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# Pre-collapse motion of the February 2021 Chamoli rock–ice avalanche, Indian Himalaya

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**Abstract.** Landslides are a major geohazard that cause thousands of fatalities every year. Despite their importance, identifying unstable slopes and forecasting collapses remains a major challenge. In this study, we use the 7 February 2021 Chamoli rock–ice avalanche as a data-rich example to investigate the potential of remotely sensed datasets for the assessment of slope stability. We investigate imagery over the 3 decades preceding collapse and assess the precursory signs exhibited by this slope prior to the catastrophic collapse. We evaluate monthly slope motion from 2015 to 2021 through feature tracking of high-resolution optical satellite imagery. We then combine these data with a time series of pre- and post-event digital elevation models (DEMs), which we use to evaluate elevation change over the same area. Both datasets show that the  $26.9 \times 10^6 \text{ m}^3$  collapse block moved over 10 m horizontally and vertically in the 5 years preceding collapse, with particularly rapid motion occurring in the summers of 2017 and 2018. We propose that the collapse results from a combination of snow loading in a deep headwall crack

and permafrost degradation in the heavily jointed bedrock. Despite observing a clear precursory signal, we find that the timing of the Chamoli rock–ice avalanche could likely not have been forecast from satellite data alone. Our results highlight the potential of remotely sensed imagery for assessing landslide hazard in remote areas, but that challenges remain for operational hazard monitoring.

## 1 Introduction

### 1.1 Landslide hazard

Landslides are a major geohazard that cause thousands of deaths each year (Petley, 2012; Froude and Petley, 2018). Mitigating landslide hazard is a major challenge facing geoscientists and hazard managers. Evaluating landslide hazard is challenging due to the wide range of source conditions and the varying temporal scales at which the driving processes

interact. Landslides are also associated with a wide range of short- to long-term triggers, ranging from earthquakes to water flow, or simple weaknesses in the rock, which further complicates their forecasting and process understanding (van Westen et al., 2006).

Ground-based observations of displacement (e.g., with GNSS/GPS – global navigation satellite system/Global Positioning System – or ground-based radar interferometry), tilt (e.g., inclinometers), pressure (e.g., piezometers), and other variables can be useful in monitoring landslide progression (e.g., Uhlemann et al., 2016). When observed, landslide precursory signs may be used to forecast a failure time or improve monitoring (Federico et al., 2012; Fukuzono, 1985; Intrieri et al., 2019; Wegmann et al., 2003). In many cases the nature and magnitude of these precursory signs precludes their detection in the absence of sensitive equipment. In situ observations can be sensitive to even small changes in slope properties and are therefore valuable for forecasting a destabilization (Sättele et al., 2015; Stähli et al., 2015). However, ground-based observations have important limitations: (i) prior knowledge of a potential slope instability is required in order for the correct instrumentation to be installed in the right locations; (ii) the landslide source region may be inaccessible, preventing the installation of in situ monitoring equipment; (iii) monitoring systems can be prohibitively expensive and require highly specialized expertise for data evaluation; and (iv) the area that can be monitored is generally limited to individual hillslopes. Altogether, ground-based monitoring techniques are useful for landslide monitoring in many cases but are insufficient for monitoring large regions or where a priori knowledge is lacking.

An increase in satellite data availability and resolution has promoted remote sensing as an alternative or complementary landslide detection and monitoring tool (e.g., Kirschbaum et al., 2019; Dille et al., 2021). Satellite remote sensing may lack the precision of some ground-based monitoring techniques, but it can provide a low-cost (for the end user) and easily accessible way to monitor vast and inaccessible terrain at daily to weekly revisit times and 0.3 to 30 m spatial resolution. Qualitative visual analysis of satellite imagery allows for the rapid identification of surface changes that may be associated with slope instabilities or the initiation of landslide motion. Further quantitative processing of satellite imagery enables the monitoring of horizontal and vertical land motions – for example via feature tracking or the stereographic generation of digital elevation models (DEMs; Shean et al., 2020; Dai et al., 2020a; Dille et al., 2021). Interferometric synthetic-aperture radar (InSAR) can provide millimeter-to centimeter-resolution line of sight displacements (e.g., Handwerger et al., 2019; Jacquemart and Tiampo, 2021; Manconi et al., 2018). Growing archives of high-resolution, open-access Earth observation data remain largely untapped for landslide monitoring. In this study we use the data-rich 7 February 2021 Chamoli rock–ice avalanche as a case study for the remote identification of landslide precursory signs.

We first introduce landslide hazards in the Himalaya with a specific focus on the Chamoli event and then offer a general overview of remote sensing of slope instabilities. Next, we explain the methods used in the current study and present and discuss the results.

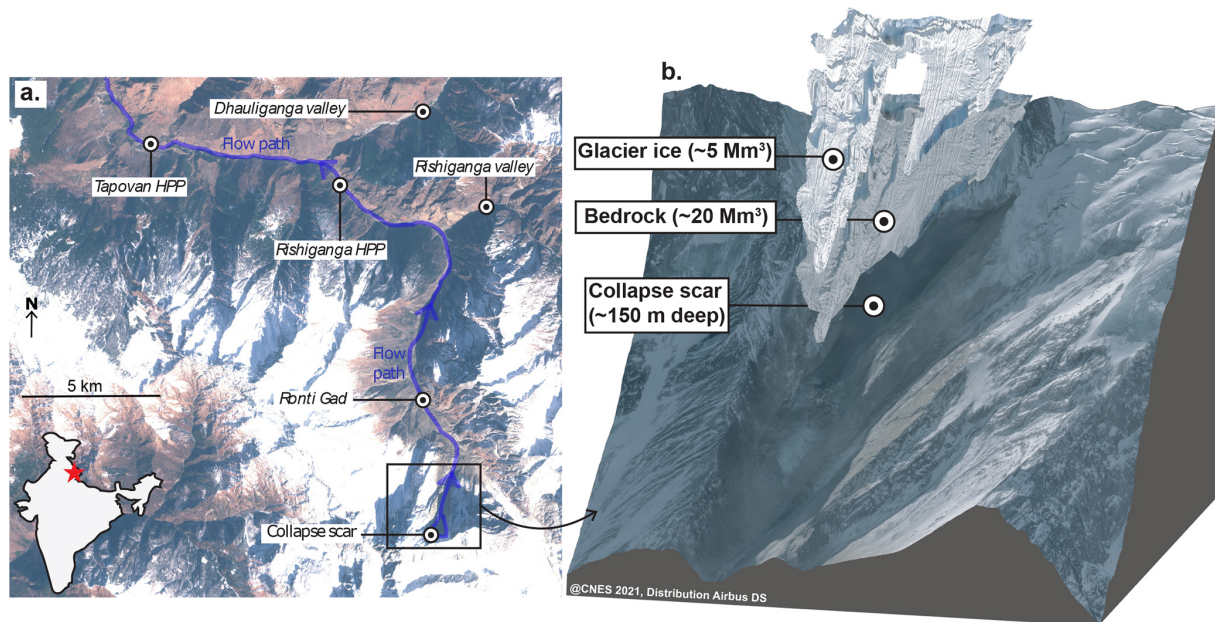
## 1.2 Landslide hazard and risk in the Himalaya

Landslides occur in high-mountain areas all over the world, and the risk is greatest in areas where high topographic relief intersects with high population densities or infrastructure – which is the case across much of the Himalayan region. Over 50 million people live directly within the Himalaya, with a further 700 million living within associated watersheds (Dimri et al., 2019). A combination of extreme topographic relief, regular tectonic activity, high seasonal rainfall intensities, and steep slopes make the Himalaya particularly susceptible to landslides (Kirschbaum et al., 2019).

In recent decades, several factors have contributed to raising landslide risk across the region: first, climatic warming has driven rapid thinning and retreat of Himalayan glaciers – which are currently losing over 10 Gt of mass per year (e.g., Kääb et al., 2012; Brun et al., 2017; Shean et al., 2020; Jakob et al., 2021; Hugonnet et al., 2021). Glacier retreat may contribute to a range of factors conducive to landslides, including a reduction in slope buttressing and an increase in meltwater availability (Holm et al., 2004; Fischer et al., 2006; Huggel et al., 2012; Kos et al., 2016; Coe et al., 2018; Dai et al., 2020a; Glueer et al., 2020). In addition to glacier retreat, permafrost degradation has also been documented to reduce slope stability (Gruber and Haeblerli, 2007; Allen et al., 2011; Fischer et al., 2012; Krautblatter et al., 2013; Haeblerli et al., 2017; Magnin et al., 2019; IPCC, 2019; Patton et al., 2019; Deline et al., 2021). Second, increasing populations, economic growth, and infrastructure development in high-mountain valleys have greatly expanded the potential consequences of landslides. This second point is apparent for the Chamoli disaster, in which the majority of deaths occurred at hydropower plants that were recently built or were under construction (Shugar et al., 2021). Other factors, including changes in precipitation patterns (e.g., Li et al., 2018; Kirschbaum et al., 2020) and land use (Cummins, 2019) may also contribute to evolving landslide hazard potential and associated risk.

