

Photodesorption and Photochemistry of Titan's Aerosol Analogs

Benjamin Fleury, Murthy S Gudipati, Isabelle Couturier-Tamburelli, Nathalie

Carrasco

► To cite this version:

Benjamin Fleury, Murthy S Gudipati, Isabelle Couturier-Tamburelli, Nathalie Carrasco. Photodesorption and Photochemistry of Titan's Aerosol Analogs. European Planetary Science Congress 2018, Sep 2018, Berlin (DE), Germany. pp.EPSC2018-659. insu-04444680

HAL Id: insu-04444680 https://insu.hal.science/insu-04444680

Submitted on 7 Feb 2024

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers. L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.



Photodesorption and Photochemistry of Titan's Aerosol Analogs

Benjamin Fleury(1), **Murthy S. Gudipati** (1), Isabelle Couturier-Tamburelli (2), Nathalie Carrasco(3) (1) Science Division, Jet Propulsion Laboratory, California Institute of Technology, 4800 Oak Grove Drive, Pasadena, California 91109, USA; (2) Aix-Marseille Université, CNRS, PIIM, UMR 7345, 13013 Marseille, France ; (3) Université Versailles St-Quentin, Sorbonne Universités, UPMC Univ. Paris 06, CNRS/INSU, LATMOS-IPSL, 11 Blvd. d'Alembert, 78280 Guyancourt, France. (gudipati@jpl.nasa.gov)

Abstract

Titan's lower stratosphere and troposphere (lower atmosphere) receives UV-depleted longer wavelength photons. These photons are not energetic enough to cause photochemical transformations of molecules in the gas phase. However, aerosols with longer wavelength absorption could transfer that energy to other molecules resulting indirect photoinduced transformations.

1. Introduction

While Titan's upper atmosphere is photochemically very active, due to the aerosols and clouds it is often assumed that Titan's lower atmosphere is photochemically inert. However, our recent works has clearly shown that even at wavelengths that pass through the upper atmosphere unattenuated (longer than 300 nm) into the lower atmosphere, photochemistry continues to happen (Couturier-Tamburelli et al. 2014; Couturier-Tamburelli, Pietri, & Gudipati 2015; Couturier-Tamburelli et al. 2018; Gudipati et al. 2013). Acetylene, the third most abundant organic molecule in Titan's atmosphere and most abundant unsaturated hydrocarbon, cannot be photoexcited at >300 nm wavelengths, either in the gas-phase or in condensed-phase such as clouds or ice. Our goal is to understand what would happen if acetylene condenses on larger aerosols that could absorb photons reaching the lower atmosphere. Would acetylene become reactive? Indeed, our results indicate it does.

2. Results

We irradiated pure acetylene ice deposited on sapphire window at 50 K using highly de-focused (to avoid multiphoton processes) laser at 355 nm. We did not find any significant change in the infrared and ultraviolet absorption spectra of acetylene. Subsequently, we deposited ~ 20 nm thick acetylene ice film on a ~ 600 nm thick Titan's aerosol analog coated sapphire window (produced through discharge). When irradiated at 355 nm, we found significant depletion of acetylene based on infrared spectra. We have repeated similar studies at longer wavelengths, whereby the depletion effect was not as dramatic.

During these studies we have also monitored gasphase molecular abundancies using quadrupole mass spectrometry. After careful experimentation and data analysis, we recognized that two processes work in parallel when acetylene coated Titan's aerosol analogs are bombarded with >300 nm photons: (a) photodesorption of acetylene, which accounts to about 25% of the total acetylene loss. Remaining 75% acetylene should have been chemically bonded with the aerosol.

Acknowledgements

This work was carried out at the Jet Propulsion Laboratory, California Institute of Technology, was under a contract with the National Aeronautics and Space Administration (NASA). Funds were provided through NASA Solar System Workings Program.

References

Couturier-Tamburelli, I., Gudipati, M. S., Lignell, A., Jacovi, R., & Pietri, N. 2014, Icar, 234, 81

Couturier-Tamburelli, I., Pietri, N., & Gudipati, M. S. 2015, Astronomy & Astrophysics, 578, A111

Couturier-Tamburelli, I., Pietri, N., Le Letty, V., Chiavassa, T., & Gudipati, M. 2018, ApJ, 852, 117

Gudipati, M. S., Jacovi, R., Couturier-Tamburelli, I., Lignell, A., & Allen, M. 2013, Nature Communications, 4, 1648