Maki, J. N., Ghaemi, F. T., et al.
 (2012). Curiosity's Mars Hand Lens Imager (MAHLI) Investigation. *Space Science Reviews*, *170*(1–4), 259–317. https://doi.org/10.1007/s11214-012-9910-4
- Edgett, K. S., Banham, S. G., Bennett, K. A., Edgar, L. A., Edwards, C. S., Fairén, A. G., et al.
 (2020). Extraformational sediment recycling on Mars. *Geosphere*, *16*(6), 1508–1537.
 https://doi.org/10.1130/GES02244.1
- Eggleton, R. A., Foudoulis, C., & Varkevisser, D. (1987). Weathering of Basalt: Changes in
 Rock Chemistry and Mineralogy. *Clays and Clay Minerals*, *35*(3), 161–169.
 https://doi.org/10.1346/CCMN.1987.0350301
- Ehlmann, B. L., Edgett, K. S., Sutter, B., Achilles, C. N., Litvak, M. L., Lapotre, M. G. A., et al.
 (2017). Chemistry, mineralogy, and grain properties at Namib and High dunes, Bagnold
 dune field, Gale crater, Mars: A synthesis of Curiosity rover observations. *Journal of*
- 1458 *Geophysical Research: Planets*, *122*(12), 2510–2543.
- 1459 https://doi.org/10.1002/2017JE005267
- Eilers, P. H. C., & Goeman, J. J. (2004). Enhancing scatterplots with smoothed densities. *Bioinformatics (Oxford, England)*, 20(5), 623–8.
 https://doi.org/10.1093/bioinformatics/btg454
- 1463 Ewing, R. C., Lapotre, M. G. A., Lewis, K. W., Day, M., Stein, N., Rubin, D. M., et al. (2017).
- Sedimentary processes of the Bagnold Dunes: Implications for the eolian rock record of
 Mars. *Journal of Geophysical Research: Planets*, *122*(12), 2544–2573.
- 1466 https://doi.org/10.1002/2017JE005324

- Fedo, C. M., McGlynn, I. O., & McSween, H. Y. (2015). Grain size and hydrodynamic sorting 1467 1468 controls on the composition of basaltic sediments: Implications for interpreting martian soils. Earth and Planetary Science Letters, 423, 67–77. 1469 1470 https://doi.org/10.1016/j.epsl.2015.03.052
- Fedo, C. M., Grotzinger, J. P., Gupta, S., Fraeman, A., Edgar, L., Edgett, K., et al. (2018). 1471 1472 Sedimentology and Stratigraphy of the Murray Formation, Gale Crater, Mars. 49th Lunar 1473 and Planetary Science Conference 19-23 March, 2018, Held at The Woodlands, Texas LPI Contribution No. 2083, Id.2078, 49. Retrieved from 1474
- http://adsabs.harvard.edu/abs/2018LPI....49.2078F
- 1475
- Fenton, L. K., Gullikson, A. L., Hayward, R. K., Charles, H., & Titus, T. N. (2019). The Mars 1476 1477 Global Digital Dune Database (MGD3): Global patterns of mineral composition and bedform stability. Icarus, 330, 189-203. https://doi.org/10.1016/J.ICARUS.2019.04.025 1478
- 1479 Forni, O., Maurice, S., Gasnault, O., Wiens, R. C., Cousin, A., Clegg, S. M., et al. (2013). Independent component analysis classification of laser induced breakdown spectroscopy 1480 spectra. Spectrochimica Acta Part B: Atomic Spectroscopy, 86, 31–41. 1481 1482 https://doi.org/10.1016/j.sab.2013.05.003
- 1483 Forni, O., Gaft, M., Toplis, M. J., Clegg, S. M., Maurice, S., Wiens, R. C., et al. (2015). First detection of fluorine on Mars: Implications for Gale Crater's geochemistry. Geophysical 1484 1485 Research Letters, 42(4), 1020–1028. https://doi.org/https://doi.org/10.1002/2014GL062742
- 1486 Forni, O., E. Dehouck, A. Cousin, C. C. Bedford, G. David, S. P. Schwenzer, J. C. Bridges et al. "Elevated Fluorine Abundances Below the Siccar Point Unconformity: Implications for 1487 Fluid Circulation in Gale Crater." In Lunar and Planetary Science Conference, no. 2548, p. 1488 1489 1503.2021.
- Fraeman, A. A., Ehlmann, B. L., Arvidson, R. E., Edwards, C. S., Grotzinger, J. P., Milliken, R. 1490 1491 E., et al. (2016). The stratigraphy and evolution of lower Mount Sharp from spectral, morphological, and thermophysical orbital data sets. Journal of Geophysical Research: 1492 Planets, 121(9), 1713–1736. https://doi.org/10.1002/2016JE005095 1493
- 1494 Fraeman, A. A., Edgar, L. A., Rampe, E. B., Thompson, L. M., Frydenvang, J., Fedo, C. M., et al. (2020). Evidence for a Diagenetic Origin of Vera Rubin Ridge, Gale Crater, Mars: 1495 1496 Summary and Synthesis of Curiosity's Exploration Campaign. Journal of Geophysical 1497 Research: Planets, 125(12), 2020–006527. https://doi.org/10.1029/2020JE006527
- 1498 Frydenvang, J., Gasda, P. J., Hurowitz, J. A., Grotzinger, J. P., Wiens, R. C., Newsom, H. E., et 1499 al. (2017). Diagenetic silica enrichment and late-stage groundwater activity in Gale crater, 1500 Mars. Geophysical Research Letters, 44(10), 4716–4724. https://doi.org/10.1002/2017GL073323 1501
- 1502 Gasda, P. J., Anderson, R. B., Cousin, A., Forni, O., Clegg, S. M., Ollila, A., et al. (2021). 1503 Quantification of manganese for ChemCam Mars and laboratory spectra using a multivariate model. Spectrochimica Acta Part B: Atomic Spectroscopy, 181, 106223. 1504 https://doi.org/10.1016/j.sab.2021.106223 1505
- Gasda, P. J., Comellas, J., Essunfeld, A., Das, D., Bryk, A. B., Dehouck, E., Schwenzer, S. P., 1506 Crossey, L., Herkenhoff, K., Johnson, J. R., Newsom, H., Lanza, N. L., Rapin, W., Goetz, 1507 W., Meslin, P.-Y., Bridges, J. C., Anderson, R., David, G., Turner, S. M. R., Thorpe, M. T., 1508
- 1509 Kah, L., Frydenvang, J., Kronyak, R., Caravaca, G., Ollila, A., Le Mouelic, S., Nellessen,
- 1510 M., Hoffman, M., Fey, D., Cousin, A., Wiens, R. C., Clegg, S. M., Maurice, S., Gasnault,
- 1511 O., Delapp, D., Reyes-Newell, A. (this issue), Overview of the Morphology and Chemistry 1512 of Diagenetic features in the clay-rich Glen Torridon unit of Gale crater, Mars. *JGR*:
- 1513 Planets.
- Gellert, R., & Clark, B. C. (2015). In Situ Compositional Measurements of Rocks and Soils with
 the Alpha Particle X-ray Spectrometer on NASA's Mars Rovers. *Elements*, *11*(1), 39–44.
 https://doi.org/10.2113/gselements.11.1.39
- Greeley, R., Lancaster, N., Lee, S., & Thomas, P. (1992). Martian aeolian processes, sediments,
 and features. *Mars*, 730–766.
- Grotzinger, J. P., Crisp, J., Vasavada, A. R., Anderson, R. C., Baker, C. J., Barry, R., et al.
 (2012). Mars Science Laboratory Mission and Science Investigation. *Space Science Reviews*, *170*(1), 5–56. https://doi.org/10.1007/s11214-012-9892-2
- Grotzinger, J. P., Sumner, D. Y., Kah, L. C., Stack, K., Gupta, S., Edgar, L., et al. (2014). A
 habitable fluvio-lacustrine environment at Yellowknife Bay, Gale crater, Mars. *Science*, *343*(6169), 1242777. https://doi.org/10.1126/science.1242777
- Grotzinger, J. P., Gupta, S., Malin, M. C., Rubin, D. M., Schieber, J., Siebach, K., et al. (2015).
 Deposition, exhumation, and paleoclimate of an ancient lake deposit, Gale crater, Mars.
 Science, 350(6257), aac7575–aac7575. https://doi.org/10.1126/science.aac7575
- Hausrath, E. M., Ming, D. W., Peretyazhko, T. S., & Rampe, E. B. (2018). Reactive transport
 and mass balance modeling of the Stimson sedimentary formation and altered fracture zones
 constrain diagenetic conditions at Gale crater, Mars. *Earth and Planetary Science Letters*,
 491, 1–10. https://doi.org/10.1016/J.EPSL.2018.02.037
- Hobbs, S. W., Paull, D. J., & Bourke, M. C. (2010). Aeolian processes and dune morphology in
 Gale Crater. *Icarus*, 210(1), 102–115. https://doi.org/10.1016/J.ICARUS.2010.06.006
- Horgan, B. H. N., Johnson, J. R., Fraeman, A. A., Rice, M. S., Seeger, C., Bell III, J. F., et al.
 (2020). Diagenesis of Vera Rubin Ridge, Gale Crater, Mars, From Mastcam Multispectral
 Images. *Journal of Geophysical Research: Planets*, *125*(11), e2019JE006322.
 https://doi.org/https://doi.org/10.1029/2019JE006322
- James Sara, M. (2017). Dyngjusandur sand sheet, Iceland, as a depositional analog to the
 Stimson Fm. in Gale Crater, Mars. *Theses and Dissertations*.
 https://doi.org/10.17077/etd.25rc5gg2
- Jerolmack, D. J., Mohrig, D., Grotzinger, J. P., Fike, D. A., & Watters, W. A. (2006). Spatial
 grain size sorting in eolian ripples and estimation of wind conditions on planetary surfaces:
 Application to Meridiani Planum, Mars. *Journal of Geophysical Research E: Planets*, *111*(12), 12–14. https://doi.org/10.1029/2005JE002544
- Jia, C., Wu, L., Chen, Q., Ke, P., Yoreo, J. J. De, & Guan, B. (2020). Structural evolution of
 amorphous calcium sulfate nanoparticles into crystalline gypsum phase. *CrystEngComm*,
 22(41), 6805–6810. https://doi.org/10.1039/D0CE01173H
- Jiang, Q., & Yang, X. (2019). Sedimentological and Geochemical Composition of Aeolian
 Sediments in the Taklamakan Desert: Implications for Provenance and Sediment Supply