## 1.3 The 2021 Chamoli hazard cascade

During the morning of 7 February 2021, a  $26.9 \times 10^6 \text{ m}^3$  wedge of rock and ice detached from the north face of the peak of Ronti, a 5500 m elevation summit in the Uttarakhand Himalaya (Fig. 1). This wedge then dropped around 1800 m to the Ronti Gad valley floor, where it continued down valley towards the Rishiganga and Dhauliganga rivers and transformed into a debris flow (Shugar et al., 2021; Cook et al., 2021). The collapse block was composed of approx-



**Figure 1.** The 7 February Chamoli rock–ice avalanche. **(a)** The path of the collapse, along with key locations (HPP refers to hydropower plant) annotated on a 10 February 2021 Sentinel-2 image. **(b)** Three-dimensional visualization showing the post-collapse scar with reconstruction of the overlying bedrock and glacier ice. © CNES 2021 (Centre national d'études spatiales), distribution Airbus DS (Defence and Space).

imately 80 % bedrock and 20 % glacier ice (Shugar et al., 2021). Frictional heat generation calculations suggest that most or almost all of the glacier ice melted during the 3400 m drop from the source to Tapovan hydropower station (Shugar et al., 2021). This melting of the ice, combined with major sediment deposition at the confluence of the Ronti Gad and Rishiganga, increased the initial rock–ice avalanche's water content and converted it into a highly mobile debris flow which left 204 people missing or killed and destroying two hydropower stations.

## 1.4 Remote-sensing techniques

### 1.4.1 Feature tracking

Optical feature tracking is a versatile technique which can be used to track surface motion by evaluating the relative position of features or patterns in repeat imagery. Feature tracking has been applied to a variety of problems, including tracking post-seismic ground deformation (e.g., Leprince et al., 2007), quantifying glacier flow velocities (e.g., Bind-schadler and Scambos, 1991; Heid and Käab, 2012; Millan et al., 2019; Van Wyk de Vries and Wickert, 2021), or measuring landslide displacements (e.g., Behling et al., 2014; Aryal et al., 2012; Lucieer et al., 2014; Peppia et al., 2017; Manconi et al., 2018; Darvishi et al., 2018; Jia et al., 2020; Dai et al., 2020a; Dille et al., 2021). The accuracy of feature tracking is limited by the spatial resolution of the imagery and the magnitude of displacements: in best-case scenarios displacement maps may reach a precision of  $\sim 0.1$  pixels

(Leprince et al., 2007; Dille et al., 2021). For openly available Sentinel-2 imagery, this means that displacements of less than 1 m cannot be detected from a single image pair. Commercial imagery datasets are available with higher spatial resolution (e.g., 3 m for Planet Dove CubeSat, 0.3 m for WorldView), which may resolve smaller displacements (e.g., Stumpf et al., 2014), but are expensive to procure and process and do not offer systematic repeat coverage. Similarly, displacements may be mapped from high-resolution UAV imagery (Peppia et al., 2017), but this requires prior knowledge of the hazardous area and targeted acquisition campaigns.

### 1.4.2 InSAR

Satellite-based InSAR is a powerful tool for detecting small changes at Earth's surface from space. It has been widely used to quantify ground displacements caused by processes such as earthquakes (e.g., Massonnet et al., 1993; Barba-Sevilla et al., 2018), groundwater extraction (e.g., Samsonov and d'Oreye, 2017; Motagh et al., 2017), volcanic unrest (e.g., Rosen et al., 1996; Tiampo et al., 2017), or landslides (e.g., Manconi et al., 2018; Handwerger et al., 2019; Dai et al., 2020b; Mondini et al., 2021; Jacquemart and Tiampo, 2021). By measuring the shift of the radar phase relative to earlier measurements of the same features, InSAR can provide measurements of ground deformation at millimeter and centimeter scales. Active radar sensors can image Earth's surface through clouds and darkness, a major advantage over passive optical sensors particularly during the monsoon season (e.g., Massonnet and Feigl, 1998). However, leveraging

InSAR data for the detection and assessment of mass movements is challenging. The oblique viewing geometry of radar satellites means that radar data can be rendered useless in areas of steep topography due to the effects of shadowing, foreshortening, and layover (Massonnet and Feigl, 1998; Wasowski and Bovenga, 2014). In the case of rapid displacements that surpass the phase-aliasing thresholds or dramatic changes in the surface cover or geometry, a loss of interferometric coherence can prohibit the quantification of (the full) ground deformation (Manconi, 2021). Despite these drawbacks, many studies have shown that InSAR can be successfully applied to assess the stability of slopes even in high-relief terrain (e.g., Manconi et al., 2018; Handwerger et al., 2019; Bekaert et al., 2020; Jacquemart and Tiampo, 2021).

### 1.4.3 Stereo-DEM generation

Stereo-DEM generation uses two or more overlapping optical images that were acquired at the same time from different viewing angles to reconstruct surface topography. Photogrammetric principles can then be used to derive DEMs from these images. These approaches can now be used to generate detailed DEM products over large spatial areas using very high-resolution satellite stereo imagery (e.g., Korona et al., 2009; Morin et al., 2016; Shean et al., 2016; Porter et al., 2018; Howat et al., 2019). However, most very high-resolution imagery is not open source and is expensive to procure, limiting its use. Increased availability of this commercial imagery and/or new open-source stereo-imagery satellites would provide many new opportunities for hazard monitoring. Repeat DEMs obtained at different time periods can provide precise estimates of surface elevation change associated with many processes, including glacier change (e.g., Brun et al., 2017; Willis et al., 2018; Zheng et al., 2019; Shean et al., 2020), snow accumulation/melt (e.g., Deschamps-Berger et al., 2020; McGrath et al., 2019; Bhushan et al., 2021), volcanic deformation (e.g., Bisson et al., 2021; Schaefer et al., 2012), and landslide or debris flow events (e.g., van Westen and Lulie Getahun, 2003). In particular for landslide research, previous studies have used DEM-differencing techniques to identify geomorphic changes (e.g., Corsa et al., 2022) and precursory motion (e.g., Higman et al., 2018; Dai et al., 2020a) on timescales ranging from years (e.g., Shugar et al., 2021; Geertsema et al., 2022) to decades (e.g., Higman et al., 2018; Lacroix et al., 2020).

## 1.5 Objectives

The objective of this study is to use the 7 February 2021 Chamoli rock–ice avalanche as a detailed case study to assess the potential and limitations of satellite-based landslide monitoring. We first assess the pre-collapse conditions of the 7 February 2021 Chamoli rock–ice avalanche and then inter-

pret them in a broader context. Our research question is the following.

Can the pre-collapse remotely sensed datasets be used automatically to identify the location or timing of the 7 February 2021 Chamoli rock–ice avalanche, and is it possible for these monitoring techniques to be upscaled for hazard monitoring on a regional or global scale?

## 2 Methods

We used a range of datasets and processing workflows to investigate the pre-collapse conditions of the Chamoli rock–ice avalanche:

1. optical satellite imagery (Landsat and Sentinel-2) was used to investigate visible changes in the collapse region over the years to decades prior to the rock–ice avalanche;
2. feature tracking of optical satellite imagery (Sentinel-2, Planet, Cartosat-1, and SPOT 7) was used to derive horizontal displacements;
3. Sentinel-1 C-band radar imagery was used to calculate interferometric synthetic-aperture radar (InSAR) displacement maps;
4. digital elevation models (DEMs) from optical satellite stereo imagery (WorldView-1/2/3, GeoEye-1, Pléiades-HR (high resolution), SPOT 7, and Cartosat-1) were used to derive vertical changes.

