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Rubidium isotopic fractionation during magmatic processes and the composition of the bulk silicate Earth

**Baoliang Wang^{1, *}, Frederic Moynier¹, Matthew G. Jackson², Fang Huang³, Xia
Hu³, Sæmundur Ari Halldórsson⁴, Wei Dai¹, Gabriel Devos¹**

¹Université Paris Cité, Institut de Physique du Globe de Paris, CNRS, 1 Rue Jussieu,
75005 Paris, France

²Department of Earth Science, University of California, Santa Barbara 93106, USA

³CAS Key Laboratory of Crust-Mantle Materials and Environments, School of Earth
and Space Sciences, University of Science and Technology of China, Hefei, Anhui
230026, China

⁴Nordic Volcanological Center, Institute of Earth Sciences, University of Iceland,
Reykjavík, Iceland

* Corresponding author (email: bwang@ipgp.fr).

1 **Abstract**

2 Rubidium is a moderately volatile element with high incompatibility and fluid
3 mobility. Its stable isotopes have great potential in tracing various geological
4 processes. For example, lunar rocks are isotopically heavier than terrestrial ones,
5 suggesting volatile loss by evaporation during or following the formation of the Moon.
6 However, these studies rely on a poorly constrained estimate for the composition of
7 the Earth's mantle and a poor understanding of high-temperature processes which may
8 act to fractionate stable Rb isotopes. It is therefore important to precisely characterize
9 different rock types that sample the Earth's mantle as well as to evaluate the
10 importance of key isotopic fractionation processes. In order to address these issues,
11 we established a high precision analytical method for Rb isotopic measurements using
12 the Nu Sapphire CC-MC-ICP-MS (collision-cell multi-collector inductively coupled
13 plasma mass spectrometer). In addition, we present a series of Rb isotopic data of
14 volcanic rocks from Hekla volcano (Iceland) and MORB (mid-ocean ridge basalt)
15 samples. We show that our method returns a high Rb sensitivity ($\sim 500 \text{ V}/\mu\text{g}\cdot\text{g}^{-1}$ for
16 ^{85}Rb) and a long-term reproducibility of 0.03‰ on $\delta^{87}\text{Rb}$ (the permil deviation of the
17 $^{87}\text{Rb}/^{85}\text{Rb}$ ratio from the SRM 984 standard). This method uses a 2 ng/g Rb solution
18 for analyses, allowing us to consume about 10 times less Rb to achieve similar or
19 better precision than previous studies. Using this method, seven geostandards and one
20 synthetic standard return Rb isotopic data consistent with previous work. Twenty-one
21 Hekla volcanic rocks, spanning compositions from basalt to rhyolite, show limited Rb
22 isotopic variation, with $\delta^{87}\text{Rb}$ values varying from -0.17‰ to -0.07‰ , demonstrating

23 that magmatic evolution has an insignificant effect on Rb isotope ratios. A set of
24 MORB samples (n=15) from different mid-ocean ridges also span a limited Rb
25 isotopic variation, displaying a range similar to the Hekla rock suite (-0.19 to -0.02‰).
26 Combining our new data together with previously reported OIB data gives an average
27 $\delta^{87}\text{Rb}$ value of $-0.12 \pm 0.08\text{‰}$ (2SD, n=25), representing the current best estimate of
28 the mantle's isotopic composition. Considering the $\delta^{87}\text{Rb}$ values of the upper
29 continental crust ($-0.14 \pm 0.01\text{‰}$, 2SE, n=73), as inferred from recent measurements
30 of granites, loess and sediments, and assuming this value represents the whole crust,
31 the revised Rb isotopic composition of the bulk silicate Earth (and by extension the
32 bulk Earth, assuming no Rb partitioned into the core) is $-0.13 \pm 0.06\text{‰}$ (2SD).

33

34 **1. Introduction**

35 Rubidium (Rb) is a univalent lithophile alkali metal element. Due to its high
36 incompatibility, Rb is enriched in crustal materials compared to the mantle. The
37 estimated Rb abundances of the lower, middle, and upper continental crust are 11 $\mu\text{g/g}$,
38 65 $\mu\text{g/g}$ and 84 $\mu\text{g/g}$, respectively (Rudnick and Gao, 2003), while that in the
39 primitive mantle is only 0.6 $\mu\text{g/g}$ (McDonough and Sun., 1995), and the depleted
40 upper mantle is 0.05 $\mu\text{g/g}$ (Workman and Hart, 2005). Rubidium is a fluid-mobile
41 element (Han et al., 2021; X. Hu et al., 2022) and it is moderately volatile with a 50%
42 condensation temperature (T_c^{50}) of 800 K (Lodders, 2003). Similar to other
43 moderately volatile elements (MVE) such as K, Cu, Zn, Cd, Cl, S and Se (Luck et al.,
44 2005; Wombacher et al., 2008; Paniello et al. 2012; Boyce et al., 2015, 2018; Wing

45 and Farquhar, 2015; Pringle et al., 2017; Barnes et al., 2019; Tian et al., 2019, 2021;
46 Gargano et al., 2020; Van Kooten et al., 2020; Vollstaedt et al., 2020; Y. Hu et al.,
47 2022; Paquet et al. 2022), Rb stable isotopes show promise for studying mechanisms
48 of volatile depletion processes in the inner Solar System.

49 Rubidium has two naturally occurring isotopes: the stable isotope ^{85}Rb (72.17%)
50 and the radioactive isotope ^{87}Rb (27.83%) (Berglund and Wieser, 2011). However,
51 ^{87}Rb is also considered as a stable isotope due to its long half-life (4.97×10^{10} y, where
52 ^{87}Rb decays to ^{87}Sr) (Nebel et al., 2011b). Previous applications of Rb stable isotopes
53 mainly focused on cosmochemical processes. For example, tektites were analyzed for
54 Rb isotopes in order to explore the influence of impact melting (Nebel et al., 2005).
55 Nebel et al. (2011a) reported Rb isotopic compositions of chondrites and observed
56 resolvable variations on the order of $\pm 1\text{‰}$ $\delta^{87}\text{Rb}$ in some meteorites. Pringle and
57 Moynier (2017) established a high-precision analytical method and reported Rb
58 isotopic data of a series of terrestrial and extraterrestrial samples, including chondrites,
59 achondrites and lunar samples. The results show significant Rb isotopic differences
60 among these samples, especially lunar samples, which exhibit Rb isotopic
61 compositions that are heavier than the Earth's mantle by ~ 0.2 permil. Subsequent
62 measurements employing a different analytical method, have also found heavy Rb
63 isotopic signatures in lunar rocks compared to terrestrial samples (Nie and Dauphas,
64 2019). All these studies show that Rb isotope can be used as a powerful tool to study
65 planetary evolution.

66 However, these planetary studies rely on the Earth's mantle composition
67 estimated by Pringle and Moynier (2017), and very few studies have applied Rb
68 isotopes to terrestrial samples. Pringle and Moynier (2017) estimated the Rb isotopic
69 composition of the mantle ($\delta^{87}\text{Rb} = -0.12 \pm 0.06\text{‰}$, 2SD) based on only 6 samples,
70 including one mid-ocean ridge basalt (MORB) from the Atlantic ridge, two ocean
71 island basalts (OIB) (Galapagos and Hawaii), one continental flood basalt (BCR-2),
72 one andesite (AGV-2) and one granite (GS-N). X. Hu et al. (2022) constrained the
73 composition of the upper continental crust (UCC) to be $-0.14 \pm 0.01\text{‰}$ (2SE, n=73)
74 based on an extensive sample set of granites, loess and sediments. A few additional
75 igneous geostandards have also been analyzed for Rb stable isotopes (Zhang et al.,
76 2018; Nie and Dauphas, 2019; Nie et al., 2021) (see Table 1).

77 Since the estimation of the Rb isotopic composition of the mantle is a key
78 parameter for inter-planetary comparison, the limited set of data obtained from
79 mantle-derived materials calls for a more comprehensive characterization. The very
80 low Rb contents ($\sim 0.05 \mu\text{g/g}$; Workman and Hart, 2005) makes it relatively difficult
81 to directly analyze depleted peridotites (e.g., more than 200 mg of samples are needed
82 with the highest sensitivity to date, i.e., $250 \text{ V}/\mu\text{g}\cdot\text{g}^{-1}$, Pringle and Moynier, 2017). In
83 this regard, MORBs are obvious targets as they are generated via partial melting of
84 depleted upper mantle materials, thus providing a straightforward way to constrain the
85 composition of this important reservoir, assuming no isotopic fractionation is
86 generated during partial melting (Hofmann, 2003). Furthermore, magmatic evolution
87 is known to fractionate stable isotopes of various isotopic systems (e.g., Fe, Teng et

88 al., 2008; Sn, Badullovich et al., 2017) and it is important to evaluate this effect on Rb
89 isotopes. Previous Rb isotopic analyses of geostandards of different compositions
90 found that Rb tends to become lighter gradually from mafic to felsic rocks (Zhang et
91 al., 2018; Hu et al., 2021). These isotopic differences observed among the
92 geostandards are relatively minor, typically within the range of approximately 0.2‰.
93 This variability may also be attributed to the mantle source of the analyzed lavas.
94 However, it is worth noting that geostandards are mass-produced and may become
95 contaminated during production as suggested by studies of Mo isotopes (Willbold and
96 Elliott 2017). Therefore, they may not be ideal for accurately estimating magmatic
97 processes on Rb isotopes. In order to better determine the behavior of Rb isotopes
98 during magmatic evolution, it is therefore necessary to investigate a co-genetic rock
99 suite that is largely controlled by magmatic differentiation from a common parental
100 melt. The Hekla volcano produced eruptive materials that range from basalt to
101 rhyolite (Thorarinsson, 1967) and the Hekla suite has been used extensively in recent
102 years to study isotopic fractionation during igneous processes responsible for
103 formation of high-Si igneous rocks for various isotopic systems (e.g., Fe, Li,
104 Schussler et al., 2009; Si, Savage et al., 2011; V, Prytulak et al., 2016; Zn, Chen et al.,
105 2013; Tl, Prytulak et al., 2017; Ti, Deng et al., 2019; K, Tuller-Ross et al., 2019; Cl,
106 Ranta et al., 2021; Ba, Bai et al., 2022). Hekla is, therefore, well suited for evaluating
107 the behavior of Rb during igneous processes ultimately forming high-Si igneous rocks.

108 Here, we developed a new analytical method and report the first high precision
109 Rb isotopic data using a Nu Sapphire CC-MC-ICP-MS (collision-cell multi-collector

110 inductively coupled plasma mass spectrometer). We present Rb isotopic compositions
111 for a variety of samples, including 21 igneous rocks from Hekla volcano, and 15
112 MORBs from different mid-ocean ridge segments, as well as seven different igneous
113 geostandards. Based on these data, we assess the effect of magmatic processes on Rb
114 isotopes and estimate the average Rb isotopic composition of the mantle. Combining
115 our findings from the mantle with the extensive crustal dataset previously reported by
116 X. Hu et al. (2022), a more precise constraint on the Rb isotopic composition of the
117 bulk silicate Earth (BSE) is provided in this study.

118

119 **2. Samples**

120 **2.1. Geostandards**

121 In order to test the accuracy of our new method, a number of geostandards with
122 compositions ranging from basalt to granite were selected because they have
123 previously been analyzed for their Rb isotopic compositions. The geostandards
124 comprise BHVO-2 (basalt, United States Geological Survey, USGS), BCR-2 (basalt,
125 USGS), W-2a (diabase, USGS), AGV-2 (andesite, USGS), GSP-2 (granodiorite,
126 USGS), GS-N (granite, Centre de Recherches Pétrographiques et Géochimiques,
127 CRPG), and GA (granite, CRPG). We also mixed pure NIST SRM 984 solution (the
128 Rb isotopic standard) with PCC-1 (peridotite, USGS) to be treated as a synthetic
129 standard to evaluate the accuracy of the measurements.

