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# Diurnal variations in oxygen and nitrogen isotopes of atmospheric nitrogen dioxide and nitrate: implications for tracing $\text{NO}_x$ oxidation pathways and emission sources

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**Abstract.** The oxygen ( $\Delta^{17}\text{O}$ ) and nitrogen ( $\delta^{15}\text{N}$ ) isotopic compositions of atmospheric nitrate ( $\text{NO}_3^-$ ) are widely used as tracers of its formation pathways, precursor (nitrogen oxides ( $\text{NO}_x$ )  $\equiv$  nitric oxide (NO) + nitrogen dioxide ( $\text{NO}_2$ )) emission sources, and physico-chemical processing. However, the lack of observations on the multi-isotopic composition of  $\text{NO}_2$  perpetuates significant uncertainties regarding the quantitative links between the isotopic composition of  $\text{NO}_x$  and  $\text{NO}_3^-$ , which ultimately may bias inferences about  $\text{NO}_3^-$  formation processes and the distribution of sources, particularly in winter urban atmospheres. We report here on the first simultaneous atmospheric observations of  $\Delta^{17}\text{O}$  and  $\delta^{15}\text{N}$  in  $\text{NO}_2$  ( $n = 16$ ) and  $\text{NO}_3^-$  ( $n = 14$ ). The measurements were carried out at sub-daily ( $\sim 3$  h) resolution over 2 non-consecutive days in an Alpine city in February 2021. A strong diurnal signal is observed in both  $\text{NO}_2$  and  $\text{NO}_3^-$  multi-isotopic composition.  $\Delta^{17}\text{O}$  of  $\text{NO}_2$  and  $\text{NO}_3^-$  ranges from 19.6‰ to 40.8‰ and from 18.3‰ to 28.1‰, respectively. During the day and night, the variability in  $\Delta^{17}\text{O}(\text{NO}_2)$  is mainly driven by the oxidation of NO by ozone, with a substantial contribution from peroxy radicals in the morning.  $\text{NO}_3^-$  mass balance equations, constrained by observed  $\Delta^{17}\text{O}(\text{NO}_2)$ , suggest that during the first day of sampling, most of the  $\text{NO}_3^-$  was formed locally from the oxidation of  $\text{NO}_2$  by hydroxyl radicals by day and via heterogeneous hydrolysis of dinitrogen pentoxide at night. For the second day, calculated and observed  $\Delta^{17}\text{O}(\text{NO}_3^-)$  do not match, particularly daytime values; the possible effects on  $\Delta^{17}\text{O}(\text{NO}_3^-)$  of a Saharan dust event that occurred during this sampling period and of winter boundary layer dynamics are discussed.  $\delta^{15}\text{N}$  of  $\text{NO}_2$  and  $\text{NO}_3^-$  ranges from  $-10.0$ ‰ to 19.7‰ and from  $-4.2$ ‰ to 14.9‰, respectively. Consistent with theoretical predictions of N isotope fractionation, the observed variability in  $\delta^{15}\text{N}(\text{NO}_2)$  is explained by significant post-emission equilibrium N fractionation. After accounting for this effect, vehicle exhaust is found to be the primary source of  $\text{NO}_x$  emissions at the sampling site.  $\delta^{15}\text{N}(\text{NO}_3^-)$  is closely linked to  $\delta^{15}\text{N}(\text{NO}_2)$  variability, bringing further support to relatively fast and local  $\text{NO}_x$  processing. Uncertainties in current N fractionation factors during  $\text{NO}_2$  to  $\text{NO}_3^-$  conversion are underlined. Overall, this detailed investigation highlights the potential and necessity of simultaneously using  $\Delta^{17}\text{O}$  and  $\delta^{15}\text{N}$  in  $\text{NO}_2$  and  $\text{NO}_3^-$  in order to better constrain quantitative inferences about the sources and formation chemistry of  $\text{NO}_3^-$  in urban environments in winter.

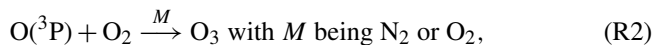
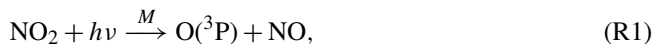
## 1 Introduction

Despite extensive efforts in emission controls in recent decades, global anthropogenic emissions of nitrogen oxides ( $\text{NO}_x \equiv$  nitrogen monoxide (NO) + nitrogen dioxide ( $\text{NO}_2$ )) remain more than 2 orders of magnitude higher than before the industrial revolution (Hoesly et al., 2018). Atmospheric nitrate ( $\text{NO}_3^- \equiv$  nitric acid ( $\text{HNO}_3$ ) + particulate nitrate ( $p\text{-NO}_3^-$ )) is the main end product of  $\text{NO}_x$  oxidation and a key component of fine particulate matter (PM), which adversely affects human health (WHO, 2021) and contributes to climate change (Masson-Delmotte et al., 2021).  $\text{NO}_3^-$  can be transported far from emission sources and can be removed from the atmosphere through dry and wet deposition within hours to days (Alexander et al., 2020; Park et al., 2004). The additional input of this “reactive” nitrogen ( $N_r$ ) into natural environments is known to have detrimental consequences, particularly regarding biodiversity and water quality (Galloway et al., 2008; Vitousek et al., 1997). It is therefore important to have a comprehensive understanding of  $\text{NO}_x$  emission sources and oxidation processes, on which effective air quality and climate change mitigation strategies rely (e.g. Bauer et al., 2007; Huang et al., 2014; Shah et al., 2018; Tsimpidi et al., 2008; Wang et al., 2013, 2020).

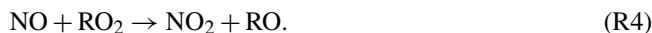
Several studies have noted that the response of  $\text{NO}_3^-$  concentration in air to  $\text{NO}_x$  emission reduction is contrasting, particularly in winter (e.g. Shah et al., 2018; Tørseth et al., 2012; Wang et al., 2020; Zhou et al., 2019). This is because a variety of factors controls the  $\text{NO}_x$  conversion efficiency and the  $\text{NO}_3^-$  content in PM, including precursor emission sources, complex multiphase chemical reactions with other reactive species, and environmental conditions (e.g. temperature, relative humidity, solar radiation) (Zhang et al., 2015). It remains difficult to assess the contribution of each parameter to the non-linear  $N_r$  chemistry, which is partly driven by close links between changes in aerosol acidity, gas–particle partitioning, and atmospheric oxidation capacity (Shah et al., 2018; Fu et al., 2020; Shi et al., 2019; K. Li et al., 2021). Ozone ( $\text{O}_3$ ) and hydroxyl radicals (OH) (Finlayson-Pitts and Pitts, 2000) are the major oxidants in the atmosphere whose chemical cycles are largely controlled by solar radiation. As a result, there are significant diurnal and seasonal variations in  $\text{NO}_x$  chemistry (e.g. Prabhakar et al., 2017; Alexander et al., 2020). Notably,  $\text{NO}_3^-$  formation is generally dominated by homogeneous OH oxidation and heterogeneous  $\text{O}_3$  chemistry during the day and summer and during the night and winter, respectively (Alexander et al., 2020). However, assessing the relative contributions of individual formation channels, together with their sensitivity to environmental parameters, is not straightforward and requires extensive in situ observations combined with modelling tools (e.g. Alexander et al., 2020; Brown et al., 2006; Newsome and Evans, 2017; Xue, 2022; Prabhakar et al., 2017).

Upon release into the atmosphere,  $\text{NO}_x$ , mainly emitted as NO, undergoes oxidation to form  $\text{NO}_2$ . During the day,

a rapid photochemical equilibrium is established between NO and  $\text{NO}_2$ , known as the “photostationary state” (PSS; Leighton, 1961), via key interconversion reactions (Reactions R1–R3):



This cycle can be disturbed by peroxy radicals ( $\text{RO}_2 \equiv$  hydroperoxyl radical ( $\text{HO}_2$ ) + methyl peroxy radical ( $\text{CH}_3\text{O}_2$ )) via typically Reaction (R4):



Note that in polluted atmospheres where  $\text{NO}_x$  mixing ratios often exceed parts-per-billion (ppb) levels, Reaction (R4) followed by Reactions (R1)–(R2) leads to the formation of  $\text{O}_3$  (Crutzen, 1979). Although the role of  $\text{RO}_2$  in the  $\text{NO}_x$  oxidation is crucial in  $\text{O}_3$  formation and  $\text{NO}_x$  oxidation rate, measuring the  $\text{RO}_2$  mixing ratio remains challenging due to the need for state-of-the-art instrumentation coupled with photochemical models to establish chemical budgets (e.g. Ren et al., 2006; Tan et al., 2018). While NO is usually oxidised relatively quickly into  $\text{NO}_2$  in summer due to the high levels of  $\text{O}_3$  and solar radiation, the shorter day length and lower temperature in winter result in a contrasting  $\text{NO}_x$  cycling. In particular, the formation of a temperature inversion at the surface can trap pollutants emitted close to the surface in a shallow layer for hours to days (e.g. Llargeron and Staquet, 2016; Olofson et al., 2009). Under those conditions, it is not uncommon for  $\text{O}_3$  levels to be very low due to quasi-complete titration by NO, which can have further impacts on the atmospheric oxidation capacity. However, pronounced  $\text{O}_3$  pollution episodes may also arise in winter in highly polluted areas, such as in oil-producing regions due to intense volatile organic compound (VOC) emissions (Edwards et al., 2014) or in China, where pollution control strategies mainly target  $\text{NO}_x$  while VOC emissions remain more or less constant (Ren et al., 2022). In urban areas,  $\text{NO}_2$  is generally mainly removed from the atmosphere by reaction with OH during the day via Reaction (R5) (Dentener and Crutzen, 1993):



$\text{NO}_2$  can also react with  $\text{O}_3$  to form nitrate radicals ( $\text{NO}_3$ ) via Reaction (R6):



However,  $\text{NO}_3$  is rapidly photolysed during the day, regenerating  $\text{NO}_2$  (Wayne et al., 1991). Another important  $\text{NO}_3$  loss reaction is that with NO in polluted environments (Brown and Stutz, 2012). At night, without photolytic activity and with lower precursor emissions, the lifetime of

$\text{NO}_3$  radicals substantially increases.  $\text{NO}_3$  reacts with  $\text{NO}_2$  to form dinitrogen pentoxide ( $\text{N}_2\text{O}_5$ ; Reaction R7), which then undergoes heterogeneous hydrolysis to form  $\text{HNO}_3$  (Reaction R8):



Reaction (R7) is temperature dependent, so  $\text{N}_2\text{O}_5$  can eventually decompose to reform  $\text{NO}_2$  and  $\text{NO}_3$ , with the  $\text{N}_2\text{O}_5/\text{NO}_3$  ratio being negatively correlated with temperature.  $\text{N}_2\text{O}_5$  is an important nocturnal sink for  $\text{NO}_x$ , notably in winter in urban atmospheres due to high aerosol loads and low temperatures. However, the efficiency of Reaction (R8) is difficult to determine because it strongly depends on parameters such as the aerosol surface density and its chemical composition (Brown et al., 2006), which are not often well characterised. In addition,  $\text{NO}_3$  can react with hydrocarbons to produce  $\text{HNO}_3$ , which could significantly contribute to the formation of  $\text{NO}_3^-$  in industrialised regions with high hydrocarbon emissions (Brown et al., 2011). It is estimated that the Reactions (R1) to (R8) lead to the formation of 82 % of  $\text{NO}_3^-$  near the surface on a global scale (Alexander et al., 2020). In polluted environments, the respective contributions of Reaction (R5) (OH pathway) and Reactions (R6)–(R8) ( $\text{N}_2\text{O}_5$  pathway) are more contrasted and are still debated (e.g. Chan et al., 2021; Fu et al., 2020). In addition, the reaction of  $\text{N}_2\text{O}_5$  with chlorine on aerosols can contribute to  $\text{NO}_3^-$  production in urban atmospheres (Thornton et al., 2010), with further impacts on  $\text{O}_3$  production in continental polluted atmospheres in winter (X. Wang et al., 2019). Other reactions, such as those involving halogen and organic intermediates, may become significant for  $\text{NO}_3^-$  production in specific regions, such as in polar, oceanic, and coastal areas (Alexander et al., 2020; Penkett et al., 2007; Savarino et al., 2013; Simpson et al., 2015).

To help better constrain the atmospheric  $N_r$  chemistry and budget, the last 3 decades has seen a growing interest in stable oxygen (O) and nitrogen (N) isotopes, notably in  $\text{NO}_3^-$  (Elliott et al., 2019; Savard et al., 2018). The isotopic composition is reported as an isotopic enrichment ( $\delta$ ) with respect to a reference material, defined as  $\delta = (R_{\text{sample}}/R_{\text{reference}} - 1)$ , and expressed in per mill (‰).  $R$  refers to the elemental abundance ratio of the heavy isotope to the light isotope (e.g.  $^{18}\text{O}/^{16}\text{O}$ ,  $^{17}\text{O}/^{16}\text{O}$ ,  $^{15}\text{N}/^{14}\text{N}$ ) in the sample and in an international isotopic reference material (Vienna Standard Mean Ocean Water for O, Li et al., 1988; atmospheric  $\text{N}_2$  for N, Mariotti, 1984). A powerful tool to help trace the relative importance of different  $\text{NO}_x$ -to- $\text{NO}_3^-$  oxidation pathways is through the use of the  $^{17}\text{O}$  excess ( $\Delta^{17}\text{O} = \delta^{17}\text{O} - 0.52 \times \delta^{18}\text{O}$ ).  $\Delta^{17}\text{O}$  is transferred to  $\text{NO}_3^-$  by  $\text{O}_3$ , which possesses a very unique  $\Delta^{17}\text{O}$  ( $26.2 \pm 1.3$ ‰; Vicars and Savarino, 2014) due to mass-independent fractionation during its formation process (Thiemens, 2006). In comparison, the  $\Delta^{17}\text{O}$  of other atmospheric oxidants such as OH is near zero due to

isotopic exchange with atmospheric water vapour (Dubey et al., 1997). Similarly, as the isotopic anomaly of atmospheric  $\text{O}_2$  is very close to 0‰ (Barkan and Luz, 2003) and since  $\text{RO}_2$  is mostly produced by the reactions  $\text{R} + \text{O}_2$  and  $\text{H} + \text{O}_2$ ,  $\Delta^{17}\text{O}$  of  $\text{RO}_2$  can be considered negligible (Alexander et al., 2020). Therefore,  $\Delta^{17}\text{O}$  in  $\text{NO}_3^-$  represents a unique tracer of the  $\text{O}_3$  implication in its formation processes that can provide valuable constraints on the relative contributions of individual reactions (e.g. Morin et al., 2011; Alexander et al., 2009; Michalski et al., 2003). By a simple mass balance calculation of O atoms in  $\text{NO}_3^-$ , the  $^{17}\text{O}$  excess of  $\text{NO}_3^-$  produced by an individual  $\text{NO}_2$ -to- $\text{NO}_3^-$  conversion process  $i$  ( $\Delta^{17}\text{O}(\text{NO}_3^-)_i$ ) can be expressed as

$$\Delta^{17}\text{O}(\text{NO}_3^-)_i = \frac{2}{3} \times \Delta^{17}\text{O}(\text{NO}_2) + \frac{1}{3} \times \Delta^{17}\text{O}(\text{add. O})_i, \quad (1)$$

where  $\Delta^{17}\text{O}(\text{NO}_2)$  is the  $^{17}\text{O}$  excess of atmospheric  $\text{NO}_2$  and  $\Delta^{17}\text{O}(\text{add. O})_i$  is the transferrable  $^{17}\text{O}$  excess of the oxidant responsible for the conversion of  $\text{NO}_2$  into  $\text{NO}_3^-$  (Michalski et al., 2003). From Eq. (1), if  $\Delta^{17}\text{O}(\text{NO}_2)$  is constrained, one can derive individual  $\Delta^{17}\text{O}$  transfer to  $\text{NO}_3^-$  relative to an  $i$  conversion process and compare this value with observed  $\Delta^{17}\text{O}(\text{NO}_3^-)$ .