- Mechanisms. Journal of Geophysical Research: Earth Surface, 124(5), 1217–1237.
 https://doi.org/10.1029/2018JF004990
- Kocurek, G., & Havholm, K. G. (1993) Eolian Sequence Stratigraphy -- A Conceptual
 Framework: Chapter 16: Recent Developments in Siliciclastic Sequence Stratigraphy.
 AAPG Special Volumes, *169*, 393–409.
- Kocurek, G., & Ewing, R. C. (2012). Source-to-sink: An earth/mars comparison of boundary
 conditions for Eolian dune systems. *SEPM Special Publications*, *102*, 151–168.
 https://doi.org/10.2110/pec.12.102.0151
- Kronyak, R. E., Kah, L. C., Miklusicak, N. B., Edgett, K. S., Sun, V. Z., Bryk, A. B., &
 Williams, R. M. E. (2019b). Extensive Polygonal Fracture Network in Siccar Point group
 Strata: Fracture Mechanisms and Implications for Fluid Circulation in Gale Crater, Mars. *Journal of Geophysical Research: Planets*, 0(0). https://doi.org/10.1029/2019JE006125
- Kronyak, R. E., Kah, L. C., Edgett, K. S., VanBommel, S. J., Thompson, L. M., Wiens, R. C., et
 al. (2019a). Mineral-Filled Fractures as Indicators of Multigenerational Fluid Flow in the
 Pahrump Hills Member of the Murray Formation, Gale Crater, Mars. *Earth and Space Science*, 6(2), 238–265. https://doi.org/10.1029/2018EA000482
- Lapotre, M. G. A., & Rampe, E. B. (2018). Curiosity's Investigation of the Bagnold Dunes, Gale
 Crater: Overview of the Two-Phase Scientific Campaign and Introduction to the Special
 Collection. *Geophysical Research Letters*, 45(19), 10,200-10,210.
 https://doi.org/10.1029/2018GL079032
- Lasue, J., Cousin, A., Meslin, P. -Y., Mangold, N., Wiens, R. C., Berger, G., et al. (2018).
 Martian Eolian Dust Probed by ChemCam. *Geophysical Research Letters*, 45(20), 10,968-10,977. https://doi.org/10.1029/2018GL079210
- Malin, M. C., Edgett, K. S., Dietrich, W. E., Gupta, S., Sumner, D. Y., Wiens, R. C., et al.
 (2000). Sedimentary Rocks of Early Mars. *Science*, *290*(5498), 1927–1937.
 https://doi.org/10.1126/science.290.5498.1927
- Malin, M. C., Ravine, M. A., Caplinger, M. A., Ghaemi, F. T., Schaffner, J. A., Maki, J. N., et al.
 (2017). The Mars Science Laboratory (MSL) Mast cameras and Descent imager:
 Investigation and instrument descriptions. *Earth and Space Science*, 4(8), 506–539.
 https://doi.org/10.1002/2016EA000252
- Mangold, N., Baratoux, D., Arnalds, O., Bardintzeff, J.-M., Platevoet, B., Grégoire, M., & Pinet,
 P. (2011). Segregation of olivine grains in volcanic sands in Iceland and implications for
 Mars. *Earth and Planetary Science Letters*, 310(3–4), 233–243.
- 1583 https://doi.org/10.1016/J.EPSL.2011.07.025
- Mangold, N., Thompson, L. M., Forni, O., Williams, A. J., Fabre, C., Le Deit, L., et al. (2016).
 Composition of conglomerates analyzed by the Curiosity rover: Implications for Gale
 Crater crust and sediment sources. *Journal of Geophysical Research: Planets*, *121*(3), 353–387. https://doi.org/10.1002/2015JE004977
- Mangold, N., Schmidt, M. E., Fisk, M. R., Forni, O., McLennan, S. M., Ming, D. W., et al.
 (2017). Classification scheme for sedimentary and igneous rocks in Gale crater, Mars. *Icarus*, 284, 1–17. https://doi.org/10.1016/J.ICARUS.2016.11.005