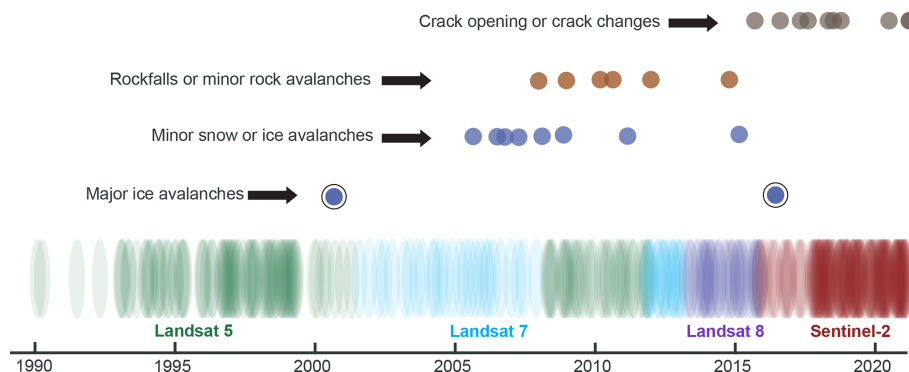
### 2.1 Qualitative observations of slope change

We investigated 3 decades of pre-collapse optical satellite imagery to gain a preliminary understanding of pre-landslide changes. We documented changes in the north-facing slope of the peak of Ronti, which sourced the February 2021 rock–ice avalanche, using all available data from Landsat 5 (TM, Thematic Mapper), Landsat 7 (ETM+, Enhanced Thematic Mapper Plus), Landsat 8 (OLI, Operational Land Imager), and Sentinel-2 with a cloud cover of less than 60%. We focused our observations on surface changes, including deformation and fracturing, and rock or ice avalanches originating from the collapsed block or surrounding area.

Our ability to detect change is limited by the spatial resolution of the imagery used (15–30 m for Landsat and 10 m for Sentinel-2). We examined a 31-year (1990–2021) time series of satellite imagery (Fig. 2), including 122 Landsat 5 images, 43 Landsat 7 images, 34 Landsat 8 images, and 155 Sentinel-2 images. A full list of images is provided in the Supplement, along with a brief description of any anomalous features.

#### 2.1.1 Optical feature tracking

We used feature tracking with a range of medium- (10 m) to high-resolution (2.5 m) satellite imagery to evaluate the



**Figure 2.** Timeline of images analyzed for change at the Chamoli site prior to the 7 February 2021 collapse, with major events or changes seen over this period. The first image (6 February 1990) was taken 31 years before the collapse, and the last image (5 February 2021) was taken 2 d before the collapse.

pre-collapse motion of the north slope of the peak of Ronti. We used two different feature-tracking toolboxes: GIV for rapid processing of Sentinel-2 data (Glacier Image Velocimetry; Van Wyk de Vries and Wickert, 2021) and autoRIFT (Autonomous Repeat Image Feature Tracking; Lei et al., 2021) for processing Planet data as part of a pipeline also including orthorectification of imagery. Both GIV and autoRIFT are based on three core components: a pre-processing module which applies one or more filters (described below) to images to enhance distinct surface features for tracking, a multipass two-dimensional image correlator, and a post-processing module to identify and filter erroneous displacement values (Van Wyk de Vries and Wickert, 2021; Lei et al., 2021). The GIV toolbox performs image cross correlation in the frequency domain, while autoRIFT performs the cross correlation in the spatial domain. Using GIV, we pre-processed the imagery using an orientation filter and ran the cross correlation with a reducing window size from 20 to 5 pixels and a window overlap of 50 %. In autoRIFT we pre-processed the imagery with a Laplacian filter and used adaptive window sizes between 32 and 64 pixels with a skip rate of 8 pixels for the cross correlation.

We calculated velocities using all available Sentinel-2 images through February 2021, excluding any images with a local cloud cover greater than 60 % (based on the L1-C QA – Level-1C Quality Assessment – band cloud mask). A total of 155 images were available, for a total of 5237 image pairs with a time separation between 50 and 500 d. We processed these image pairs using GIV. We also resampled the velocity time series to monthly resolution (see Fig. 4c–g) using a weighted-averaging scheme described in Van Wyk de Vries and Wickert (2021), with weights based on the proportion of the time period within a given month.

We also downloaded all PlanetScope Dove Classic (four-band) Level-1B imagery with less than 20 % cloud cover acquired between January 2020 and January 2021. We processed 4701 image pairs using autoRIFT with a time separation of 100 to 350 d. The near-infrared (NIR) band from

the L1-B images was orthorectified on the 2015 pre-event reference DEM (Bhushan and Shean, 2021), and the systematic median offset (computed over static, non-glacierized surfaces) was removed from each pairwise surface displacement map in both east–west and north–south directions. Despite the higher product resolution (3 vs. 10 m Sentinel-2 images) and use of a high-resolution DEM for improved orthorectification, the Planet velocity maps had a high random background noise. We attribute this to spurious correlation over surfaces with varying shadow cover due to steep slopes and changing illumination, as the images were captured by different satellites during different times of the day/year. To compensate for this higher background noise, we chose a higher minimum temporal separation (100 d, compared to 50 d for Sentinel-2) between Planet image pairs when calculating time-averaged velocity maps. We also calculated displacements (using both GIV and autoRIFT) on one pair of high-resolution Cartosat-1 images (October 2017 to November 2018).

We used these velocity data to evaluate whether the collapsed block moved prior to collapse – with a null hypothesis that the block moved no more than the surrounding “stable” (non-glacierized) bedrock. We tracked the motion of a medial bedrock ridge (Fig. 1) to measure the motion of the underlying bedrock, which is independent from the flow of the overlying glaciers. We divided the collapse block into three different regions alongside a zone of stable ground and create a time series of average displacement for each zone.

## 2.2 InSAR maps

We processed Sentinel-1 data from ascending and descending orbit tracks 56 and 63, respectively, to investigate whether the precursory motion of the collapse block could have been detected in radar interferograms. All radar data were downloaded from the Alaska Satellite Facility Distributed Active Archive Center (ASF DAAC). Because the descending track is heavily affected by layover artifacts, we

only performed the full processing with data from the ascending orbit. We processed the data with the InSAR Scientific Computing Environment (ISCE; Rosen et al., 2012), removed the topographic phase using the 2015 pre-event WorldView DEM (Bhushan and Shean, 2021) composite (resampled to 8 m), and masked out all pixels with an interferometric coherence of less than 0.3. Single-look complex (SLC) images were multi-looked to one and three looks in azimuth and range, respectively. We generated 108 interferograms covering the period of January 2017 to November 2020, each spanning 12 d. We manually selected the best interferograms and performed unwrapping with the Statistical-Cost, Network-Flow Algorithm for Phase Unwrapping (SNAPHU; Chen and Zebker, 2002).

### 2.3 DEM generation

We produced multiple pre-event and post-event DEM products from very high-resolution (Maxar–DigitalGlobe WorldView-1/2/3, GeoEye-1, and Airbus–CNES Pléiades; 0.3 to 0.5 m GSD, ground sample distance) and high-resolution (Airbus SPOT 7 and ISRO Cartosat-1, 1.5 m to 2.5 m GSD; Indian Space Research Organisation) satellite imagery captured between 2015 and February 2021. The DEM products were used to calculate the vertical motion of the collapse block from 2015 to 10 February 2021.

We used the NASA Ames Stereo Pipeline (ASP; Shean et al., 2016; Beyer et al., 2018) to process all of the images. For this particular study, we primarily used four products spanning two time periods: the 2015 pre-event WorldView DEM composite (Bhushan and Shean, 2021); an intermediate period, a 2018 pre-event DEM composite produced by averaging the November 2018 Cartosat-1 (Appendix A1) and December 2018 SPOT 7 (Appendix A2) DEMs; and the 10–11 February 2021 post-event composite DEM derived from Pléiades and WorldView–GeoEye stereo imagery (Shean et al., 2021). We calculated the difference between three composite DEM products to create 2015–2018, 2015–2021, and 2018–2021 DEMs of difference (DoDs). The 2015–2018 DoD provides insight into vertical changes in the hillslope prior to failure, while the 2015–2021 DoD provides the volume and geometry of the collapsed block. We calculated an empirical uncertainty estimate for each DoD using the tiling method (Berthier et al., 2016; Miles et al., 2018; Jacquemart et al., 2020).

## 3 Results

### 3.1 Qualitative observations of slope change

We identified four main types of processes in our 31-year optical satellite image time series.

1. *Major ice avalanches (January–April 2000 and September–October 2016)*. These large-volume ice

avalanches originated from the steep hanging glacier to the west of the collapse block and temporarily filled Ronti Gad with ice, snow, and sediment.