Recent studies in urban areas have attempted to interpret the variability in  $\Delta^{17}\text{O}(\text{NO}_3^-)$  in aerosols in order to quantify the relative contribution of homogeneous and heterogeneous processes to  $\text{NO}_3^-$  formation (e.g. Fan et al., 2023, 2022; He et al., 2020, 2018; Lim et al., 2022; Wang et al., 2023; Y. L. Wang et al., 2019; Kim et al., 2023; Y.-L. Zhang et al., 2022; Z. Li et al., 2022). However, to that end, it is necessary to have a clear quantitative understanding of the transfers of  $\Delta^{17}\text{O}$  in the  $N_r$  cycle. To date, due to very limited observational data, there is a lack of well-established knowledge on the dynamics of  $\Delta^{17}\text{O}$  in  $\text{NO}_2$ , the key intermediate species in the formation of  $\text{NO}_3^-$ . Consequently, strong assumptions about  $\Delta^{17}\text{O}(\text{NO}_2)$  have to be made when interpreting  $\Delta^{17}\text{O}(\text{NO}_3^-)$  measurements, which could potentially lead to biased conclusions. Notably, the most difficult regions for the interpretation of  $\text{NO}_3^-$  records are potentially polluted areas where the isotopic composition of  $\text{NO}_2$  is expected to be highly variable in space and time. Most studies typically estimate  $\Delta^{17}\text{O}(\text{NO}_2)$  during the day by assuming that an isotopic steady state (ISS) is reached between  $\text{NO}_x$  and  $\text{O}_3$ , resulting in  $\Delta^{17}\text{O}(\text{NO}_2)$  depending only on the relative contributions of different oxidants to NO oxidation. A recent study reported the first in situ observations of  $\Delta^{17}\text{O}(\text{NO}_2)$  in an urban environment (Grenoble, France) in spring (Albertin et al., 2021). Time-resolved  $\text{NO}_2$  sampling (ca. 3 h) during 24 h revealed a strong diurnal cycle in  $\Delta^{17}\text{O}(\text{NO}_2)$ , reaching ca. 40‰ during the day and decreasing down to ca. 20‰ at night. The observed  $\Delta^{17}\text{O}(\text{NO}_2)$  values and diurnal variability were consistent with its expected behaviour derived from  $\Delta^{17}\text{O}$  mass balance equations under the ISS assumption during the day. However, this first  $\Delta^{17}\text{O}(\text{NO}_2)$  dataset is very small. More measurements are needed to test

the applicability of this new isotopic tool in different environments and seasons and to assess whether the ISS is still valid. At night, the low  $\Delta^{17}\text{O}(\text{NO}_2)$  measured by Albertin et al. (2021) is consistent with the oxidation of freshly emitted NO by  $\text{O}_3$ . Nonetheless, since the timescale for the oxidation of  $\text{NO}_2$  into  $\text{NO}_3^-$  is thought to exceed the duration of the night (Alexander et al., 2020), it is also common to assume that the isotopic composition of nocturnal  $\text{NO}_2$  reflects more daytime formation and conditions of the previous days. While this assumption may hold true in remote areas (Morin et al., 2011), significant uncertainties subsist in urban areas where the nighttime  $\text{NO}_3^-$  chemistry may be more efficient. In such circumstances, the production of  $\text{NO}_3^-$  from  $\text{NO}_2$  formed at night would lead to a lower-than-expected  $\Delta^{17}\text{O}$  transfer to  $\text{NO}_3^-$ . For these reasons, the dual survey of the O isotopic composition of  $\text{NO}_2$  and  $\text{NO}_3^-$  would certainly help to accurately interpret  $\Delta^{17}\text{O}(\text{NO}_3^-)$  observations in polluted atmospheres, particularly with sampling at sub-daily timescales, which would allow us to study the diurnal dynamics of  $\Delta^{17}\text{O}(\text{NO}_2)$  and its links with  $\Delta^{17}\text{O}(\text{NO}_3^-)$ .

In addition to  $\Delta^{17}\text{O}$ ,  $\delta^{15}\text{N}$  in  $\text{NO}_3^-$  ( $\delta^{15}\text{N}(\text{NO}_3^-)$ ) can be used as a tracer of  $\text{NO}_3^-$  sources and/or chemical processing. As different  $\text{NO}_x$  emission sources often have distinct  $\delta^{15}\text{N}$  fingerprints depending on the  $\text{NO}_x$  production mechanism (Heaton, 1990; Felix et al., 2012; Fibiger and Hastings, 2016; Walters et al., 2015a, b; Yu and Elliott, 2017; Miller et al., 2018),  $\delta^{15}\text{N}(\text{NO}_3^-)$  is a potentially valuable tool to trace the origins of its gaseous precursor. However, due to N fractionation effects associated with physico-chemical processing,  $\delta^{15}\text{N}$  is altered during the conversion of  $\text{NO}_x$  to  $\text{NO}_3^-$  (Elliott et al., 2019). Therefore, the variability in  $\delta^{15}\text{N}(\text{NO}_3^-)$  can be attributed to the following: (1) a change in  $\text{NO}_x$  emission sources and (2) N isotopic fractionations between NO and  $\text{NO}_2$ , between  $\text{NO}_2$  and  $\text{NO}_3^-$ , and during the transport of  $\text{NO}_3^-$ . These effects co-exist with relative contributions varying according to environmental conditions and the mix of  $\text{NO}_x$  emissions. Numerous observations in diverse environments have emphasised the substantial influence of N fractionation effects in altering the original  $^{15}\text{N}$  composition of emitted gaseous  $\text{NO}_3^-$  precursors (e.g. Bekker et al., 2023; Chang et al., 2018; Geng et al., 2014; Luo et al., 2023; Vicars et al., 2013; J. Li et al., 2021). Although some N fractionation factors are available from calculations (Walters and Michalski, 2015) and laboratory experiments (Li et al., 2020; Walters et al., 2016), there is still a lack of observational constraints on the magnitude of the N isotopic partitioning between  $\text{NO}_x$  and  $\text{NO}_3^-$ , which could lead to biased interpretations of  $\delta^{15}\text{N}(\text{NO}_3^-)$  observations.

Following the preliminary work of Albertin et al. (2021), this study presents for the first time simultaneous measurements of atmospheric  $\text{NO}_2$  and  $\text{NO}_3^-$  multi-isotopic compositions. The sampling took place at a high temporal resolution ( $\sim 3$  h) in late February 2021 in an urban Alpine city.  $\Delta^{17}\text{O}/\delta^{15}\text{N}$  data of  $\text{NO}_2$  and  $\text{NO}_3^-$ , meteorological parameters, and atmospheric observations ( $\text{NO}$ ,  $\text{NO}_2$ ,  $\text{O}_3$ , and

PM) are collated in order to investigate diurnal  $N_r$  chemistry, N fractionation effects, and  $\text{NO}_x$  emissions. Our winter measurements extend the atmospheric  $\text{NO}_2$  multi-isotopic composition record, which is only composed of spring measurements performed during a single day by Albertin et al. (2021). The general aim of this case study is to test  $\text{NO}_2$ -based isotopic approaches for tracing the origins and fate of  $\text{NO}_x$ , for instance in urban areas on sub-daily timescales. The added value of  $\Delta^{17}\text{O}(\text{NO}_2)$  measurements in  $N_r$  chemistry studies is more critically assessed here than in Albertin et al. (2021) through the use of accurate  $\text{NO}_x$  measurements. Besides, using the isotopic theoretical framework developed previously, we explore the potential benefits of combining isotopic observations of  $\text{NO}_2$  and  $\text{NO}_3^-$  to gain a more detailed and quantitative understanding of the links between atmospheric  $N_r$  chemistry processes and variability in  $\text{NO}_2$  and  $\text{NO}_3^-$  isotopic composition. The framework used in inferring dominant  $\text{NO}_x$  emission sources from  $\text{NO}_2\delta^{15}\text{N}$  measurements is also tested.

## 2 Material and methods

### 2.1 Study site and sample collection

The study was conducted in February 2021 in Chamonix-Mont-Blanc, France ( $45^\circ 55' 21''$  N,  $6^\circ 52' 11''$  E; altitude 1035 m above sea level, m.a.s.l.). This narrow ( $\sim 2$  km wide on average in Chamonix) 23 km Alpine valley of about 12 000 inhabitants is surrounded by high-elevation mountains. The city can experience severe PM pollution events during the winter season, mainly due to wood-combustion for domestic heating and road traffic (Chazette et al., 2005; Quimbayo-Duarte et al., 2021; Weber et al., 2018; Aymoz et al., 2007). The study's sampling site was located at a CNRS (Centre National de la Recherche Scientifique) facility in a residential area, 1.2 km south of the Chamonix city centre and 1.4 km north of the Mont-Blanc tunnel. Ambient air monitoring inlets and off-line gas and aerosol samplers were installed on the facility's terrace, 3 m above ground level (m.a.g.l.). Over the campaign, the surface was partly covered with snow.

Atmospheric particles (aerosols) were collected using a high-volume sampler (Digitel<sup>®</sup>, DH77, total suspended particle inlet,  $1\text{ m}^3\text{ min}^{-1}$ ) and glass fibre (GF) filters (Whatman<sup>®</sup>, 150 mm diameter). Evaluating the collection efficiency of total  $\text{NO}_3^-$  has long been debated (e.g. Schaap et al., 2002; Appel et al., 1980), and, although not free from sampling artefacts (e.g. potential volatilisation of  $\text{HNO}_3$  after exposure to ambient air), GF filters have been used on several times to study nitrate isotopes, mainly in coastal sites (e.g. Savarino et al., 2007; Michalski et al., 2003; Morin et al., 2009; Frey et al., 2009; Morin et al., 2007a; Patris et al., 2007; Vicars et al., 2013). Under these conditions, the aerosol alkalinity is supposed to allow the collection of  $\text{HNO}_3$  (Prospero and Savoie, 1989). In our case, as the am-

bient air in Chamonix is expected to be free of sea salt, the GF filters were not washed before use in order to keep the initial NaCl coating inherited from the filter manufacturing process. Therefore, in addition to  $p\text{-NO}_3^-$ , we are confident that the high GF filter NaCl loading allowed the quantitative collection of  $\text{HNO}_3$  at our site, as has been shown previously in the literature (Appel et al., 1981; see also our reply to the comments of Reviewer 2 and related data on this specific issue, Albertin, 2023). Two field blanks were performed to evaluate the initial content of trace elements and possible contamination during handling. Atmospheric  $\text{NO}_2$  was collected concurrently to filter samples using a pre-cleaned honeycomb denuder tube coated with a mixture of 2.5 M KOH (in methanol) and ultrapure guaiacol inserted into a ChemComb<sup>®</sup> 3500 speciation cartridge (Thermo Scientific<sup>®</sup>, USA). A second coated denuder was placed in series into the cartridge to check for  $\text{NO}_2$  breakthrough. After sampling, denuders were rinsed with 10 mL of deionised water to solubilise trapped  $\text{NO}_2$ . Detailed information on the denuder sampling protocol is available in Albertin et al. (2021). Similarly to blank filters, two blank denuders were prepared. Blank filters and denuders were subjected to the same handling, storage, and analytical treatment as field samples. Filters and denuder extractions were stored and transported frozen to IGE (Grenoble, France) for analysis.

Following the objective to investigate the diurnal isotopic composition of  $\text{NO}_2$  and  $\text{NO}_3^-$ , denuder and filter samplings were conducted continuously for 24 h with sampling time steps ranging from 1.5 to 7.5 h. During the day, denuder and filter samplings were synchronised. At night, two filter samplings were performed while three sets of denuder tubes were collected. This sampling protocol was conducted on 2 non-consecutive days, from 19 February 2021 at 21:00 local time (LT) to 20 February 2021 at 21:00 LT (sampling period no. 1 denoted SP 1) and from 24 February 2021 at 07:30 LT to 25 February 2021 at 07:30 LT (sampling period no. 2 denoted SP 2).

## 2.2 Chemical and isotopic analysis

Concentrations of major ions from filter extractions in deionised water were determined by ion chromatography (Thermo Scientific<sup>®</sup> Dionex<sup>™</sup> Integriion<sup>™</sup> HPIC). Reported to the total filter surface, the  $\text{NO}_3^-$  contribution from blank filters represented on average  $(8 \pm 9)\%$  of sampled  $\text{NO}_3^-$ . Atmospheric mass concentrations (expressed in  $\mu\text{g m}^{-3}$ ) were calculated as the ratio of the total ion filter loading (corrected for the blank contribution) to the total volume of air pumped through the filter in standard temperature and pressure (STP) conditions.  $\text{NO}_2^-$  concentrations in denuder extractions were first estimated using the Griess–Saltzman reaction and UV–Vis spectrometry at 544 nm. Even though the eluted matrix can interfere with colorimetric analyses, measured concentrations on first denuder tubes were relatively well correlated with ambient  $\text{NO}_2$  measurements during atmospheric

sampling and allowed us to give indications regarding field blanks and regarding the volume needed to perform isotopic analysis.

Isotopic analyses were performed using an isotope ratio mass spectrometer (IRMS; Thermo Finnigan<sup>™</sup> MAT 253) for analyses of  $^{15}\text{N}/^{14}\text{N}$ ,  $^{17}\text{O}/^{16}\text{O}$ , and  $^{18}\text{O}/^{16}\text{O}$  in  $\text{NO}_3^-$  and  $\text{NO}_2$  samples. Briefly,  $\text{NO}_3^-$  from filter extractions was converted into gaseous  $\text{N}_2\text{O}$  by the bacterial denitrifier method (Sigman et al., 2001; Casciotti et al., 2002; Kaiser et al., 2007) in which  $\approx 100$  nmol of  $\text{NO}_3^-$  ions were injected into 2 mL of a bacteria medium (strain of *Pseudomonas aureofaciens*) under anaerobic conditions.  $\text{NO}_2$  denuder extractions were treated separately with the azide method (McIlvin and Altabet, 2005; Albertin et al., 2021) in which 2 mL of a 2 M sodium azide / acetic acid 100 % buffer was injected into  $\approx 100$  nmol of  $\text{NO}_2^-$ , allowing quantitative conversion into  $\text{N}_2\text{O}$ . For both filter and denuder extractions, ions were converted into  $\text{N}_2\text{O}$ , which was further thermally decomposed into  $\text{O}_2$  and  $\text{N}_2$  in a gold tube heated at  $850^\circ\text{C}$ . Then,  $\text{O}_2$  and  $\text{N}_2$  molecules were separated on a chromatography column and sent separately into the IRMS for the dual analysis of O and N isotopes (see Morin et al., 2009, for more details on the analytical line). The isotopic composition of  $\text{NO}_3^-$  samples was analysed in triplicate (the mean value of replicate measurements and the associated repeatability are reported in Table S2 in the Supplement). The limited number of  $\text{NO}_2$  samples did not allow for replicate measurements. From UV–Vis analysis, all  $\text{NO}_2$  samples presented a negligible blank ( $< 4\%$ ; mean of  $1.7$  nmol  $\text{mL}^{-1}$ ) except for the sample collected between 13:30 and 16:30 LT during SP 2, which showed a blank of around  $(14.0 \pm 1.4)\%$ . Therefore, the measured  $\Delta^{17}\text{O}$  of this sample was corrected for the blank effect assuming that the contaminated  $\text{NO}_2^-$  possessed a  $\Delta^{17}\text{O}$  of  $0\%$ . No correction from this blank effect was applied to the  $\delta^{15}\text{N}$  measurements of  $\text{NO}_2$  because the  $\delta^{15}\text{N}$  fingerprint of the contamination could not be characterised. This uncertainty is propagated in the calculations of Sect. 3 and considered in the discussions. Possible isotopic changes resulting from the conversion and analysis process of  $\text{NO}_3^-$  and  $\text{NO}_2$  samples were evaluated using international  $\text{NO}_3^-$  and  $\text{NO}_2^-$  isotopic reference materials, respectively (Table S3). Accuracy of the analytical method was estimated as the standard deviation ( $\sigma$ ) of the residuals between measurements of the reference materials and their expected values. In our study, average measurement uncertainties in  $\delta^{15}\text{N}$ ,  $\delta^{17}\text{O}$ ,  $\delta^{18}\text{O}$ , and  $\Delta^{17}\text{O}$  were estimated to be  $\pm 0.3\%$ ,  $\pm 0.9\%$ ,  $\pm 1.3\%$ , and  $\pm 0.4\%$ , respectively, for  $\text{NO}_3^-$  samples and  $\pm 0.3\%$ ,  $\pm 0.4\%$ ,  $\pm 0.9\%$ , and  $\pm 0.3\%$ , respectively, for  $\text{NO}_2$  samples. Detailed information about the calibration procedure can be found in Morin et al. (2009) and in Albertin et al. (2021) for  $\text{NO}_3^-$  and  $\text{NO}_2$  samples, respectively.

### 2.3 Ancillary data

During atmospheric samplings, surface  $\text{NO}_x$  mixing ratios were measured at the study site using an incoherent broadband cavity-enhanced absorption spectrometer for  $\text{NO}_2$  (IBBCEAS; Barbero et al., 2020) and an optical-feedback cavity-enhanced absorption spectrometer for  $\text{NO}$  (OFCEAS; Richard et al., 2018).  $\text{PM}$  concentrations ( $\text{PM}_{10}$  and  $\text{PM}_{2.5}$ ) were monitored by an optical particle counter (GRIMM<sup>®</sup>, EDM 164). The  $\text{O}_3$  mixing ratio was monitored at the local air quality monitoring site located a kilometre north of the sampling site (Environnement SA<sup>®</sup>, O3 42M; <https://www.atmo-auvergnerhonealpes.fr/>, last access: 5 November 2021). Surface temperature ( $T_{\text{surface}}$ ) and relative humidity (RH) were measured by a portable logger (Tinytag, TGP-4500, Gemini Data Loggers) located at the air quality monitoring site. Vertical temperatures were measured from 11 similar loggers fixed along the Planpraz cable car (45°55'39" N, 6°51'55" E) from 1098 to 2021 m a.s.l. (data obtained from personal communications with Catherine Coulaud, IGE, 2021). The  $\text{NO}_2$  photolysis rate ( $J_{\text{NO}_2}$ ) was calculated for the two sampling periods using a photochemical box model (CiTTyCAT version 2.02; Galeazzo et al., 2018; Pugh et al., 2012) using the Fast-J photolysis scheme of Wild et al. (2000) and a surface albedo fixed to 0.65, a value representative of a snow-covered surface (average value between fresh and old snow; more details can be found in Text S1 in the Supplement).

### 2.4 Interpretation framework for isotopic signals

In this section, we briefly state the key concepts and equations necessary to interpret isotopic signals measured in  $\text{NO}_2$  and  $\text{NO}_3^-$ . A more detailed description and complete equation derivations can be found in the cited references.