- Maurice, S., Wiens, R. C., Saccoccio, M., Barraclough, B., Gasnault, O., Forni, O., et al. (2012).
 The ChemCam instrument suite on the Mars Science Laboratory (MSL) rover: Science
 objectives and mast unit description. *Space Science Reviews*, *170*(1–4), 95–166.
 https://doi.org/10.1007/s11214-012-9912-2
- Maurice, S., Clegg, S. M., Wiens, R. C., Gasnault, O., Rapin, W., Forni, O., et al. (2016).
 ChemCam activities and discoveries during the nominal mission of the Mars Science
 Laboratory in Gale crater, Mars. *Journal of Analytical Atomic Spectrometry*, *31*(4), 863–
 889. https://doi.org/10.1039/C5JA00417A
- McCollom, T. M., Hynek, B. M., Rogers, K., Moskowitz, B., & Berquó, T. S. (2013). Chemical and mineralogical trends during acid-sulfate alteration of pyroclastic basalt at Cerro Negro volcano and implications for early Mars. *Journal of Geophysical Research: Planets*, *118*(9), 1719–1751. https://doi.org/10.1002/jgre.20114
- McCollom, T. M., Robbins, M., Moskowitz, B., Berquõ, T. S., Jöns, N., & Hynek, B. M. (2013).
 Experimental study of acid-sulfate alteration of basalt and implications for sulfate deposits
 on Mars. *Journal of Geophysical Research E: Planets*, *118*(4), 577–614.
 https://doi.org/10.1002/jgre.20044
- Meslin, P.-Y., Gasnault, O., Forni, O., Schröder, S., Cousin, A., Berger, G., et al. (2013). Soil
 diversity and hydration as observed by ChemCam at Gale crater, Mars. *Science (New York, N.Y.*), *341*(6153), 1238670. https://doi.org/10.1126/science.1238670
- Milliken, R. E., Grotzinger, J. P., & Thomson, B. J. (2010a). Paleoclimate of Mars as captured
 by the stratigraphic record in Gale Crater. *Geophysical Research Letters*, *37*(4).
 https://doi.org/10.1029/2009GL041870
- Milliken, R. E., Grotzinger, J. P., & Thomson, B. J. (2010b). Paleoclimate of Mars as captured
 by the stratigraphic record in Gale Crater. *Geophysical Research Letters*, *37*(4).
 https://doi.org/10.1029/2009GL041870
- Morris, R. V., Vaniman, D. T., Blake, D. F., Gellert, R., Chipera, S. J., Rampe, E. B., et al.
 (2016). Silicic volcanism on Mars evidenced by tridymite in high-SiO ₂ sedimentary rock at
 Gale crater. *Proceedings of the National Academy of Sciences*, *113*(26), 7071–7076.
 https://doi.org/10.1073/pnas.1607098113
- Morrison, S. M., Downs, R. T., Blake, D. F., Vaniman, D. T., Ming, D. W., Hazen, R. M., et al.
 (2018b). Crystal chemistry of martian minerals from Bradbury Landing through Naukluft
 Plateau, Gale crater, Mars. *American Mineralogist*, *103*(6), 857–871.
 https://doi.org/10.2138/am-2018-6124
- Morrison, S. M., Downs, R. T., Blake, D. F., Prabhu, A., Eleish, A., Vaniman, D. T., et al.
 (2018a). Relationships between unit-cell parameters and composition for rock-forming
 minerals on Earth, Mars, and other extraterrestrial bodies. *American Mineralogist*, *103*(6),
 848–856. https://doi.org/10.2138/am-2018-6123
- Morrone, C., De Rosa, R., Le Pera, E., & Marsaglia, K. M. (2017). Provenance of volcaniclastic
 beach sand in a magmatic-arc setting: An example from Lipari island (Aeolian archipelago,
 Tyrrhenian Sea). *Geological Magazine*, 154(4), 804–828.
- 1631 https://doi.org/10.1017/S001675681600042X
- 1632 Le Mouélic, S., Gasnault, O., Herkenhoff, K. E., Bridges, N. T., Langevin, Y., Mangold, N., et