130 **2.2. Hekla volcano samples**

131 In recent years, the Hekla magma suite has been extensively utilized for
132 investigating isotopic fractionation during magmatic processes (e.g., Savage et al.,
133 2011; Chen et al., 2013; Prytulak et al., 2017; Deng et al., 2019; Tuller-Ross et al.,
134 2019). A set of Hekla volcanic rocks were collected (see sample locations in Bai et al.,
135 2022) and analyzed for Rb isotope in this study. Located at the margin of the
136 propagating South Iceland Volcanic Zone, Hekla volcano is an active central volcano
137 in Iceland that rises nearly 1500 masl (meters above the sea level; Thorarinsson,
138 1967). It is well-known as a natural laboratory for investigating magmatic processes
139 and the origin of high-Si rocks (Sigmarsson et al., 1992, 2022; Geist et al., 2021). It
140 erupts a series of cogenetic volcanic products, including basalts, basaltic andesites,
141 andesites, dacites and rhyolites. At present, origin of Hekla's wide compositional
142 variability is a matter of debate. Sigmarsson et al. (1992) proposed a three-stage
143 model to explain generation of Hekla's igneous rocks from basalt to rhyolite: (1) The
144 rise of basaltic magma causes partial melting of metabasaltic crust, thus forming
145 dacitic magma. (2) The basaltic magmas undergo fractional crystallization to generate
146 basaltic andesite compositions, and the latter mixes with dacitic liquid to form
147 andesites. (3) Rhyolites are formed from dacitic magma through fractional
148 crystallization. Chekol et al. (2011) suggested that andesitic to more evolved
149 compositions were produced through AFC (assimilation and fractional crystallisation)
150 processes. These studies both advocated the contribution of crustal components to
151 high-Si rocks (i.e., assimilation and/or partial melting of the crust). By contrast, Geist
152 et al. (2021) suggested that the entire Hekla suite is formed by fractional

153 crystallization processes alone, over a period of several tens of thousands of years. A
154 more recent study argued that partial melting of the metabasaltic crust remains the
155 preferred mechanism for petrogenesis of the Hekla silicic rocks, with the time span
156 from silicic magma formation to eruption being less than a few centuries (Sigmarsson
157 et al., 2022).

158 In this study, we analyzed 9 basalts, 6 basaltic andesites, 2 andesites, 3 dacites
159 and 1 rhyolite from Hekla volcano. Some of these have previously been analyzed for
160 their Ba isotopic composition (Bai et al., 2022). The Hekla rocks studied here are
161 generally fresh, Holocene igneous rocks with little to no visible signs of secondary
162 alteration (Sigmarsson et al., 1992; Chekol et al., 2011; Geist et al., 2021).
163 Furthermore, besides some rare secondary incrustations, formed during solidification,
164 cooling and secondary degassing (Óskarsson, 1981), the vast amount of data
165 published from the Hekla suite does not indicate that secondary alteration affects the
166 chemical and isotopic composition. Thus, the Hekla suite records extensive magmatic
167 evolution of a OIB magma, from basalt through rhyolite, making this suite ideal for
168 studying the effect of magmatic processes on Rb isotopes.

169 **2.3. MORB samples**

170 Sixteen MORB samples from different mid-ocean ridge segments were collected
171 to estimate the Rb isotopic composition of the upper mantle. Among them, nine are
172 from the Atlantic Ocean, four are from the Indian Ocean and three are from the East
173 Pacific Ocean. All MORB samples in this study are fresh glasses (quenched melts),
174 which were not subjected to significant alteration as indicated by the chemical index

175 of alteration (CIA) and loss on ignition (LOI) (She et al., 2023). According to a
176 common classification scheme (Schilling, 1975; Schilling et al., 1983), MORB
177 samples can be divided into Normal-MORBs (N-MORBs, $La/Sm_N < 0.7$), Enriched-
178 MORBs (E-MORBs, $La/Sm_N > 1.8$) and Transitional-MORBs (T-MORBs, with
179 La/Sm_N between N- and E-MORBs). Importantly for the purpose of this study, these
180 MORB samples cover all the different types mentioned above (Amsellem et al., 2018;
181 Deng et al., 2018; Inglis et al., 2019), which allows us to evaluate the role of partial
182 melting and/or mantle source-derived variability on $\delta^{87}Rb$ values.

183

184 **3. Analytical methods**

185 **3.1. Sample dissolution**

186 Sample digestion and chemical purification were conducted at the Institut de
187 Physique du Globe de Paris (IPGP). 18.2 M Ω -cm ultrapure water and distilled high-
188 purity concentrated acids were used throughout the processes. For each Hekla sample,
189 whole-rock powder was made from approximately 50g of rock. For MORB samples,
190 fresh glass fragments were handpicked and powdered with an agate mortar (She et al.,
191 2023). 20-60 mg of sample powders were weighed into Savillex PFA Teflon beakers
192 to yield >10 ng Rb for isotopic measurement. A separate 50 mg of Hekla sample
193 powders were weighed for major and trace elements measurement. Samples were first
194 digested with a 3:1 volume mix of concentrated HF and HNO₃ by heating at 130 °C
195 for 48h on a hotplate. After evaporation of HF-HNO₃, the samples were sequentially
196 treated with aqua regia (concentrated HCl: HNO₃=3:1, volume ratio) and 6 mol/L HCl

197 at 120 °C to remove fluoride complexes. Finally, dried samples were re-dissolved in 1
198 mL 0.5 mol/L HNO₃ for major and trace elements measurements, or re-dissolved in 1
199 mol/L HNO₃ for subsequent column chemistry.

200 **3.2. Major and trace elements measurement**

201 Full basic geochemical data of the MORB samples has been reported in She et al.
202 (2023). Elemental concentration measurement was performed for all the Hekla
203 samples at IPGP. An aliquot of the digested solution was diluted by a factor of 15000
204 for elemental analysis. Major and trace elements were measured using an Agilent
205 quadrupole inductively coupled plasma mass spectrometer (Q-ICP-MS). Two
206 geostandards (i.e, BHVO-2 and BCR-2) were processed with the same procedures to
207 monitor the accuracy. The data are provided in the Supplementary Material.

208 **3.3. Chemical purification and Rb isotopic measurement**

209 Pringle and Moynier (2017) developed a column chemistry procedure of Rb
210 isotopes at IPGP. The procedure consists of four steps, with one column of Eichrom
211 DGA resin (50–100 µm, 1.8 mL) to remove Ca, then two columns of Bio-Rad
212 AG50W-X12 resin (200–400 mesh, 20 mL and 10 mL, respectively) to remove other
213 matrix elements, and finally one column of Bio-Rad AG50W-X8 resin (200–400 mesh,
214 1 mL) to further separate K. However, this procedure is time-consuming and requires
215 large volumes of acid. Previous studies showed that the separation of Rb and K could
216 be successfully realized by using Sr-spec resin (Zhang et al., 2018; Nie and Dauphas,
217 2019; Hu et al., 2021). Thus, based on the column chemistry established by Pringle
218 and Moynier (2017), we modified the last step by replacing it with a Sr-spec resin

219 chemistry. The full chemical purification procedure is listed in Supplementary Table 1,
220 in which the first three steps have been described in Pringle and Moynier (2017) and
221 remain the same here. In the text that follows, we provide a detailed description of
222 only the last step (i.e., Sr-spec purification).

223 After the first three steps of purification, sample elution was dried down and re-
224 dissolved in 3 mol/L HNO₃ for Sr-spec column chemistry. 1 mL Eichrom Sr-spec
225 resin (25–50 μm) was filled into a pre-cleaned homemade heat-shrink Teflon column
226 (0.46 cm inner diameter, 6.5 cm length). The resin was cleaned with 3 mL of 3 mol/L
227 HNO₃ and 2 mL of H₂O successively, and then conditioned with 3 mL 3 mol/L HNO₃
228 prior to sample loading. Sample was loaded on the resin in 0.25 mL of 3 mol/L HNO₃
229 and followed by another 0.25 mL of 3 mol/L HNO₃. The Rb was collected with 1 mL
230 3 mol/L HNO₃ (Supplementary Table 1). Compared to the AG50W-X8 resin method
231 (which required 40 mL of acid to collect Rb) that was previously used by Pringle and
232 Moynier (2017), the Sr-spec in this study used only 1 mL of acid. This greatly
233 simplifies the chemical purification process. Using the Sr-spec method, we can
234 achieve a complete separation of Rb from K (Fig. 1).

235 The recovery and blank of the whole procedure were carefully evaluated. The
236 yield of each sample was >95%, and the total procedural blank was <0.03 ng, which
237 was negligible (<0.3%) compared to the total Rb in sample (>10 ng).

238 Rubidium isotopic measurements were performed on CC-MC-ICP-MS Nu
239 Sapphire at IPGP. The Nu Sapphire is a new generation double-focusing MC-ICP-MS
240 which includes both a high energy ion pathway (traditional MC-ICP-MS) and a low

241 energy ion pathway (equipped with a collision cell). For Rb isotopic measurements,
242 we used conventional high energy mode. Sample solution was introduced through an
243 Apex Omega desolvating system using a 100 $\mu\text{L}/\text{min}$ PFA nebulizer (Elemental
244 Scientific, Omaha, NE, USA). Rubidium isotopic abundances were analyzed in low
245 resolution mode with standard Ni dry cones (“dry plasma”). The isotopes ^{85}Rb
246 and ^{87}Rb were simultaneously collected on L3 and H1 Faraday cups, respectively,
247 with $10^{11}\Omega$ resistors, respectively. Strontium-88 was collected on the H3 Faraday cup
248 ($10^{11}\Omega$ resistor) in order to evaluate ^{87}Sr intensity. The interference of ^{87}Sr on ^{87}Rb is
249 corrected by assuming an $^{87}\text{Sr}/^{88}\text{Sr}$ value of 0.085, following prior work (Nebel et al.,
250 2005; Pringle and Moynier, 2017; Zhang et al., 2018; Hu et al., 2021). A constant
251 $^{87}\text{Sr}/^{88}\text{Sr}$ ratio of 0.085 is sufficient for the Sr correction, because the ^{87}Sr interference
252 on ^{87}Rb is insignificant due to efficient Sr removal during column chemistry, and thus
253 the uncertainty on the $^{87}\text{Sr}/^{88}\text{Sr}$ ratio would have a minor impact. For example, taking
254 two extreme values of $^{87}\text{Sr}/^{88}\text{Sr}$ (0.083 and 0.089, corresponding to $^{87}\text{Sr}/^{86}\text{Sr}$ of 0.70
255 and 0.75, respectively) from the reported range in terrestrial and extraterrestrial rocks
256 and for the sample with the most Sr residue ($^{88}\text{Sr}/^{85}\text{Rb}=0.003$), would only affect the
257 correction by only $\sim 0.03\%$.

258 Relevant operating parameters of the instrument are shown in Supplementary
259 Table 2. The sensitivity of ^{85}Rb was $\sim 500 \text{ V}/\mu\text{g}\cdot\text{g}^{-1}$. Samples and standard solutions
260 were diluted to 2 ng/g using 0.5 mol/L HNO_3 for analyses. Before each analysis, a 30-
261 second blank measurement was performed in 0.5 mol/L HNO_3 . Each sample analysis
262 consisted of 40 cycles of data with an integration time of 5 seconds per cycle. In order

263 to eliminate potential cross-contamination, the sample introduction system was
264 cleaned 90 seconds using 0.5 mol/L HNO₃ following each standard and sample
265 analysis. Each sample was measured at least 3 times, bracketed by the international
266 isotopic standard NIST SRM 984. Rubidium isotopic data were reported as
267 $\delta^{87}\text{Rb} = [({}^{87}\text{Rb}/{}^{85}\text{Rb})_{\text{sample}} / ({}^{87}\text{Rb}/{}^{85}\text{Rb})_{\text{SRM984}} - 1] \times 1000$ (‰).