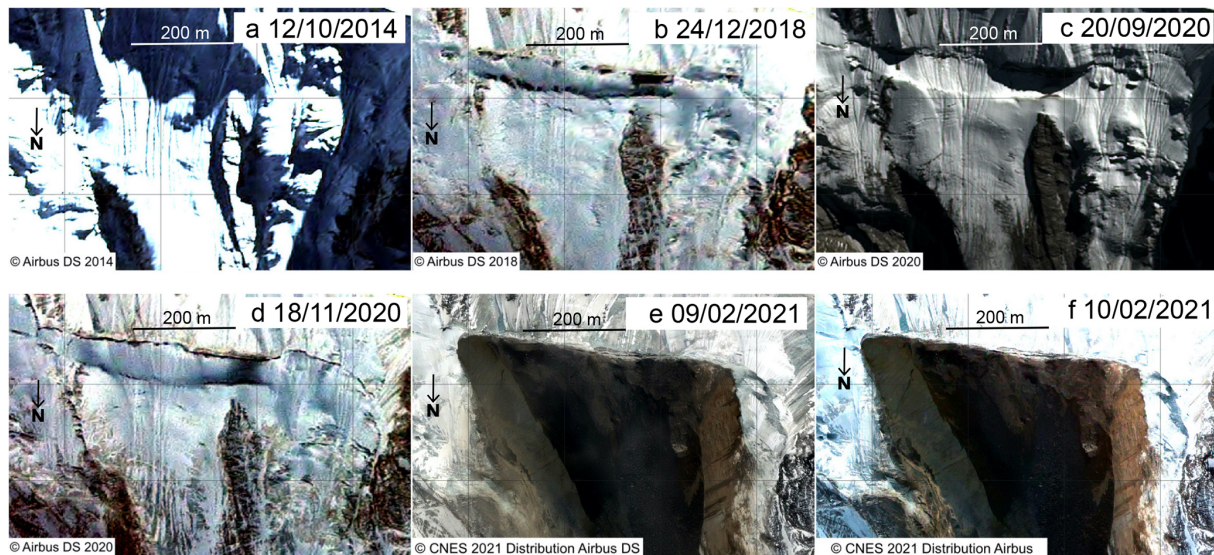
2. *Minor snow or ice avalanches (2005, 2006, 2007, 2008, 2012, and 2015)*. These smaller-volume avalanches may have originated from either the adjacent hanging glacier or the seasonal snowpack and did not appear to infill the underlying valley with any significant quantity of material (with the exception of one ~ 500 m long snow–ice deposit in May 2006).
3. *Minor landslides avalanches (2007, 2009, 2011, 2012, 2013, and 2015)*. These minor rockfalls or rock avalanches originated from the peak of Ronti or as the unconsolidated sediment on the flanks of Ronti Gad and also do not appear to have deposited major volumes of sediment.
4. *Opening and widening of cracks at the headwall of the collapse block (2016–2021)*. This is the gradual opening of a wide crack in the north face of the peak of Ronti.

We only interpret the fourth process type (crack opening) as a real sign of pre-collapse conditions. Minor rockfalls and snow–ice avalanches are a common feature of high-relief, high-slope active landscapes (e.g., Petley, 2012; Huggel et al., 2012; Vincent et al., 2015; Kirschbaum et al., 2020). The major ice avalanches represent a serious geohazard in the upper Ronti Gad but appear to relate to internal dynamics of the western hanging glacier rather than instability in the underlying bedrock. The area of these 2000 and 2016 major ice avalanches was estimated at 0.16 and 0.2 km<sup>2</sup>, with melting and/or redistribution of the resulting valley floor deposits within 3 years of the event (Shugar et al., 2021; Supplement Sect. 3.1). Regular large ice avalanches have been observed at many other hanging glaciers in active, high-mountain environments (e.g., Faillettaz et al., 2008; Vincent et al., 2015).

The conspicuous crack at the headwall of the failed block was first visible in optical imagery in March 2016, although its location roughly aligns with a pre-existing glacier crevasse – suggesting that a minor crack opening in the bedrock may have preceded this date. The crack grew until the end of 2018 and appeared to become infilled with snow over the course of 2019 and 2020 based on the amount of exposed bedrock on the rock walls. The crack widened further between 2018 and the 7 February 2021 collapse to a total width of 50–70 m but less rapidly than the opening in 2016–2018. We confirmed our observations of the crack opening with several very high-resolution (~ 0.5 m) images (Fig. 3).

### 3.2 Optical feature tracking

Feature tracking provides the most complete spatiotemporal assessment of displacement of the methods used in this study – with data coverage from late 2015 until early 2021. We used results from the Cartosat-1 image pair and the Planet



**Figure 3.** Time series of the headwall crack opening in high-resolution optical images from SPOT 7 and Pléiades-HR. Please note that the date format used in this figure is day/month/year.

archive for validation of the Sentinel-2 displacements. In all three cases, the collapsed block (most notably, the bedrock ridge at the center of this block) exhibited displacements exceeding the background noise level on stable bedrock ( $< 1 \text{ m yr}^{-1}$ ).

The horizontal velocity of the collapsed block ranged from around  $5$  to  $20 \text{ m yr}^{-1}$ , with the most rapid motion occurring in the summers of 2017 and 2018 (see Fig. 4d–g). We do not observe an increase in velocity of the collapsed block immediately prior to its failure in February 2021. The Sentinel-2 image record includes seven cloud-free images from early 2021, including one image taken 2 d prior to the collapse; therefore this lack of speedup is unlikely the result of a temporal data gap. Periods with the highest block velocity (2017–2018) correspond to periods of the greatest increase in headwall crack width – particularly the summers of 2017 and 2018. This is consistent with motion occurring on the entire collapsed block, rather than only on the glaciers or a superficial layer of rock.

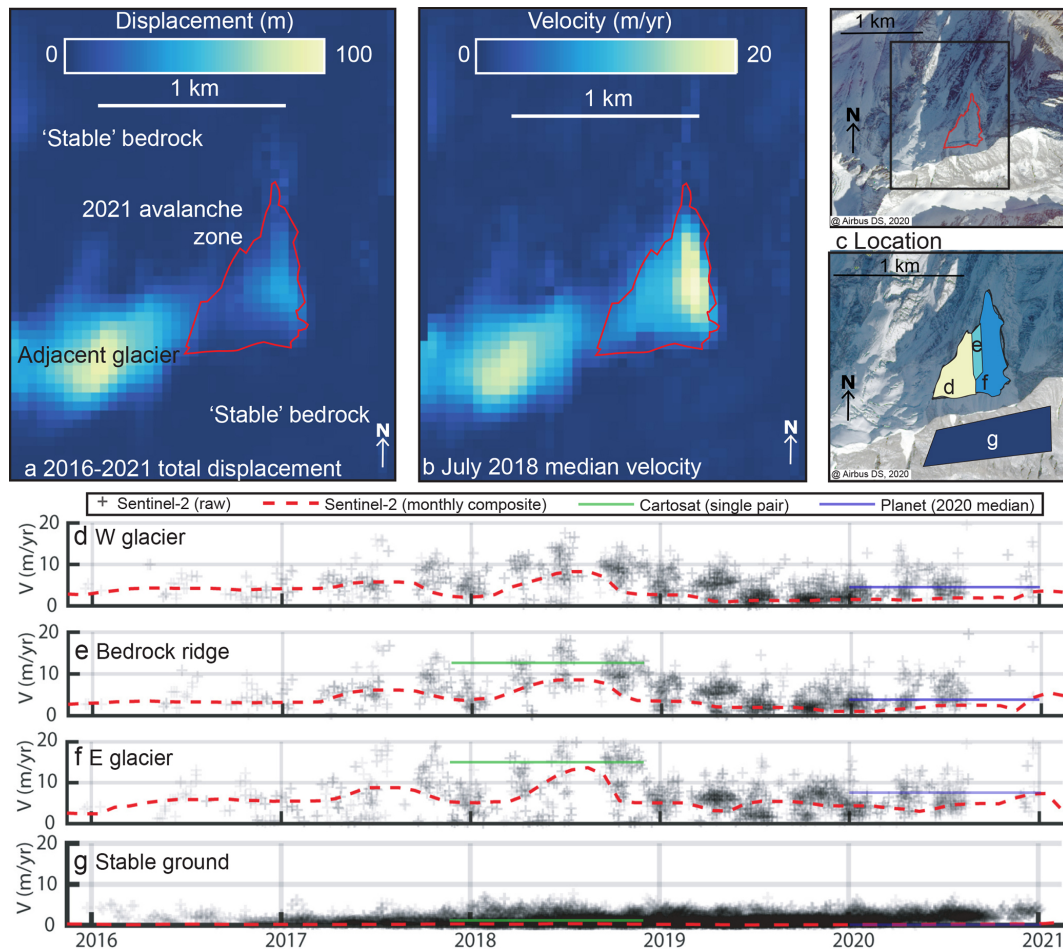
Total 2016–2021 horizontal displacements were  $\sim 20$ – $30 \text{ m}$  (Fig. 4a), which is similar in magnitude to the width of the crack as measured directly from Sentinel-2 imagery. Due to the steep topography (mean slope of  $42.6^\circ$ ), the visible horizontal motion does not account for all of the true deformation. After correcting for the viewing angle, the total block motion is  $25$ – $40 \text{ m}$ . Overall, the feature-tracking results demonstrate that the collapse block was mobile several years prior to its collapse in 2021.