- 1633al. (2015). The ChemCam Remote Micro-Imager at Gale crater: Review of the first year of1634operations on Mars. *Icarus*, 249, 93–107. https://doi.org/10.1016/J.ICARUS.2014.05.030
- Mountney, N. P., & Russell, A. J. (2004). Sedimentology of cold-climate aeolian sandsheet
 deposits in the Askja region of northeast Iceland. *Sedimentary Geology*, *166*(3–4), 223–244.
 https://doi.org/10.1016/J.SEDGEO.2003.12.007
- Muhs, D. R. (2004). Mineralogical maturity in dunefields of North America, Africa and
 Australia. *Geomorphology*, 59(1–4), 247–269.
 https://doi.org/10.1016/j.geomorph.2003.07.020
- Nesbitt, H. W., & Wilson, R. E. (1992). Recent chemical weathering of basalts. *American Journal of Science*, 292(10), 740–777.
- Newsom, H. E., Edgett, K. S., Fey, D. M., Wiens, R. C., Frydenvang, J., Banham, S. G., et al.
 (2018). A Buried Aeolian Lag Deposit at an Unconformity Between the Murray and
 Stimson Formations at Marias Pass, Gale Crater, Mars. 49th Lunar and Planetary Science *Conference 19-23 March, 2018, Held at The Woodlands, Texas LPI Contribution No. 2083, Id.2263, 49.* Retrieved from http://adsabs.harvard.edu/abs/2018LPI....49.2263N
- O'Connell-Cooper, C. D., Spray, J. G., Thompson, L. M., Gellert, R., Berger, J. A., Boyd, N. I.,
 et al. (2017). APXS-derived chemistry of the Bagnold dune sands: Comparisons with Gale
 Crater soils and the global Martian average. *Journal of Geophysical Research: Planets*,
 122(12), 2623–2643. https://doi.org/10.1002/2017JE005268
- Pettijohn, F. J. (1954). Classification of Sandstones. *The Journal of Geology*, 62(4), 360–365.
 https://doi.org/10.1086/626172
- Pokrovsky, O. S., & Schott, J. (2000). Kinetics and mechanism of forsterite dissolution at 25°C
 and pH from 1 to 12. *Geochimica et Cosmochimica Acta*, 64(19), 3313–3325.
 https://doi.org/10.1016/S0016-7037(00)00434-8
- Potter-McIntyre, S. L., Chan, M. A., & McPherson, B. J. (2014). Concretion Formation In
 Volcaniclastic Host Rocks: Evaluating the Role of Organics, Mineralogy, and Geochemistry
 On Early Diagenesis. *Journal of Sedimentary Research*, 84(10), 875–892.
 https://doi.org/10.2110/jsr.2014.58
- Potter, S. L., Chan, M. A., Petersen, E. U., Dyar, M. D., & Sklute, E. (2011). Characterization of
 Navajo Sandstone concretions: Mars comparison and criteria for distinguishing diagenetic
 origins. *Earth and Planetary Science Letters*, 301(3–4), 444–456.
 https://doi.org/10.1016/J.EPSL.2010.11.027
- Rampe, E. B., Ming, D. W., Blake, D. F., Bristow, T. F., Chipera, S. J., Grotzinger, J. P., et al.
 (2017). *Mineralogy of an ancient lacustrine mudstone succession from the Murray formation, Gale crater, Mars. Earth and Planetary Science Letters* (Vol. 471).
 https://doi.org/10.1016/j.epsl.2017.04.021
- Rampe, E. B., Lapotre, M. G. A., Bristow, T. F., Arvidson, R. E., Morris, R. V., Achilles, C. N.,
 et al. (2018). Sand Mineralogy Within the Bagnold Dunes, Gale Crater, as Observed In Situ
 and From Orbit. *Geophysical Research Letters*, 45(18), 9488–9497.
 https://doi.org/10.1029/2018GL079073
- 10/2 https://doi.org/10.1029/2010GE079075
- 1673 Rampe, E. B., Bristow, T. F., Morris, R. V., Morrison, S. M., Achilles, C. N., Ming, D. W., et al.

- 1674 (2020a). Mineralogy of Vera Rubin Ridge From the Mars Science Laboratory CheMin
 1675 Instrument. *Journal of Geophysical Research: Planets*, *125*(9), e2019JE006306.
 1676 https://doi.org/10.1029/2019JE006306
- Rampe, E.B., Yen, A., Bristow, T., Blake, D.F., Vaniman, D., Achilles, C., Chipera, S., Downs,
 R.T., Ming, D.W., Morris, R.V. and Morrison, S.M., 2020b, December. Mineralogy of the
 Greenheugh Pediment and Underlying Murray Formation from the Mars Science
 Laboratory CheMin Instrument. In *AGU Fall Meeting Abstracts* (Vol. 2020, pp. P070-09).
- Rapin, W., Meslin, P.-Y., Maurice, S., Vaniman, D., Nachon, M., Mangold, N., et al. (2016).
 Hydration state of calcium sulfates in Gale crater, Mars: Identification of bassanite veins. *Earth and Planetary Science Letters*, 452, 197–205.
 https://doi.org/10.1016/J.EPSL.2016.07.045
- Rapin, W., Meslin, P. Y., Maurice, S., Wiens, R. C., Laporte, D., Chauviré, B., et al. (2017b).
 Quantification of water content by laser induced breakdown spectroscopy on Mars. *Spectrochimica Acta Part B: Atomic Spectroscopy*, *130*, 82–100.
 https://doi.org/10.1016/J.SAB.2017.02.007
- Rapin, W., Bousquet, B., Lasue, J., Meslin, P. Y., Lacour, J. L., Fabre, C., et al. (2017a).
 Roughness effects on the hydrogen signal in laser-induced breakdown spectroscopy.
 Spectrochimica Acta Part B: Atomic Spectroscopy, 137, 13–22.
 https://doi.org/10.1016/J.SAB.2017.09.003
- Rudolph, B., K. Bennett, and M. Rice. "Sources of sand in Mt Sharp: Possible volcanic layers in
 Gale crater, Mars." *Planets* 122 (2019): 2489-2509.
- 1695
- Schröder, S., Meslin, P. Y., Gasnault, O., Maurice, S., Cousin, A., Wiens, R. C., et al. (2015).
 Hydrogen detection with ChemCam at Gale crater. *Icarus*, 249, 43–61.
 https://doi.org/10.1016/J.ICARUS.2014.08.029
- Siebach, K. L., McLennan, S. M., & Fedo, C. M. (2017). Geochemistry of the Stimson
 Sandstone, Gale Crater, Mars. 48th Lunar and Planetary Science Conference, Held 20-24
 March 2017, at The Woodlands, Texas. LPI Contribution No. 1964, Id.2499, 48. Retrieved
 from http://adsabs.harvard.edu/abs/2017LPI....48.2499S
- Smith, R. J., Rampe, E. B., Horgan, B. H. N., & Dehouck, E. (2018). Deriving Amorphous
 Component Abundance and Composition of Rocks and Sediments on Earth and Mars. *Journal of Geophysical Research: Planets*, *123*(10), 2485–2505.
 https://doi.org/10.1029/2018JE005612
- Stack, K. M., Grotzinger, J. P., Lamb, M. P., Gupta, S., Rubin, D. M., Kah, L. C., et al. (2019).
 Evidence for plunging river plume deposits in the Pahrump Hills member of the Murray
 formation, Gale crater, Mars. *Sedimentology*, *66*(5), 1768–1802.
 https://doi.org/10.1111/sed.12558
- Steele, L. J., Kite, E. S., & Michaels, T. I. (2018). Crater Mound Formation by Wind Erosion on
 Mars. *Journal of Geophysical Research: Planets*, *123*(1), 113–130.
- 1713 https://doi.org/https://doi.org/10.1002/2017JE005459
- 1714 Sutter, B., McAdam, A. C., Mahaffy, P. R., Ming, D. W., Edgett, K. S., Rampe, E. B., et al.