268

269 **4. Results**

270 The pure NIST SRM 984 solution yielded $\delta^{87}\text{Rb}$ values between -0.05‰ and
271 +0.04‰, with an average of 0.00 ± 0.03 ‰ (2SD, n=23; Fig. 2a and Supplementary
272 Table 4). Rubidium isotopic compositions of one in-house standard (IPGP-Rb, a pure
273 Rb ICP-MS solution), previously used in Pringle and Moynier (2017) and seven
274 geostandards are presented in Table 1, and literature values are also listed for
275 reference. The value of IPGP-Rb (-0.12 ± 0.02 ‰, 2SD, n=6) was identical to that ($-$
276 0.12 ± 0.08 ‰, 2SD, n=40) measured by Pringle and Moynier (2017) using a Thermo-
277 Fisher Neptune plus. The $\delta^{87}\text{Rb}$ values of geostandards all agree well with previously
278 published data (Fig. 2b). In addition, some of these geostandards are used to test the
279 effect of the measuring concentrations (Fig. 3). These results are discussed in section
280 5.1.

281 Hekla samples data are reported in Table 2. Consistent with previous studies (e.g.,
282 Sigmarsson et al., 1992, 2022; Geist et al., 2021), Rb contents of Hekla suite display
283 reasonably good correlations with SiO₂ and MgO contents as well as other
284 incompatible minor (K₂O) and trace elements (La), gradually increasing with the

285 magmatic evolution (Fig. 6a, b, c and d), which is consistent with its strong
286 incompatibility. The average Rb concentrations of these five rock types—basalts,
287 basaltic andesites, andesites, dacites, and rhyolite—are 10.4 $\mu\text{g/g}$, 27.2 $\mu\text{g/g}$, 29.2 $\mu\text{g/g}$,
288 55.7 $\mu\text{g/g}$ and 68.3 $\mu\text{g/g}$, respectively. However, irrespective of Rb contents, all Hekla
289 samples show uniform Rb isotopic composition within uncertainty, with $\delta^{87}\text{Rb}$ values
290 ranging from -0.17 to -0.07‰ (Fig. 6e and f).

291 The $(\text{La}/\text{Sm})_{\text{N}}$ of MORB samples ranges from 0.4 to 2.0 (Table 3 and Fig. 7a),
292 covering different types of N-MORB, E-MORB and T-MORB (Schilling, 1975;
293 Schilling et al., 1983). Rubidium isotopic compositions of MORBs vary from -0.19 to
294 -0.02‰, with an average of $-0.12 \pm 0.09\%$ (2SD, $n=15$). Most of the different types of
295 MORB samples exhibit similar Rb isotope signatures, except one E-MORB sample
296 (CH30 DR 17-03), which has slightly heavier Rb isotopic composition ($-0.02 \pm$
297 0.02% , 2SD) than others (Fig. 7a and b). Nevertheless, Rb isotope data of MORB
298 samples in this study shows no obvious trend with Rb and MgO contents, Na_8 and
299 Ba/Th (Fig. 7b, c, d and e).

300

301 **5. Discussion**

302 **5.1. Validating high precision Rb isotopic measurements using a Nu Sapphire**

303 **5.1.1 Accuracy and precision**

304 In order to evaluate the accuracy of our method, we first analyzed the pure NIST
305 SRM 984 solution. Self-bracketing NIST SRM 984 solutions generated the expected
306 $\delta^{87}\text{Rb}$ values around zero (within the uncertainty), with an average of $0.00 \pm 0.03\%$

307 (2SD, n=23; Fig. 2a). In addition, a synthetic sample (SRM 984+PCC-1, after Rb
308 separation following column chemistry) was also measured to test the matrix
309 separation efficiency of the chemical purification. This peridotite sample, PCC-1
310 contains almost no Rb (0.07 $\mu\text{g/g}$, Takazawa et al., 2000; Sossi et al., 2016).
311 Approximately 20 mg of PCC-1 powder was digested and then redissolved into 1mL
312 of pure SRM 984 solution which contained 25 ng of Rb. Given the mass ratio of Rb
313 contained in SRM 984 and PCC-1 (i.e., 25:1.4), 95% of the Rb of this synthetic
314 sample comes from the pure SRM 984. Due to its low Rb content, the Rb isotopic
315 composition of PCC-1 has never been measured, but assuming that its $\delta^{87}\text{Rb}$ value is
316 within $\pm 1\text{‰}$ from SRM 984 (this is reasonable because as a peridotite it should have
317 Rb isotopic characteristics comparable to that of the mantle, i.e., $\delta^{87}\text{Rb} = -0.12 \pm$
318 0.06‰ , 2SD; Pringle and Moynier, 2017), this synthetic sample should have isotopic
319 composition that is within error of the pure SRM 984 after column chemistry. The
320 results obtained validate this prediction because the average $\delta^{87}\text{Rb}$ value of the three
321 synthetic samples is $-0.01 \pm 0.03\text{‰}$ (2SD, n=3). This confirms the reliability of matrix
322 separation of our chemistry. The in-house standard IPGP-Rb solution and seven
323 geostandards were analyzed to further test the accuracy of our method (Fig. 2b). All
324 the results presented in this study are consistent with previously published data
325 (Pringle and Moynier, 2017; Zhang et al., 2018; Nie and Dauphas, 2019; Nie et al.,
326 2021; Hu et al., 2021). The long-term reproducibility of our Rb isotopic analyses is
327 0.03‰ (2SD), as estimated from the repeated measurements of AGV-2 ($\delta^{87}\text{Rb} = -0.14$
328 $\pm 0.03\text{‰}$, 2SD, n=17; Fig. 2a and Supplementary Table 4) during 17 analytical

329 sessions over 3 months (where, in each session, AGV-2 was measured at least 3 times).

330 **5.1.2 Effect of Rb concentration**

331 The Nu sapphire coupled to the Apex Omega allows us to obtain the highest
332 sensitivity ($500 \text{ V}/\mu\text{g}\cdot\text{g}^{-1}$ for ^{85}Rb) ever achieved compared with previous studies, e.g.,
333 100 to $\sim 250 \text{ V}/\mu\text{g}\cdot\text{g}^{-1}$ in Pringle and Moynier (2017), $50 \text{ V}/\mu\text{g}\cdot\text{g}^{-1}$ in Zhang et al.
334 (2018) and Nie et al. (2021), and $45 \text{ V}/\mu\text{g}\cdot\text{g}^{-1}$ in Hu et al. (2021). For the purposes of
335 evaluating the effect of measuring concentrations on accuracy and precision, and
336 examining the lower limit of the analytical concentration on the Nu Sapphire, we
337 measured IPGP-Rb, BHVO-2, GSP-2, and AGV-2 at various Rb concentrations. The
338 concentrations of samples and bracketing standards were set from 10 ng/g, 2 ng/g, 1
339 ng/g and even 0.5 ng/g. Under different conditions, even if the Rb concentration was
340 as low as 0.5 ng/g, we could obtain accurate values consistent with previously
341 reported data for these standards. However, the analytical precision decreases
342 significantly at the lowest running concentrations. For example, the precisions
343 reported for 10 ng/g and 2 ng/g are within 0.05‰ (2SD), but precision is generally
344 worse and more variable at 1 ng/g (down to 0.14‰, 2SD). At 0.5 ng/g the precision is
345 0.38‰ (2SD) (Fig. 3 and Supplementary Table 5). Given that at 2 ng/g measurements
346 still provide excellent precision, we set 2 ng/g as the analytical concentration for the
347 rest of this work. This allows us to perform Rb isotopic measurement with only a
348 small amount of sample.

349 **5.1.3 Effects of concentration and acid molarity mismatch**

350 Previous studies demonstrated that a concentration mismatch between standards

351 and samples can have a more significant impact on measurements of K, Ca and Cu
352 isotopes on the Nu Sapphire compared to Neptune or Neptune Plus, and an intensity
353 match within 1% was suggested during analysis (Moynier et al., 2021; Dai et al., 2021;
354 Wang et al., 2022). In order to test the effect of concentration mismatch on Rb
355 isotopic measurements, pure SRM 984 solutions with different Rb concentrations
356 were analyzed as unknown samples. The bracketing standard was kept at 10 ng/g,
357 while samples were run with concentrations varying from ~5 ng/g to 15 ng/g (i.e., the
358 ratio of sample/standard varied from just under 0.5 to a bit over 1.5). Unlike K and Ca
359 isotopes, Rb isotopic analyses on Nu Sapphire are insensitive to the concentration
360 mismatch (Fig. 4a and Supplementary Table 6)

361 It was also reported that Nu Sapphire measurements may be more sensitive to
362 mismatch of acid molarity between samples and bracketing standard (Moynier et al.,
363 2021; Dai et al., 2021; Chen et al., 2021; Li et al., 2022; Wang et al., 2022; Zheng et
364 al., 2022). The influence of acid molarity mismatch on Rb isotopic measurements was
365 assessed by a series of condition tests. Pure SRM 984 solutions were treated as
366 samples, and measured with the same Rb concentration (i.e., 2 ng/g) as the bracketing
367 standard. However, they were prepared with different molarities of acid. The standard
368 was diluted using 0.5 mol/L HNO₃, while samples were diluted in HNO₃ at various
369 concentrations (from 0.1 mol/L to 1.0 mol/L). The results show that acid molarity
370 mismatch has no obvious effect on the Rb isotopic measurements when analyzed on
371 the Nu Sapphire (Fig. 4b and Supplementary Table 6). Nevertheless, we use the same
372 HNO₃ (0.5 mol/L) to prepare samples and standard solutions during measurements.

373 **5.1.4 Effect of matrix elements**

374 A small amount of matrix residues after column chemistry may also influence
375 isotope measurements. For example, interferences of isobars and polyatomic ions, and
376 influences of matrix on ionization and transmission of target elements, can affect the
377 quality of the isotopic measurements (Tan and Horlick, 1987; Tanner, 1992; Chen et
378 al., 2021). In particular, previous studies have reported that some matrix elements can
379 strongly affect Rb isotopic values during mass spectrometry analysis (Zhang et al.,
380 2018; Nie and Dauphas, 2019; Hu et al., 2021). In this study, we tested these effects
381 on the Nu Sapphire, of Sr, K, Cr, Ga, Ca, Ti, Fe, Mn, Al and Na. Pure standard
382 solutions of various elements were doped into a 2 ng/g SRM 984 to obtain different
383 ratios of [element]/[Rb]. The doped solutions were measured against pure SRM 984
384 with the same Rb concentration. The results are presented in Fig. 5 and
385 Supplementary Table 7.