### 3.3 InSAR maps

Even with knowledge of the location of the failed block, the processed interferograms do not allow for a pre-collapse

identification of the instability on the peak of Ronti. Of the 108 available interferograms, roughly half exhibited a complete loss of coherence, largely due to snow cover (November through May). Good-quality interferograms are limited to summer months, and on the collapse block, coherence is only retained on the ice-free part at the bottom of the wedge. The upper, glacier-covered part of the collapse block remains decorrelated, likely due to shadowing and glacier–snow cover. Figure 5a highlights the very low radar backscatter in this zone, and Fig. 5b and c confirm the spatial agreement between the loss of coherence and glacier cover. Data gaps lower in the valley are also related to loss of coherence, possibly due to vegetation cover or moisture variability. Many interferograms are characterized by high amounts of noise, likely from variable atmospheric properties or other artifacts.

A high-quality interferogram from July 2020 (Fig. 5b) does not indicate any motion on the lower part of the collapse block in the summer prior to the failure, but it is impossible to determine whether this is consistent in other interferograms of that year due to high noise levels. Areas of high coherence are small and discontinuous on the collapse block, making it hard to determine any changes. In less steep terrain northwest of the collapse block, the motion of a rock glacier (on the order of centimeters per year) can be detected consistently in all interferograms that remain coherent over that part of the image (Fig. 5b). This highlights that the reliability and information content of InSAR velocity maps can be highly variable even across a small study area. Despite its sensitivity, InSAR is not able to provide any conclusive information about the pre-failure conditions of the collapse block in this challenging terrain.



**Figure 4.** Surface displacement and horizontal velocity from optical-image feature tracking. **(a)** The total displacement over the entire Sentinel-2 era. **(b)** A snapshot velocity during an episode of rapid displacement in summer 2018. **(d)–(g)** Time series of velocity averaged across specific zones shown in panel **(c)**. Note the episodes of rapid displacement in 2017–2018 relative to 2016 or 2020, corroborated by the Cartosat-1- and Planet-derived velocities.

### 3.4 DEM analysis

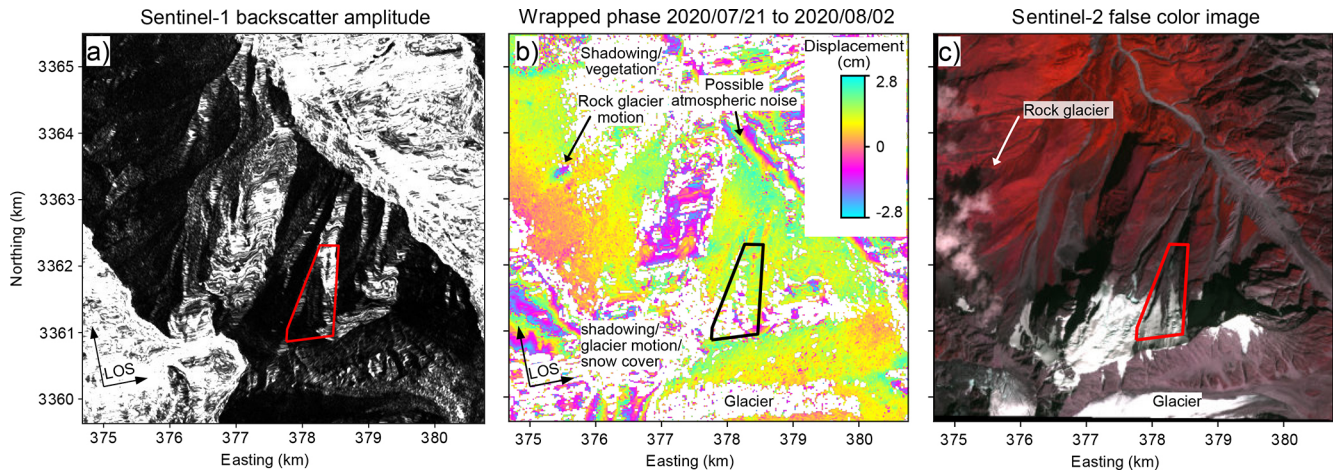
We calculated the geometry of the collapsed block, equal to the zone of negative elevation change in the 2018–2021 DoD (Fig. 6c; volume:  $26.9 \times 10^6 \text{ m}^3$ , 95 % confidence interval:  $26.5\text{--}27.3 \times 10^6 \text{ m}^3$ ; Shugar et al., 2021). The earlier DoD (2015–2018) shows a very different pattern (Fig. 6b), with a  $\sim 100 \text{ m}$  wide zone of elevation loss at the upper altitude limit of the collapsed block (“headwall crack”) and a broad zone of elevation gain over the remainder of the block (“bulge”). The magnitude of this pre-collapse elevation loss is greatest in the central and western portion of the headwall crack, while the elevation gain is most pronounced on the central and eastern portions of the bulge.

DEM analysis further confirms the results from direct image observations and feature tracking – large changes occurred on the collapsed block prior to its collapse. The zone of negative elevation change is wider than the crack as directly observed in optical imagery, which may result from

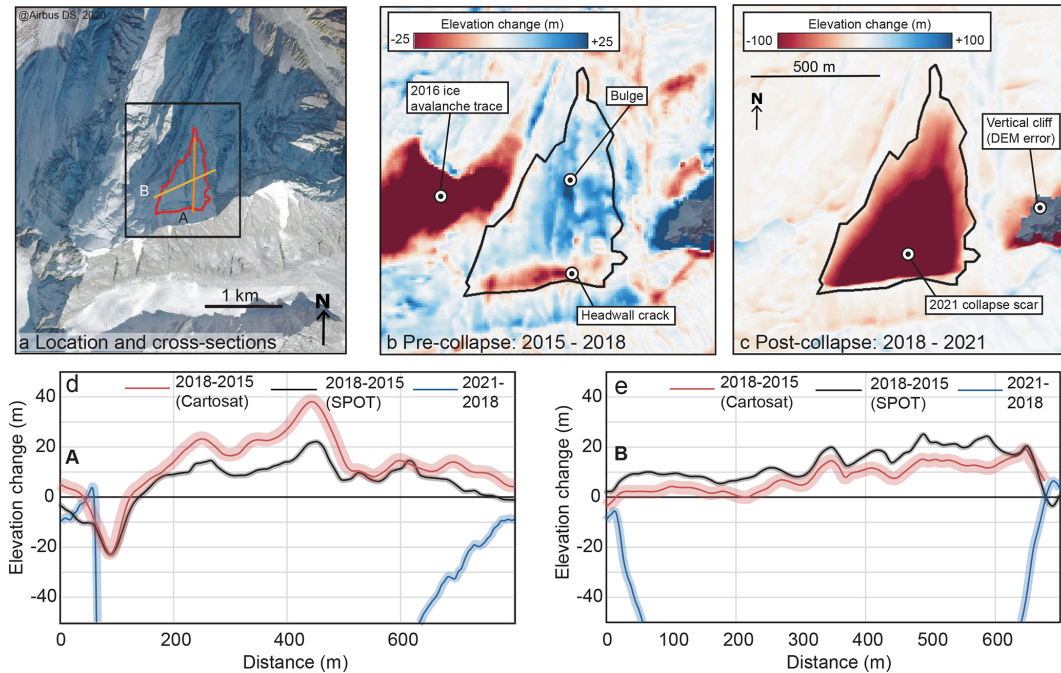
limits in the DEM resolution or partial collapse of the surrounding rock or ice into the crack.

## 4 Discussion

The pre-collapse motion of the avalanche block raises important questions about the causes and timing of the 7 February 2021 Chamoli rock–ice avalanche. In this section, we use this disaster as a case study to discuss the potential and limitations of satellite data for remote hazard monitoring. Furthermore, we explore the conditions of the collapse using our multi-dataset observations and evaluate whether the location or timing of the collapse could have been identified beforehand.