- 1715 (2017). Evolved gas analyses of sedimentary rocks and eolian sediment in Gale Crater,
- 1716 Mars: Results of the Curiosity rover's sample analysis at Mars instrument from Yellowknife
- 1717 Bay to the Namib Dune. *Journal of Geophysical Research: Planets*, *122*(12), 2574–2609. https://doi.org/10.1002/2016/E005225
- 1718 https://doi.org/10.1002/2016JE005225
- Sutter, B., McAdam, A. C., Ming, D. W., Archer, P. D., Thompson, L. M., Stern, J. C.,
 Eigenbrode, J. L., Mahaffy, P. R. (2022) "Constraining alteration processes along the Siccar
 Point Unconformity, Gale crater: Results from the Sample Analysis at Mars instrument",
 53rd Lunar and Planetary Science Conference, Woodlands, TX, USA.
- Thomas, M., Clarke, J. D. A., & Pain, C. F. (2005). Weathering, erosion and landscape processes
 on Mars identified from recent rover imagery, and possible Earth analogues. *Australian Journal of Earth Sciences*, 52(3), 365–378. https://doi.org/10.1080/08120090500134597
- Thomas, N. H., Ehlmann, B. L., Anderson, D. E., Clegg, S. M., Forni, O., Schröder, S., et al.
 (2018). Characterization of Hydrogen in Basaltic Materials With Laser-Induced Breakdown
 Spectroscopy (LIBS) for Application to MSL ChemCam Data. *Journal of Geophysical*
- 1729 *Research: Planets*, *123*(8), 1996–2021.
- 1730 https://doi.org/https://doi.org/10.1029/2017JE005467
- Thomas, N. H., Ehlmann, B. L., Rapin, W., Rivera-Hernández, F., Stein, N. T., Frydenvang, J.,
 et al. (2020). Hydrogen Variability in the Murray Formation, Gale Crater, Mars. *Journal of Geophysical Research: Planets*, *125*(9), e2019JE006289.
 https://doi.org/10.1029/2019JE006289
- Thomson, B. J., Bridges, N. T., Milliken, R., Baldridge, A., Hook, S. J., Crowley, J. K., et al.
 (2011). Constraints on the origin and evolution of the layered mound in Gale Crater, Mars
 using Mars Reconnaissance Orbiter data. *Icarus*, 214(2), 413–432.
 https://doi.org/10.1016/J.ICARUS.2011.05.002
- Thompson, L. M., O'Connell-Cooper, C., Berger, J. A., Yen, A., APXS Team. (this issue).
 Widespread alteration at the base of the Siccar Point unconformity and evidence for a
 regional potassic source rock at Gale crater: Exploration of the Murray formationGreenheugh pediment cap rock contact with the Alpha Particle X-ray spectrometer. *JGR*:
- 1743 Planets.
- Thorpe, M. T., Bristow, T. F., Rampe, E. B., Tosca, N. J., Grotzinger J. P., Fedo, C. M., Chipera,
 S. J., Downs, G., Achilles, C. N., Blake, D. F., Castle, N., Craig, P., Des Marais, D. J.,
 Downs, R. T., Hazen, R. M., Ming, D. W., Morris, R. V., Morrison, S. M., Treiman, A. H.,
 Tu, V., Vaniman, D. T., Yen, A. S., Bryk, A. B., Bennett, K. A., Fox, V. K., Siebach, K. L.,
 Fraeman, A. A., Vasavada, A. R. (this issue). The Mineralogy an dSedimentary History of
 the Glen Torridon Region, as detailed by the Mars Science Laboratory CheMin Instrument. *JGR: Planets*.
- Tosca, N. J., McLennan, S. M., Lindsley, D. H., & Schoonen, M. A. A. (2004). Acid-sulfate
 weathering of synthetic Martian basalt: The acid fog model revisited. *Journal of Geophysical Research E: Planets*, *109*(5), 5003. https://doi.org/10.1029/2003JE002218
- Treiman, A. H., Bish, D. L., Vaniman, D. T., Chipera, S. J., Blake, D. F., Ming, D. W., et al.
 (2016a). Mineralogy, provenance, and diagenesis of a potassic basaltic sandstone on Mars:
 CheMin X-ray diffraction of the Windjana sample (Kimberley area, Gale Crater). *Journal of Geophysical Research. Planets*, *121*(1), 75–106. https://doi.org/10.1002/2015JE004932