386 Strontium correction with a fixed $^{87}\text{Sr}/^{88}\text{Sr}$ value of 0.085 is effective only when
387 there is a very small amount of Sr residue in the Rb solution. The result of our Sr
388 doping test showed that accurate $\delta^{87}\text{Rb}$ values could be obtained when $\text{Sr}/\text{Rb} \leq 0.005$
389 (i.e., $^{88}\text{Sr}/^{85}\text{Rb}$ intensity ≤ 0.006). In order to minimize the impact of the isobaric
390 interference of ^{87}Sr on ^{87}Rb , all natural samples in this study had $^{88}\text{Sr}/^{85}\text{Rb}$ intensity \leq
391 0.003 (for most samples the ratio was ≤ 0.002 or lower) after chemical purification,
392 and therefore the Sr correction did not affect the quality of the $\delta^{87}\text{Rb}$ data. Because Rb
393 and K have very similar geochemical behavior, a challenge of our method is to
394 achieve complete separation of Rb and K. Previous studies have shown that K affects
395 Rb isotopic measurements, but to different degrees between different instruments. For

396 example, Pringle and Moynier (2017) (Neptune Plus with quartz cyclonic spray
397 chamber or Apex) and Zhang et al. (2018) (Neptune Plus with quartz dual cyclonic-
398 Scott spray chamber) found that the K/Rb ratios must be maintained below 10 and 25,
399 respectively, to obtain accurate $\delta^{87}\text{Rb}$ values. Nie and Dauphas (2019) (Neptune with
400 quartz dual cyclonic-Scott spray chamber) found that it was sufficient to keep it below
401 50. A more recent study found that the isotopic measurements were very sensitive to
402 K residue (Neptune Plus with quartz dual cyclonic-Scott spray chamber; Hu et
403 al.,2021) with the $\delta^{87}\text{Rb}$ values shifted when $\text{K/Rb} > 1.5$. Thus, it is necessary to
404 evaluate the effect of K on Rb isotopic measurements on the Nu Sapphire coupled to
405 an Apex Omega. Our doping test results showed that Rb isotopic analyses were not
406 affected by K as long as the K/Rb ratio ≤ 20 . Chromium and Ga were doped here to
407 test the effect of some possible impurities, since they are eluted very closely to Rb
408 according to the elution curves presented in Pringle and Moynier (2017). The
409 influence was negligible as long as Cr/Rb was ≤ 10 and Ga/Rb was ≤ 20 . Other matrix
410 elements (such as Ca, Ti, Fe, Mn, Al and Na) affected Rb isotopic measurements to
411 different extent. $\delta^{87}\text{Rb}$ values were modified when Ti/Rb was ≥ 20 , Fe/Rb was ≥ 20
412 and Al/Rb was ≥ 10 , while Ca, Na and Mn had no obvious influence within the range
413 of concentration ratios set in our doping tests (i.e., $[\text{element}]/[\text{Rb}] \leq 50$).

414 **5.2. Magmatic processes do not fractionate Rb isotopes**

415 Igneous products of the Hekla volcano have been widely used to investigate
416 behaviors of different isotopic systems during magmatic processes. For example, Th,
417 Fe, Si, B and V isotopic compositions of Hekla igneous rocks correlate with SiO_2

418 contents, suggesting that magmatic evolution has a significant influence on these
419 isotope systems (Sigmarsson et al., 1992; Rose-Koga and Sigmarsson, 2008;
420 Schuessler et al., 2009; Prytulak et al., 2016; Savage et al., 2011). Moreover, highly
421 light Cl isotopic values observed in high-Si rocks from Hekla have been interpreted to
422 reflect the assimilation of magmatic brine (Ranta et al., 2021). However, no isotopic
423 variation was observed in the same set of samples for Li, Zn, Mo, Tl, K and Ba
424 isotopes (Schuessler et al., 2009; Chen et al., 2013; Yang et al., 2015; Prytulak et al.,
425 2017; Tuller-Ross et al., 2019; Bai et al., 2022). A set of igneous rocks from the Hekla
426 volcano were analyzed in this study in order to explore the Rb isotopic behaviors
427 during magmatic evolution. In the following sections, we will systematically evaluate
428 different magmatic processes which have the potential to fractionate Rb isotopes.

429 **5.2.1 Assimilation and/or exsolution of magmatic fluids**

430 As Rb is a fluid-mobile element, hydrothermal fluid activities may lead to its
431 isotopic fractionation. For example, X. Hu et al. (2022) reported a highly heavy Rb
432 isotopic signature (-0.29‰ to 1.20‰) in Kampa leucogranites compared to the
433 available data on sedimentary materials of the UCC (-0.31‰ to 0.07‰). After ruling
434 out the influences of source heterogeneity, partial melting, fractional crystallization
435 and chemical weathering, this highly heavy Rb isotopic signature was attributed to the
436 involvement of magmatic fluids (X. Hu et al., 2022). This is consistent with previous
437 studies that have also shown that Kampa leucogranite magma was modified by fluids
438 during its evolution (Liu et al., 2020; Huang et al., 2021). Therefore, the influence of
439 assimilation and/or exsolution of magmatic fluids need to be considered.

440 Previous chlorine isotopes study indicates that the assimilation of magmatic
441 brine plays an important role in generating light chlorine isotopic characteristics
442 evident in Hekla rhyolites (Ranta et al., 2021). However, we do not observe any
443 abnormal signals in high-Si rocks from Hekla, either in terms of Rb concentration or
444 isotopic composition. For example, Rb contents follow a consistent evolution trend
445 when plotted against SiO₂ and MgO contents (Fig. 6a and b). Besides, compared with
446 associated basalts ($\delta^{87}\text{Rb}_{\text{average}} = -0.12 \pm 0.06\text{‰}$, 2SD, n=8) and basaltic andesites
447 ($\delta^{87}\text{Rb}_{\text{average}} = -0.15 \pm 0.06\text{‰}$, 2SD, n=6), dacites ($\delta^{87}\text{Rb}_{\text{average}} = -0.15 \pm 0.03\text{‰}$, 2SD,
448 n=3) and one rhyolite sample ($\delta^{87}\text{Rb} = -0.17 \pm 0.04\text{‰}$) have similar Rb isotopic
449 compositions. This similarity indicates that the influence of magmatic brine on Rb
450 isotope in our samples is not obvious, if indeed magmatic brines are concentrated in
451 rhyolites relative to basalts and dacites. In addition, fluid exsolution is also unlikely to
452 affect Rb isotopic signature of Hekla rocks, given the identical Li isotopic
453 composition of the different Hekla lavas, which indicates that exsolution of fluids
454 from melt is not significant (Schuessler et al., 2009).

455 **5.2.2 Degassing**

456 Rubidium is a moderately volatile element with a T_c^{50} of 800 K (Lodders, 2003).
457 Previous studies on extraterrestrial samples indicate that volatile loss during
458 evaporation may lead to Rb isotopic fractionation (Pringle and Moynier, 2017; Nie
459 and Dauphas, 2019). Hekla lavas experience significant loss of F, Cl and S during
460 eruptions (Moune et al., 2007; Ranta, et al., 2021). Therefore, Rb isotope fractionation
461 caused by volatile loss during volcanic degassing is possible.

462 While the Rb isotopic compositions of Hekla rocks are relatively uniform and
463 non-systematic, the total variation observed ($\sim 0.10\%$) exceeds the precision of our
464 method (0.03%). Rubidium isotopes can therefore be used to estimate the amount of
465 Rb loss during volcanic degassing at Hekla. Rubidium is evaporated as RbCl gas
466 during degassing, and the theoretical Rayleigh fractionation coefficient (α) can be
467 calculated as the square root of the mass difference between $^{85}\text{RbCl}$ and $^{87}\text{RbCl}$ [i.e.,
468 $(120.3648/122.3622)^{0.5} = 0.9838$]. Based on this, we can estimate a maximum of 0.6%
469 of Rb is lost during volcanic degassing, which is similar to the loss amount of K
470 (0.8~1.3%) estimated by Tuller-Ross et al. (2019). However, it should be noted that
471 this estimation is based on the assumption that the Rb isotopic variation of Hekla
472 rocks is entirely due to degassing. As a result, the actual loss of Rb during volcanic
473 degassing may be much less than our estimated value. This is supported by the fact
474 that the Rb contents are consistent with the behavior of other non-volatile
475 incompatible elements (e.g., La) during magma evolution (Fig. 6d), indicating that
476 there may be no significant loss of Rb following magma evolution at Hekla. Therefore,
477 we conclude that volatile loss associated with petrogenesis of the Hekla rock suite is
478 unlikely to fractionate Rb isotopes.

479 5.2.3 Magmatic evolution

480 The Hekla basaltic to intermediate rocks are commonly considered to be
481 controlled by fractional crystallization process (Sigmarsson et al., 1992, 2022; Geist et
482 al., 2021). Currently, there is an ongoing debate regarding the petrogenesis of the
483 silicic rocks from Hekla. For example, studies of short-lived radiogenic isotopes

484 (Sigmarsson et al., 1992, 2022; Chekol et al., 2011) and B and Cl stable isotopes
485 (Rose-Koga and Sigmarsson, 2008; Ranta et al., 2021) suggest that crustal anatexis
486 and assimilation are important processes for the formation of dacites, from which
487 rhyolites are generated via fractional crystallization. Geist et al. (2021) suggested a
488 “fractional crystallization alone” model, arguing that the evolution of magma at Hekla
489 was primarily driven by fractionation crystallization based on trends in trace element
490 ratios, as well as oxygen isotopic data, with which they excluded significant
491 contribution from the crust. The Hekla dacites, together with other types of rocks,
492 show a good correlation between their Rb concentrations and SiO₂ as well as MgO
493 contents, strictly following a magmatic differentiation trend (Fig. 6a and b). However,
494 an influence of the crust cannot be ruled out, as partial melting of metabasaltic crust to
495 form dacites also fits the consistent trend of elemental Rb. For example, we can
496 roughly evaluate the contribution of the crust melting by using the equation:
497 $C_1/C_0=F^{(D-1)}$, where C_1 is the Rb content of the melt (i.e., dacites), C_0 is Rb content of
498 parent material (i.e., metabasalts), F is the degree of partial melting, and D is bulk
499 partition coefficient of Rb. Using the average Rb concentration of dacites (54.8 μg/g)
500 and basalts (10.5 μg/g) to represent C_1 and C_0 respectively. And assuming that the
501 distribution coefficient of Rb during partial melting of the metabasalts is between 0
502 and 0.2 (which is the highest Rb partitioning of the dominant phenocryst phases in
503 Hekla rocks according to Geist et al., 2021), it is estimated that the dacites can be
504 formed by 13% to 19% partial melting of metabasalts. This is consistent with previous

505 estimates using elemental Th and U (10~20%, Sigmarsson et al., 1992, 2022) as well
506 as Tl (16%, Prytulak et al., 2017).