#### 2.4.1 $\Delta^{17}\text{O}$ mass balance equations

Because  $\text{NO}_2$  and  $\text{NO}_3^-$  loss processes do not fractionate in terms of the oxygen mass-independent anomaly and considering that each source reaction induces a transfer of  $\Delta^{17}\text{O}$  to  $\text{NO}_2$  and  $\text{NO}_3^-$ , one considers the mass conservation of  $\Delta^{17}\text{O}$  during fractionation processes in the  $N_r$  cycle. Hence, one can implement  $\Delta^{17}\text{O}$  in the general mass balance equation of  $\text{NO}_2$  and  $\text{NO}_3^-$ . An overall expression of the time derivative of  $\Delta^{17}\text{O}$  in the species  $X$  ( $\Delta^{17}\text{O}(X)$ , with  $X$  being  $\text{NO}_2$  or  $\text{NO}_3^-$ ) is derived as a function of its deviation from  $\Delta^{17}\text{O}$  transferred through each production channel  $i$  ( $P_i$ ) ( $\Delta^{17}\text{O}_i(X)$ ), weighted according to the relative contributions of the production channels (Vicars et al., 2013):

$$\frac{d}{dt}(\Delta^{17}\text{O}(X)) = \frac{1}{\tau(X)} \times \sum_i \frac{P_i}{\sum_i P_i} \times (\Delta^{17}\text{O}_i(X) - \Delta^{17}\text{O}(X)), \quad (2)$$

where  $P_i$  expresses reaction rate constant times of the atmospheric concentrations of reacting species and  $\tau$  is the atmospheric lifetime of the species  $X$  at steady state ( $\tau = [X]/\sum_i P_i$ , with  $[X]$  being the atmospheric mixing ratio of the species  $X$ ).

During the day, the rapid photochemical cycling of  $\text{NO}_x$  (Reactions R1–R4) leads to an isotopic equilibrium between  $\text{NO}$  and  $\text{NO}_2$ ; i.e.  $\Delta^{17}\text{O}(\text{NO}) \approx \Delta^{17}\text{O}(\text{NO}_2)$  (Michalski et al., 2014). Therefore, using the steady-state approximation and considering  $\text{NO} + \text{O}_3$  (Reaction R3) and  $\text{NO} + \text{RO}_2$  (Reaction R4) the main sources of  $\text{NO}_2$  at our site, the overall daytime  $\Delta^{17}\text{O}$  in  $\text{NO}_2$  can be expressed by

$$\Delta^{17}\text{O}_{\text{day}}(\text{NO}_2) \approx T_{\text{NO}+\text{O}_3} \times \Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2), \quad (3)$$

where  $\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2)$  is the  $\text{O}_3$  isotopic anomaly transferred to  $\text{NO}$  through Reaction (R3) (Savarino et al., 2008).  $T_{\text{NO}+\text{O}_3}$ , sometimes named  $A$  in the literature, represents the proportion of  $\text{O}$  atoms originating from  $\text{O}_3$  in  $\text{NO}_2$ , hence the relative importance of Reaction (R3) in the conversion of  $\text{NO}$  into  $\text{NO}_2$  (Michalski et al., 2003; Morin et al., 2007b; Albertin et al., 2021):

$$T_{\text{NO}+\text{O}_3} = \frac{k_{\text{NO}+\text{O}_3}[\text{O}_3]}{k_{\text{NO}+\text{O}_3}[\text{O}_3] + k_{\text{NO}+\text{RO}_2}[\text{RO}_2]}, \quad (4)$$

where  $k_{\text{NO}+\text{O}_3}$  and  $k_{\text{NO}+\text{RO}_2}$  are the kinetic constants of Reactions (R3) and (R4), respectively. The kinetic constants used in this study are listed in Table A1 in Appendix A. At night, considering that (1)  $\Delta^{17}\text{O}(\text{NO}) \approx 0\%$  ( $\text{NO}_x$  emission without  $\text{NO}_2$  recycling), (2) there is no  $\Delta^{17}\text{O}$  equilibrium between  $\text{NO}$  and  $\text{NO}_2$  (no photochemical cycling), and (3)  $\text{O}_3$  is the main oxidant of  $\text{NO}$  (no nighttime production of  $\text{RO}_2$ ),  $\Delta^{17}\text{O}(\text{NO}_2)$  is determined by the  $\Delta^{17}\text{O}$  transfer via Reaction (R3) and by the nighttime residuals of  $\text{NO}_2$  formed during the previous daytime hours (Albertin et al., 2021) following

$$\Delta^{17}\text{O}_{\text{night}}(\text{NO}_2) \approx x \times \Delta^{17}\text{O}_{\text{day}}(\text{NO}_2) + \frac{(1-x)}{2} \times (\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2) + \Delta^{17}\text{O}(\text{NO})), \quad (5)$$

where  $x$  is the  $\text{NO}_2$  formed during the day as a fraction of the total  $\text{NO}_2$  measured at night.

At our sampling site, we hypothesise that Reaction (R5) (OH pathway) and Reactions (R6)–(R8) ( $\text{N}_2\text{O}_5$  pathway) are the main daytime and nighttime  $\text{NO}_3^-$  production channels, respectively. At steady state, from Eq. (2), we derive general expressions for  $\Delta^{17}\text{O}$  in  $\text{NO}_3^-$  during the day and night, associated with the OH and  $\text{N}_2\text{O}_5$  pathways, respectively

(Alexander et al., 2020):

$$\Delta^{17}\text{O}_{\text{day}}(\text{NO}_3^-) \approx \frac{2}{3} \times \Delta^{17}\text{O}_{\text{day}}(\text{NO}_2), \quad (6)$$

$$\Delta^{17}\text{O}_{\text{night}}(\text{NO}_3^-) \approx \frac{2}{3} \times \Delta^{17}\text{O}_{\text{night}}(\text{NO}_2) + \frac{1}{6} \times \Delta^{17}\text{O}_{\text{NO}_2+\text{O}_3}(\text{NO}_3), \quad (7)$$

where  $\Delta^{17}\text{O}_{\text{NO}_2+\text{O}_3}(\text{NO}_3)$  is the  $^{17}\text{O}$  excess transfer from  $\text{O}_3$  to  $\text{NO}_3$  during Reaction (R6) (Berhanu et al., 2012). Without wet scavenging, dry deposition is the main sink of  $\text{NO}_3^-$  (Park et al., 2004). Assuming a mean  $\text{NO}_3^-$  deposition velocity of  $0.5 \text{ cm s}^{-1}$  (mean value of the dry-deposition velocities of  $\text{HNO}_3$  and  $p\text{-NO}_3^-$ ; Zhang et al., 2009), and considering the maximum daytime and minimum nighttime boundary layer heights of 500 and 100 m a.g.l., respectively (estimations based on measured vertical temperature profiles; Fig. S2 in the Supplement), the estimated residence time of  $\text{NO}_3^-$  against dry deposition can reach up to 28 h during the day and 6 h at night (Table B1 in Appendix B). Therefore, on sub-daily timescales, the  $^{17}\text{O}$  excess in  $\text{NO}_3^-$  during the day is more likely to reflect a combination of daytime and nighttime production processes than that during the night. Note that our estimated residence times for  $\text{NO}_3^-$  against dry deposition are upper limits as they represent the time required to reduce by a factor  $e$  the concentration of  $\text{NO}_3^-$  present at the top of the boundary layer;  $\text{NO}_3^-$  close to the surface would have a much shorter residence time.

#### 2.4.2 Nitrogen isotopic fractionation effects

Each source of  $\text{NO}_x$  generates a  $\delta^{15}\text{N}$  fingerprint which depends on the type and conditions (temperature and pressure) of combustion and on the type of fuel (e.g. coal, oil, gas) (Heaton, 1990; Felix et al., 2012; Fibiger and Hastings, 2016; Walters et al., 2015a, b; Yu and Elliott, 2017; Miller et al., 2018). The mean  $\delta^{15}\text{N}$  of  $\text{NO}_x$  ( $\delta^{15}\text{N}(\text{NO}_x)$ ) emitted in the atmosphere results from the sum of each  $\text{NO}_x$  emission  $\delta^{15}\text{N}$  fingerprint weighted by their relative contribution to the total  $\text{NO}_x$  emissions. Once in the atmosphere,  $\text{NO}_x$  is subjected to oxidation processes and isotopic exchanges that alter the initial  $\delta^{15}\text{N}(\text{NO}_x)$ . As a result,  $\delta^{15}\text{N}$  in  $\text{NO}_2$  and in  $\text{NO}_3^-$  is a complex function of both the  $\delta^{15}\text{N}$  signature of  $\text{NO}_x$  emissions and N isotopic effects. These latter can be categorised into three groups: (1) the equilibrium isotope effect (EIE), (2) the kinetic isotope effect (KIE), and (3) the photochemical isotope fractionation effect (PHIFE) (Miller and Yung, 2000; Young et al., 2002). The magnitude of these isotopic effects is quantified as the  $^{15}\text{N}$  enrichment factor ( $\epsilon$ ), which is defined as  $(\alpha - 1)$ , where  $\alpha$  represents the N isotopic fractionation factor.

A general expression for  $\delta^{15}\text{N}(\text{NO}_2)$  can be derived as a function of a factor  $F_N$ , which represents the overall N isotopic fractionation effects between  $\text{NO}_x$  emissions and  $\text{NO}_2$  (expressed in ‰), the fraction of  $\text{NO}_2$  with respect to

$\text{NO}_x$  ( $f_{\text{NO}_2} = [\text{NO}_2]/[\text{NO}_x]$ ), and  $\delta^{15}\text{N}(\text{NO}_x)$  (Albertin et al., 2021; Li et al., 2020):

$$\delta^{15}\text{N}(\text{NO}_2) = F_N \times (1 - f_{\text{NO}_2}) + \delta^{15}\text{N}(\text{NO}_x). \quad (8)$$

Therefore, the  $^{15}\text{N}$  isotopic shift between  $\delta^{15}\text{N}(\text{NO}_2)$  and  $\delta^{15}\text{N}(\text{NO}_x)$  is given by

$$\begin{aligned} \delta^{15}\text{N}(\text{NO}_2) - \delta^{15}\text{N}(\text{NO}_x) &= \Delta^{15}(\text{NO}_2 - \text{NO}_x) \\ &= F_N \times (1 - f_{\text{NO}_2}). \end{aligned} \quad (9)$$

Physico-chemical processes between  $\text{NO}$  and  $\text{NO}_2$  can preferentially promote or deplete  $^{15}\text{N}$  in  $\text{NO}_2$  with respect to emissions of  $\text{NO}_x$  (i.e.  $\delta^{15}\text{N}(\text{NO}_2) \neq \delta^{15}\text{N}(\text{NO}_x)$ ). The importance of this fractionation shift is modulated by the factor  $(1 - f_{\text{NO}_2})$ . When  $\text{NO}$  is almost entirely converted into  $\text{NO}_2$  ( $f_{\text{NO}_2} \approx 1$ ), N fractionation effects can be neglected (i.e.  $\delta^{15}\text{N}(\text{NO}_2) \approx \delta^{15}\text{N}(\text{NO}_x)$ ).

From samples collected at Jülich, Germany, Freyer et al. (1993), observed for the first time the linear relation described by Eq. (8) and set the theoretical framework to interpret  $\delta^{15}\text{N}$  variabilities in atmospheric  $\text{NO}_2$ . They showed that the observed seasonal variation in  $\delta^{15}\text{N}(\text{NO}_2)$  was driven by N fractionation effects (represented in the  $F_N$  factor) caused by photochemistry and isotopic equilibrium. Based on this work, as well as that of Li et al. (2020), Albertin et al. (2021) derived an expression of  $F_N$  during the day assuming that the  $\text{NO}\text{--}\text{NO}_2$  system is in isotopic equilibrium (steady state):

$$(F_N)_{\text{day}} \approx \frac{\alpha_{\text{LCIE}}^* A_{\text{day}}^* + (\alpha_{\text{EIE}(\text{NO}_2/\text{NO})} - 1)}{A_{\text{day}}^* + 1}, \quad (10)$$

with  $\alpha_{\text{LCIE}}^* = \alpha_{\text{KIE}(\text{NO}+\text{O}_3)} - \alpha_{\text{PHIFE}}$

$$\text{and } A_{\text{day}}^* = \frac{J_{\text{NO}_2}}{k_{\text{NO}+\text{NO}_2} [\text{NO}]},$$

where  $\alpha_{\text{LCIE}}^*$  is the fractionation factor of combined KIE and PHIFE (LCIE stands for Leighton cycle isotope effect) and  $\alpha_{\text{EIE}(\text{NO}_2/\text{NO})}$  is the EIE fractionation factor between  $\text{NO}$  and  $\text{NO}_2$ .  $\alpha_{\text{EIE}(\text{NO}_2/\text{NO})}$  and  $\alpha_{\text{KIE}(\text{NO}+\text{O}_3)}$  are temperature dependent and can be calculated following the theoretical approach of Walters and Michalski (2015) (Table D1 in Appendix D). From calculations based on the zero-point energy of  $^{15}\text{NO}_2$  and the absorption cross section of  $^{14}\text{NO}_2$ ,  $\alpha_{\text{PHIFE}}$  is estimated to vary between 1.0020 and 1.0042 for a range of solar zenith angles between 90 and 0° (Fang et al., 2021). In this study we use a mean value of  $\alpha_{\text{PHIFE}}$  at 1.0031.  $A_{\text{day}}^*$  is defined as the ratio of the  $\text{NO}_2$  lifetime with respect to isotopic exchanges over the daytime  $\text{NO}_2$  chemistry lifetime.  $J_{\text{NO}_2}$  is the  $\text{NO}_2$  photolysis rate,  $k_{\text{NO}+\text{O}_3}$  is the rate constant of Reaction (R3), and  $k_{\text{NO}+\text{NO}_2}$  is the rate constant of the isotopic exchange  $^{15}\text{NO}_2 + ^{14}\text{NO} \rightarrow ^{14}\text{NO}_2 + ^{15}\text{NO}$ . During the day,  $\Delta^{15}(\text{NO}_2 - \text{NO}_x)$  varies according to the environmental conditions. In low- $\text{NO}_x$  conditions (e.g. remote and polar regions)  $\Delta^{15}(\text{NO}_2 - \text{NO}_x)$  is predicted to be controlled by



LCIE factors ( $A_{\text{day}}^* \gg 1$ ), whereas an EIE-dominated regime ( $A_{\text{day}}^* \ll 1$ ) is expected in polluted environments (high- $\text{NO}_x$  conditions). At night,  $J_{\text{NO}_2}$  and  $\alpha_{\text{PHIFE}}$  are null and  $A_{\text{night}}^*$  is defined as the ratio of the NO lifetime with respect to isotopic exchange with  $\text{NO}_2$  to the NO chemical lifetime at night ( $A_{\text{night}}^* = \frac{k_{\text{NO}+\text{O}_3}[\text{O}_3]}{k_{\text{NO}+\text{NO}_2}[\text{NO}_2]}$ ). In this study, we consider only one particular case with  $A_{\text{night}}^* \ll 1$ , which means that isotopic exchanges are much faster than NO oxidation by  $\text{O}_3$ . In this scenario, KIEs are negligible compared to EIEs and  $(F_{\text{N}})_{\text{night}}$  can be expressed as

$$(F_{\text{N}})_{\text{night}} \approx \frac{(\alpha_{\text{EIE}(\text{NO}_2/\text{NO})} - 1)}{\alpha_{\text{EIE}(\text{NO}_2/\text{NO})}}. \quad (11)$$

The complete derivation of Eqs. (10) and (11) is given in Albertin et al. (2021).

KIE and EIE are also expected during the conversion of  $\text{NO}_2$  to  $\text{NO}_3^-$ . The  $^{15}\text{N}$  partitioning associated with isotopic equilibrium between  $\text{N}_2\text{O}_5$  and  $\text{NO}_2$  (Reaction R7) can be theoretically computed as a function of temperature (Walters and Michalski, 2015; Table D1). At 298 K, if N isotopic equilibrium is reached,  $\text{N}_2\text{O}_5$  is predicted to have  $\delta^{15}\text{N}$  values 27.6‰ higher than  $\text{NO}_2$ . Considering that the  $\text{NO}_2/\text{NO}_3^-$  isotopic fractionation through the  $\text{N}_2\text{O}_5$  pathway is solely controlled by EIE,  $\text{NO}_3^-$  is therefore expected to be enriched in  $^{15}\text{N}$  relative to  $\text{NO}_2$ . However, to date, no experimental study has reported on  $^{15}\text{N}$  partitioning between atmospheric  $\text{NO}_2$  and  $\text{NO}_3^-$ , and the fractionation factors are still being debated (Freyer, 1991; Fang et al., 2021). The dominant  $\text{NO}_x$ -to- $\text{NO}_3^-$  conversion processes considered in this study, along with corresponding  $\Delta^{17}\text{O}$  transfer factors and the known  $^{15}\text{N}$  enrichment factors at 298 K (determined from both experimental and computational studies), are illustrated in Fig. C1 in Appendix C.

### 3 Results and discussion

#### 3.1 Temporal variations in general atmospheric observations

Surface temperatures during SP 1 and SP 2 show similar values, with a marked diurnal cycle (from  $-2$  to  $16^\circ\text{C}$ ; Fig. 1). A positive temperature gradient with altitude is observed from the late evening to morning. Surface temperature rises around midday and reaches a maximum at around 15:00 LT, resulting in a negative temperature gradient with altitude. In deep alpine valleys, the diurnal variability in surface air temperature is strongly influenced by the temporal evolution of the boundary layer structure, particularly in winter with the formation of a surface layer inversion (Whiteman, 1982). As previously observed in Chamonix (Chazette et al., 2005), the nocturnal surface layer inversion regularly thickens during the night of the sampling periods. After sunrise, air masses warm up until the nocturnal inversion layer breaks down in the late morning. Observed RH behaviour relatively corre-

lates with the increase in temperature during the day, showing a rapid decrease between the mid-morning and early afternoon (from 96 % to 23 % and from 96 % to 30 % for SP 1 and SP 2, respectively; Fig. 1).