1758 1759 1760 1761	 Treiman, A. H., Bish, D. L., Vaniman, D. T., Chipera, S. J., Blake, D. F., Ming, D. W., et al. (2016b). Mineralogy, provenance, and diagenesis of a potassic basaltic sandstone on Mars: CheMin X-ray diffraction of the Windjana sample (Kimberley area, Gale Crater). <i>Journal of Geophysical Research: Planets</i>, 121(1), 75–106. https://doi.org/10.1002/2015JE004932
1762	Vaniman, D. T., Bish, D. L., Ming, D. W., Bristow, T. F., Morris, R. V., Blake, D. F., et al.
1763	(2014). Mineralogy of a Mudstone at Yellowknife Bay, Gale Crater, Mars. <i>Science</i> ,
1764	343(6169), 1243480–1243480. https://doi.org/10.1126/science.1243480
1765	Vasavada, A. R., Grotzinger, J. P., Arvidson, R. E., Calef, F. J., Crisp, J. A., Gupta, S., et al.
1766	(2014). Overview of the Mars Science Laboratory mission: Bradbury Landing to
1767	Yellowknife Bay and beyond. <i>Journal of Geophysical Research: Planets</i> , 119(6), 1134–
1768	1161. https://doi.org/10.1002/2014JE004622
1769 1770 1771 1772 1773	 Watkins, J. A., Grotzinger, J., Stein, N., Banham, S. G., Gupta, S., Rubin, D., et al. (2016). Paleotopography of Erosional Unconformity, Base of Stimson Formation, Gale Crater, Mars. 47th Lunar and Planetary Science Conference, Held March 21-25, 2016 at The Woodlands, Texas. LPI Contribution No. 1903, p.2939, 47, 2939. Retrieved from http://adsabs.harvard.edu/abs/2016LPI47.2939W
1774 1775 1776 1777	 Wiens, R. C., Maurice, S., Barraclough, B., Saccoccio, M., Barkley, W. C., Bell III, J. F., et al. (2012). The ChemCam Instrument Suite on the Mars Science Laboratory (MSL) Rover: Body Unit and Combined System Tests. <i>Space Science Reviews</i>, 170(1–4), 167–227. https://doi.org/10.1007/s11214-012-9902-4
1778	Wiens, R. C., Maurice, S., Lasue, J., Forni, O., Anderson, R. B., Clegg, S., et al. (2013). Pre-
1779	flight calibration and initial data processing for the ChemCam laser-induced breakdown
1780	spectroscopy instrument on the Mars Science Laboratory rover. <i>Spectrochimica Acta Part</i>
1781	<i>B: Atomic Spectroscopy</i> , 82, 1–27. https://doi.org/10.1016/j.sab.2013.02.003
1782	Wiens, R. C. (2013a) MSL MARS CHEMCAM LIBS SPECTRA 4/5 RDR V1.0 [Data set].
1783	NASA Planetary Data System. https://doi.org/10.17189/1519485
1784	Wiens, R. C. (2013b) MSL MARS CHEMCAM REMOTE MICRO-IMAGER CAMERA 5
1785	RDR V1.0 [Data set]. NASA Planetary Data System. https://doi.org/10.17189/1519494
1786	Wiens, R. C., Edgett, K. S., Stack, K. M., Dietrich, W. E., Bryk, A. B., Mangold, N., et al.
1787	(2020). Origin and composition of three heterolithic boulder- and cobble-bearing deposits
1788	overlying the Murray and Stimson formations, Gale Crater, Mars. <i>Icarus</i> , 350, 113897.
1789	https://doi.org/10.1016/j.icarus.2020.113897
1790	Willetts, B. B., Rice, M. A., & Swaine, S. E. (1982). Shape effects in aeolian grain transport.
1791	Sedimentology, 29(3), 409–417. https://doi.org/10.1111/j.1365-3091.1982.tb01803.x
1792	Yen, A. S., Ming, D. W., Vaniman, D. T., Gellert, R., Blake, D. F., Morris, R. V., et al. (2017).
1793	Multiple stages of aqueous alteration along fractures in mudstone and sandstone strata in
1794	Gale Crater, Mars. <i>Earth and Planetary Science Letters</i> , 471, 186–198.
1795	https://doi.org/10.1016/J.EPSL.2017.04.033
1796	Yen, A. S., Morris, R. V., Ming, D. W., Schwenzer, S. P., Sutter, B., Vaniman, D. T., et al.
1797	(2021). Formation of Tridymite and Evidence for a Hydrothermal History at Gale Crater,
1798	Mars. <i>Journal of Geophysical Research: Planets</i> , 126(3), e2020JE006569.
1799	https://doi.org/10.1029/2020JE006569

- 1800 Zhu, B., & Yang, X. (2009). Chemical Weathering of Detrital Sediments in the Taklamakan
- 1801 Desert, Northwestern China. *Geographical Research*, 47(1), 57–70.
- 1802 https://doi.org/10.1111/J.1745-5871.2008.00555.X
- Iso3 Zimbelman, J. R., & Foroutan, M. (2020). Dingo Gap: Curiosity Went Up a Small Transverse
 Aeolian Ridge and Came Down a Megaripple. *Journal of Geophysical Research: Planets*,
- 1805 *125*(12), e2020JE006489. https://doi.org/10.1029/2020JE006489
- 1806

1807 Figure Headings

1808 **Figure 1**. Map of the Stimson formation localities investigated to date along the Curiosity

1809 rover's traverse. The labelled, colored units show the present-day extent of the Stimson

1810 formation (Siccar Point group) at the Emerson Plateau (blue), Naukluft Plateau (green), Murray

1811 Buttes (yellow), and Greenheugh pediment capping unit (orange). The solid white line shows the

- 1812 traverse, and the yellow waypoints mark key drill sites in the Mt Sharp group mudstones and
- 1813 Siccar Point group Stimson formation sandstones. The black streaks oriented from the NE to the
- 1814 SW derive the active Bagnold dune field.

1815 **Figure 2**. Stratigraphic column of the geological units investigated by the MSL Curiosity rover

- 1816 in Gale crater, Mars. The Stimson formation sandstones are represented as tan-colored blocks.
- 1817 Dots are drill hole samples with those discussed in this paper labelled as BK (Buckskin), Big Sky
- 1818 (BS), Greenhorn (GH), Lubango (LB), Okoruso (OK), Oudam (OU), Hutton (HU), Edinburgh
- 1819 (EB). Mt Sharp group drill samples are labelled purple, Stimson formation drills are labelled red.
- 1820 Image credit is courtesy of the MSL sedimentology and stratigraphy group.
- 1821 **Figure 3.** Mastcam mosaics of the Stimson formation at A) the top of the Greenheugh pediment
- 1822 (mcam14517, sol 2702), B) side view of the Greenheugh pediment (mcam14007, sol 2683), C)
- 1823 Williams outcrop at the Emerson plateau showing cross-bedding and a vertical light-toned halo
- 1824 (mcam004777, sol 1087), D) Naukluft plateau at the unconformity (dashed white line) of the
- 1825 Murray formation and the Stimson formation (mcam0597, sol 1275). MAHLI images of the 1826 bedrock drill targets for the Stimson formation at each locality are also provided in the insets: a)
- bedrock drill targets for the Stimson formation at each locality are also provided in the insets: a)
 is of the Edinburgh target of the Edinburgh interval (sol 2703, after DRT), b) The Machir Bay
- target of the Ladder interval (sol 2699, after DRT), c) The target Big Sky (sol 1114, after DRT),
- and d) the target Okoruso (sol 1330, after DRT). Image credits for the Mastcam mosaics:
- 1830 NASA/JPL-Caltech/MSSS.
- **Figure 4.** ChemCam RMI image mosaics, MAHLI images, and Navcam images showing the diversity of alteration features within the Stimson formation at each locality.
- 1833 **Figure 5. A.** Distribution of ChemCam classifications between each Stimson locality where each
- 1834 count equals to one ChemCam observation point. Bedrock low totals is the number of bedrock
- observation points that had total sum of oxides <90 wt%. **B.** Distribution of ChemCam
- 1836 classifications between each Greenheugh stratigraphic interval (excluding the basal Stimson
- 1837 unit). Observation points that targeted float rocks and soils at these localities were not included.
- 1838 Sampling bias will have affected the number of ChemCam points on alteration features relative
- to those of bedrock, particularly for the silica-rich halos, as they were intentionally prioritized in
- 1840 the Emerson and Naukluft plateaus.
- 1841 **Figure 6.** Histograms of ChemCam major element oxides for the bulk rock dataset at each
- 1842 Stimson locality. These data exclude those outside the 90-105 wt% range for total sum of oxides.
- 1843 Binning was selected based on the Minitab default setting which defines the histogram intervals
- 1844 according to the number of observations and the maximum and minimum values in the dataset.
- 1845 For our dataset, the number of intervals range between 21 and 34. The black vertical line shows
- 1846 the mean value of the major element oxide at each locality. The y-axis shows the number of data