507 In cases where traditional major and trace elements analyses fail to distinguish
508 between competing petrogenesis models, emerging isotope systems may hold
509 potential for providing new insights. Therefore, we aim to investigate whether Rb
510 isotopes can offer new insights into the ongoing debate over the petrogenesis of Hekla.
511 The anhydrous phenocryst assemblage of Hekla lavas is dominated by plagioclase,
512 olivine, clinopyroxene, orthopyroxene, magnetite, apatite, ilmenite and zircon
513 (Sigmarsson et al., 1992, 2022; Geist et al., 2021). Based on partition coefficients
514 from Geist et al. (2021), we anticipate that only a limited amount of Rb would enter
515 these phases during fractional crystallization. This is consistent with what is observed
516 for K (Tuller-Ross et al., 2019). Since Rb and K are both highly incompatible
517 elements and the concentration of these elements correlate in the Hekla samples (Fig.
518 6c), they must exhibit very similar geochemical behavior during magmatic evolution.
519 Tuller-Ross et al. (2019) proposed that K-bearing minerals were absent when solid
520 phases crystallized from melt at Hekla. Rubidium mainly exists in the form of
521 isomorphic substitution of K, thus Rb would not get incorporated into solid phases
522 during magmatic evolution. This is also supported by the good correlation between Rb
523 and La (Fig. 6d) as well as other trace elements such as Tl, Ba, U, Th and Pb
524 (Sigmarsson et al., 1992; Prytulak et al., 2017; Geist et al., 2021). If considering that
525 the Hekla system undergoes fractional crystallization alone as proposed by Geist et al.
526 (2021), the magmatic process would not lead to significant Rb isotopic fractionation,

527 given that limited Rb enters the solid phases during fractional crystallization. For
528 comparison, we also evaluate the impact in the case of crustal involvement during
529 magmatic evolution. Prytulak et al. (2017) suggested that the contributions of crustal
530 components were highly unlikely to be low temperature altered crust, pelagic clays
531 and/or Fe-Mn sediments, based on the homogeneity of the Tl isotopic compositions
532 across the range of magmatic products at Hekla. A recent study shows that the Rb
533 isotopic composition of the upper continental crust ($-0.14 \pm 0.01\text{‰}$, 2SE, n=73; X. Hu
534 et al., 2022) is close to the mantle's composition ($-0.12 \pm 0.06\text{‰}$, 2SD, n=6; Pringle
535 and Moynier., 2017). The possible similarity of the Rb isotopic composition between
536 the crust and mantle implies that even if the origin of dacites is influenced by the
537 crustal contribution, they still have the same Rb isotopic composition as less evolved
538 rocks. The two competing models (i.e., assimilation and/or partial melting of the crust
539 versus fractional crystallization alone, Sigmarsson et al., 1992, 2022; Chekol et al.,
540 2011; Geist et al., 2021), accordingly, both agree with the absence of resolvable Rb
541 isotopic variation between the dacites ($\delta^{87}\text{Rb}_{\text{average}} = -0.15 \pm 0.03\text{‰}$, 2SD, n=3) and
542 other less evolved rock types (average $\delta^{87}\text{Rb}$ value of $-0.12 \pm 0.06\text{‰}$, 2SD, n=8 for
543 basalts and $-0.15 \pm 0.06\text{‰}$, 2SD, n=6 for basaltic andesites). Therefore, the ongoing
544 petrogenesis debate is not resolved with Rb isotopes.

545 The origin of dacites is still controversial. However, despite this uncertainty, the
546 uniformity of $\delta^{87}\text{Rb}$ values in the Hekla rock suite suggests that magmatic processes
547 do not significantly modify the Rb isotopic characteristics of the source melt. This is
548 consistent with the prediction made by Zeng et al. (2019) through first-principle

549 calculations. They estimated that the Rb isotopic equilibrium fractionation between
550 silicate melt and mineral would not exceed 0.01‰ at 900~1000°C, which is within
551 our analytical uncertainty. It is, however, worth noting that Hekla is an anhydrous,
552 tholeiitic system, and the behavior for Rb isotope could be different if a hydrous (e.g.,
553 biotite) or K-rich phases are involved during magma evolution.

554 **5.3. Rubidium isotopic composition of the Earth's mantle**

555 Magma emplaced at mid-ocean ridges makes up a major part of terrestrial
556 magma production (White and Klein, 2014; Nan et al., 2022; Wu et al., 2023).
557 MORBs are generated by partial melting of the mantle (Hofmann, 2003; White and
558 Klein, 2014). According to Workman and Hart (2005), the bulk distribution
559 coefficient $D_{\text{solid/melt}}$ of Rb during mantle partial melting can be low as 0.00001, and
560 MORBs correspond to melting degrees between 6% and 20%. This means that during
561 the partial melting of the mantle, more than 99% of Rb tends to enter the basaltic melt,
562 and therefore we can safely assume that Rb isotopic fractionation does not occur
563 during partial melting. The lack of correlation between Rb isotopic compositions and
564 the indicator of partial melting degree Na_8 (which is defined as $Na_8 = Na_2O + 0.373 \times$
565 $MgO - 2.98$; Klein and Langmuir, 1987) of MORB samples in this study (Fig. 7d) is
566 consistent with this assumption. In addition, Rb isotopic fractionation is very limited
567 during magmatic evolution based on no obvious correlation between $\delta^{87}Rb$ and MgO
568 contents (Fig. 7c) as well as our observation of Hekla volcanic rocks. Therefore, the
569 Rb isotopic characteristics of MORB samples can reflect the composition of their
570 mantle source.

571 In this study, various MORB samples with a broad geographical distribution
572 were analyzed for Rb isotopes. All but one of these samples span a limited range of
573 Rb isotopic compositions (-0.19 to -0.07‰) despite different localities. One exception
574 is the E-MORB sample, CH30 DR 17-03, which is found to have a slightly heavier
575 Rb isotopic composition ($\delta^{87}\text{Rb} = -0.02 \pm 0.02\text{‰}$, Table 3 and Fig. 7a). Previous studies
576 proposed that formation of E-MORBs may be related to the involvement of deep
577 recycling of crustal materials (Hofmann, 2003; Elliott et al., 2006). Although a recent
578 study estimated the Rb isotopic composition of the upper continental crust is $-0.14 \pm$
579 0.01‰ (2SE, n=73; X. Hu et al., 2022), low temperature alteration process would lead
580 to significant Rb isotopic fractionation (Zhang et al., 2021) thus surface sediments or
581 altered oceanic crust (AOC) may have abnormal Rb isotopic characteristics.
582 Recycling of these materials may cause the inhomogeneity of the local mantle sources
583 of MORBs. This may account for the heavier Rb isotopic characteristic associated
584 with higher Rb content observed in CH30 DR 17-03 (Fig. 7b). However, the addition
585 of abnormal Rb was unlikely via fluid, since extremely high Ba/Th, an indicator of
586 fluid enrichment, is not observed in CH30 DR 17-03 (Fig. 7e). It is therefore,
587 necessary to further test if Rb isotopes can be used to trace subduction recycling into
588 the mantle, for example by targeting altered oceanic crust or sediments. In addition, it
589 is worth noting that N-MORBs show more scattered Rb isotopic compositions
590 compared to T-MORBs (Fig. 7a). A similar scatter for Sn isotopes was also observed
591 in this set of samples (She et al., 2023). The scatter was unlikely caused by secondary
592 modification, as the MORBs in this study did not undergo significant alteration

593 according to the CIA and LOI values (She et al., 2023). Thus, it is more likely related
594 to the diversity of sampling sites as suggested by She et al. (2023), which span
595 locations as diverse as the Atlantic, Indian, and East Pacific Ocean. Nevertheless, Rb
596 isotopic compositions of all MORB samples cluster within a narrow range. Based on
597 the average $\delta^{87}\text{Rb}$ value of these MORBs, the Rb isotopic composition of the upper
598 mantle is estimated at $-0.12 \pm 0.09\text{‰}$ (2SD, n=15). This value is consistent with the
599 previous estimate ($-0.12 \pm 0.06\text{‰}$, 2SD; n=6) from Pringle and Moynier (2017). If we
600 assume that Rb isotopic compositions of OIBs also reflect the characteristics of their
601 mantle source, we can combine MORBs and Hekla basalts in this study with the OIB
602 data reported in the literature (Pringle and Moynier, 2017) to estimate the Rb isotopic
603 composition of the mantle. This compilation yields a $\delta^{87}\text{Rb}$ value = $-0.12 \pm 0.08\text{‰}$
604 (2SD, n=25; Fig. 8).

605 **5.4. Rubidium isotopic composition of the Bulk Silicate Earth**

606 As mentioned above, Rb is a strongly incompatible element, and it can be
607 calculated that the continental crust hosts ~ 55% of the Earth's Rb budget (Rudnick
608 and Gao, 2003; Pringle and Moynier, 2017). Thus, the contribution of the continental
609 crust actually plays a major role in the Rb isotopic composition of the BSE. As mantle
610 partial melting and magmatic processes do not produce significant Rb isotopic
611 fractionation (this study), the continental crust should have similar Rb isotopic
612 characteristics to that of the mantle. This is consistent with the similarity between the
613 estimated Rb isotopic composition of the upper continental crust ($-0.14 \pm 0.01\text{‰}$, 2SE,
614 n=73; X. Hu et al., 2022) and our estimation of the mantle's composition ($-0.12 \pm$

615 0.08‰, 2SD; n=25). Given the relative proportions of the crust and mantle, the Rb
616 isotopic composition of the BSE can be estimated through a mass balance calculation:
617 $\delta^{87}\text{Rb}_{\text{BSE}} = f \times \delta^{87}\text{Rb}_{\text{crust}} + (1-f) \times \delta^{87}\text{Rb}_{\text{mantle}}$, where f (~55%, Rudnick and Gao, 2003;
618 Pringle and Moynier, 2017) is the proportion of BSE's Rb budget hosted by the
619 continental crust. The rubidium isotopic composition of the continental crust
620 ($\delta^{87}\text{Rb}_{\text{crust}}$) is assumed to be represented by the average value of the UCC (-0.14 ±
621 0.09‰, 2SD, n=73; X. Hu et al., 2022), and $\delta^{87}\text{Rb}_{\text{mantle}}$ to be equal to -0.12 ± 0.08‰
622 (2SD, n=25) as discussed above. Using this approach, the Rb isotopic composition of
623 the BSE is estimated to be -0.13 ± 0.06‰ (2SD; propagated error).

624

625 **6. Conclusions**

626 In this study, we developed a high precision Rb isotope analytical method on
627 CC-MC-ICP-MS Nu Sapphire. Seven geostandards as well as a synthetic sample
628 (PCC-1+SRM 984) were measured. $\delta^{87}\text{Rb}$ values reported here for these standards are
629 all in good agreement with reference data, confirming the accuracy of the method.
630 The long-term reproducibility of Rb isotopic analyses is 0.03‰ (2SD) at 2 ng/g
631 running concentration. Based on this high precision method, a series of igneous rocks
632 from Hekla volcano and 15 MORB samples with various localities were measured.
633 Hekla volcanic rocks exhibit very uniform Rb isotopic composition (from -0.17 to -
634 0.07‰) over a wide range of bulk-rock compositions, ranging from basalts to
635 rhyolites, indicating that Rb isotopic fractionation during magmatic evolution is not
636 significant. Similarly, MORB samples also have a narrow Rb isotopic variation range

637 (-0.19 to -0.02‰) regardless of their various localities. Combining basaltic samples
638 (MORBs and Hekla basalts) in this study and previously published OIB data, the Rb
639 isotopic composition of the mantle is estimated at $-0.12 \pm 0.08\text{‰}$ (2SD, n=25). The
640 average composition of the BSE is evaluated to be $-0.13 \pm 0.06\text{‰}$ (2SD), and this is a
641 more precise estimate that is consistent with previous studies.

642

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652

653

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886 **Figure 1.** Elution curves for Rb and K of our Sr-spec column chemistry. The
887 homemade heat-shrink Teflon columns (0.46 cm inner diameter, 6.5 cm length)

888 contain 1 mL of Eichrom Sr-spec resin (25–50 μm) and the elution are realized using
889 3 mol/L HNO_3 . Relevant data see Supplementary Table 3.

890 **Figure 2.** Rubidium isotopic data of standards. (a): The $\delta^{87}\text{Rb}$ of pure NIST SRM 984
891 solution and AGV-2. The long-term reproducibility of Rb isotopic analyses is 0.03‰
892 (2SD) based on the repeated measurements of AGV-2; (b): Rubidium isotopic
893 compositions of one in-house standard IPGP-Rb solution and seven geostandards. The
894 data are reported as $\delta^{87}\text{Rb}$ relative to the SRM 984 solution. Data from the literature
895 are reported for comparison.

896 **Figure 3.** Effect of Rb testing concentrations on accuracy and precision. Rubidium
897 concentrations decrease from 10 to 0.5 ng/g for a pure Rb standard as well as for three
898 rock standards (BHVO-2, AGV-2, GSP-2). Relevant data are reported in
899 Supplementary Table 5.