During SP 1 and SP 2, the mixing ratios of NO,  $\text{NO}_2$ , and  $\text{O}_3$  exhibit diurnal patterns (Fig. 1) that are typical in Chamonix in late February (Fig. S3) and more generally in urban areas (Mayer, 1999). The highest NO level is observed in the morning, peaking at around 10:30 LT (82 and  $152 \text{ nmol mol}^{-1}$  for SP 1 and SP 2, respectively). The concurrent increase in  $\text{NO}_2$  and decrease in  $\text{O}_3$  (down to  $1 \text{ nmol mol}^{-1}$ ) can be attributed to Reaction (R3) ( $\text{NO} + \text{O}_3$ ).  $\text{NO}_x$  decreases in the late morning, likely due to a combination of lower emissions,  $\text{NO}_2$  oxidation, and the dilution effect. Meanwhile,  $\text{O}_3$  gradually recovers to ca.  $30 \text{ nmol mol}^{-1}$ , a typical winter background air level in Europe (Gaudel et al., 2018). Due to local emissions,  $\text{NO}_x$  increases again from 16:00 LT, resulting in  $\text{O}_3$  titration, and subsequently to an increase in  $\text{NO}_2$  (up to  $40 \text{ nmol mol}^{-1}$ ). After 18:30 LT, NO remains low until the morning, and  $\text{NO}_2$  decreases slowly until midnight, stalls around  $10 \text{ nmol mol}^{-1}$ , and then rises again at 05:30 LT. After the late-afternoon titration,  $\text{O}_3$  gently recovers and stays relatively low throughout the night, likely due to a titration effect from nocturnal NO emissions which are confined in the surface layer inversion.

Both SP 1 and SP 2 show diurnal variations in PM mass concentrations (Fig. 1), with morning and evening peaks related to local emissions from traffic and home heating (Aymoz et al., 2007).  $\text{PM}_{10}$  concentrations display an additional increase at midday following the breakdown of the temperature inversion. At 12:30 LT, while  $\text{PM}_{10}$  concentrations increase moderately during SP 1 to reached  $38.6 \mu\text{g m}^{-3}$ , a sharp increase to  $119 \mu\text{g m}^{-3}$  is observed during SP 2. Then  $\text{PM}_{10}$  concentrations decrease during the afternoon of SP 1 but remain high during the afternoon of SP 2 until the surface inversion layer forms. On average,  $\text{PM}_{10}$  concentration is 3 times higher during SP 2 ( $59.4 \pm 37.6 \mu\text{g m}^{-3}$ ) than during SP 1 ( $20.6 \pm 10.2 \mu\text{g m}^{-3}$ ). The considerable increase in  $\text{PM}_{10}$  concentrations between SP 1 and SP 2 is likely to be explained by a Saharan dust episode that started on 23 February (Figs. S3, S4, and S5). Saharan dust deposition is a well-known phenomenon in the Alps and is characterised by a sudden increase in coarse particles, mainly composed of alumino-silicates as well as calcium and potassium (Angelisi and Gaudichet, 1991; Delmas, 1994; Di Mauro et al., 2019; Goudie and Middleton, 2001; Greilinger et al., 2018; Schwikowski et al., 1995; Sodemann et al., 2006).

The  $\text{NO}_3^-$  mass concentration varies from 0.3 to  $3.4 \mu\text{g m}^{-3}$ , with an average of  $(0.9 \pm 0.6) \mu\text{g m}^{-3}$  for SP 1 and of  $(1.2 \pm 0.9) \mu\text{g m}^{-3}$  for SP 2 (Fig. 1). During both sampling periods,  $\text{NO}_3^-$  concentration is within the range of previous observations made in Chamonix in winter (Allard, 2018).  $\text{NO}_3^-$  shows a distinctive peak at  $3.4 \mu\text{g m}^{-3}$  during SP 2 between 10:30 and 13:30 LT, correlated with the

PM<sub>10</sub> surge. During transport, dust can undergo heterogeneous uptake and conversion of gases on its surface, leading to the inclusion of secondary species such as NO<sub>3</sub><sup>-</sup>, sulfate, and ammonium (Usher et al., 2003). NO<sub>3</sub><sup>-</sup> on dust results mainly from HNO<sub>3</sub> uptake and heterogeneous reactions of N<sub>2</sub>O<sub>5</sub> (see Usher et al., 2003, for a review and also references therein). Mineral dust is believed to significantly contribute to NO<sub>3</sub><sup>-</sup> formation and size distribution, particularly in regions close to dust emission sources (Karydis et al., 2016). However, the origin of NO<sub>3</sub><sup>-</sup> during SP 2 at our site remains unclear and could be attributed to the advection of both nitrated-dust particles formed through heterogeneous processes during transport and anthropogenic fine particles (Aymoz et al., 2004).

Figure 2 shows the temporal evolution of measured  $\Delta^{17}\text{O}$  and  $\delta^{15}\text{N}$  of NO<sub>2</sub> and NO<sub>3</sub><sup>-</sup> in Chamonix during the two sampling periods (SP 1 and SP 2). All isotopic data used in this study are reported in Tables S1 and S2. In the following analysis, first we describe  $\Delta^{17}\text{O}(\text{NO}_2)$  measurements and give interpretation in light of NO<sub>x</sub> chemistry cycling (Sect. 3.2). Then, an analysis of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  measurements is proposed, aggregating daytime and nighttime periods and comparing them with  $\Delta^{17}\text{O}(\text{NO}_3^-)$  estimates derived from  $\Delta^{17}\text{O}(\text{NO}_2)$  measurements and  $\Delta^{17}\text{O}$  mass balance for major chemical processes (Sect. 3.3). In light of these results and atmospheric conditions during SP 1 and SP 2, sub-daily  $\Delta^{17}\text{O}(\text{NO}_3^-)$  dynamics are investigated. In Sect. 3.4, N fractionation effects in the NO<sub>x</sub> cycle are quantified and the dominant NO<sub>x</sub> emission source is identified. The dynamics of  $\delta^{15}\text{N}(\text{NO}_3^-)$  are also described, and its use to trace NO<sub>x</sub> emission sources and oxidation processes is discussed.

### 3.2 $\Delta^{17}\text{O}$ of NO<sub>2</sub> and NO<sub>x</sub> diurnal cycling

Over the course of SP 1 and SP 2,  $\Delta^{17}\text{O}(\text{NO}_2)$  shows a large diurnal variability (from 19.6‰ to 40.8‰) with a weighted mean  $\pm 1$  standard deviation of  $(25.2 \pm 7.1)\%$ .  $\Delta^{17}\text{O}(\text{NO}_2)$  values during the day (07:30–18:00 LT,  $(28.5 \pm 7.3)\%$ ) are significantly higher ( $p$  value = 0.002,  $n = 16$ ) than during the night (18:00–07:30 LT,  $(20.8 \pm 1.0)\%$ ). By day,  $\Delta^{17}\text{O}(\text{NO}_2)$  follows a similar increasing trend during SP 1 and SP 2, reaching a respective maximum of 40.8‰ between 13:30–16:30 LT and 35.0‰ between 10:30–13:30 LT. For both sampling periods, after sunset,  $\Delta^{17}\text{O}(\text{NO}_2)$  stabilises between 21:00 and 07:30 LT at ca. 20‰. Using the same sampling methodology in a mid-latitude urban area in spring, Albertin et al. (2021) reported very similar  $\Delta^{17}\text{O}(\text{NO}_2)$  values over the course of a day (20.5‰–39.2‰), following a comparable diurnal pattern. As presented above (Sect. 2.4.1), according to the ISS (isotopic steady state) framework, the variability in  $\Delta^{17}\text{O}(\text{NO}_2)$  reflects changes in the relative contributions of Reaction (R3) (NO + O<sub>3</sub>) and Reaction (R4) (NO + RO<sub>2</sub>) to the overall production of NO<sub>2</sub>. At our site,  $\Delta^{17}\text{O}(\text{NO}_2)$  drops rapidly during the 16:30–18:00 LT interval to 23.3‰ and 20.9‰ during SP 1 and SP 2, respectively.

Since the isotope recycling rate in the NO<sub>x</sub>–O<sub>3</sub> system is driven at first order by  $J_{\text{NO}_2}$  (Michalski et al., 2014), due to low solar radiation between 16:30–18:00 LT at our site, such a rapid drop of  $\Delta^{17}\text{O}(\text{NO}_2)$  suggests that sampled NO<sub>2</sub> might not be at ISS anymore when the NO<sub>2</sub> photolysis is very slow, notably at the end of the day in winter. Therefore, we only consider that ISS holds between 07:30 and 16:30 LT in order to avoid the questionable end-of-the-day measurements in our ISS-based analysis. At night (i.e. no ISS), the observed drop of  $\Delta^{17}\text{O}(\text{NO}_2)$  in the early evening reflects the rapid replacement of NO<sub>2</sub> formed during the day by NO<sub>2</sub> produced during the night via the conversion of freshly emitted NO, in line with Eq. (5). Then, high NO<sub>2</sub> throughout the night, along with relatively low O<sub>3</sub>, supports observations of low  $\Delta^{17}\text{O}(\text{NO}_2)$  at night (i.e.  $\approx 20\%$ ).

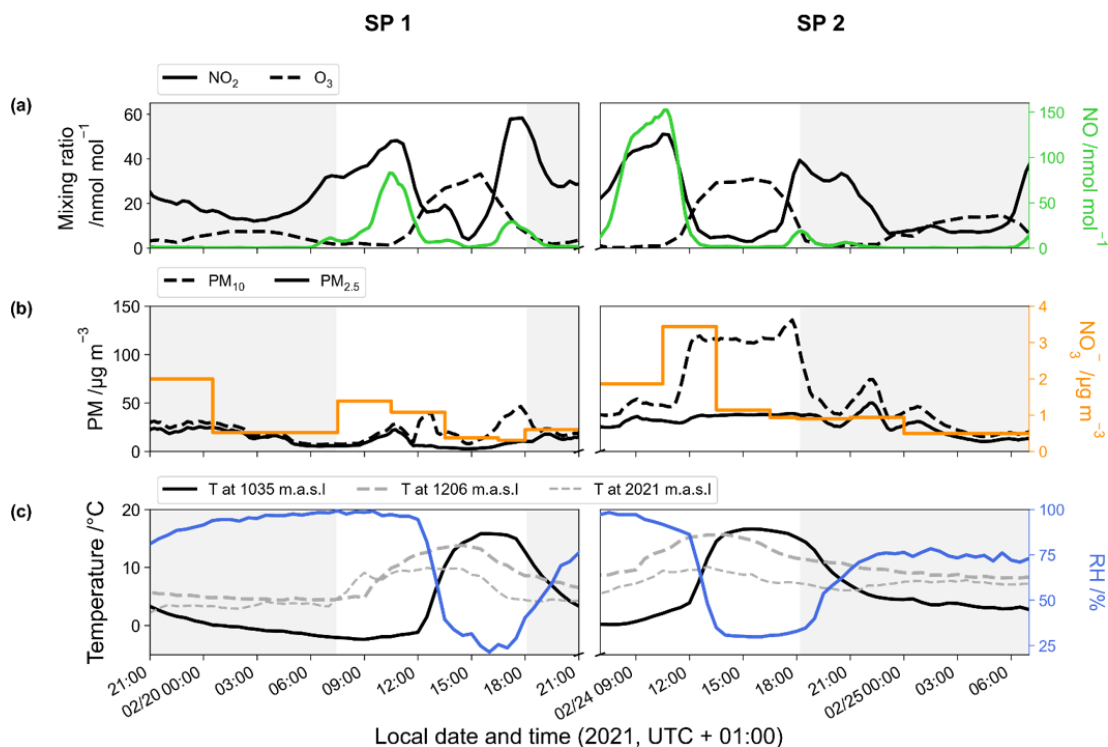
Using Eq. (3), we derive from  $\Delta^{17}\text{O}(\text{NO}_2)$  observations the relative contribution of Reaction (R3) (NO + O<sub>3</sub>) to Reaction (R4) (NO + RO<sub>2</sub>) in the formation of NO<sub>2</sub> ( $T_{\text{NO}+\text{O}_3}$ ):

$$T_{\text{NO}+\text{O}_3} = \frac{\Delta^{17}\text{O}_{\text{day}}(\text{NO}_2)}{\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2)}. \quad (12)$$

Between 07:30 and 16:30 LT,  $T_{\text{NO}+\text{O}_3}$  varies from 0.55 to 1.00 (Table 1), with a mean of 0.88 and 0.75 for SP 1 and SP 2, respectively. The NO + O<sub>3</sub> pathway is dominant between 13:30 and 16:30 LT, corresponding to the time when O<sub>3</sub> is highest (Fig. 1). In contrast, the maximum contribution for the NO + RO<sub>2</sub> pathway is observed between 07:30 and 10:30 LT, when NO levels are high and rising continuously. Interestingly, previous studies have reported a high sensitivity of RO<sub>2</sub> to changes in NO<sub>x</sub>, particularly at high NO<sub>x</sub> levels (Ren et al., 2006; Stone et al., 2012). Sources of RO<sub>2</sub> in wintertime are mainly driven by the production of OH radicals from HONO photolysis, alkene ozonolysis, and formaldehyde photolysis (Tan et al., 2018). During winter, HONO plays a crucial role in NO<sub>x</sub>/O<sub>3</sub>/RO<sub>2</sub> chemistry, particularly in the morning, as its photolysis can potentially accelerate daytime oxidation processes, leading to increased RO<sub>2</sub> production (Alicke et al., 2003; Aumont et al., 2003). Direct emissions from vehicle exhaust could be a significant source of VOCs and HONO at our site (Brulfert et al., 2005; Gu et al., 2019; Kirchstetter et al., 1996; Kurtenbach et al., 2001; Liu et al., 2023). Heterogeneous processes acting on ground surfaces and aerosols can also contribute to HONO formation (Aumont et al., 2003). In addition, snowpack releases may also be a potential source of HONO (Grannas et al., 2007), as detected in Paris after a snow event, which could significantly impact the urban OH budget (Michoud et al., 2015).

Following the approach of Albertin et al. (2021), combining Eqs. (3) and (4) allows us to derive the RO<sub>2</sub> mixing ratio from observed  $\Delta^{17}\text{O}(\text{NO}_2)$  and the O<sub>3</sub> mixing ratio following

$$[\text{RO}_2] = \frac{k_{\text{NO}+\text{O}_3}[\text{O}_3]}{k_{\text{NO}+\text{RO}_2}} \left( \frac{\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2)}{\Delta^{17}\text{O}_{\text{day}}(\text{NO}_2)} - 1 \right). \quad (13)$$



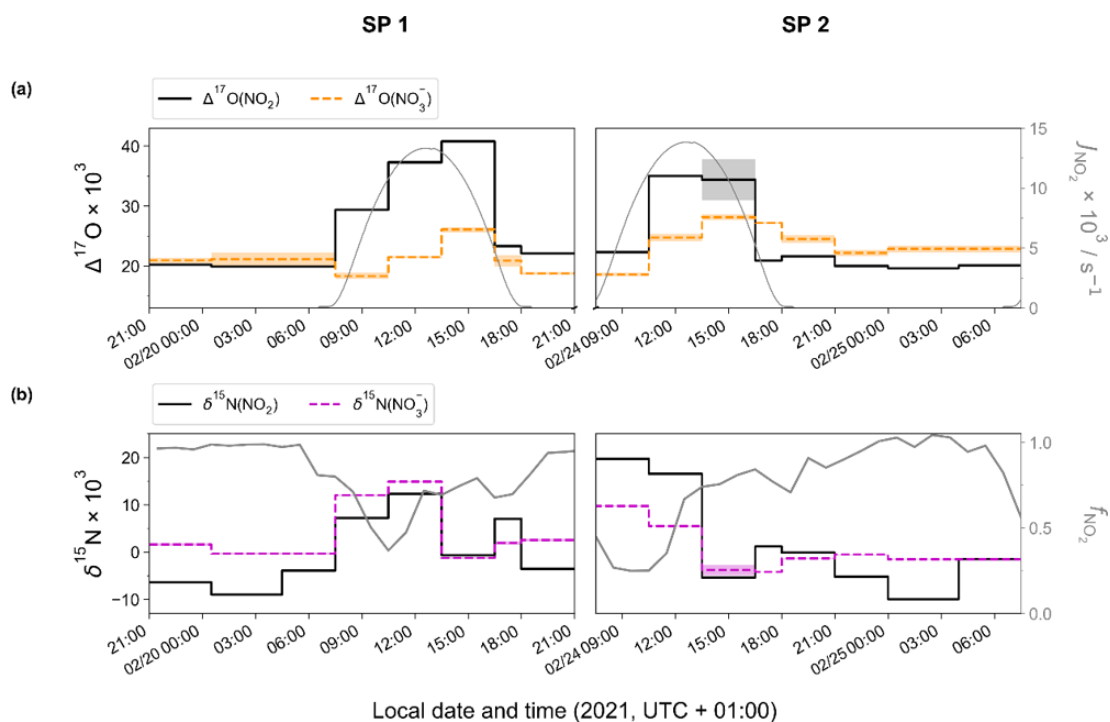
**Figure 1.** Temporal evolution of the 1 h rolling mean of (a)  $\text{NO}_2$  (black line),  $\text{O}_3$  (dashed line), and  $\text{NO}$  (green line) mixing ratios; (b)  $\text{PM}$  (dashed line for  $\text{PM}_{10}$  and solid line for  $\text{PM}_{2.5}$ ) and  $\text{NO}_3^-$  (orange line) mass concentrations; and (c) temperature at the surface (black line), at 1206 m (dashed grey line), and at 2021 m (light dashed grey line) and surface relative humidity (blue line). Data were collected during the two sampling periods (SP 1 and SP 2) in Chamonix. Grey-backdrop-shaded areas represent the nighttime (sunset to sunrise).