points in each bin (frequency). The histogram for MnO was restricted to 1 wt% but has outliers 1847 1848 to higher values which skews the mean, particularly for Greenheugh.

Figure 7. Boxplots of geochemical variation associated with the unconformity for the bedrock 1849

datasets of the underlying Carolyn Shoemaker formation and overlying Stimson formation. FeO_T 1850

1851 and TiO₂ showed little variation and were not included in the plots. Dashed lines show the mean

- composition of the Stimson formation at Greenheugh (red) and the Carolyn Shoemaker 1852
- formation (blue). Dotted lines show the uncertainty around the respective sample means through 1853
- 1854 2x the standard error.

1855 Figure 8. A. H peak area data distribution for the Greenheugh pediment capping unit according to the different stratigraphic intervals. The concretion-rich Hutton interval has the lowest peak H 1856

1857 abundance while the non-concretionary Edinburgh and Ladder Intervals have the higher H

abundance. B. F variation with distance from the unconformity for the Stimson formation and the 1858

underlying Mt Sharp group. Alteration features (halos, concretions, raised ridges, veins) and 1859 alteration-free bedrock are shown. C. Annotated Mastcam image (mcam 14083) of the side of

1860

1861 the Greenheugh pediment showing the division between the concretionary Glean Beag Interval and the Edinburgh Interval that has fewer concretions concentrated towards the base. Image 1862

credit for the Mastcam mosaic used in C: NASA/JPL-Caltech/MSSS. 1863

Figure 9. Dendrogram of the Cluster analysis results for Stimson across all localities. 1864

Figure 10. Pie charts showing the distribution of the clusters across Stimson localities. The red 1865 and blue wedges correspond to the relative abundances of mafic minerals present at each locality. 1866

Figure 11. Biplots showing the geochemistry of the Stimson formation at the Greenheugh 1867

pediment capping unit compared to that of the earlier Stimson formations, the geochemical 1868

1869 compositions of the clusters, and the chemical compositions of the primary igneous minerals

identified in the Big Sky drilled sample of the Stimson sandstone (Yen et al., 2017), and the 1870

Gobabeb scooped sample of the active aeolian Bagnold dunes (Achilles et al., 2017). Bulk 1871 1872 Stimson compositions are shown as density contours in addition to a linear regression model

(dashed black line). Cluster compositions also show 1 standard deviation. A. A 1873

 $Al_2O_3+Na_2O+K_2O$ vs MgO + FeO_T plot to show the variation in chemistry that relates to mafic 1874

and felsic minerals. B. A MnO vs MgO plot to show the unusual chemistry of Cluster 5 1875

1876 Figure 12. Annotated ChemCam RMI mosaics of targets A. Mowry (sol 1080) of the Emerson

Plateau, B. Bentiaba (sol 1320) of the Naukluft plateau, C. Lowther Hills (sol 2697) of the 1877

Gleann Beag interval, and **D**. Dunedin (sol 2720) of the Edinburgh Interval. Colored circles show 1878

1879 the cluster memberships of the associated ChemCam observation points.

1880 Figure 13: A. and B. show geochemical plots of the ChemCam targets immediately above and

1881 below the unconformity. Each point is an observation point composition with green points

related to the Carolyn Shoemaker formation (target names are Ariundle, Shannochie, 1882

Corriecravie, Hutton), and orange and gray points related to the Stimson formation (target names 1883

are Huttons section and Gleann Beag). Black arrow shows the geochemical trend relating to Ca-1884

sulfates. The solid line represents a mixing line between the average composition of the Stimson 1885

1886 formation at the Greenheugh pediment and the bulk Carolyn Shoemaker formation (Dehouck et al., *this issue*). Percentages along the mixing line show the percentage of Stimson (e.g., 25% is

1888 25% Stimson, 75% Murray). Dashed lines represent 1 standard deviation of this mixing line. C.

and **D**. show the MAHLI images of Stimson targets near the unconformity at the Greenheugh

1890 pediment capping unit. Yellow arrows point to angular mudstone clasts in the Stimson formation.

1891 White arrows point to light-toned, spherical grains that may also relate to rounded mudstone

1892 clasts (Banham et al., *this issue*).

Figure 14. Mastcam image of light-toned alteration features in the underlying Murray formation
pooling at the unconformity with the Stimson formation and breaching through the fractures.
Image was taken at the Naukluft plateau (mcam05931, sol 1267). Image credit: NASA/JPLCaltech/MSSS.

1897**Table 1.** Basic statistics (mean, standard error of the mean (SE Mean), standard deviation

1898 (StDev), minimum (Min), median (Med), and maximum (Max)) for the bulk rock data at each

1899 locality. Mean and median values are highlighted based on whether they represent the largest

1900 value (red) or the lowest value (blue), taking into consideration the standard error of the mean.

Both standard deviation and error are provided as standard deviation is representative of the

dispersion of the individual data values to the mean, while the standard error of the meanconsiders the number of points in the dataset and measures how far the sample mean of the data

is from the true population.

Table 2. Mean composition, standard error of the mean (+/-), and standard deviation (std) of the
derived clusters. Colors highlight the highest abundances (red), the lowest (blue), and the
gradation in between for each of the five clusters.

1908

Figure 1.