**Figure 5.** (a) Sentinel-1 radar backscatter amplitude from ascending orbit track 56. (b) Wrapped-phase ( $0\pi$  to  $2\pi$  or 2.8 cm) interferogram from July 2020. (c) Corresponding false-color image. Large areas of low coherence (masked as white) and patchy coverage illustrate the complexities of InSAR monitoring in high-alpine terrain. The avalanche block is outlined in black and red. Image: Sentinel-2. Please note that the date format used in this figure is year/month/day.



**Figure 6.** Elevation change of the avalanche zone pre- and post-collapse. (a) SPOT 7 true-color composite image from September 2020 with the location of cross-sections A and B for context. (b, c) DoD maps at different times. (d, e) Two cross-section profiles (a) across the DoD. The composite 2018 DEM was used for panel (b). Cross-sectional uncertainties are assigned for an area equal to the length of the section line multiplied by the pixel size. Note that the elevation change bounds are different in panels (b) and (c).

#### 4.1 Future perspectives: remote-sensing-based hazard monitoring

Our work on the Chamoli avalanche took place after the collapse, with the full knowledge of the position of the avalanche source. This work is useful for better understanding the conditions of the slope collapse. However, to be di-

rectly useful for hazard monitoring and prevention, these techniques must identify avalanche locations and sizes before – rather than after – they occur. The key questions therefore remain: would it have been possible to identify the Chamoli landslide prior to its collapse using the methods used in our study, and can these methods be applied elsewhere to identify future failures?

The qualitative analyses of optical satellite images, feature-tracking results, and DEM analysis all indicate that precursory signs of slope failure were detectable. Satellite images show a crack growing over the 5 years prior to failure (Fig. 3), feature tracking reveals tens of meters of horizontal displacement of the collapsed block, and DEM differences show tens of meters of vertical elevation change over the collapsed block. Combining this information with background knowledge about the district of Chamoli, such as the extreme relief, steep slopes, and historic avalanches, it would in principle have been possible to identify this as an unstable slope with high collapse potential.

While the data are sufficient to identify precursory signs of this rock–ice avalanche after the collapse has already occurred, there are important limitations to their use for automated hazard monitoring and pre-collapse detection of unstable slopes. The first key limitation is the very low signal-to-noise ratio of feature-tracking horizontal velocity maps, DEM-derived elevation change maps, and InSAR velocity maps in the steep terrain most susceptible to slope failure. For feature tracking, the noise level of the composite 2016–2021 mean velocity maps is low ( $< 1 \text{ m yr}^{-1}$ ). However, the background noise level (as evaluated over stable bedrock) of individual velocity maps is much higher – and in some cases comparable to the magnitude of the signal ( $\sim 5\text{--}20 \text{ m yr}^{-1}$ ). For the DEMs, artifacts range from meters to tens of meters in scale, and additional “noise” is introduced by real elevation changes from glacier and snowpack change (Fig. 7b). While these issues with false positives can be mitigated, this is challenging without knowing the signal of interest.

InSAR, while also being susceptible to false positives, is additionally prone to false negatives. The north-facing aspect of the peak of Ronti provides a twofold challenge: the illumination of the slope is limited (low backscatter, Fig. 5a), and any motion – assuming it is largely in the direction of the steepest slope – is oriented in the direction in which the radar instrument is least sensitive. Additionally, the non-glacierized area of the collapse wedge is small, making it challenging to identify fringe patterns amongst the noise. Furthermore, with the largest velocities reaching tens of meters per year, the InSAR measurements are prone to phase aliasing and underestimation of the true displacement. Sentinel-1 InSAR would not have provided an adequate tool for monitoring in this case, even with knowledge of the location of the instability.

The second key limitation is that none of the datasets produced in this work could predict the timing of collapse. While most methods pick up precursory signs of slope failure, these begin almost 5 years prior to eventual collapse. The largest measured changes did not occur immediately prior to failure but rather preceded failure by around 3 years. Mao et al. (2022) propose that abrupt growth of the “summit” crack (which they term bergschrund) and lateral cracks is visible in the 5 February 2021 Sentinel-2 image (2 d prior to collapse). This summit crack was, however, already promi-

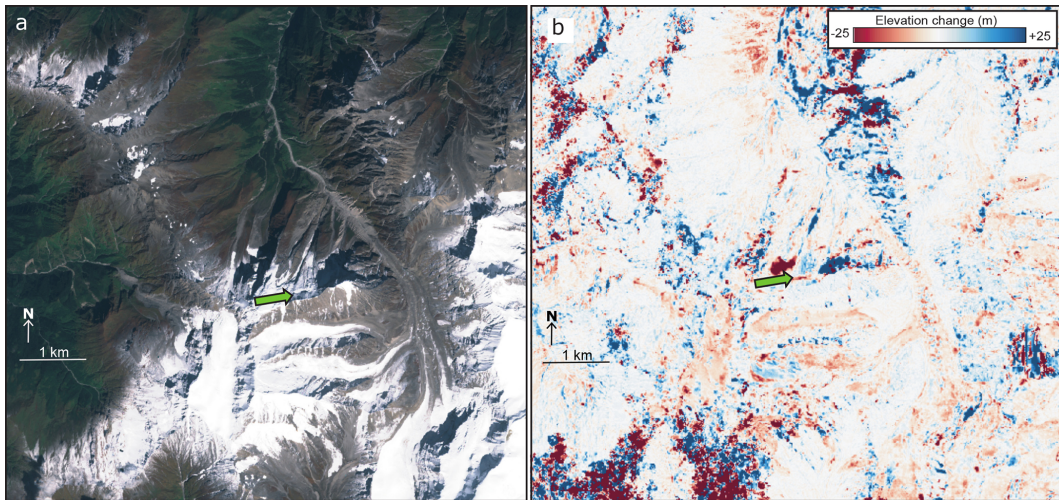
nently visible and growing over the 5 years prior to collapse (e.g., Fig. 2), and the proposed lateral cracks do not precisely correspond to the margins of the eventual collapse zone. In addition, the proposed crack growth episode visible was not associated with any detectable peak in displacement (Fig. 4). Even with the knowledge that the collapse occurred on 7 February 2021, signs pointing to an imminent collapse in remote-sensing data from late 2020 or early 2021 are highly ambiguous. Tiwari et al. (2022) show that seismic signals were detectable from the Chamoli block as early as 2.5 h prior to collapse. While they do not explore how easily detectable the pre-collapse signals would have been without prior knowledge of the event, they do note that the  $\sim 15$  min between initial collapse and impact of the Tapovan dam would have enabled evacuation if a local early-warning system had been in place.

One final limitation is related to the immense size of hazardous areas relative to the scale of hazards themselves. The Chamoli collapsed block had an area of around  $0.25 \text{ km}^2$ , while the Himalaya cover over  $0.5 \times 10^6 \text{ km}^2$ . Any methods aimed at automatically detecting hazards prior to their occurrence must have a low “false-positive” (identified as a hazard in the database but not of real concern) rate or any resulting database will be populated primarily with these incorrectly flagged regions. This becomes a major challenge when considering the high incidence of noise and artifacts in feature-tracking-derived displacement or DoD maps (e.g., Fig. 7). In addition, any hazard detection methods involving manual intervention, for instance an expert assessment of crack growth across a time series of optical satellite images, is not feasible on a large scale and will be limited to previously identified zones of high hazard. In addition, the DEMs and elevation change maps used in this study were generated from imagery not accessible in open-source archives. Changes in the accessibility of commercial data or the launch of new, open-access, stereo-imagery satellites would facilitate the use of elevation change in large-scale geohazard monitoring.

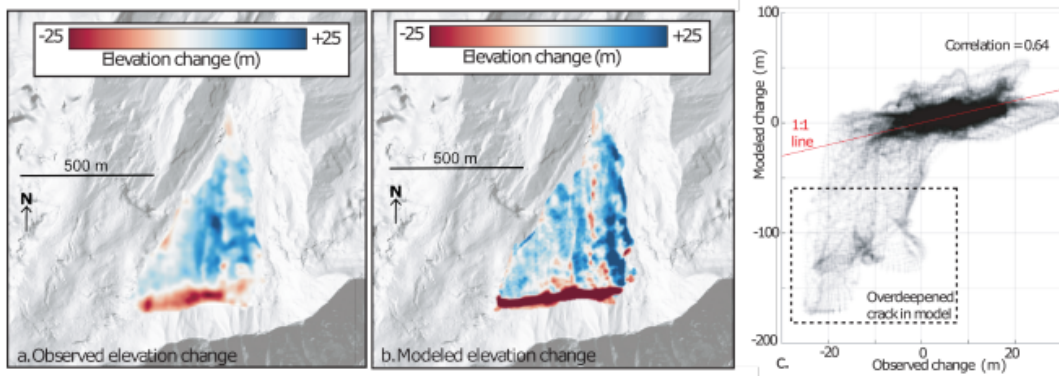
## 4.2 Three-dimensional block motion

We examined the three-dimensional motion of the collapse block as a first step towards understanding the Chamoli rock–ice avalanche collapse mechanism(s). Rotation and translation are the two primary modes of landslide motion (e.g., Záruba and Mencl, 2014), with each having a distinct surface displacement pattern. We used a combination of horizontal displacement (feature tracking), vertical displacement (2015–2018 DoD), collapse block thickness (2018–2021 DoD), and post-landslide topography to calculate the dominant mode of pre-collapse motion for the Chamoli collapse block.