Between 07:30 and 16:30 LT, we estimate an average  $\text{RO}_2$  mixing ratio at our site of  $(0.88 \pm 0.88) \text{ pmol mol}^{-1}$  and  $(4.92 \pm 5.16) \text{ pmol mol}^{-1}$  during SP 1 and SP 2, respectively (Table 1). Studies conducted in urban winter environments reported  $\text{RO}_2$  measurements at a few  $\text{pmol mol}^{-1}$  (Ren et al., 2006; Emmerson et al., 2005; Tan et al., 2018; Kanaya et al., 2007), in good agreement with our estimations. Similarly,  $\text{RO}_2$  mixing ratios derived by Albertin et al. (2021) from observed  $\Delta^{17}\text{O}(\text{NO}_2)$  in spring (mean of  $(13.8 \pm 11.2) \text{ pmol mol}^{-1}$ ) were also found to be in line with studies conducted in the same season. The fact that our isotopic inference of  $\text{RO}_2$  mixing ratios carried out in two different seasons (winter and spring) are both comparable to direct in situ  $\text{RO}_2$  measurements confirms the sensitivity of our method in probing the  $\text{NO}_x/\text{O}_3/\text{RO}_2$  chemical dynamics. We think that our method may be very valuable in deciphering oxidation processes of  $\text{N}_r$  species, down to sub-daily temporal scales. Nonetheless, we recognise that without concurrent in situ measurements of  $\text{RO}_2$  and  $\Delta^{17}\text{O}(\text{NO}_2)$ , it is not possible to validate unambiguously the quantitative estimation of  $\text{RO}_2$  levels with our method.

At this point, it is important to recall that the choice of the  $\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2)$  in Eq. (12) is of particular importance for quantifying  $T_{\text{NO}+\text{O}_3}$  (as for  $\text{RO}_2$ ). In the literature,  $\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2)$  varies between 35‰ and 41‰

**Table 1.**  $T_{\text{NO}+\text{O}_3}$  and  $\text{RO}_2$  mixing ratio (mean value  $\pm$  overall uncertainty) derived from the isotopic measurements.

Sampling interval (start–end)	$T_{\text{NO}+\text{O}_3}$	$\text{RO}_2$ ( $\text{pmol mol}^{-1}$ )
SP 1		
20 Feb 07:30–20 Feb 10:30	$0.72 \pm 0.01$	$0.86 \pm 0.75$
20 Feb 10:30–20 Feb 13:30	$0.91 \pm 0.01$	$1.77 \pm 0.36$
20 Feb 13:30–20 Feb 16:30	$1.00 \pm 0.01$	$0.00 \pm 0.91$
Mean	0.88	0.88
SD	0.14	0.88
SP 2		
24 Feb 07:30–24 Feb 10:30	$0.55 \pm 0.01$	$0.58 \pm 1.67$
24 Feb 10:30–24 Feb 13:30	$0.86 \pm 0.01$	$3.56 \pm 0.50$
24 Feb 13:30–24 Feb 16:30	$0.84 \pm 0.08$	$10.63 \pm 6.75$
Mean	0.75	4.92
SD	0.18	5.16



**Figure 2.** Temporal evolution of observed (a)  $\Delta^{17}\text{O}$  and (b)  $\delta^{15}\text{N}$  of atmospheric  $\text{NO}_2$  (top and bottom solid black line) and  $\text{NO}_3^-$  (top orange and bottom dashed magenta line) in Chamonix (length of horizontal line denotes sampling period; shaded area denotes overall analytical error). The  $\text{NO}_2$  photolysis rate ( $J_{\text{NO}_2}$ , top grey line) is from CiTTYCAT box-model output.  $f_{\text{NO}_2}$  ( $f_{\text{NO}_2} = [\text{NO}_2]/([\text{NO}_2] + [\text{NO}])$ ; bottom grey line) is calculated from the hourly mean mixing ratio of  $\text{NO}$  and  $\text{NO}_2$ .

(Michalski et al., 2003; Savarino et al., 2016; Vicars et al., 2012; Y.-L. Zhang et al., 2022; Y. Li et al., 2022). This relatively wide range of values is partly a result of some confusion in defining  $\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2)$ , the  $\Delta^{17}\text{O}$  transfer from  $\text{O}_3$  to  $\text{NO}$ . Indeed, the term  $\Delta^{17}\text{O}(\text{O}_3^*)$  is sometimes erroneously used as the transfer function of  $\Delta^{17}\text{O}$  from  $\text{O}_3$  to  $\text{NO}_2$  during Reaction (R3).  $\Delta^{17}\text{O}(\text{O}_3^*)$  is actually defined as  $\Delta^{17}\text{O}$  of  $\text{O}_3$  terminal atoms and is also named  $\Delta^{17}\text{O}(\text{O}_3)_{\text{term}}$  in the literature. As the  $\Delta^{17}\text{O}$  in  $\text{O}_3$  is borne by its terminal atoms,  $\Delta^{17}\text{O}(\text{O}_3^*) = 1.5 \times \Delta^{17}\text{O}(\text{O}_3)_{\text{bulk}}$ . However,  $\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2)$  can be equal to  $1.5 \times \Delta^{17}\text{O}(\text{O}_3)_{\text{bulk}}$  if only terminal atoms of  $\text{O}_3$  react with  $\text{NO}$ . But laboratory experiments by Savarino et al. (2008) did show that  $\text{O}_3$  does react with  $\text{NO}$  not solely with its terminal atoms but also, to a small but significant extent, with its central atom (probability of  $(8 \pm 5)\%$  for the abstraction of central atoms during the reaction  $\text{NO} + \text{O}_3$ ). Consequently,  $\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2)$  is slightly lower than  $1.5 \times \Delta^{17}\text{O}(\text{O}_3)_{\text{bulk}}$  and the  $\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2)$  expression determined by Savarino et al. (2008) should be used:  $\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2) = 1.18 \pm 0.07 \times \Delta^{17}\text{O}(\text{O}_3)_{\text{bulk}} + (6.6 \pm 1.5)\%$ .

Assuming that their maximum measured daytime  $\Delta^{17}\text{O}(\text{NO}_2)$  reflects the conversion of  $\text{NO}$  to  $\text{NO}_2$  only through Reaction (R3) (i.e.  $T_{\text{NO}+\text{O}_3} = 1$ ), Albertin et al. (2021) derived a  $\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2)$  value of  $39.2\%$  from Eq. (3). Given the respective analytical uncertainties

(around  $\pm 1\%$ ), their value is in very good agreement with the maximum daytime value of  $40.8\%$  we observed in Chamonix. Similarly to Albertin et al. (2021), assuming that the highest daytime  $\Delta^{17}\text{O}(\text{NO}_2)$  value at our site corresponds to  $T_{\text{NO}+\text{O}_3} \approx 1$  leads to  $\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2) = 40.8\%$ . Using the experimental  $\Delta^{17}\text{O}_{\text{NO}+\text{O}_3}(\text{NO}_2)$  transfer function determined by Savarino et al. (2008), we estimate a bulk  $^{17}\text{O}$  excess of  $\text{O}_3$  ( $\Delta^{17}\text{O}(\text{O}_3)_{\text{bulk}}$ ) at  $(29.0 \pm 2.2)\%$ . This value is consistent with the range of direct  $\Delta^{17}\text{O}(\text{O}_3)_{\text{bulk}}$  measurements at mid-latitudes (mean of  $(26.2 \pm 1.3)\%$ ; Vicars and Savarino, 2014), although falling at the upper end of the range. Interestingly, Vicars and Savarino (2014) reported a significant peak in  $\Delta^{17}\text{O}(\text{O}_3)_{\text{bulk}}$  during February–March in Grenoble, France (located 120 km southwest of Chamonix), based on year-round measurements with  $\Delta^{17}\text{O}(\text{O}_3)_{\text{bulk}}$  values  $2\%$ – $3\%$  higher than the annual mean of  $26.2\%$ . Although the cause of increased values during this period is unknown, our derived  $\Delta^{17}\text{O}(\text{O}_3)_{\text{bulk}}$  matches remarkably well the February–March measurements reported by Vicars and Savarino (2014).

It is worth pointing out that a more accurate calibration of  $\Delta^{17}\text{O}(\text{NO}_2)$  measurements is desirable. There is currently no internationally accepted nitrite salt standard with a positive  $\Delta^{17}\text{O}$ . Nevertheless, we are rather confident in our present calibration methodology. Indeed, the  $\Delta^{17}\text{O}(\text{NO}_2)$  values measured in Chamonix closely align with previous obser-

variations in Grenoble. In both studies, the maximum (daytime) and minimum (nighttime)  $\Delta^{17}\text{O}(\text{NO}_2)$  measurements conform to the expected values derived from the  $\Delta^{17}\text{O}$  theoretical framework (based on well-established  $\text{NO}_x$  chemistry and  $\Delta^{17}\text{O}(\text{O}_3)$  measurements) when  $\text{O}_3$  overwhelmingly dominates the  $\text{NO}$ -to- $\text{NO}_2$  conversion. Therefore, at this stage, we do not consider that  $\Delta^{17}\text{O}(\text{NO}_2)$  calibration is an issue for this study. However, as we cannot completely rule out a small bias in our calibration, we have manufactured enriched nitrite salts and are presently working on refined new nitrite salt standards. Note that the rapid exchange of nitrite O isotopes in aqueous solution (Casciotti et al., 2007) is a challenge for inter-laboratory data comparisons. If new measurements of nitrite  $\Delta^{17}\text{O}$  references are found to differ from our current calibration, we plan to publish a correction to the paper with updated values and any potential implications.

### 3.3 Interpretation of $\Delta^{17}\text{O}$ in atmospheric nitrate

Over the two sampling periods,  $\Delta^{17}\text{O}(\text{NO}_3^-)$  varies significantly (from 18.3‰ to 28.1‰), with a weighted mean of  $(22.5 \pm 3.1)\%$ . While  $\Delta^{17}\text{O}(\text{NO}_2)$  values are relatively similar during the two sampling periods,  $\Delta^{17}\text{O}(\text{NO}_3^-)$  values are systematically higher during SP 2 than during SP 1, except during the 07:30–10:00 LT interval.  $\Delta^{17}\text{O}(\text{NO}_3^-)$  in Chamonix is in the same range as that of most previous observations in urban environments (9‰–44‰; e.g. Kim et al., 2023; Wang et al., 2023; Fan et al., 2023; Y.-L. Zhang et al., 2022; Lim et al., 2022; Z. Li et al., 2022) but lower than most values measured during the cold season, which are typically  $> 25\%$ . Unlike  $\Delta^{17}\text{O}(\text{NO}_2)$ , daytime and nighttime  $\Delta^{17}\text{O}(\text{NO}_3^-)$  values at our site are not significantly different ( $p$  value  $> 0.05$ ,  $n = 14$ ). Similarly, from 12 h resolved sampling in winter Beijing, He et al. (2018) found no significant difference between daytime and nocturnal  $\Delta^{17}\text{O}(\text{NO}_3^-)$  and suggest that each sample reflects  $\text{NO}_3^-$  produced during both the day and the night. From high-time-resolution (3 h) aerosol sampling in winter Beijing, Y.-L. Zhang et al. (2022) reported  $\Delta^{17}\text{O}(\text{NO}_3^-)$  values of between 23.4‰ and 39.3‰, with higher values observed at night ( $(31.0 \pm 2.6)\%$ ) than during the day ( $(29.3 \pm 3.0)\%$ ). This diurnal behaviour of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  was attributed to the changes in the branching ratio of nocturnal and photochemical reactions on  $\text{NO}_3^-$  formation. In Chamonix, the range of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  values is very different from Y.-L. Zhang et al. (2022) observations, with consistently lower values and a distinct diurnal tendency. However, in the cases of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  measurements at sub-daily temporal scale, the atmospheric lifetime of  $\text{NO}_x$  and  $\text{NO}_3^-$  is critical for comparing  $\Delta^{17}\text{O}(\text{NO}_3^-)$  records from one site to another. Pollutant levels and atmospheric conditions between Chamonix and Beijing are very different, notably in winter when Asian urban areas can experience severe haze pollution episodes with  $\text{NO}_3^-$  mass concentration exceeding  $70 \mu\text{g m}^{-3}$ , which is over 10 times higher than in

Chamonix (Lim et al., 2022; He et al., 2018; Y.-L. Zhang et al., 2022). In such conditions, PM can reach several hundreds of  $\mu\text{g m}^{-3}$  for several days, which can significantly impact atmospheric processes involved in the formation of secondary species. Aside from the intrusion of Saharan dust during SP 2, the pollutant level in Chamonix is indicative of a moderately polluted region, with significant diurnal variations.

#### 3.3.1 Steady-state evaluation of $\Delta^{17}\text{O}(\text{NO}_3^-)$

To investigate the factors influencing the variability in  $\Delta^{17}\text{O}(\text{NO}_3^-)$  at our site, one compares observed  $\Delta^{17}\text{O}(\text{NO}_3^-)$  with estimated values of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  derived from  $\Delta^{17}\text{O}$  mass balance and observed  $\Delta^{17}\text{O}(\text{NO}_2)$ , assuming the OH and  $\text{N}_2\text{O}_5$  pathways dominate the formation of  $\text{NO}_3^-$  at our site. Therefore, calculated  $\Delta^{17}\text{O}(\text{NO}_3^-)$  reflects the theoretical  $\Delta^{17}\text{O}$  transfer during the oxidation of  $\text{NO}_2$  to  $\text{NO}_3^-$  at our site through the dominant chemical process during the day (i.e. OH pathway) and at night (i.e.  $\text{N}_2\text{O}_5$  pathway).

As presented, during the day, we consider the conversion of  $\text{NO}_2$  into  $\text{NO}_3^-$  to be predominantly influenced by Reaction (R5) (OH pathway). Hence, the theoretical corresponding  $^{17}\text{O}$  excess transfer to  $\text{NO}_3^-$  is estimated using Eq. (6) and observed  $\Delta^{17}\text{O}(\text{NO}_2)$  between 07:30 and 18:00 LT ( $n = 3$  per sampling period). Then, in order to estimate a daytime average value of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  which is representative of the potential for the formation of  $\text{NO}_3^-$  from surface  $\text{NO}_2$  by the OH pathway, each calculated  $\Delta^{17}\text{O}(\text{NO}_3^-)$  is weighted by the product  $[\text{NO}_2] \times J_{\text{NO}_2}$  (the diurnal variability in the OH mixing ratio is assumed to follow the diurnal  $J_{\text{NO}_2}$  variation; Liu et al., 2021). Finally, an overall mean daytime  $\Delta^{17}\text{O}(\text{NO}_3^-)$  for SP 1 and SP 2 is estimated by taking the sum of the weighted calculated values ( $= \Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-)$ ). The same approach is used during the night, assuming that the conversion of  $\text{NO}_2$  into  $\text{NO}_3^-$  is dominated by Reactions (R6)–(R8) ( $\text{N}_2\text{O}_5$  pathway). Equation (7) and observed  $\Delta^{17}\text{O}(\text{NO}_2)$  between 18:00 and 07:30 LT ( $n = 3$  per sampling day) are used to estimate  $\Delta^{17}\text{O}(\text{NO}_3^-)$ . Each calculated  $\Delta^{17}\text{O}(\text{NO}_3^-)$  is weighted by the product  $[\text{NO}_2] \times [\text{O}_3]$  (i.e.  $\text{NO}_3^-$  production rate) and summed to estimate a mean nighttime  $\Delta^{17}\text{O}(\text{NO}_3^-)$  for SP 1 and SP 2. The  $^{17}\text{O}$  excess transferred from  $\text{O}_3$  to  $\text{NO}_2$  during Reaction (R6) ( $\Delta^{17}\text{O}_{\text{NO}_2+\text{O}_3}(\text{NO}_3)$ ) is fixed at 44.7‰. This value is set according to the transfer function reported by Berhanu et al. (2012), whereby  $\Delta^{17}\text{O}_{\text{NO}_2+\text{O}_3}(\text{NO}_3) = (1.23 \pm 0.19) \times \Delta^{17}\text{O}(\text{O}_3)_{\text{bulk}} + (9.02 \pm 0.99)$  and  $\Delta^{17}\text{O}(\text{O}_3)_{\text{bulk}} = 29.0\%$  (see Sect. 2.4.1). We compare hereafter  $\Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-)$  with the weighted day and night averages of observed  $\Delta^{17}\text{O}(\text{NO}_3^-)$  at our site. During the day,  $\Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-)$  is compared with  $\Delta^{17}\text{O}(\text{NO}_3^-)$  observations averaged between 07:30 and 18:30 LT ( $n = 3$ ). At night,  $\Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-)$  is compared with  $\Delta^{17}\text{O}(\text{NO}_3^-)$  observations averaged between 18:30 and 07:30 LT.

