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## **Author Correction: Evidence for auroral influence on Jupiter's nitrogen and oxygen chemistry revealed by ALMA**

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# Author Correction: Evidence for auroral influence on Jupiter's nitrogen and oxygen chemistry revealed by ALMA

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In the version of the article originally published, all column density values were a factor of 10 too low. Consequently, “two orders” has been changed to “one order” in the abstract and the text, and the values have been corrected on the y axes of Fig. 3 and the color scales in Fig. 4. In addition, in the section “The spatial distribution of CO and HCN”, the CO and HCN masses reported in the text were a factor of 10 too low and the loss factors for CO and HCN were a factor of 10 too high; they have been corrected accordingly. Specifically, CO is now shown to have a meridionally uniform column density of  $1.86 \pm 0.52 \times 10^{16} \text{ cm}^{-2}$  and the total mass of SL9-derived CO is  $5.47 \pm 0.26 \times 10^{14} \text{ g}$ . This corresponds to a loss factor of  $0.9 \pm 0.3$  since 1995–1998. HCN has a uniform column density of  $22.6 \pm 5.7 \times 10^{13} \text{ cm}^{-2}$ , which is one order of magnitude lower than that measured 6.5 years after the SL9 impacts. The total mass of HCN is  $5.0 \pm 0.1 \times 10^{12} \text{ g}$ . This corresponds to a loss factor of  $5.0 \pm 3.0$  when compared with the 1995–1998 period, or  $12.0 \pm 3.5$  when compared with the value derived in 2000 from the Cassini flyby data. The conclusions of the paper remain unchanged.

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