We compared our observations of vertical and horizontal slope displacement to a synthetic displacement, with the hypothesis that all of the observed change could be explained by translation. We used the 2018–2021 DoD to calculate the



**Figure 7.** Optical satellite image (a; Sentinel-2, 28 September 2020) and DoD (b; 2018–2015) of the Chamoli collapse site. Note the large number of steep slopes, complex terrain, and high noise levels in the DoD. In order to be useful for hazard prevention, these methods need to be able to identify potentially hazardous slopes without prior knowledge about collapses (e.g., green arrow shows the location of the headwall crack).



**Figure 8.** (a) Observed and (b) modeled elevation change of the Chamoli landslide block prior to collapse. The modeled scenario (b) is based on 20 m of pure downslope translation. (c) A scatterplot comparing each observed and modeled pixel.

thickness and location of the mobile block and translated this using the direction of the steepest slope ( $\sim$  NNE) and a displacement magnitude of 20 m. The displacement magnitude is chosen to match our observed horizontal displacement from feature tracking and is consistent with the findings of Qi et al. (2021). We then calculated the difference in surface elevation between the original and translated surfaces.

Figure 8 shows a comparison between the observed change in surface elevation of the landslide block and the modeled change. The pattern of elevation change is similar for the observed and modeled cases – both exhibit a deep crack, bulging in the lower collapse zone, and greater elevation gain on this bulge to the east relative to the west. The two-dimensional correlation score is 0.64, with the greatest model–data difference at the headwall crack, which is as much as 150 m deeper in the model case. These results are

consistent with the Chamoli collapsed block moving downslope by translation in the years prior to collapse.

### 4.3 A possible avalanche-triggering mechanism

A viable triggering mechanism for the Chamoli landslide must explain both the lag between the initial instability and collapse and the timing of the collapse – in the middle of the winter. Syn-collapse seismic signals show that there was no seismic trigger for the collapse (Pandey et al., 2021; Shugar et al., 2021; Cook et al., 2021). Nearby meteorological stations and reanalysis data reveal heavy snowfall and a  $\sim$  5 K positive temperature anomaly in the week preceding collapse, as well as a temperature inversion in the valley (e.g., Pandey et al., 2021; Dandabathula et al., 2021; Zhou et al., 2021; Shugar et al., 2021). In the longer term, this region has warmed an estimated 0.014 (Zhou et al., 2021) to

$0.033 \text{ K yr}^{-1}$  (Shrestha et al., 2021) since 1980, for a total warming of 0.4 to 0.9 K over the past 3 decades.

Zhou et al. (2021) and Dandabathula et al. (2021) propose that this sudden temperature increase may have triggered the collapse, and Rana et al. (2021) associate it with lubrication of pre-existing fractures via melting of fresh snow. Kropáček et al. (2021) and Pandey et al. (2021) suggest that loading from heavy snowfall may have contributed to the failure. Despite the positive temperature anomaly, temperatures at the collapse altitude ( $\sim 5000 \text{ m}$ ) would have been below the freezing point on the day of collapse, and liquid water would not have been present at the surface (Shugar et al., 2021; Dandabathula et al., 2021). Positive summer temperatures (Shrestha et al., 2021) and a steep surface slope of the collapse block will have prevented strong cumulative surface loading of the collapse block through snow deposition. Existing hypotheses do not provide strong mechanistic links between observed meteorological changes and the slope failure.

The stability of a slope can be described by the balance between two terms: driving forces ( $F_D$ ) and resistive forces ( $F_R$ ). Driving forces are primarily gravitational, while resistive forces are primarily related to slope cohesion and friction (Appendix B). For a detached wedge such as the Chamoli collapse block, dominant resistive forces are likely friction along the margins and base of the collapsed block. The balance between these two forces is known as the factor of safety FS:

$$\text{FS} = \frac{F_R}{F_D}. \quad (1)$$

A slope is considered unstable when its factor of safety falls below 1 (e.g., Záruba and Mencl, 2014; Das and Sivakugan, 2016).

The Chamoli collapse area is composed of heavily jointed bedrock (e.g., Shugar et al., 2021). A pure translational pre-collapse motion is consistent with a collapse block basal shear plane along a single bedding plane. High-resolution post-collapse satellite imagery also suggests that the detachment occurred along a bedding plane. This failure plane may have been superficially weakened by freeze–thaw fracturing (Qi et al., 2021; Kropáček et al., 2021; Shrestha et al., 2021) or at greater depth by changes in permafrost conditions (e.g., Gruber and Haeberli, 2007; Krautblatter et al., 2013). The surface velocity peaks in summers 2017 and 2018 suggest that surface meltwater may have reached into the later failure surface. Meltwater infiltration may directly impact friction ( $F_D$ ) and, in a delayed way, also alter ground temperatures through advection of heat and release of latent heat upon refreezing. Gruber and Haeberli (2007) note that advection-driven melt of permafrost thaw corridors may drive destabilization of large volumes of rock. Deep permafrost thaw may occur over long timescales (e.g., Gruber and Haeberli, 2007; Krautblatter et al., 2013) and provides one potential explanation for the 5-year lag between initial instability and collapse.

The deep headwall crack provides accommodation space for cumulative snow accumulation and loading and also limits the melting of accumulated snow by reducing its surface exposure. Observations of elevation change over 2015–2018 show the opening of a crack at least 25 m deep at the collapse block headwall (Fig. 8a), although DEMs may underestimate the true depth of the crack due to the viewing angle, slope geometry, and stereo-DEM processing parameters. The purely translation model of block motion (Fig. 8b) suggests that the true crack depth would have been closer to 150 m. Snow, ice, or rock debris loading within a headwall crack would exert a horizontal force on the collapse block. This horizontal force (“push”) acts to reduce the factor of safety both by directly increasing the driving force of the collapse block and reducing the angle between the driving force vector and slope direction (equivalent to an increase in slope; see Appendix B).

Accumulation of snow or ice in the crack is visible in optical satellite imagery, with additional input from snow–ice avalanches from the overlying slope (e.g., Fig. 3b–d). A storm in the days preceding the 7 February collapse brought substantial snowfall to the Chamoli region, with local snowfall estimates ranging from 8.5 to 48 mm water equivalent of precipitation (Shugar et al., 2021; estimates from local weather stations and the Weather Research and Forecasting model). We use these data to calculate the potential range of snow loading on the collapsed block, which is equivalent to a slope-parallel force of 7000–40 000 kN (Appendix A3). Considering the total precipitation between crack initiation (March 2016) and collapse (February 2021) this rises to  $6.3 \times 10^9$  to  $9.9 \times 10^9 \text{ N}$ , i.e., 2%–3% of the total driving force of the collapse block.

In the absence of in situ instrumentation and observations, it may not be possible to determine the exact cause of the failure at Chamoli. Nevertheless, we propose a mechanism which is compatible with both the lag between initial instability and collapse and the timing of the eventual collapse. Snow and ice loading in the headwall crack would progressively increase the driving force of the collapse block, while meltwater infiltration and permafrost degradation in a bedrock fracture would steadily reduce its resistive forces (basal friction). The combination of these two processes would reduce the factor of safety and pre-condition the block for failure, with the early February positive temperature anomaly and loading from snowfall providing a final driver for mid-winter collapse.