At night during SP 1, observed  $\Delta^{17}\text{O}(\text{NO}_3^-)$  and  $\Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-)$  are in good agreement ( $\Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-) - \Delta^{17}\text{O}(\text{NO}_3^-) = \Delta^{17}(\text{NO}_3^-_{\text{calc}} - \text{NO}_3^-_{\text{obs}}) = 0.9\%$ ), suggesting a local and rapid (< 12 h) conversion of  $\text{NO}_2$  into  $\text{NO}_3^-$  via the  $\text{N}_2\text{O}_5$  pathway. During the day, observed  $\Delta^{17}\text{O}(\text{NO}_3^-)$  is 0.5% higher than  $\Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-)$ , also suggesting that  $\text{NO}_3^-$  is formed locally during the day for oxidation of surface  $\text{NO}_2$  through the OH pathway. Small differences between observed and calculated  $\Delta^{17}\text{O}$  of  $\text{NO}_3^-$  during the day and night could be explained by the presence of  $\text{NO}_3^-$  residues formed during the previous night and day, respectively, which are not considered in the calculations since they do not account for the  $\text{NO}_3^-$  lifetime. In contrast to SP 1,  $\Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-)$  during SP 2 is significantly lower than the mean observed  $\Delta^{17}\text{O}(\text{NO}_3^-)$ , particularly during the day with a  $\Delta^{17}(\text{NO}_3^-_{\text{calc}} - \text{NO}_3^-_{\text{obs}})$  of  $-6.4\%$ . The significant gap between observed and calculated  $\Delta^{17}\text{O}(\text{NO}_3^-)$  suggests a different origin and/or formation process of  $\text{NO}_3^-$  during SP 2 compared to SP 1. Although less important than during the day,  $\Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-)$  values for SP 2 at night are lower by 2.2% compared to the observed value. This small shift can be explained by residuals of enriched daytime  $\text{NO}_3^-$ . It is important to point out that, although the  $\text{NO}_2$  sample collected on 24 February between 13:30 and 16:30 LT presents an important blank (ca. 14%), ambient  $\text{NO}_2$  is low during the sampling period (mean of  $(4.5 \pm 1.8) \text{ nmol mol}^{-1}$ ). Therefore, as each  $\Delta^{17}\text{O}$  value used to estimate  $\Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-)$  is weighted by the mean ambient  $\text{NO}_2$  mixing ratio over the sampling period, the uncertainty related to this blank has little influence on the daily average of  $\Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-)$ . Given the low  $\Delta^{17}(\text{NO}_3^-_{\text{calc}} - \text{NO}_3^-_{\text{obs}})$  during SP 1, observed  $\Delta^{17}\text{O}(\text{NO}_3^-)$  can be explained by the local and rapid (< 12 h) oxidation of  $\text{NO}_2$ , dominated by the OH and  $\text{N}_2\text{O}_5$  pathway during the day and night, respectively. However, in contrast to SP 1, the  $^{17}\text{O}$  excess measured in  $\text{NO}_3^-$  during the day of SP 2 cannot be fully constrained by the oxidation of surface  $\text{NO}_2$  through the OH pathway, suggesting that the formation mechanisms of  $\text{NO}_3^-$  are different between SP 1 and SP 2 and/or the presence of  $\text{NO}_3^-$  not formed locally during SP 2. Below we examine the changes in the sub-daily dynamics of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  between SP 1 and SP 2 in light of atmospheric observations.

### 3.3.2 $\Delta^{17}(\text{NO}_3^-)$ sub-daily dynamics

Between 07:30–10:30 LT,  $\Delta^{17}\text{O}(\text{NO}_3^-)$  is very similar during SP 1 (18.3%) and SP 2 (18.6%). Nonetheless, in the following sampling time step (i.e. between 10:30–13:30 LT),  $\Delta^{17}\text{O}(\text{NO}_3^-)$  is significantly different between SP 1 (21.5%) and SP 2 (24.7%). Intriguingly, between 07:30–13:30 LT,  $\Delta^{17}\text{O}(\text{NO}_2)$  during SP 2 (26.9%) is lower compared to SP 1 (32.8%). As a consequence, if one considers the notion that  $\text{NO}_3^-$  is formed from the oxidation of local  $\text{NO}_2$  through identical pathways during SP 1 and SP 2, observed  $\Delta^{17}\text{O}(\text{NO}_3^-)$  should be lower during SP 2 than during

SP 1. Therefore, the more pronounced increase in observed  $\Delta^{17}\text{O}(\text{NO}_3^-)$  during the 10:30–12:30 LT interval of SP 2 suggests a different origin and/or formation channel of  $\text{NO}_3^-$ , as mentioned previously. One more piece of evidence is that, during this period of time,  $\text{PM}_{10}$  and  $\text{NO}_3^-$  levels increase significantly during SP 2, alongside the disruption of the inversion layer (depicted in Fig. 1). It can be inferred that this rise in  $\text{PM}_{10}$  is mostly due to the presence of Saharan dust. The simultaneous increase in  $\text{NO}_3^-$  and of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  corroborates the hypothesis that this  $\text{NO}_3^-$  was not formed from the oxidation of ambient  $\text{NO}_2$ . Furthermore, such an increase in  $\Delta^{17}\text{O}(\text{NO}_3^-)$  can only be supported by the oxidation of  $\text{NO}_2$  through the  $\text{N}_2\text{O}_5$  pathway, which is not expected to be important during the day due to the rapid photolysis of  $\text{NO}_3$  and its titration by NO (Brown and Stutz, 2012).

Interestingly, aerosol samplings conducted at various heights (8, 120, and 260 m a.g.l.) in Beijing, China, revealed a positive vertical gradient of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  in winter, from on average 29% to 33% (Fan et al., 2022). In summer, the  $\Delta^{17}\text{O}(\text{NO}_3^-)$  values at the three altitudes were very similar. This increase in  $\Delta^{17}\text{O}(\text{NO}_3^-)$  with altitude in winter was believed to result from a stratification of  $\text{NO}_2$ -to- $\text{NO}_3^-$  oxidation processes due to low vertical mixing and elevated surface  $\text{NO}_x$  emissions. However, the authors did not consider the potential variability in  $\Delta^{17}\text{O}(\text{NO}_2)$  with altitude, which can be substantial in urban areas at night as low  $\Delta^{17}\text{O}(\text{NO}_2)$  results from surface NO oxidation. For our study, we propose an alternative interpretation of the vertical variability in  $\Delta^{17}\text{O}(\text{NO}_3^-)$ , where  $\Delta^{17}\text{O}(\text{NO}_2)$  is considered the main driver. During the formation of the nocturnal boundary layer,  $\text{NO}_2$  formed during the day can be trapped above the surface layer in the nocturnal residual layer (NRL). This  $\text{NO}_2$  has a high  $\Delta^{17}\text{O}$  because it was formed during the previous daytime hours under the ISS framework (Eq. 3). Throughout the night, this highly enriched  $\text{NO}_2$  (ca. 37%, which is the average of the maximum  $\Delta^{17}\text{O}(\text{NO}_2)$  during SP 1 and SP 2) can be converted to  $\text{NO}_3^-$  via the  $\text{N}_2\text{O}_5$  pathway, hence leading to a substantial  $\Delta^{17}\text{O}$  transfer to  $\text{NO}_3^-$  at around 32%, which is in the range of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  observed by Fan et al. (2022) in winter. In the meantime, NO emitted at the surface during the night can be converted to  $\text{NO}_2$  by  $\text{O}_3$ , with a  $\Delta^{17}\text{O}$  transfer of ca. 20% (Eq. 5 with  $x = 0$ ). This low enriched  $\text{NO}_2$  can be further oxidised to  $\text{NO}_3^-$  by the  $\text{N}_2\text{O}_5$  pathway, which results in a  $\Delta^{17}\text{O}$  transfer at around 21%. This  $\text{NO}_2$  with a low  $\Delta^{17}\text{O}$  is very likely to be formed only at the surface during the night in areas experiencing important  $\text{NO}_x$  emissions (Michalski et al., 2014). Furthermore, surface  $\text{NO}_2$  with low  $\Delta^{17}\text{O}$  is not expected to be transported aloft as it is formed in the surface inversion layer during the night. Therefore,  $\text{NO}_3^-$  formed in the NRL during winter nights may be more enriched than the  $\text{NO}_3^-$  formed concurrently at the surface, regardless of the  $\text{NO}_2$  oxidation process involved. When the inversion layer breaks during the following day, the  $\text{NO}_3^-$  that was formed in the NRL during the night is mixed with the  $\text{NO}_3^-$  formed at the surface, resulting

**Table 2.** Mean observed  $\Delta^{17}\text{O}$  data of  $\text{NO}_2$  ( $\Delta^{17}\text{O}(\text{NO}_2)$ ) and  $\text{NO}_3^-$  ( $\Delta^{17}\text{O}(\text{NO}_3^-)$ ) in Chamonix and mean calculated  $\Delta^{17}\text{O}$  of  $\text{NO}_3^-$  ( $\Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-)$ ) using Eqs. (6) and (7) in the day and at night, respectively, constrained with observed  $\Delta^{17}\text{O}(\text{NO}_2)$ . Calculated day and night values were weighted by  $[\text{NO}_2] \times J_{\text{NO}_2}$  and  $[\text{NO}_2] \times [\text{O}_3]$ , respectively

		$\Delta^{17}\text{O}(\text{NO}_2)$ (‰)	$\Delta^{17}\text{O}(\text{NO}_3^-)$ (‰)	$\Delta^{17}\text{O}_{\text{calc}}(\text{NO}_3^-)$ (‰)	$\Delta^{17}(\text{NO}_3^-_{\text{calc}} - \text{NO}_3^-_{\text{obs}})$
Daytime (07:30–18:00)	SP 1	$30.0 \pm 7.3$	$23.0 \pm 3.1$	$22.5 \pm 4.6$	–0.5
	SP 2	$26.1 \pm 6.9$	$23.9 \pm 3.8$	$17.5 \pm 4.6$	–6.4
Nighttime (18:00–07:30)	SP 1	$21.2 \pm 1.1$	$20.5 \pm 1.1$	$21.4 \pm 0.7$	0.9
	SP 2	$20.8 \pm 1.0$	$23.2 \pm 1.0$	$21.0 \pm 0.6$	–2.2

in an increase in the overall surface  $\Delta^{17}\text{O}$ . In this scenario, the presence of the Saharan dust during SP 2 may have increased the  $\text{NO}_3^-$  loading aloft by promoting heterogeneous processes acting on aerosol surfaces in the vicinity of Chamonix. Hence,  $\text{NO}_2$  stratification at night could explain the observed increase in  $\Delta^{17}\text{O}$  of  $\text{NO}_3^-$  at the surface following the collapse of the nocturnal inversion layer. However, we cannot determine whether the enriched  $\text{NO}_3^-$  was formed in the vicinity of Chamonix and/or transported to our site by Saharan dust.

Although the exact nature of the high  $^{17}\text{O}$  excess measured in  $\text{NO}_3^-$  during SP 2 remains unclear, boundary layer dynamics is thought to play a significant role in the variability in  $\Delta^{17}\text{O}(\text{NO}_3^-)$  at the surface due to the stratification of  $\text{NO}_2$ . Therefore, a wider consideration of such factors should be explored to avoid possible over-interpretation of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  variabilities at the surface, especially in urban areas experiencing significant boundary layer dynamics in winter and high surface emissions of  $\text{NO}_x$  at night. Measuring  $\Delta^{17}\text{O}(\text{NO}_2)$  at various altitudes could provide better insights into the vertical dynamics of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  and subsequently quantitative information on  $\text{NO}_3^-$  production processes.

### 3.4 Nitrogen isotopic compositions

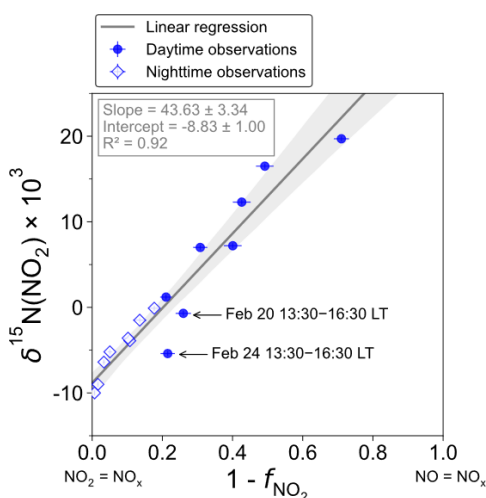
#### 3.4.1 N fractionation effects in the $\text{NO}_x$ cycle

Over the two sampling periods,  $\delta^{15}\text{N}(\text{NO}_2)$  shows substantial diurnal variability (from  $-10.0\%$  to  $19.7\%$ ,  $n = 16$ ) with a weighted mean of  $(4.0 \pm 9.1)\%$ . In contrast, Albertin et al. (2021) reported a weak diurnal fluctuation in  $\delta^{15}\text{N}(\text{NO}_2)$  in spring in Grenoble, in a narrow range from about  $-12\%$  to  $-10\%$ . In summer in an urban–suburban location, Walters et al. (2018) also observed a wide range of  $\delta^{15}\text{N}(\text{NO}_2)$  values; however, unlike in our study, these are almost consistently negative (from  $-31.4\%$  to  $0.4\%$ ) with an overall mean at  $(-11.4 \pm 6.9)\%$ . As shown in Eq. (8), fluctuations in  $\delta^{15}\text{N}(\text{NO}_2)$  reflect changes in  $\text{NO}_x$  emission sources and/or N fractionation effects, these latter being weighted by  $1 - f_{\text{NO}_2}$ ; i.e. the more  $\text{NO}_x$  is in the form of NO, the greater the N fractionation effects (see Sect. 2.4). Hence, in the previous works of Albertin et al. (2021) and Wal-

ters et al. (2018), due to high  $f_{\text{NO}_2}$  ( $> 0.7$ ), isotope effects were small ( $< 2.7\%$ ) and  $\delta^{15}\text{N}(\text{NO}_2)$  was mostly driven by changing contributions of  $\text{NO}_x$  emission sources. At our site,  $f_{\text{NO}_2}$  shows a wider range, from 0.3 to 1.0, suggesting significant N isotopic fractionation effects, with minimum and maximum contributions corresponding to the highest and lowest observed value of  $\delta^{15}\text{N}(\text{NO}_2)$ , respectively. This pronounced seasonal behaviour of N isotope fractionation effects within the  $\text{NO}_x$  cycle has previously been outlined in a seminal study of Freyer et al. (1993). Overall, compared with summer, lower  $f_{\text{NO}_2}$  during winter months due to lower  $\text{O}_3$  concentrations and higher  $\text{NO}_x$  emissions favours EIE between NO and  $\text{NO}_2$ , which also has a higher fractionation factor due to the lower temperatures (see Appendix D: “Equilibrium N fractionation factors”). Besides, this seasonal fluctuation in  $f_{\text{NO}_2}$  can be expected to be observed on smaller timescales, typically on the diurnal scale in urban areas where NO is generally fully oxidised into  $\text{NO}_2$  at night due to a lower  $\text{NO}_x$  emission rate, resulting in higher  $f_{\text{NO}_2}$  at night than during the day, as observed in our sampling (Fig. 2).

Figure 3 shows the linear dependence of  $\delta^{15}\text{N}(\text{NO}_2)$  on  $(1 - f_{\text{NO}_2})$  over the two sampling periods, indicating the significant influence of atmospheric processes that alter the N isotopic distribution during the conversion of  $\text{NO}_x$  into  $\text{NO}_2$ . The linear regression gives a slope and an intercept of about  $(43.6 \pm 3.3)\%$  and  $(-8.8 \pm 1.0)\%$ , respectively. According to Eqs. (10) and (11), the linearity between daytime (07:30–18:00 LT) and nighttime (18:00–07:30 LT) values suggests that EIE dominates the N fractionation processes between  $\text{NO}_x$  and  $\text{NO}_2$ . The influence of LCIE during the day could explain the greater variability around the linear fit in the daytime observations.

The relative importance of EIE and LCIE in the N fractionation between emitted  $\text{NO}_x$  and  $\text{NO}_2$  is assessed by calculating the  $A^*$  factor during the day (07:30–18:00 LT;  $A^*_{\text{day}}$ ) and night (18:00–07:30 LT;  $A^*_{\text{night}}$ ) (Table 3).  $A^*_{\text{day}}$  and  $A^*_{\text{night}}$  expressions are given in Sect. 2.4.2. Overall, during  $\text{NO}_2$  sampling intervals, the  $A^*$  values are small (mean  $\pm 1$  standard deviation:  $0.21 \pm 0.51$ ) and reflect an EIE-dominated regime with high  $\text{NO}_x$  (Li et al., 2020). It is interesting to note that the highest  $A^*$  values are observed between 13:30 and



**Figure 3.** Correlation plot of  $\delta^{15}\text{N}$  of atmospheric  $\text{NO}_2$  vs.  $(1 - f_{\text{NO}_2})$  from observations in Chamonix in February 2021.  $f_{\text{NO}_2}$  values are averaged over the collection period of each  $\text{NO}_2$  sample. The grey shading is the 95 % confidence interval. White diamonds and blue dots represent the nighttime (18:00–07:30 LT) and daytime (07:30–18:00 LT) observations, respectively. The linear regression is plotted over the nighttime and daytime observations.

16:30 LT and correspond to the two data points in Fig. 3 that lie outside the 95 % confidence interval of the regression line. These results suggest that EIE is the dominant N fractionation process between  $\text{NO}_x$  and  $\text{NO}_2$  during both the day and the night ( $A^* < 0.46$ ), with the exception of mid-afternoon when LCIE competes with EIE ( $A^* > 0.46$ ).

To quantify the overall N fractionation effect ( $F_{\text{N}}$ ) between  $\text{NO}_x$  and  $\text{NO}_2$ , we dissociate the two samples collected between 13:30 and 16:30 LT into a different group (group no. 1 denoted GP 1) from the other samples (group no. 2 denoted GP 2).  $F_{\text{N}}$  of GP 1 and GP 2 is calculated using Eq. (10) (which combines LCIE and EIE regimes) and Eq. (11) (which considers only the EIE regime), respectively. Calculated  $F_{\text{N}}$  values are reported in Table 3, and data used for calculations can be found in Sect. 2.4.2 and in the Supplement. Calculated  $F_{\text{N}}$  is significantly different between GP 1 and GP 2, with a mean of 16.4 % and 42.3 %, respectively. The close match between the calculated average  $F_{\text{N}}$  of GP 2 and the observed  $F_{\text{N}}$  ( $(43.6 \pm 3.3)\%$ ; slope of the regression line in Fig. 3) provides strong evidence for the reliability of Eq. (11), as well as the expression of  $\alpha_{\text{EIE}(\text{NO}_2/\text{NO}_x)}$  used therein, to accurately describe the N fractionation between  $\text{NO}_x$  emissions and  $\text{NO}_2$  at our site and, hence, to describe most of the variability in  $\delta^{15}\text{N}(\text{NO}_2)$  measurements. This result holds significant importance in confirming the theoretical N isotopic fractionation framework used in prior research studies. It is also important to stress the influence of LCIE effects for GP 1, highlighting the high dependency of  $\delta^{15}\text{N}(\text{NO}_2)$  on local environmental conditions. According to the  $A^*$  factor, a greater influence of LCIE in the mid-

afternoon could have contributed to the outlying of the two samples collected between 13:30 and 16:30 LT (GP 1). However, as mentioned above, the sample collected on 24 February between 13:30 and 16:30 LT has a significant blank. Therefore, it cannot be confirmed with certainty that the reason this sample falls outside the 95 % confidence interval of the regression line is solely due to LCIE. Nevertheless, the overall conclusion that EIE dominates the variability in  $\delta^{15}\text{N}(\text{NO}_2)$  at our site is not affected by this uncertainty.