900 **Figure 4.** Effects of concentration and acid molarity mismatch on Rb isotopic
901 measurements. Samples and standards are both pure SRM984 solutions. (a): the
902 bracketing standard was kept at 10 ng/g, samples with different Rb concentrations
903 were prepared to obtain the sample/standard concentration ratio varying from 0.5 to
904 1.5; (b): samples and bracketing standard were prepared with different molarities of
905 acid. Samples were dissolved in HNO_3 from 0.1 mol/L to 1.0 mol/L, while the
906 standard was diluted in 0.5 mol/L HNO_3 .

907 **Figure 5.** Matrix effects on Rb isotopic measurements evaluated by doping the
908 solutions with various elements. Rubidium concentrations were set to 2 ng/g, and

909 [element]/[Rb] represents the concentration ratio of different elements to Rb, e.g.,
910 [K]/[Rb]=20 corresponds to K of 40 ng/g.

911 **Figure 6.** Rubidium elemental and isotopic characteristics of Hekla volcanic rocks.
912 (a): Rb concentration plotted against SiO₂ content; (b): Rb concentration plotted
913 against MgO content; (c): Rb concentration plotted against K₂O content; (d): Rb
914 concentration plotted against La content; (e): $\delta^{87}\text{Rb}$ vs. SiO₂ content; (f): $\delta^{87}\text{Rb}$ vs. Rb
915 concentration.

916 **Figure 7.** Correlations between $\delta^{87}\text{Rb}$ and (a) (La/Sm)_N, (b) Rb concentration, (c)
917 MgO content, (d) Na₈ (indicator of partial melting degree which is defined as
918 Na₈=Na₂O + 0.373 × MgO – 2.98; Klein and Langmuir, 1987), and (e) Ba/Th for
919 MORB samples.

920 **Figure 8.** Average Rb isotopic composition of the Earth's mantle ($\delta^{87}\text{Rb}=-0.12 \pm$
921 0.08% , 2SD) estimated based on MORBs and Hekla basalts in this study as well as
922 previously reported OIB (Pringle and Moynier, 2017).