Bagnold 57 Buckskin TI'9 - Big Sky Emerson Plateau sols 992-1154 orusc 1223 - Gobabeb

aph Peak

Naukluft Plateau sols 1279-1352

Murray Buttes sols 1414-1455

Bradouny grown

•1659 - Ogunquit Beach

Vera Rubin ridge 2058 - Duluth ى 2224 Highfie^{ld} 2264 - Rock Hall

> lutton Edinburgh

> > Greenheugh pediment sols 2693-2733

250 500 1000 meters 1500 2000 ō

Figure 2.



Figure 3.



Figure 4.



Figure 5.



Figure 6.



Figure 7.



Figure 8.



Figure 9.







Figure 10.



Figure 11.







Sol 2694 – Target Huttons_Section Context view from ~ 25 cm standoff



Figure 12.



Figure 13.


OCluster 1 OCluster 2 OCluster 3 OCluster 4 OCluster 5 (mafic) (mafic avg) (felsic avg) (high MnO) (felsic) Figure 14.



		90-105 wt% totals									
Variables	Locality	Mear	SE Mear	StDev	Min	Med	Max				
SiO ₂	Emerson	46. 1	0.2	3.0	38.9	945.6	57.1				
	Naukluft	47.2	0.2	3.6	37.6	647.3	61.7				
	Greenheugh	44.9	0.2	3.1	35.6	644.5	59.0				
TiO ₂	Emerson	1.0	0.0	0.3	0.5	0.9	2.7				
	Naukluft	0.9	0.0	0.2	0.6	0.9	2.0				
	Greenheugh	0.9	0.0	0.1	0.6	0.9	1.7				
Al ₂ O ₃	Emerson	11.3	0.2	2.9	5.7	10.8	22.5				
	Naukluft	12.1	0.2	2.9	3.3	12.0	19.6				
	Greenheugh	10.3	0.2	2.9	4.6	9.6	21.0				
FeO⊤	Emerson	19.2	0.1	2.3	6.6	19.5	26.0				
	Naukluft	19.3	0.1	2.4	7.8	19.4	32.1				
	Greenheugh	20.12	20.3	3.7	0.9	20.3	41.2				
MgO	Emerson	8.1	0.2	2.9	0.9	7.7	21.6				
	Naukluft	6.8	0.2	2.5	2.0	6.7	22.0				
	Greenheugh	8.4	0.2	2.8	2.1	8.1	17.4				
CaO	Emerson	6.2	0.1	2.0	1.6	6.2	17.2				
	Naukluft	6.0	0.1	1.7	1.5	6.0	11.5				
	Greenheugh	5.0	0.1	1.8	1.0	4.9	11.1				
Na ₂ O	Emerson	2.9	0.1	0.8	1.5	2.7	8.1				
	Naukluft	3.8	0.0	0.7	1.5	3.2	6.0				
	Greenheugh	3.1	0.1	0.8	1.7	2.9	8.0				
K ₂ O	Emerson	0.6	0.0	0.5	0.1	0.4	2.7				
	Naukluft	0.6	0.0	0.4	0.1	0.4	2.4				
	Greenheugh	0.6	0.0	0.3	0.1	0.5	1.8				
MnO	Emerson	0.2	0.0	0.2	0.0	0.2	2.6				

Naukluft 0.1 0.0 0.1 0.0 0.1 0.3 Greenheugh 0.3 0.0 0.6 0.1 0.2 5.5 Sum of Oxides Emerson 95.4 0.2 3.5 90.0 95.1 104.2 Naukluft 96.1 0.2 3.6 90.0 95.6 104.8 Greenheugh 93.5 0.2 2.8 90.0 92.8 102.2								
Greenheugh 0.3 0.0 0.6 0.1 0.2 5.5 Sum of OxidesEmerson 95.4 0.2 3.5 90.0 95.1 104.2 Naukluft 96.1 0.2 3.6 90.0 95.6 104.8 Greenheugh 93.5 0.2 2.8 90.0 92.8 102.2	Naul	duft	0.1	0.0	0.1	0.0	0.1	0.3
Sum of Oxides Emerson 95.4 0.2 3.5 90.095.1 104.2 Naukluft 96.1 0.2 3.6 90.0 95.6 104.8 Greenheugh 93.5 0.2 2.8 90.0 92.8 102.2	Gree	nheugh	0.3	0.0	0.6	0.1	0.2	5.5
Naukluft <mark>96.1</mark> 0.2 3.6 90.0 <mark>95.6</mark> 104.8 Greenheugh <mark>93.5</mark> 0.2 2.8 90.0 <mark>92.8</mark> 102.2	Sum of OxidesEme	rson	95.4	0.2	3.5	90.0	95.1	104.2
Greenheugh 93.5 0.2 2.8 90.0 92.8 102.2	Naul	duft	96.1	0.2	3.6	90.0	95.6	104.8
	Gree	nheugh	93.5	0.2	2.8	90.0	92.8	102.2

Cluster	SiO ₂	+/- std	TiO ₂	+/-	Std	Al ₂ O ₃	+/- std	FeO _T	+/- std	MgO	+/- sto	l CaO	+/- 9	std	Na ₂ O	+/- std	K ₂ C	0+/- std	MnO	+/- std
1	43.9	0.7 2.8	0.8	0.0	0.1	6.9	0.4 1.7	21.4	0.7 3.0	14.7	1.1 4.1	3.6	0.4	1.6	2.3	0.1 0.6	0.3	0.1 0.2	0.7	0.2 0.7
2	43.9	0.1 1.8	1.0	0.0	0.2	9.7	0.1 1.8	20.6	0.1 2.1	8.6	0.1 2.4	\$ 5.6	0.1	1.8	2.7	0.0 0.4	0.4	0.0 0.2	0.2	0.0 0.0
3	48.5	0.2 2.8	0.9	0.0	0.2	13.1	0.1 2.5	18.7	0.1 2.0	6.5	0.1 1.9	9 6.0	0.1	1.8	3.4	0.0 0.7	0.7	0.0 0.4	0.1	0.0 0.1
4	54.5	0.8 3.2	0.9	0.1	0.2	18.0	0.6 2.4	9.8	1.0 3.7	2.5	0.3 1.3	5.7	0.7	2.6	5.7	0.3 1.3	1.0	0.2 0.6	0.1	0.0 0.0
5	42.8	1.2 2.6	0.8	0.0	0.1	10.3	1.4 3.2	21.7	1.1 2.4	8.8	1.9 4.2	2 4.0	0.9 2	2.0	3.3	0.3 0.7	1.0	0.3 0.6	2.9	0.2 0.4