The  $\delta^{15}\text{N}$  shift in  $\text{NO}_2$  relative to emitted  $\text{NO}_x$  ( $\Delta^{15}(\text{NO}_2 - \text{NO}_x)$ ) is calculated for individual  $\text{NO}_2$  samples using the mean ambient temperature during each sampling period. The mean atmospheric  $\delta^{15}\text{N}$  of  $\text{NO}_x$  ( $\delta^{15}\text{N}(\text{NO}_x)$ ) is then estimated by subtracting the  $\Delta^{15}(\text{NO}_2 - \text{NO}_x)$  value from the observed  $\delta^{15}\text{N}(\text{NO}_2)$  value.  $\Delta^{15}(\text{NO}_2 - \text{NO}_x)$  and  $\delta^{15}\text{N}(\text{NO}_x)$  estimates are reported in Table 3.  $\Delta^{15}(\text{NO}_2 - \text{NO}_x)$  varies greatly over the two sampling periods (from 0.7 ‰ to 30.7 ‰) with a mean value of ca. 9 ‰ (mean of GP 1 and GP 2).  $\delta^{15}\text{N}(\text{NO}_x)$  shows much less variability with an overall mean at  $(-7.8 \pm 1.9)\%$  (mean of GP 1 and GP 2), in very good agreement with the value derived from the regression relationship ( $-8.8\%$ ; intercept of the regression line in Fig. 3). Therefore, there appears to be little variation in  $\text{NO}_x$  emission sources at our site, and the wide variability in  $\delta^{15}\text{N}(\text{NO}_2)$  is mainly driven by important equilibrium post-emission isotopic effects.

### 3.4.2 $\text{NO}_x$ emission sources derived from $\delta^{15}\text{N}(\text{NO}_2)$

To identify the main source of  $\text{NO}_x$  that contributes to the calculated  $\delta^{15}\text{N}(\text{NO}_x)$  values at our site, Fig. 4 displays the temporal variation in  $\delta^{15}\text{N}(\text{NO}_x)$  obtained from individual  $\text{NO}_2$  samples (dashed horizontal line) and the  $\delta^{15}\text{N}$  range for different  $\text{NO}_x$  emission sources (coloured bands) such as for coal combustion ( $(19.5 \pm 2.3)\%$  for power plants with selective catalytic reduction technology; Felix et al., 2012; Elliott et al., 2019), fossil gas combustion ( $(-16.5 \pm 1.7)\%$ ; Walters et al., 2015a), and fertilised soils ( $(-33.8 \pm 12.2)\%$ ; Miller et al., 2018). The  $\delta^{15}\text{N}$  of  $\text{NO}_x$  released during biomass combustion is primarily driven by the  $\delta^{15}\text{N}$  of the biomass burnt (Fibiger and Hastings, 2016). We estimate an average  $\delta^{15}\text{N}$  of biomass combustion  $\text{NO}_x$  at  $(-0.1 \pm 1.3)\%$ , using the empirical relationship of Chai et al. (2019) (which was derived from the combustion of several North American wood species), and an average  $\delta^{15}\text{N}$  of biomass at  $(-2.8 \pm 2.0)\%$  representative of temperate forests (Martinelli et al., 1999). Regarding road traffic emissions, we have to stress that  $\delta^{15}\text{N}$  values reported in the literature are rather variable mainly because N fractionations during the process of  $\text{NO}_x$  production can vary depending on the type of fuel used, the type of vehicle, the presence of an emission control system, and the time of commuting (Ammann et al., 1999; Felix and Elliott, 2014; Heaton, 1990; Miller et al., 2017; Walters et al., 2015b; Zong et al., 2020, 2017). We use here the mean vehicle-emitted  $\delta^{15}\text{N}(\text{NO}_x)$  value given



**Table 3.** Summary table of data used to estimate the N isotopic fractionation between  $\text{NO}_x$  emissions and  $\text{NO}_2$  at our site ( $\Delta^{15}\text{N}(\text{NO}_2 - \text{NO}_x)$ ) and to derive  $\text{NO}_x$  emissions  $\delta^{15}\text{N}$ -fingerprint ( $\delta^{15}\text{N}(\text{NO}_x)$ ). The data reported are the mean values for each  $\text{NO}_2$  sampling period (mean value  $\pm$  absolute uncertainty).

Sampling interval (start–end)	$A^{*a}$	$F_{\text{N}}^b$	$f_{\text{NO}_2}$	$\Delta^{15}\text{N}(\text{NO}_2 - \text{NO}_x)^c$ (‰)	$\delta^{15}\text{N}(\text{NO}_x)$ (‰)
<b>GP 1</b>					
20 Feb 13:30–20 Feb 16:30	$0.46 \pm 0.08$	$25.00 \pm 2.27$	$0.74 \pm 0.02$	$6.5 \pm 0.9$	$-7.2 \pm 0.9$
24 Feb 13:30–24 Feb 16:30	$2.09 \pm 0.39$	$6.90 \pm 1.97$	$0.78 \pm 0.02$	$1.5 \pm 0.4$	$-6.9 \pm 0.6$
Mean	1.27	16.39	0.76	4.1	-7.1
SD	1.14	12.58	0.03	3.5	0.2
<b>GP 2</b>					
19 Feb 21:00–20 Feb 00:30	$0.03 \pm 0.01$	$43.06 \pm 0.17$	$0.97 \pm 0.01$	$1.4 \pm 0.1$	$-7.8 \pm 0.3$
20 Feb 00:30–20 Feb 04:30	$0.07 \pm 0.01$	$43.51 \pm 0.18$	$0.98 \pm 0.01$	$0.7 \pm 0.1$	$-9.7 \pm 0.3$
20 Feb 04:30–20 Feb 07:30	$0.02 \pm 0.01$	$43.78 \pm 0.22$	$0.89 \pm 0.01$	$4.7 \pm 0.4$	$-8.6 \pm 0.5$
20 Feb 07:30–20 Feb 10:30	$0.05 \pm 0.01$	$43.92 \pm 0.18$	$0.60 \pm 0.02$	$17.6 \pm 1.0$	$-10.4 \pm 1.1$
20 Feb 10:30–20 Feb 13:30	$0.10 \pm 0.02$	$43.06 \pm 0.18$	$0.57 \pm 0.02$	$18.4 \pm 1.0$	$-6.1 \pm 1.0$
20 Feb 16:30–20 Feb 18:00	$0.03 \pm 0.01$	$39.97 \pm 0.18$	$0.69 \pm 0.02$	$12.3 \pm 0.8$	$-5.3 \pm 0.8$
20 Feb 18:00–20 Feb 21:00	$0.01 \pm 0.01$	$41.75 \pm 0.19$	$0.90 \pm 0.01$	$4.3 \pm 0.5$	$-7.9 \pm 0.6$
24 Feb 07:30–24 Feb 10:30	$0.01 \pm 0.01$	$43.21 \pm 0.18$	$0.29 \pm 0.02$	$30.7 \pm 0.9$	$-11.9 \pm 0.9$
24 Feb 10:30–24 Feb 13:30	$0.07 \pm 0.01$	$41.95 \pm 0.18$	$0.51 \pm 0.02$	$20.6 \pm 1.0$	$-4.1 \pm 1.1$
24 Feb 16:30–24 Feb 18:00	$0.16 \pm 0.03$	$39.80 \pm 0.16$	$0.79 \pm 0.02$	$8.4 \pm 0.6$	$-7.2 \pm 0.7$
24 Feb 18:00–24 Feb 21:00	$0.01 \pm 0.01$	$40.88 \pm 0.18$	$0.82 \pm 0.02$	$7.2 \pm 0.6$	$-7.3 \pm 0.6$
24 Feb 21:00–25 Feb 00:00	$0.03 \pm 0.02$	$42.20 \pm 0.19$	$0.95 \pm 0.01$	$2.1 \pm 0.3$	$-7.3 \pm 0.4$
25 Feb 00:00–25 Feb 04:00	$0.19 \pm 0.03$	$42.48 \pm 0.18$	$0.99 \pm 0.01$	$0.3 \pm 0.1$	$-10.3 \pm 0.3$
25 Feb 04:00–25 Feb 07:30	$0.09 \pm 0.01$	$42.69 \pm 0.17$	$0.86 \pm 0.02$	$5.8 \pm 1.0$	$-7.3 \pm 1.0$
Mean	0.06	42.31	0.77	9.6	-7.9
SD	0.06	1.32	0.21	9.1	2.0

<sup>a</sup> Calculated from  $A_{\text{day}}^*$  between 07:30–18:00 LT and from  $A_{\text{night}}^*$  between 18:00–07:30 LT ( $A_{\text{day}}^*$  and  $A_{\text{night}}^*$  expressions are given in Sect. 2.4.2). <sup>b</sup> Calculated from Eq. (10) for GP 1 and from Eq. (11) for GP 2. <sup>c</sup> Calculated from Eq. (9).

by Song et al. (2022) at  $(-7.1 \pm 4.1)\text{‰}$ , calculated from 181 measurements reported in the literature.

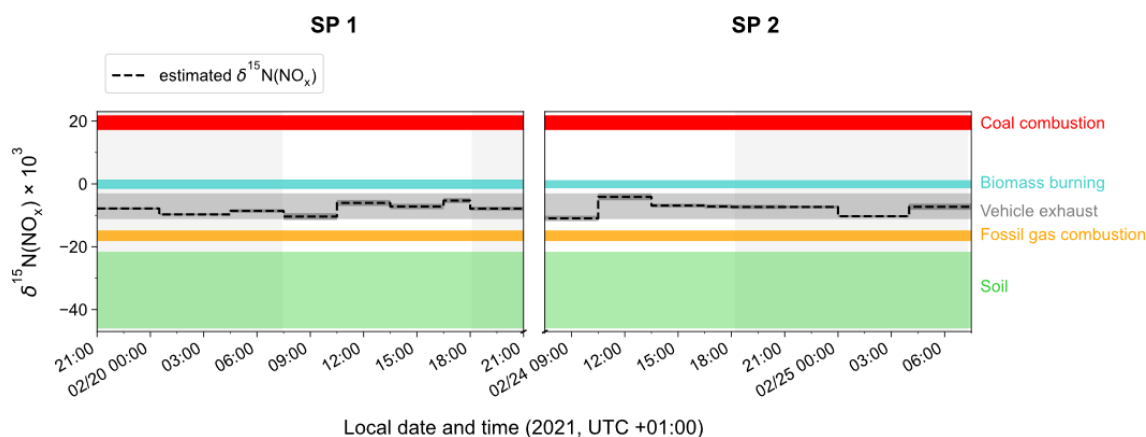
As previously noted, the values of estimated  $\delta^{15}\text{N}(\text{NO}_x)$  show much less variability than  $\delta^{15}\text{N}(\text{NO}_2)$ , with no significant differences observed between daytime and nighttime values. The values of  $\delta^{15}\text{N}(\text{NO}_x)$  range from  $-11.0\text{‰}$  to  $-4.1\text{‰}$ , and despite the associated uncertainty, they are consistent with the  $\delta^{15}\text{N}$  range of  $\text{NO}_x$  emissions from vehicle exhaust. The two sampling periods show similar  $\delta^{15}\text{N}(\text{NO}_x)$  values with a slight diel variability. The estimated small variation in  $\delta^{15}\text{N}(\text{NO}_x)$  throughout the day can be attributed to the temporal changes in the  $\delta^{15}\text{N}$  signature of mobile  $\text{NO}_x$  sources. It has been shown that  $\text{NO}_x$  emitted by cold engines has a lower  $\delta^{15}\text{N}$  signature compared to  $\text{NO}_x$  emitted from warm engines (Walters et al., 2015b). Hence, the early morning drop in  $\delta^{15}\text{N}(\text{NO}_x)$  could be attributed to the influence of  $\text{NO}_x$  emitted from cold engines. As the day progresses, the time of commuting increases and therefore  $\delta^{15}\text{N}(\text{NO}_x)$  tends to be less negative. Conversely, during the night, the slow  $\delta^{15}\text{N}(\text{NO}_x)$  decline could be due to the replacement of  $\text{NO}_x$  from vehicle exhaust by  $\text{NO}_x$  emitted by fossil gas combustion, which is commonly used in Chamonix for home heat-

ing. Although biomass burning used for home heating would also tend to increase  $\delta^{15}\text{N}(\text{NO}_x)$  during the day, it is unlikely to contribute more during the day than at night.

According to local  $\text{NO}_x$  emission inventories (Atmo-Auvergne-Rhône-Alpes, 2018; ORCAE, 2022), road transport is responsible for 64 % of  $\text{NO}_x$  emissions, ahead of heating-oil and fossil gas combustion. Despite the consistency between our results and existing inventories, the significant variability in the  $\delta^{15}\text{N}$  signature of  $\text{NO}_x$  emissions from vehicle exhaust precludes a reliable quantitative source apportionment of  $\text{NO}_x$  emissions from our estimated  $\delta^{15}\text{N}(\text{NO}_x)$ . Furthermore, the lack of information on the exact  $\delta^{15}\text{N}$  signature of  $\text{NO}_x$  emitted from heating-oil combustion could also contribute to the potential bias of the emission source apportionment.

### 3.4.3 Interpretation of $\delta^{15}\text{N}(\text{NO}_3^-)$ observations

$\delta^{15}\text{N}(\text{NO}_3^-)$  also exhibits substantial variability during the day, ranging from  $-1.3\text{‰}$  to  $14.9\text{‰}$  and from  $-4.2\text{‰}$  to  $9.7\text{‰}$  during SP 1 and SP 2, respectively. At night,  $\delta^{15}\text{N}(\text{NO}_3^-)$  is less variable, with an overall mean of



**Figure 4.** Time evolution of  $\delta^{15}\text{N}(\text{NO}_x)$  (dashed black line) estimated from  $\delta^{15}\text{N}(\text{NO}_2)$  observations in Chamonix after correction of N fractionation effects (length of horizontal line denotes sampling period; dark-grey-shaded area denotes overall calculation error bar). Coloured shaded areas represent the standard deviation of the mean  $\delta^{15}\text{N}$  value of individual  $\text{NO}_x$  emission source (coal combustion in red, biomass burning in blue, vehicle exhaust in grey, fossil gas in orange, and soil emissions in green). Grey-backdrop-shaded areas represent the nighttime (sunset to sunrise).

( $1.4 \pm 1.2$ )‰ and ( $-1.1 \pm 0.4$ )‰ during SP 1 and SP 2, respectively.  $\delta^{15}\text{N}(\text{NO}_3^-)$  is within the range of observations reported in urban areas (He et al., 2020; W. Zhang et al., 2022). A similar diurnal pattern was observed in samples collected during a cruise along the Californian coast in spring 2010 (Vicars et al., 2013), and isotopic exchanges between NO and  $\text{NO}_2$  during the day were found to be the primary driver of the diel variability. In the previous section, we demonstrated that there is a significant  $^{15}\text{N}$  partitioning between  $\text{NO}_x$  emissions and  $\text{NO}_2$ , the latter being enriched in  $^{15}\text{N}$  compared to  $\text{NO}_x$  emissions. Interestingly, important enrichments in  $^{15}\text{N}$  are also observed in  $\text{NO}_3^-$ .

As described above, at night during SP 1 and SP 2,  $\delta^{15}\text{N}(\text{NO}_2)$  is close to  $\delta^{15}\text{N}(\text{NO}_x)$  due to small N fractionation effects. However, between 18:00–07:30 LT,  $\text{NO}_3^-$  is enriched in  $^{15}\text{N}$  relative to  $\text{NO}_2$  by +6.3‰ and +1.4‰ on average during SP 1 and SP 2, respectively (Table 4). If we assume that, at night,  $\text{NO}_3^-$  is formed mainly by the conversion of surface  $\text{NO}_2$  via the  $\text{N}_2\text{O}_5$  pathway, then the difference between  $\delta^{15}\text{N}(\text{NO}_3^-)$  and  $\delta^{15}\text{N}(\text{NO}_2)$  ( $\Delta^{15}(\text{NO}_3^- - \text{NO}_2)$ ) should reflect the N enrichment factor associated with this oxidation process. It is likely that an isotopic equilibrium is established between  $\text{NO}_2$ ,  $\text{NO}_3$ , and  $\text{N}_2\text{O}_5$ , hence affecting the partitioning of  $^{15}\text{N}$  between  $\text{NO}_2$  and  $\text{NO}_3^-$  produced at night (Walters and Michalski, 2016). Neglecting KIE associated with the  $\text{N}_2\text{O}_5$  pathway and using the expression of the EIE fractionation factor between  $\text{N}_2\text{O}_5$  and  $\text{NO}_2$  given by Walters and Michalski (2015) (Appendix D) constrained with the mean nighttime temperature at our site, the isotopic composition of  $\text{NO}_3^-$  is expected to be enriched in  $^{15}\text{N}$  by about 29‰ compared to  $\text{NO}_2$ . This estimated  $^{15}\text{N}$  enrichment is about 3 times higher than the observed  $\Delta^{15}(\text{NO}_3^- - \text{NO}_2)$  at our site. As daytime  $\text{NO}_3^-$  exhibits higher  $\delta^{15}\text{N}$  values than during the night, it is not possible for daytime residuals at

night to account for the lower-than-predicted fractionation effect between  $\text{NO}_2$  and  $\text{NO}_3^-$ . These results highlight the importance of improving our understanding of the  $^{15}\text{N}$  fractionation between  $\text{NO}_2$  and  $\text{NO}_3^-$  associated with the  $\text{N}_2\text{O}_5$  pathway. This could be achieved in an atmospheric simulation chamber that allows us to reproduce individual processes in controlled conditions. The  $^{15}\text{N}$  isotopic enrichment of  $\text{NO}_2$  and  $\text{NO}_3^-$  collected from 07:30 to 18:00 LT shows a very contrasted distribution between SP 1 and SP 2, with a respective average  $\Delta^{15}(\text{NO}_3^- - \text{NO}_2)$  of  $-0.4$ ‰ and  $-10.0$ ‰ (Table 4). Although subjected to significant uncertainties (Fan et al., 2019), the OH pathway is often associated with a KIE of  $-3$ ‰ (Freyer, 1991), which is at odds with our observations. Similarly to the  $\text{N}_2\text{O}_5$  pathway, there is an important need to better estimate the fractionation factor associated with the OH pathway.