Figure 1

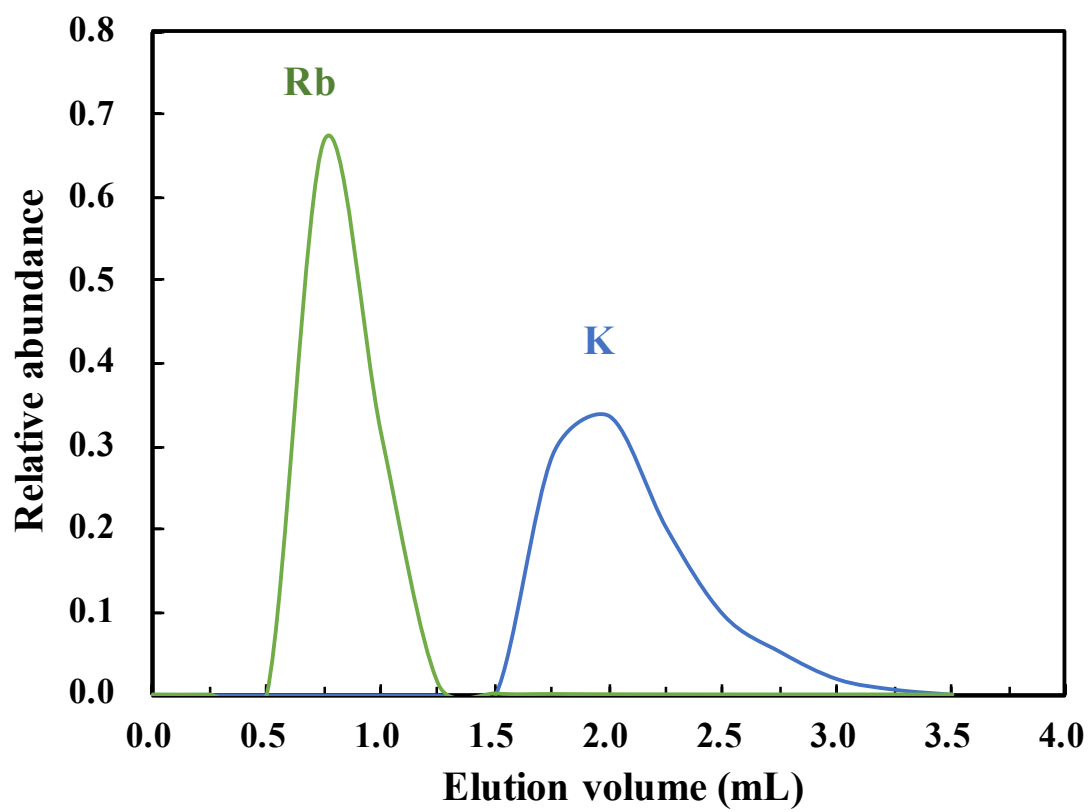


Figure 2

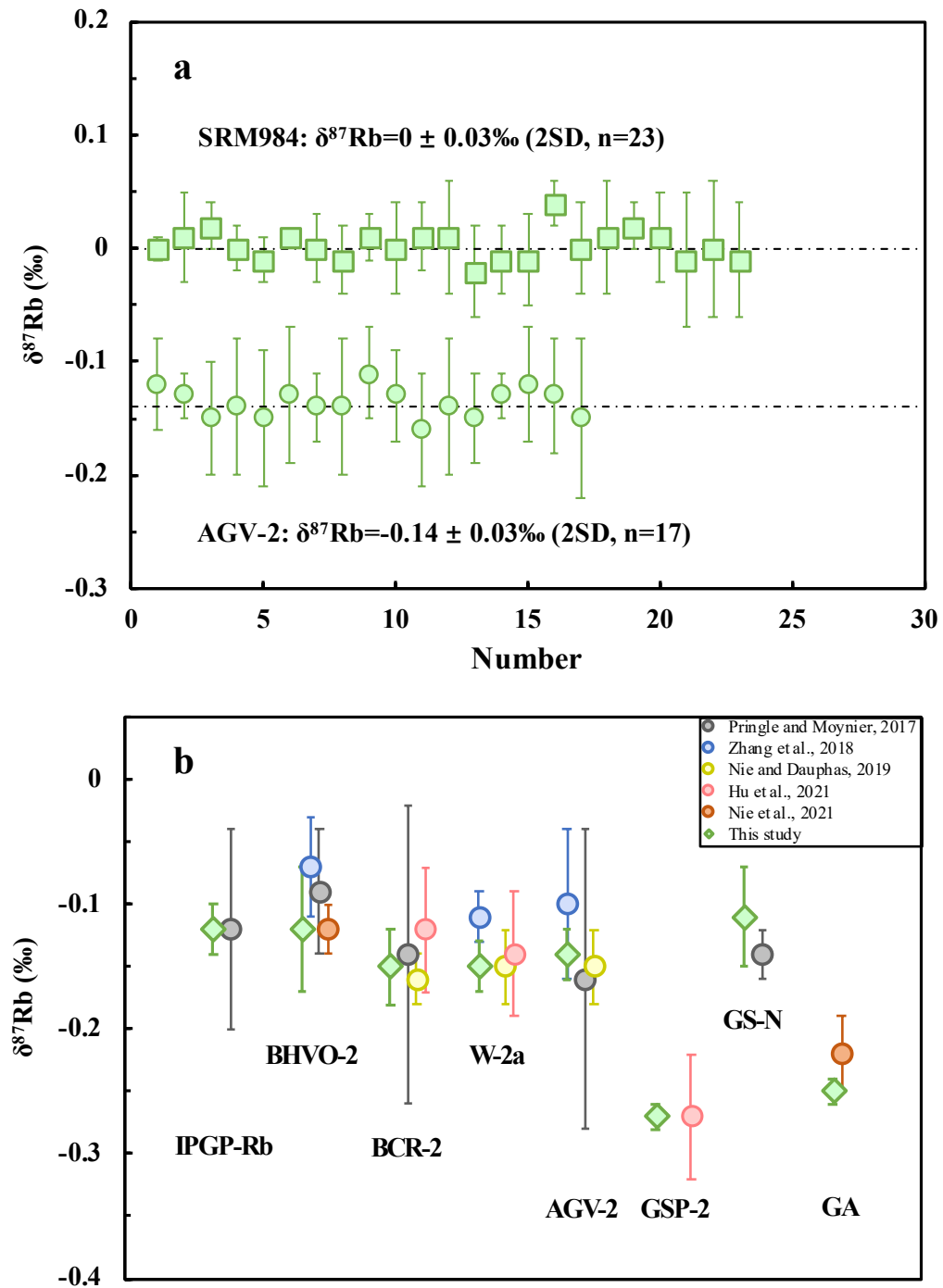


Figure 3

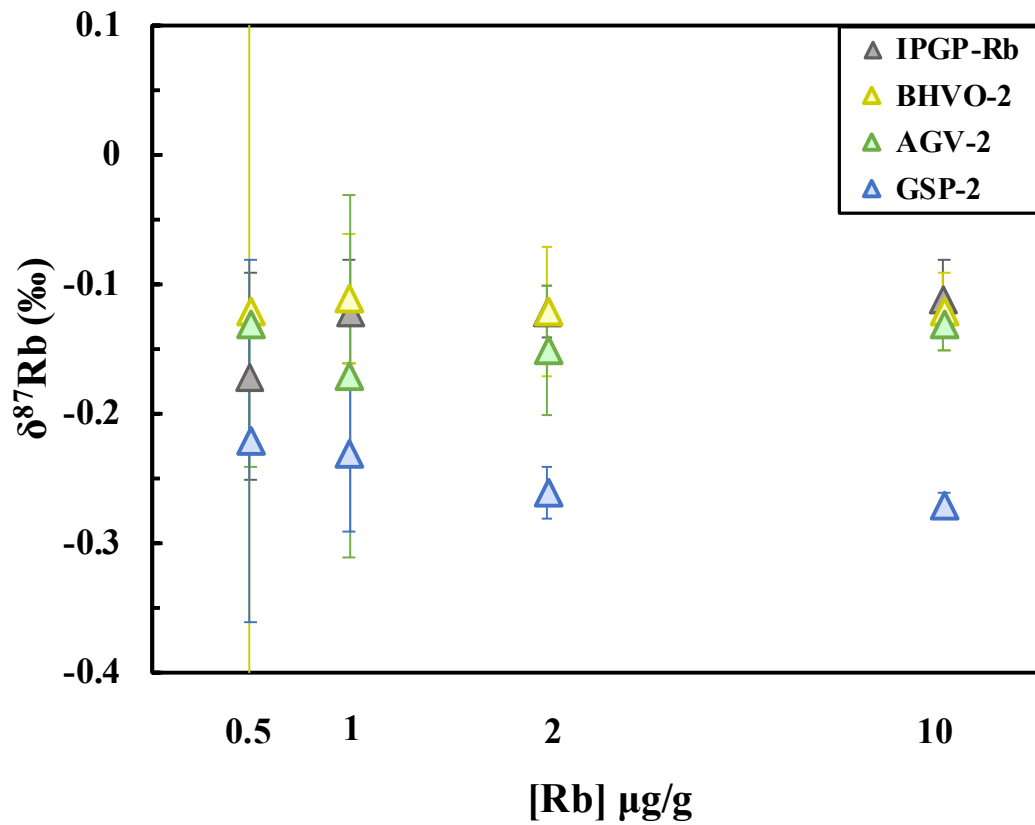


Figure 4

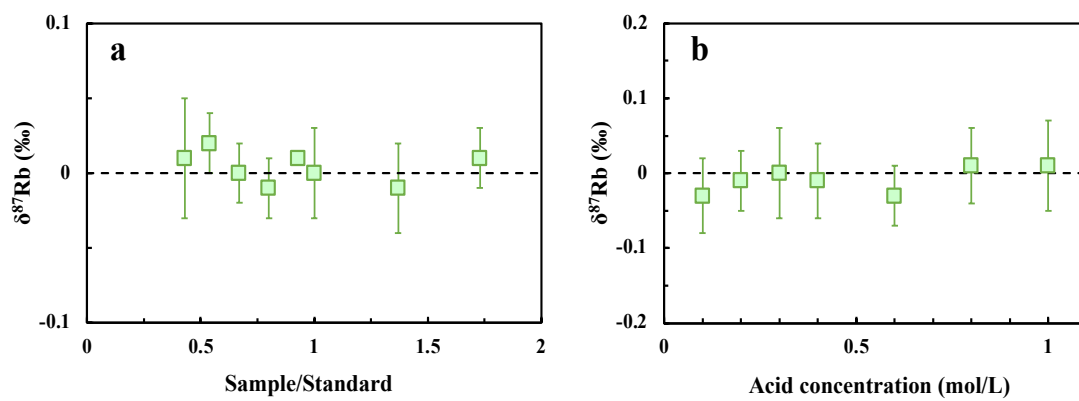


Figure 5

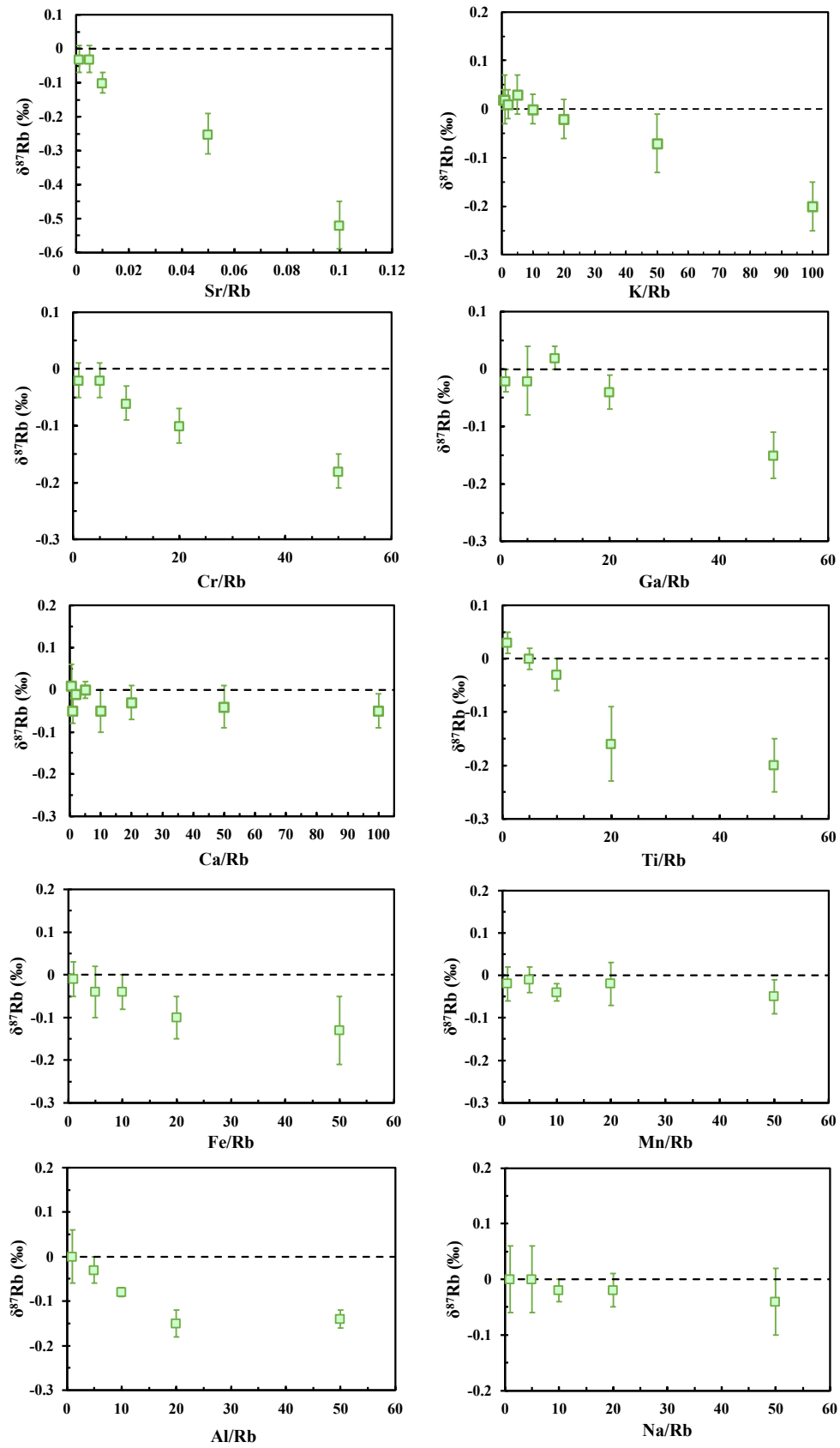


Figure 6

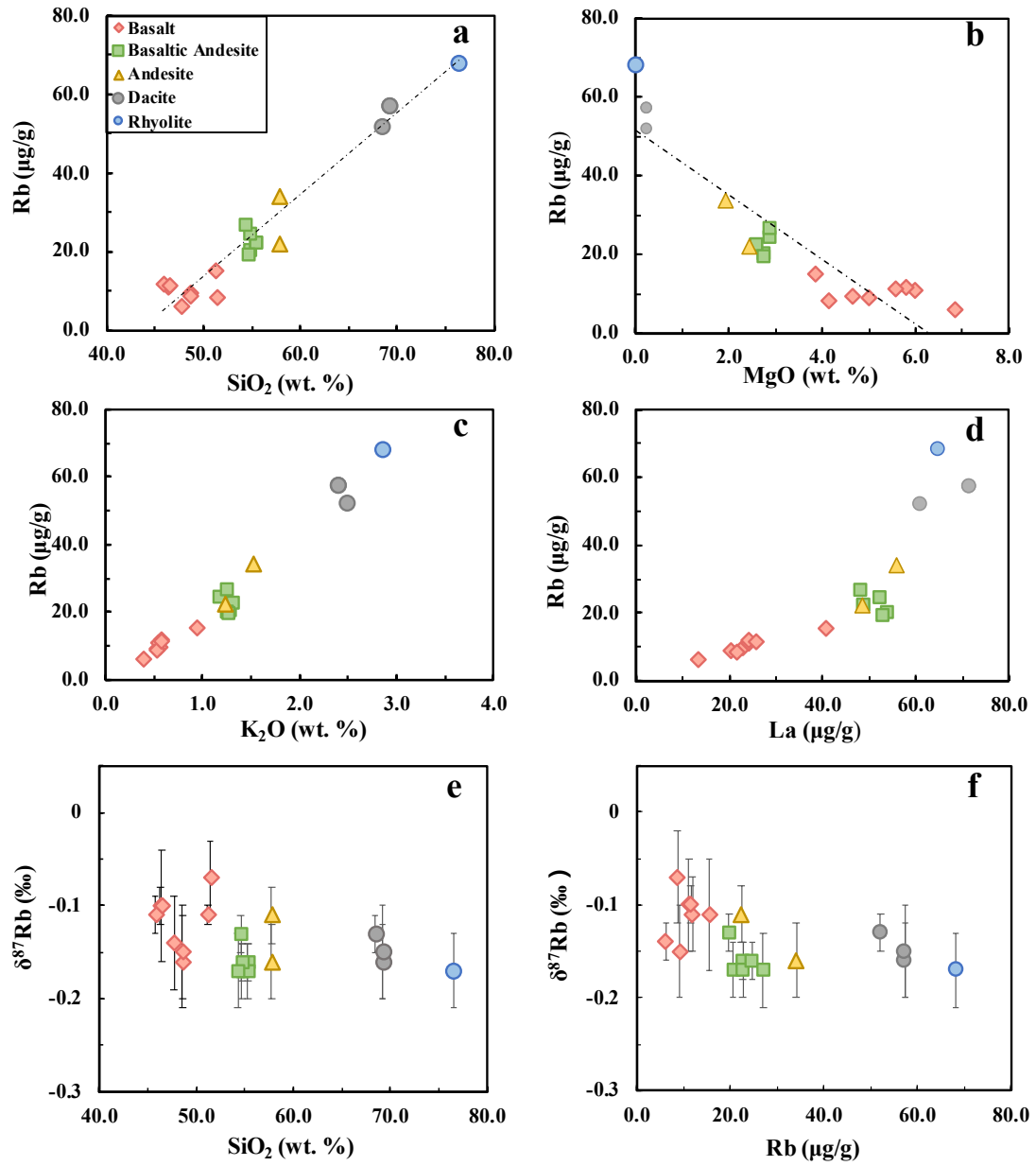


Figure 7

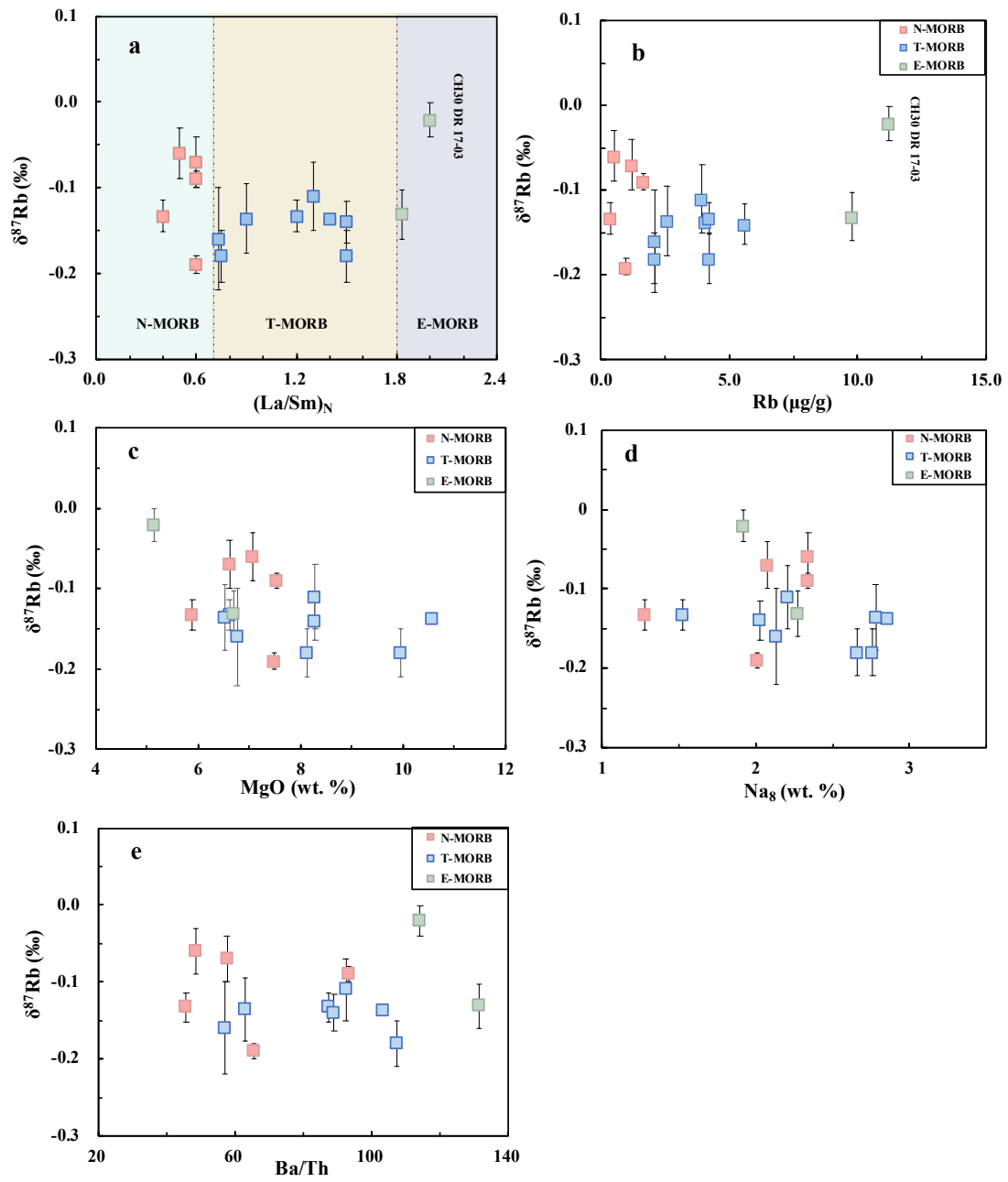


Figure 8

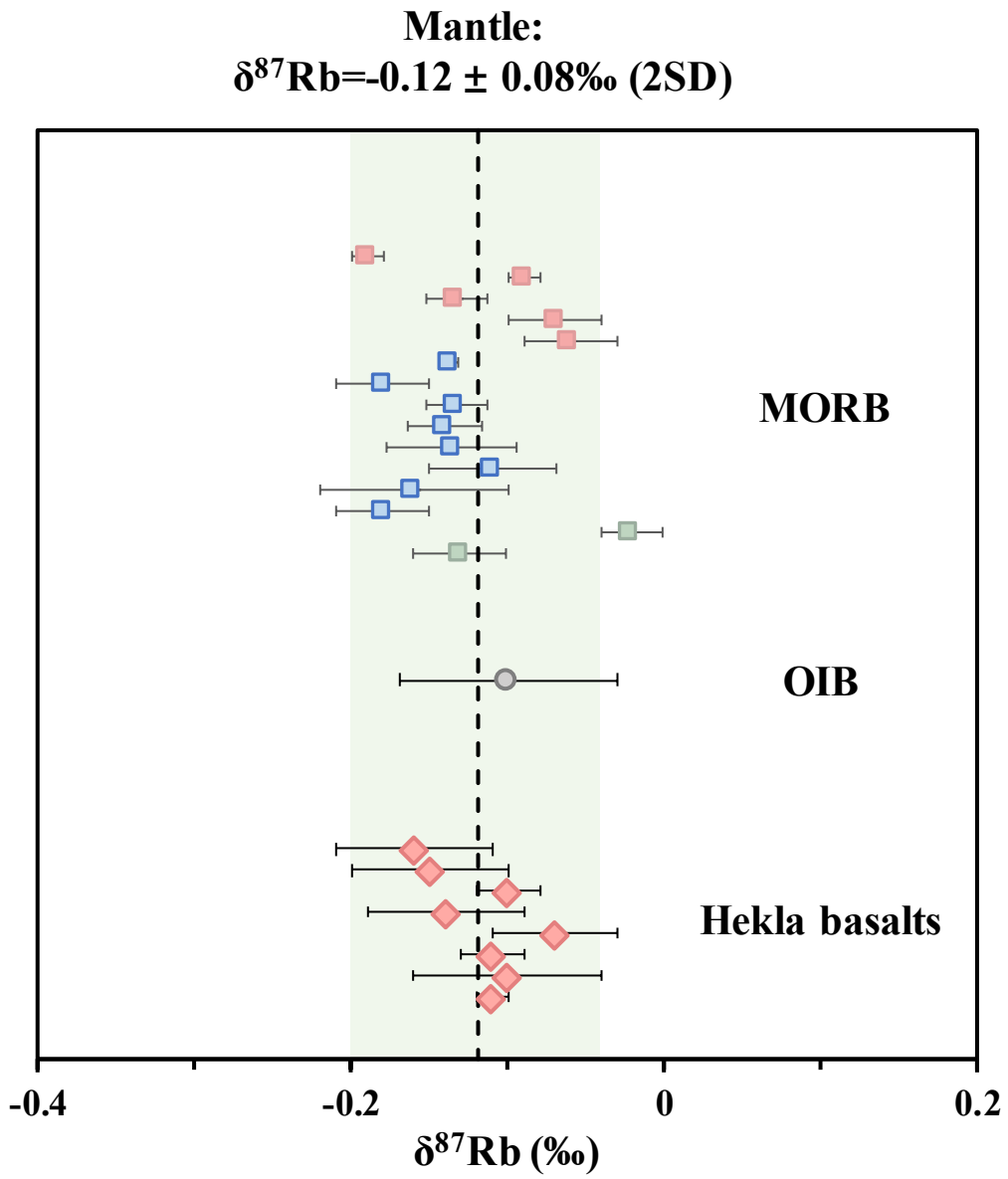


Table 1. Rubidium isotopic compositions of standards.

Standard	<i>This study</i>		<i>Reference value</i>		Literature
	$\delta^{87}\text{Rb}$ (‰)	2SD	$\delta^{87}\text{Rb}$ (‰)	2SD	
IPGP-Rb	-0.12	0.02	-0.12	0.08	Pringle and Moynier (2017)
BHVO-2	-0.12	0.05	-0.12	0.02	Nie et al. (2021)
			-0.09	0.05	Pringle and Moynier (2017)
			-0.07	0.04	Zhang et al. (2018)
BCR-2	-0.15	0.03	-0.12	0.03	Hu et al. (2021)
			-0.14	0.12	Pringle and Moynier (2017)
			-0.16	0.02	Nie and Dauphas (2019)
W-2a	-0.15	0.02	-0.14	0.05	Hu et al. (2021)
			-0.11	0.02	Zhang et al. (2018)
			-0.15	0.03	Nie and Dauphas (2019)
AGV-2	-0.14	0.02	-0.16	0.12	Pringle and Moynier (2017)
			-0.10	0.06	Zhang et al. (2018)
			-0.15	0.03	Nie and Dauphas (2019)
GSP-2	-0.27	0.01	-0.27	0.02	Hu et al. (2021)
GS-N	-0.11	0.04	-0.14	0.02	Pringle and Moynier, 2017
GA	-0.25	0.01	-0.22	0.03	Nie et al. (2021)

Table 2. Rubidium isotopic compositions of Hekla volcanic rocks.

Sample	Type	SiO ₂ (wt. %)	MgO (wt. %)	K ₂ O (wt. %)	La (µg/g)	Rb (µg/g)	δ ⁸⁷ Rb (‰)	2SD ^a	n
ICE 20-22	Basalt	48.59	4.66	0.56	22.7	9.8	-0.16	0.05	4
ICE 20-25	Basalt	48.60	5.00	0.54	20.1	9.2	-0.15	0.05	4
ICE 20-28	Basalt	46.27	6.00	0.55	24.1	11.2	-0.10	0.02	3
ICE 20-31	Basalt	47.71	6.83	0.39	13.1	6.2	-0.14	0.05	4
ICE 20-36	Basalt	51.44	4.16	0.54	21.4	8.7	-0.07	0.04	3
ICE 20-37	Basalt	45.80	5.80	0.58	24.0	11.9	-0.11	0.02	3
ICE 20-41	Basalt	46.44	5.58	0.58	25.8	11.6	-0.10	0.06	4
ICE 20-45	Basalt	51.20	3.84	0.95	40.5	15.5	-0.11	0.01	3
ICE 20-19	Basaltic Andesite	54.70	2.75	1.28	53.9	20.6	-0.17	0.03	3
ICE 20-23	Basaltic Andesite	55.34	2.58	1.31	48.6	22.7	-0.16	0.02	3
ICE 20-23 (Replicate ^b)	Basaltic Andesite	55.34	2.58	1.31	48.6	22.7	-0.17	0.03	3