Overall, forecasting the 7 February 2021 Chamoli rock–ice avalanche prior to its occurrence from remotely sensed datasets would have been very challenging and certainly not routine work using well-established methods. Current image resolution, characteristics, and processing algorithms result in noise levels on a similar order to the signal itself. Only the joint interpretation of feature-tracking results, DEM differences, and satellite images reveals clear precursory signs of slope instability. In addition, none of the data in this study are able to adequately forecast the timing of collapse. As such,

current archives of satellite images do not currently appear to be practical for forecasting individual events. At the same time, this should not prevent remote monitoring of hazardous zones, particularly when adjacent to vulnerable areas. Every slope failure will exhibit a different range of pre-collapse signals, and new instabilities might be recognized in some cases. Even though the forecasting of individual events remains a challenge, these data have value for identifying zones of the highest risk for in situ monitoring or the installation of early-warning systems (Cook et al., 2021).

Feature tracking, DEM difference, and InSAR datasets can be processed and analyzed on a regional or even global scale – and in many cases pre-processed datasets are already available online (e.g., Morin et al., 2016; Gardner et al., 2018; Crosetto et al., 2020; Dai et al., 2020a; Provost et al., 2022). While these pre-processed datasets are not generally produced for slope stability monitoring, they can be used to improve hazard maps and reduce landslide-related damage. Future advances in Earth observation satellite capabilities and processing algorithms will improve the quality of such products.

## 5 Conclusions

The deadly 7 February 2021 Chamoli rock–ice avalanche was initiated by the failure of more than  $25 \times 10^6 \text{ m}^3$  of rock and ice high in the Uttarakhand Himalaya. We investigated the conditions of the avalanche source zone over the decades preceding collapse through a combination of optical and radar satellite images. We used feature tracking to calculate horizontal slope displacements and differenced photogrammetrically generated DEMs to investigate vertical displacements. We showed that the collapsed block moved 20–30 m prior to its collapse, with the most rapid motion occurring around 3 years prior to failure. Comparison between our datasets and modeled displacement maps shows that the motion occurred primarily via downslope translation, opening up a deep crack at the headwall. A combination of permafrost degradation and snow and ice debris loading within this headwall crack may explain both the lag between initial instability and collapse and the mid-winter timing of the collapse. Finally, we assessed the potential of these datasets and approaches for monitoring other unstable slopes. While they are effective at identifying precursory signals at a known collapse site, it remains very challenging to predict such collapses with sufficient levels of confidence in high-mountain areas.

### Appendix A: Pre-event DEM mosaics

#### A1 Cartosat-1 (2017–2018)

We procured four Cartosat-1 stereo pairs from October 2017 and November 2018 (data sheet in the Supplement) to com-

pute DEMs for an intermediate period between 2015 and 2021. Initial assessments of the Cartosat-1 products revealed high stereo-ray intersection errors ( $> 100 \text{ m}$ ) and offsets from reference elevation models ( $\sim 400 \text{ m}$ ), indicative of poor relative and absolute accuracy of the vendor-supplied RPC (rational polynomial coefficient) models. To address these issues we employed ASP's "bundle\_adjust" utility on all of the eight overlapping images and the corresponding RPC models using similar techniques as described in Bhushan et al. (2021) and Dehecq et al. (2020). The bundle adjustment procedure matches similar features between all input overlapping images and minimizes their reprojection error by updating the RPC camera with translation and rotation parameters. Using the updated RPC model obtained after bundle adjustment, we generated a draft DEM from one of the four pairs using the default ASP settings and aligned it to a filtered and masked version of the HMA (High Mountain Asia) 8 m DEM mosaic v2 (Shean, 2021). The alignment matrix was used to further update the self-consistent RPC model output from bundle adjustment, ensuring improved absolute geolocation accuracy. Following this, the input images were orthorectified at their native resolution of 2.5 m using the 30 m Copernicus DEM (converted to ellipsoidal heights), and stereo processing (correlation and triangulation) was performed for all of the four input pairs using the settings described in Shean and Bhushan (2021).

The Cartosat-1 DEMs were posted at 10 m resolution with the UTM 44N (Universal Transverse Mercator) projection and heights above the WGS84 (World Geodetic System 1984) ellipsoid. Consequently, the DEMs were co-registered to the HMA 8 m DEM mosaic v2 (Shean, 2021) over non-glacierized surfaces using a two-step procedure: ASP's "pc\_align" followed by Nuth and Kääb (2011) alignment implemented in Shean et al. (2019) to remove any residual horizontal and vertical offsets in the final output DEMs.

#### A2 SPOT 7 (2018)

We also derived a DEM from the 24 December 2018 SPOT 7 stereo pair using ASP's Semi-Global Matching correlator and other settings similar to those described in Lacroix (2016) and Deschamps-Berger et al. (2020). The final output DEM was posted at a resolution of 10 m with the UTM 44N projection and heights above the WGS84 ellipsoid. The DEM was co-registered to the HMA 8 m DEM mosaic v2 (Shean, 2021) over non-glacierized surfaces to ensure consistency with all the DEM products derived in this study.

### Appendix B: Factor of safety calculations for the Chamoli block

The factor of safety FS is calculated from the balance driving and resistive forces (e.g., Záruba and Mencl, 2014; Das and Sivakugan, 2016):

$$FS = \frac{F_R}{F_D} = \frac{AC + Mg \cos(\alpha) \tan(\phi)}{Mg \sin(\alpha)}, \quad (\text{B1})$$

where  $A$  is slip surface area,  $C$  is cohesion,  $M$  is the mass of the unstable region,  $g$  is gravity,  $\alpha$  is slope, and  $\phi$  is the friction angle. A system may be considered unstable when the factor of safety falls below 1.

Introducing an additional horizontal force  $F_H$  modifies this balance in two ways: firstly by increasing the driving force and secondly by altering the angle between the driving force vector and resistive forces vector as

$$FS = \frac{AC + Mg \cos(\alpha + \alpha') \tan(\phi)}{F_H + Mg \sin(\alpha + \alpha')}. \quad (\text{B2})$$

The change in angle of the driving force vector  $\alpha'$  is then given by  $\alpha' = \arctan\left(\frac{F_H}{Mg \sin(\alpha)}\right)$ . In our situation, for a given mass accumulated in the headwall crack  $M_C$ , we have  $F_H = M_C \sin(\alpha)$ .

The pre-event storm brought 8.5 to 48 mm water equivalent of precipitation (Shugar et al., 2021; estimates from local weather stations and the Weather Research and Forecasting model). We may use this data to calculate possible loading of this snow on the collapsed block – considering a 500 m long, 70 m wide crack with a 500 m long and fed by a 180 m wide avalanche zone. Assuming that all of the snowfall was channeled into the crack, total loading  $M_C$  would be equal to

$$M_C = A_A \cdot P \cdot \rho_P \cdot g, \quad (\text{B3})$$

with  $A_A$  being the accumulation area feeding the crack,  $P$  being precipitation (in meters), and  $\rho_P$  being the density of the precipitation. Total snow loading in the headwall crack associated with this single precipitation event would therefore be 10 000–60 000 kN, equivalent to a slope-parallel horizontal force of 7000–40 000 kN.

IMGERG (Integrated Multi-satellite Retrievals for GPM, Global Precipitation Measurement) precipitation data suggest that around  $9 \pm 2$  m of precipitation fell in the collapse area between crack initiation in 2016 and collapse in 2021. Using the same calculation, maximum snow load in the headwall crack is equal to  $8.6\text{--}13.5 \times 10^9$  N, equivalent to a slope-parallel horizontal force of  $6.3\text{--}9.9 \times 10^9$  N. For reference, the estimated total driving force of the collapse block, composed of  $21 \times 10^6$  m<sup>3</sup> of rock and  $6 \times 10^6$  m<sup>3</sup> of ice, is  $\sim 4.0 \times 10^{11}$  N.

*Code and data availability.* All code used in this study is openly available online: GIV at <https://doi.org/10.5281/zenodo.4904544> (Van Wyk de Vries, 2021), autoRIFT at <https://github.com/nasa-jpl/autoRIFT> (Kennedy, 2022), ASP at <https://github.com/NeoGeographyToolkit/StereoPipeline> (Alexandrov, 2022), and ISCE (Interferometric synthetic aperture radar Scientific Computing Environment) at <https://github.com/isce-framework/isce2> (Burns, 2022). The 2015 pre-event DEM is available at <https://doi.org/10.5281/zenodo.4554647> (Bhushan and Shean, 2021), and the 2021 post-event DEM is at <https://doi.org/10.5281/zenodo.4558692> (Shean et al., 2021).

*Supplement.* The supplement related to this article is available online at: <https://doi.org/10.5194/nhess-22-3309-2022-supplement>.

*Author contributions.* All authors designed the study and conducted the research. MVWDV wrote the paper, with input from all co-authors. The final version has been approved by all co-authors.

*Competing interests.* The contact author has declared that none of the authors has any competing interests.

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