There are significant differences in  $\Delta^{15}(\text{NO}_3^- - \text{NO}_2)$  between SP 1 and SP 2, providing further evidence that  $\text{NO}_3^-$  collected during these two periods has undergone different formation processes and/or originates from different sources of  $\text{NO}_2$ . In addition, possible fractionation associated with phase change between  $\text{HNO}_3$  and  $p\text{-NO}_3$  during transport of Saharan dust could influence the  $\delta^{15}\text{N}$  of  $\text{NO}_3^-$  collected during SP 2. However, given the lack of knowledge about N fractionation factors between  $\text{NO}_2$  and  $\text{NO}_3^-$  and our limited dataset, we cannot conclude on whether the changes in the distribution of  $\text{NO}_3^-$  isotopes during SP 2 result from changes in the phase distribution of  $\text{NO}_3^-$  or in  $\text{NO}_2$  oxidation processes.

#### 4 Summary and implications

This study reports the first simultaneous measurements and analysis of  $\Delta^{17}\text{O}$  and  $\delta^{15}\text{N}$  in  $\text{NO}_2$  and  $\text{NO}_3^-$ . The samplings

**Table 4.** Mean observed  $\delta^{15}\text{N}$  data of  $\text{NO}_2$  ( $\delta^{15}\text{N}(\text{NO}_2)$ ) and  $\text{NO}_3^-$  ( $\delta^{15}\text{N}(\text{NO}_3^-)$ ), calculated atmospheric  $\delta^{15}\text{N}$  of  $\text{NO}_x$  ( $\delta^{15}\text{N}(\text{NO}_x)$ ), and  $\delta^{15}\text{N}$  shift between  $\delta^{15}\text{N}(\text{NO}_3^-)$  and  $\delta^{15}\text{N}(\text{NO}_2)$  ( $\Delta^{15}(\text{NO}_3^- - \text{NO}_2)$ ).

		$\delta^{15}\text{N}(\text{NO}_2)$ (‰)	$\delta^{15}\text{N}(\text{NO}_3^-)$ (‰)	$\delta^{15}\text{N}(\text{NO}_x)$ (‰)	$\Delta^{15}\text{N}(\text{NO}_3^- - \text{NO}_2)$ (‰)
Daytime (07:30–18:00)	SP 1	$7.4 \pm 4.7$	$7.0 \pm 6.7$	$-9.9 \pm 2.9$	-0.4
	SP 2	$14.0 \pm 13.9$	$4.0 \pm 6.4$	$-10.8 \pm 2.1$	-10.0
Nighttime (18:00–07:30)	SP 1	$-5.1 \pm 2.3$	$1.4 \pm 1.2$	$-9.0 \pm 0.8$	6.3
	SP 2	$-2.5 \pm 4.2$	$-1.1 \pm 0.4$	$-9.9 \pm 1.9$	1.4

were conducted at high temporal resolution ( $\sim 3$  h) in Chamonix, the French Alps, over 2 distinct days in late February 2021. The isotopic signals of both  $\text{NO}_2$  and  $\text{NO}_3^-$  show substantial diurnal variabilities, which are investigated in the light of local meteorological parameters and atmospheric observations ( $\text{NO}$ ,  $\text{NO}_2$ ,  $\text{O}_3$ , and  $\text{PM}$ ).

The observed variability in  $\Delta^{17}\text{O}(\text{NO}_2)$  can be well explained using  $\Delta^{17}\text{O}$  mass balance equations and corroborates the analysis of previous observations carried out in Grenoble, the French Alps, over a single day in spring (Albertin et al., 2021). On average, the high levels of  $\text{NO}_2$  at our site are primarily driven by the oxidation of local  $\text{NO}$  emissions by  $\text{O}_3$ . The observed diurnal variability in  $\Delta^{17}\text{O}(\text{NO}_2)$  appears to be consistent with the diurnal variability expected in the  $\text{NO}_x/\text{O}_3/\text{RO}_2$  chemistry, with  $\text{RO}_2$  levels of the order of  $\text{pmol mol}^{-1}$ , which is in agreement with the range of direct winter  $\text{RO}_2$  measurements reported in the literature.  $\text{RO}_2$  is thought to contribute significantly to the formation of  $\text{NO}_2$  in the early morning under high- $\text{NO}_x$  conditions, which is in line with the effective morning production of radical species reported in urban areas in winter. At night,  $\Delta^{17}\text{O}(\text{NO}_2)$  reflects the nocturnal oxidation of surface  $\text{NO}$  emissions by  $\text{O}_3$ . These results provide additional evidence that  $\Delta^{17}\text{O}(\text{NO}_2)$  measurements represent valuable constraints in the study of the reactive  $\text{NO}_x$  chemistry, down to the sub-daily temporal scales.

A clear linear relationship is found between  $\delta^{15}\text{N}(\text{NO}_2)$  and the  $\text{NO}_2/\text{NO}_x$  ratio, indicating significant post-emission N fractionation effects. Theoretical N isotopic fractionation factors between  $\text{NO}$  and  $\text{NO}_2$  at equilibrium and fractionation factors derived from the isotopic observations are found to be in good agreement, providing further support for the N isotopic fractionation theoretical framework commonly applied to the Leighton cycle. Observed  $\delta^{15}\text{N}(\text{NO}_2)$  corrected for N fractionation effects allows us to estimate the overall  $\delta^{15}\text{N}$  signature of ambient  $\text{NO}_x$  at our site. Based on the existing  $\delta^{15}\text{N}$  fingerprints of different  $\text{NO}_x$  emission sources, the main contribution at our site is very likely to be vehicle exhaust, which is confirmed by local emission inventories.

We use  $\Delta^{17}\text{O}$  mass balance equations of  $\text{NO}_3^-$  constrained by observed  $\Delta^{17}\text{O}(\text{NO}_2)$  to assess whether  $\text{NO}_3^-$  could originate locally from the oxidation of  $\text{NO}_2$  at our site. During the first day of sampling,  $\Delta^{17}\text{O}$  records of  $\text{NO}_2$  and  $\text{NO}_3^-$

support the local oxidation of  $\text{NO}_2$  to  $\text{NO}_3^-$  by  $\text{OH}$  radicals during the day and via the heterogeneous hydrolysis of  $\text{N}_2\text{O}_5$  during the night. The second day of sampling was affected by a Saharan dust event, accompanied by notable changes in the isotopic composition of  $\text{NO}_3^-$ . We propose that the formation of a surface inversion layer at night could have influenced the vertical distribution of  $\Delta^{17}\text{O}(\text{NO}_2)$  and resulted in a positive gradient of  $\Delta^{17}\text{O}(\text{NO}_3^-)$  with altitude, independently of the local  $\text{NO}_2$ -to- $\text{NO}_3^-$  conversion processes near the surface. In such a scenario, the presence of Saharan dust could have promoted heterogeneous  $\text{NO}_2$  oxidation, leading to higher  $\Delta^{17}\text{O}$  in  $\text{NO}_3^-$  formed aloft. The latter would have then mixed with the  $\text{NO}_3^-$  formed near the surface when the inversion broke up during the day. Although still uncertain, the influence of the boundary layer dynamics on the distribution of  $\Delta^{17}\text{O}$  in  $\text{NO}_3^-$  should be investigated in the future, notably for urban areas in winter.

The combined analysis of the first concurrent observations of  $\delta^{15}\text{N}$  in  $\text{NO}_2$  and  $\text{NO}_3^-$  highlights persistent uncertainties in the current estimates of the N fractionation factors associated with  $\text{NO}_2$  and  $\text{NO}_3^-$  conversion processes. However,  $\delta^{15}\text{N}(\text{NO}_3^-)$  records need to be corrected for N fractionation effects if they are to be used to trace back the  $\delta^{15}\text{N}$  fingerprint of the primary  $\text{NO}_x$  emission sources. Detailed simulation chamber experiments could provide more kinetic data on the various N fractionation processes in order to better exploit  $\delta^{15}\text{N}(\text{NO}_3^-)$  records to identify and quantify of the sources of reactive nitrogen.

The present thorough investigation of the  $\Delta^{17}\text{O}$  and  $\delta^{15}\text{N}$  in  $\text{NO}_2$  and  $\text{NO}_3^-$  highlights (1) the potential to use sub-daily  $\Delta^{17}\text{O}$  and  $\delta^{15}\text{N}$  records to trace the sources and formation chemistry of  $\text{NO}_3^-$ , (2) the importance of measuring the  $\text{NO}_2$  isotopic composition to avoid misinterpretation of  $\text{NO}_3^-$  isotopic records, and (3) the persistent knowledge gaps that prevent a complete assessment of the factors driving the variability in  $\text{NO}_3^-$  isotopic records. In most studies, the  $\text{NO}_3^-$  isotopic composition is interpreted on the basis of estimates of the isotopic composition of its precursor gases, assuming that both the chemistry of  $\text{NO}_2$  (including its conversion to  $\text{NO}_3^-$ ) and N isotopic fractionation effects are known. However, these assumptions are subject to very significant uncertainties, mainly in urban atmospheres. Hence, given the recent development of a method for measuring the multi-

isotopic composition of  $\text{NO}_2$ , the accuracy and validity of the current interpretation framework of  $\text{NO}_3^-$  isotopic records should be tested in various environments. Such investigation can be performed by simultaneously collecting  $\text{NO}_2$  and  $\text{NO}_3^-$ , as done here. We recommend using this combined isotopic record in order to avoid biased interpretations of  $\text{NO}_3^-$  isotopic records, particularly in urban areas during winter, preferably at high temporal resolution ( $< 24$  h). In addition, the vertical distribution of  $\text{NO}_2$  and  $\text{NO}_3^-$  isotopic composition should be documented in order to explore the possible role of the boundary layer dynamics in the variability in  $\text{NO}_2$  and  $\text{NO}_3^-$  isotopic composition observed at the surface.

## Appendix A: Reaction chemical rate

**Table A1.** Kinetic constants used in this study.

Reactions	Rate constants ( $\text{cm}^3 \text{mol}^{-1} \text{s}^{-1}$ )	References
$\text{NO} + \text{O}_3 \rightarrow \text{NO}_2 + \text{O}_2$	$k_{\text{NO}+\text{O}_3} = 1.4 \times 10^{-12} \exp(-1310(K)/T)$	Atkinson et al. (2004)
$\text{NO} + \text{RO}_2 \rightarrow \text{NO}_2 + \text{RO}$	$k_{\text{NO}+\text{RO}_2} = 2.3 \times 10^{-12} \exp(360(K)/T)$	Atkinson et al. (2006)
$\text{NO}_2 + \text{O}_3 \xrightarrow{M} \text{NO}_3 + \text{O}_2$	$k_{\text{NO}_2+\text{O}_3} = 1.4 \times 10^{-13} \exp(-2470(K)/T)$	Atkinson et al. (2004)
$^{15}\text{NO}_2 + ^{14}\text{NO} \rightarrow ^{14}\text{NO}_2 + ^{15}\text{NO}$	$k_{\text{NO}+\text{NO}_2} = 8.14 \times 10^{-14}$	Sharma et al. (1970)

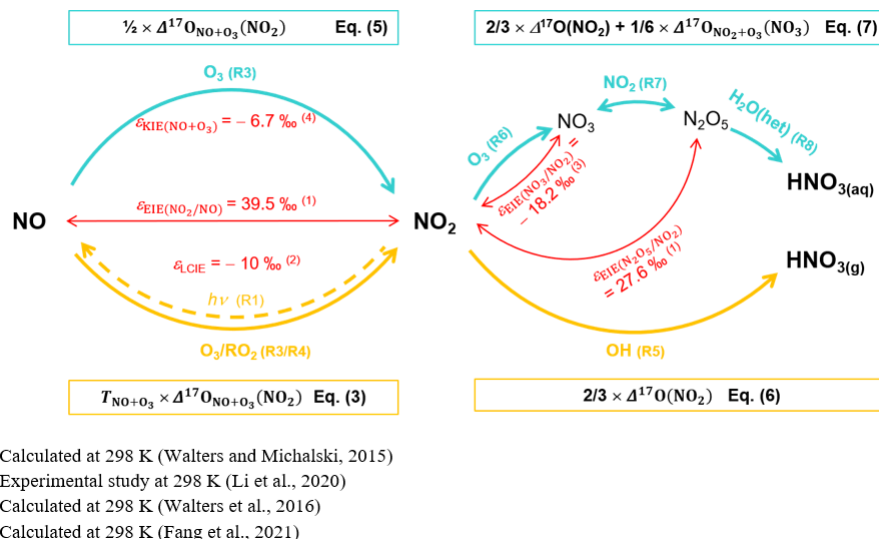
## Appendix B: Atmospheric lifetime of $\text{NO}_2$ and $\text{NO}_3^-$

**Table B1.** Mean daytime (07:30–18:00 LT) and nighttime (18:00–07:30 LT) atmospheric lifetime of  $\text{NO}_2$  ( $\tau_{\text{NO}_2}$ ) and  $\text{NO}_3^-$  ( $\tau_{\text{NO}_3^-}$ ) and dry-deposition constant ( $k_d = V_d \times \text{BLH}$ , where  $V_d$  is the dry-deposition velocity and BLH is the boundary layer height).

	$\tau_{\text{NO}_2}^{\text{a}}$	$\tau_{\text{NO}_3^-}^{\text{b}}$	$k_d(\text{NO}_2)$ ( $\text{s}^{-1}$ )	$k_d(\text{NO}_3^-)$ ( $\text{s}^{-1}$ )
Daytime (07:30–18:00)	5.1 min	27.8 h	$0.5 \times 10^{-5}$	$1.0 \times 10^{-5}$
Nighttime (18:00–07:30)	10.0 h	5.6 h	$2.5 \times 10^{-5}$	$5.0 \times 10^{-5}$

<sup>a</sup> Atmospheric lifetime relative to photolysis during the day (dry deposition and reaction  $\text{NO}_2 + \text{OH}$  are negligible) and to dry deposition and oxidation via  $\text{O}_3$  during the night. <sup>b</sup> Atmospheric lifetime relative to dry deposition. The boundary layer is fixed at 500 m during the day and at 100 m during the night. The dry-deposition velocity ( $V_d$ ) is fixed at 0.25 and  $0.50 \text{ cm s}^{-1}$  for  $\text{NO}_2$  and  $\text{NO}_3^-$ , respectively (Holland et al., 1999; Zhang et al., 2009).

### Appendix C: Schematic of the N reactive cycle and associated $\Delta^{17}\text{O}$ transfers and N enrichment factors



**Figure C1.** Adapted from Elliott et al. (2019). Sketch of dominant daytime (thick yellow arrows) and nighttime (thick cyan arrows)  $\text{NO}_x$ -to- $\text{NO}_3^-$  conversion processes and associated quantified N fractionation effects at 298 K (thin red arrows and text) and  $\Delta^{17}\text{O}$  transfers (yellow and cyan boxes).

### Appendix D: Equilibrium N fractionation factors

**Table D1.** Calculated regression coefficients for the N isotope exchange between  $\text{NO}_2/\text{NO}$  and  $\text{N}_2\text{O}_5/\text{NO}_2$  over the temperature range of 150 to 450 K (Walters and Michalski, 2015) and for the N kinetic fractionation for the reaction  $\text{NO} + \text{O}_3$  over the temperature range of 220 to 320 K (Fang et al., 2021).

$(\alpha_{\text{EIE}(X/Y)} - 1) \times 1000 = \frac{A}{T^4} \times 10^{10} + \frac{B}{T^3} \times 10^8 + \frac{C}{T^2} \times 10^6 + \frac{D}{T} \times 10^4$				
$X/Y$	$A$	$B$	$C$	$D$
$\text{NO}_2/\text{NO}$	3.847	-7.680	6.003	-0.118
$\text{N}_2\text{O}_5/\text{NO}_2$	1.004	-2.525	2.718	0.135
$(\alpha_{\text{KIE}(X+Y)} - 1) \times 1000 = A \times \exp(B/T)$				
$X + Y$	$A$	$B$		
$\text{NO} + \text{O}_3$	0.982	3.352		

**Data availability.** Data presented in this article are included in the Supplement.

**Supplement.** The supplement related to this article is available online at: <https://doi.org/10.5194/acp-24-1361-2024-supplement>.

**Author contributions.** Grants obtained by KL and JS funded the project. AB, RG, QF, and IV performed the calibration of IBBCEAS and OFCEAS instruments, data acquisition, and post-processing. SA conducted the sampling and laboratory analysis. NC provided technical support for isotopic mass spectrometry analysis. The study was designed as part of SA's PhD thesis supervised by SB and JS. The paper was written by SA under the supervision of JS and SB and with contributions from all co-authors.

**Competing interests.** The contact author has declared that none of the authors has any competing interests.

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