ICE 20-34	Basaltic Andesite	54.59	2.76	1.27	52.7	19.7	-0.13	0.02	3
ICE 20-40	Basaltic Andesite	54.72	2.88	1.18	52.0	24.7	-0.16	0.02	3
ICE 20-44	Basaltic Andesite	54.27	2.87	1.25	48.0	27.0	-0.17	0.04	3
ICE 20-24	Andesite	57.76	2.42	1.24	48.3	22.4	-0.11	0.03	3
ICE 20-35	Andesite	57.77	1.91	1.53	55.9	34.1	-0.16	0.04	3
ICE 20-32	Dacite	69.21	0.24	2.41	71.6	57.4	-0.16	0.04	3
ICE 20-32 (Replicate ^b)	Dacite	69.21	0.24	2.41	71.6	57.4	-0.15	0.05	3
ICE 20-39	Dacite	68.43	0.24	2.49	61.0	52.1	-0.13	0.02	3
ICE 20-33	Rhyolite	76.52	0.02	2.87	64.8	68.3	-0.17	0.04	3

^a2SD = 2 times the standard deviation of the population of three repeat measurements of a sample solution;

^b Replicate = repeat chemical purification and measurement of different aliquots of a stock solution.

Table 3. Rubidium isotopic compositions of MORB samples.

Sample	Location	MgO ^a (wt. %)	(La/Sm) _N ^a	Na ₈ ^a (wt. %)	Ba/Th ^a	Rb ^a (μg/g)	δ ⁸⁷ Rb (‰)	2SD	n
CYP78 12-34	East Pacific Ridge	7.53	0.6	2.34	93.2	1.7	-0.09	0.01	4
SEARISE1 DR04	East Pacific Ridge	6.61	0.6	2.08	57.6	1.2	-0.07	0.03	4
SEARISE2 DR03	East Pacific Ridge	7.06	0.5	2.34	48.4	0.5	-0.06	0.03	4
CYP74 31-35	Mid Atlantic Ridge	10.56	1.4	2.86	103.3	4.0	-0.14	0.01	3
ARP 73 13 4	Mid Atlantic Ridge	9.95	1.5	2.66	107.2	4.2	-0.18	0.03	3
ARP 74 9 13C COLUMN	Mid Atlantic Ridge	6.61	1.2	1.53	87.3	4.2	-0.13	0.02	3
ARP 74 7 8 SURFACE	Mid Atlantic Ridge	8.28	1.5	2.03	88.9	5.6	-0.14	0.02	3
DSDP-52-417D-55-2	Mid Atlantic Ridge	5.89	0.4	1.28	45.6	0.4	-0.13	0.02	3
CH30 DR 17-03	Mid Atlantic Ridge	5.15	2.0	1.92	114.0	11.2	-0.02	0.02	3
CH30 DR 17-03 (Replicate)	Mid Atlantic Ridge	5.14	2.0	1.92	114.0	11.2	0.01	0.04	3
CH31-DR12-137	Mid Atlantic Ridge	8.28	1.3	2.21	92.3	3.9	-0.11	0.04	4
EW9309 3D-1g	Mid Atlantic Ridge	6.69	1.8	2.28	131.3	9.8	-0.13	0.03	3
MD57 D6-7	Central Indian Ridge	8.13	0.8	2.76	83.3	3.0	-0.18	0.03	4
MD57 D3-3	Central Indian Ridge	7.48	0.6	2.01	65.4	1.0	-0.19	0.01	3
MD34 D1	Central Indian Ridge	6.52	0.9	2.78	62.8	2.6	-0.14	0.04	3
MD57 D13	Central Indian Ridge	6.76	0.7	2.13	56.9	2.1	-0.16	0.06	3

^a Major and trace elements data is from She et al. (2023).